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Opracowanie utworu pod tytułem:

"Zimne fermiony" w ramach kursu zaawansowanego, organizowanego w dniach 31.08 – 25.09.09 będącego kontynuacją szkoleń z zakresu eksploatacji i zarządzania dużą infrastrukturą badawczą organizowanego przez Narodowe Laboratorium Technologii Kwantowych



INNOWACYJNA GOSPODARKA
NARODOWA STRATEGIA SPÓJNOŚCI

UNIA EUROPEJSKA
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ROZWOJU REGIONALNEGO



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Kurs letni: inżynieria kwantowa – Zimne fermiony

PIOTR DEUAR

ULTRACOLD FERMI GASES

WARSZAWA

24 SEPTEMBER 2009

PLAN

MORNING:
9⁰⁰

- BASIC PROPERTIES OF FERMI GASES
- IDEAL GAS IN A BOX
- IN A TRAP : THOMAS-FERMI APPROXIMATION

11⁰⁰ TOMASZ KAWALEC

AFTERNOON
15⁰⁰

- BCS SUPERFLUIDITY
- OTHER TOPICS :
 - SPIN-POLARISED GAS ?
 - BEC-BCS CROSSOVER ?
 - FFLO ?
 - DIPOLES ?

- USEFUL REFERENCES : - S. Giorgini et al. Reviews of Modern Physics
VOLUME 80 PAGE 1215 (2008)
- KRYSZTOF SACHA : KONDENSAT BOSEGO-EINSTEINA



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Real Basics

	BOSONS	FERMIONS
SPIN	0, 1, 2, ...	$\frac{1}{2}, \frac{3}{2}, \frac{5}{2}, \dots$
EXCHANGE	$\Psi(x_1, x_2) = \Psi(x_2, x_1)$	$\Psi(x_1, x_2) = -\Psi(x_2, x_1)$
COMMUTATION	$[\hat{\Psi}(x), \hat{\Psi}(y)] = \delta(x-y)$	$[\hat{\Psi}(x), \hat{\Psi}(y)]_+ = \delta(x-y)$
SUPERFLUID?	IN IDEAL OR INTERACTING GAS	IN INTERACTING GAS ONLY
PARTICLES/ORBITAL	0, 1, ... ∞	0, 1

Related to properties of relativistic field theory by the

Spin-statistics theorem (Markus Fierz 1939)
rederived Pauli 1940

ANTI-COMMUTATOR
 $[A, B]_+ = AB + BA$

QUESTION: What are some fermions that you can make a gas out of?



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Pauli exclusion

- Suppose we have two fermions with positions x & y
& one orbital ϕ

- if both particles are in ϕ then wavefunction is:

$$\Psi = \phi(x) \otimes \phi(y)$$

$$= -\phi(y) \otimes \phi(x)$$

$$\rightarrow 2\phi(x) \otimes \phi(y) = 0 \quad \text{for all } x \text{ \& } y$$

$$\rightarrow \phi(x) = 0 \text{ or } \phi(y) = 0 \text{ for all } x \text{ \& } y$$

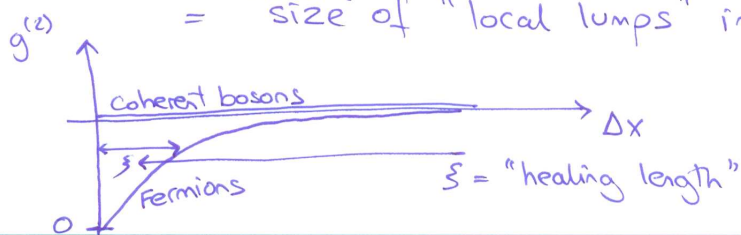
$$\rightarrow \phi = 0 \rightarrow \text{NO SUCH STATE}$$

Effect 1: Repulsion between particles

two-body (density-density) correlation function

$$g^{(2)}(\Delta x) = \frac{\langle \hat{\psi}^\dagger(x) \hat{\psi}(x+\Delta x) \hat{\psi}^\dagger(x+\Delta x) \hat{\psi}(x) \rangle}{n(x) n(x+\Delta x)}$$

= shape & size of "local lumps" in the gas



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nature

Vol 445 | 25 January 2007 | doi:10.1038/nature05513

LETTERS

Comparison of the Hanbury Brown–Twiss effect for bosons and fermions

T. Jeltes¹, J. M. McNamara¹, W. Hogervorst¹, W. Vassen¹, V. Krachmalnicoff², M. Schellekens², A. Perrin³, H. Chang², D. Boiron², A. Aspect² & C. I. Westbrook²

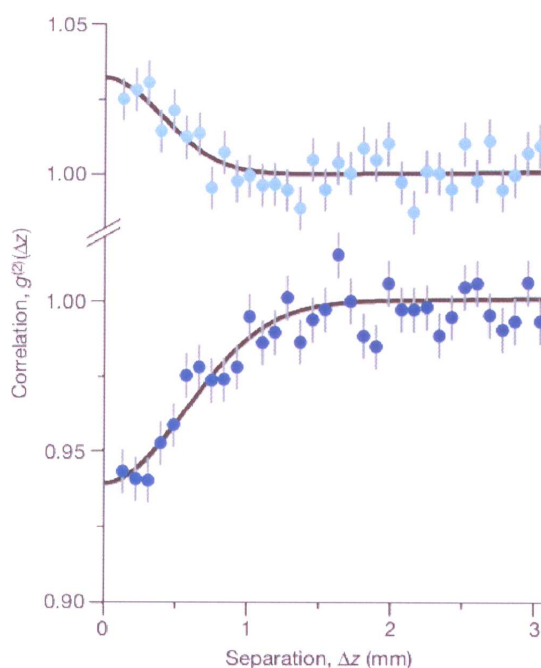


Figure 2 | Normalized correlation functions for $^4\text{He}^+$ (bosons) in the upper plot, and $^3\text{He}^+$ (fermions) in the lower plot. Both functions are measured at the same cloud temperature ($0.5\ \mu\text{K}$), and with identical trap parameters. Error bars correspond to the square root of the number of pairs in each bin. The line is a fit to a gaussian function. The bosons show a bunching effect, and the fermions show antibunching. The correlation length for $^3\text{He}^+$ is expected to be 33% larger than that for $^4\text{He}^+$ owing to the smaller mass. We find $1/e$ values for the correlation lengths of $0.75 \pm 0.07\ \text{mm}$ and $0.56 \pm 0.08\ \text{mm}$ for fermions and bosons, respectively.



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Effect 2: COLLISIONS SUPRESSED

- ~~part~~ - in dilute gas the size of the interatomic potential is much less than the healing length ξ
 - ↳ atoms find it hard to get close enough to collide
 - elastic collisions suppressed
- three-body collisions are similarly unlikely
 - ↳ 3-body loss strongly reduced in comparison with Bose gas
- few collisions mean slow thermalisation
 - ↳ difficult to cool!
 - ↳ use buffer gas cooling
- attractive interactions are necessary for BCS superfluidity (more on that later)
 - ↳ need two different "species" of fermion or long-range interactions (dipoles)

Effect 3: Fermi sphere (see below)



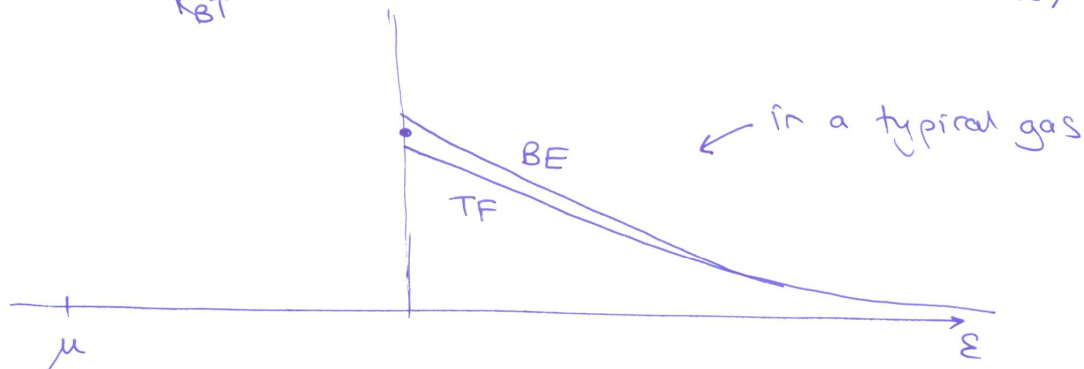
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Q: why is fermi / Bose statistics irrelevant for everyday gases?

Bose-Einstein
Fermi-Dirac } distribution $n(\epsilon) = \frac{1}{\exp[(\epsilon - \mu)/k_B T] \pm 1}$

mean number of particles per orbital with energy ϵ

if $\frac{\epsilon - \mu}{k_B T} \gg 1$ have Boltzmann distribution $n(\epsilon) \sim e^{-\frac{\mu - \epsilon}{k_B T}}$



consider the ideal gas: $\epsilon = \frac{p^2}{2m} \geq 0$

$$\mu = \left. \frac{\partial E}{\partial N} \right|_{E,V} = k_B T \log \left[\frac{n}{n_Q} \right]$$

$\frac{n}{n_Q}$ = "phase-space density" $n_Q = \left(\frac{m k_B T}{2\pi \hbar^2} \right)^{3/2}$



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"typical" gas - consider air

$$T \sim 300 \text{ K}$$

$$m \sim 30 u = 5 \times 10^{-26} \text{ kg}$$

$$k_B = 1.38 \times 10^{-23} \text{ J/K}$$

$$\hbar = 1.054 \times 10^{-34} \text{ J}\cdot\text{s}$$

$$\rightarrow n_Q \sim 1.62 \times 10^{32} \text{ m}^{-3}$$

$$\text{density } n \sim \frac{N_A}{22.4 \times 10^{-3} \text{ m}^3} \sim 2.69 \times 10^{25} \text{ m}^{-3}$$

$$\therefore \text{phase-space density} = 0.000000166$$

$$\rightarrow \frac{\mu}{k_B T} \approx -15.6$$

thus $\frac{E - \mu}{k_B T} > 15.6$ which is very far from any differences in statistics

Degenerate Fermi gas : $\frac{n}{n_Q} \approx 1$

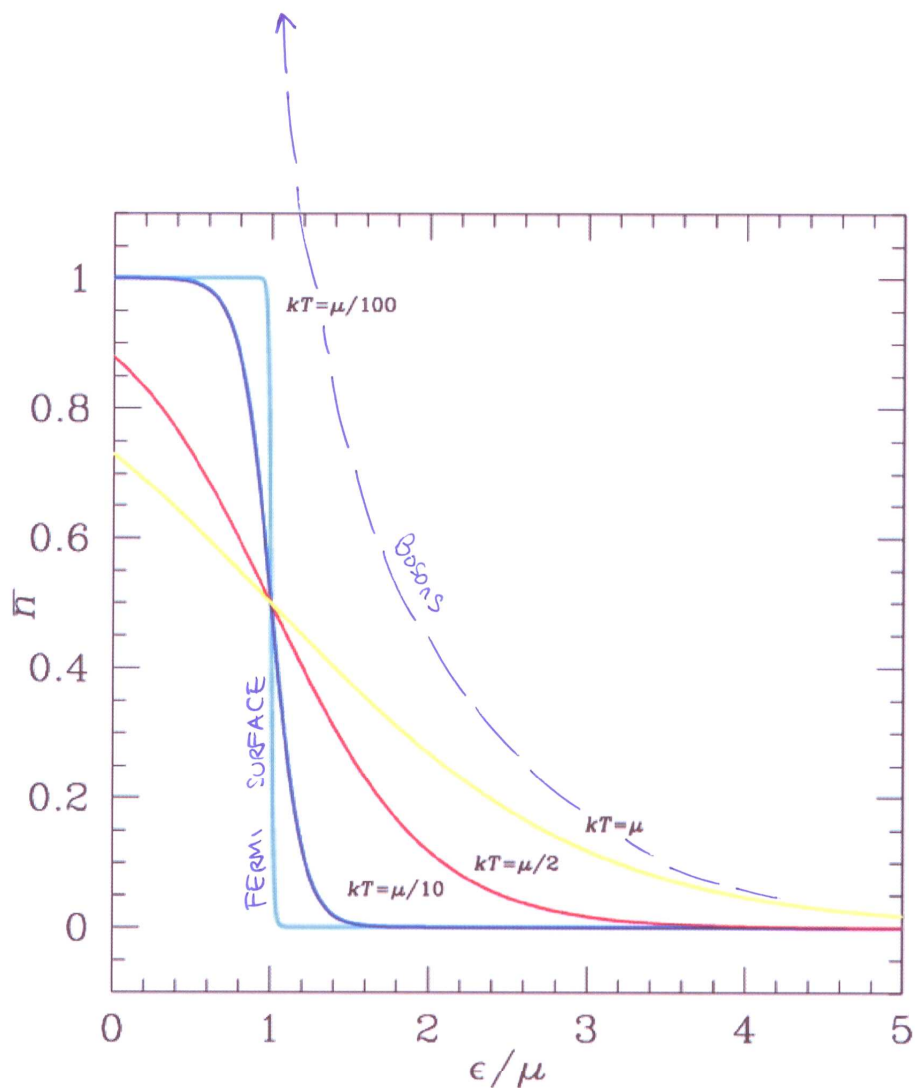
here $\mu > 0$ and a "Fermi sphere"
& "Fermi surface" appear

at the fermi energy $E_F \approx \mu$

as $T \rightarrow 0$



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IDEAL FERMION GAS IN A BOX $T \rightarrow 0$

Box length L , 3D, periodic boundary conditions

eigenstates $\phi_{\vec{k}} \sim e^{i\vec{k} \cdot \vec{x}}$

$$\varepsilon(\vec{k}) = \frac{p^2}{2m} = \frac{\hbar^2 k^2}{2m}$$

here: $\vec{k} = [k_x, k_y, k_z]$

$$k_j = n_j \Delta k$$

$$n_j = \dots, -2, -1, 0, 1, 2, \dots$$

$$\Delta k = \frac{2\pi}{L}$$

From Fermi-Dirac distribution at $T \rightarrow 0$

$$\text{occupation} = \begin{cases} 1 & \text{if } \varepsilon(\vec{k}) < \mu \\ 0 & \text{if } \varepsilon(\vec{k}) > \mu \end{cases}$$

Want to find Fermi energy $E_F = \mu = k_B T_F$ (Fermi temp.)
as a function of density etc.

To have N fermions: $N = 1 \times \sum_{\substack{\vec{k} \\ \varepsilon(\vec{k}) < \mu}} \cdot$
occupation modes

let \vec{k} go to polar coordinates k, θ, ϕ

$$N \approx \frac{1}{(\Delta k)^3} \int_{\varepsilon(\vec{k}) < \mu} d^3 \vec{k}$$

$$= \frac{V}{(2\pi)^3} \int_0^{k_F} k^2 dk \int_0^{2\pi} d\theta \int_0^\pi \sin\theta d\theta$$

where $\mu = \frac{\hbar^2 k_F^2}{2m}$ defines fermi momentum k_F

$$N = \frac{V k_F^3}{6\pi^2}$$



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hence, density $n = \frac{k_F^3}{6\pi^2} \rightarrow k_F^2 = (6\pi^2 n)^{2/3}$

$$\rightarrow \mu = \frac{\hbar^2 k_F^2}{2m} = \frac{\hbar^2}{2m} (6\pi^2 n)^{2/3} = k_B T_F$$

compare this to the BEC critical temperature in a box:

$$k_B T_c = \frac{\hbar^2}{2m} \left[\frac{8\pi\sqrt{\pi}}{g(3/2)} n \right]^{2/3}$$

→ serious quantum gas effects occur for temperatures scaling as $\propto n^{2/3}$

→ for both $\left\{ \begin{array}{l} \text{high } n \\ \text{low } m \end{array} \right.$ are good for raising critical temperature.



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IN A TRAP – THOMAS FERMI APPROXIMATION

Allows one to obtain the 1-particle density and distribution in momentum-space if interactions are small compared to μ

LOCAL DENSITY APPROXIMATION (the original TF approx as opposed to the Boson case)

$$H = \frac{\vec{p}^2}{2m} + U(\vec{r}) \quad \leftarrow \therefore \text{interaction should be small!}$$

$$\text{harmonic trap } U(\vec{r}) = \frac{m}{2} \left\{ \omega_x^2 x^2 + \omega_y^2 y^2 + \omega_z^2 z^2 \right\}$$

ASSUMPTION 1: chemical potential is the same regardless of where in the cloud the atom is added.

Note: in the trap $\mu = (6N)^{1/3} \hbar \omega_{ho}$ Different from uniform gas!
 $\omega_{ho} = [\omega_x \omega_y \omega_z]^{1/3}$

this is obtained similarly to uniform gas, but the summing over states is more complicated.

ASSUMPTION 2: locally the gas behaves like a uniform gas $\rightarrow U(\vec{r})$ must be "smooth enough".

\therefore LOCAL FERMI SPHERE with local Fermi Energy

$$\mu = E_F(\vec{r}) + U(\vec{r})$$





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since locally the gas behaves like the uniform ideal gas

$$E_F(\vec{r}) = \frac{\hbar^2 k_F(\vec{r})^2}{2m}, \quad n(\vec{r}) = \frac{k_F(\vec{r})^3}{6\pi^2}$$

combining

$$n(\vec{r}) = \frac{1}{6\pi^2} \left[\frac{2m}{\hbar^2} E_F(\vec{r}) \right]^{3/2}$$

$$= \frac{1}{6\pi^2} \left[\frac{2m}{\hbar^2} (\mu - U(\vec{r})) \right]^{3/2}$$

For the harmonic trap

$$n(\vec{r}) = \frac{1}{6\pi^2} \frac{(2m)^{3/2}}{\hbar^3} \left[\mu - \frac{m}{2} \sum_i \omega_i^2 x_i^2 \right]^{3/2}$$

- which can be anisotropic.

- is zero for $U(\vec{r}) > \mu$ like BEC T-F profile

- is $\propto (\mu - U(\vec{r}))^{3/2}$ unlike BEC which has $\propto (\mu - U(\vec{r}))^1$

However, the momentum distribution is always isotropic
(unlike the BEC)

to see this:

As always, occupation of orbitals is given by Fermi-Dirac distribution. However now, there ~~is~~ is a local distribution at each \vec{r} due to the local density approximation

$$n(\vec{k}, \vec{r}) = \frac{1}{\exp\left\{ \left[\frac{\hbar^2 k^2}{2m} + U(\vec{r}) - \mu \right] / k_B T \right\} + 1}$$



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Consider space split up into little boxes of volume ΔV
length l

the gas is \approx uniform in each box

here momentum is discretized into values with
spacing $\Delta k = \frac{2\pi}{l}$

the number of particles with momenta between k and $k + \Delta k$
in the entire system is

$$\tilde{n}(k)(\Delta k)^3 = \sum_{\vec{k} \in [\vec{k}, \vec{k} + \Delta \vec{k}]} \times \sum_{\vec{r}} \times n(k, r)$$

Volume in k -space
of sides Δk in all 3 directions

Sum over boxes

hence

$$\begin{aligned} \tilde{n}(k) &\approx \frac{1}{(\Delta k)^3} \times \sum_{\vec{k} \in [\vec{k}, \vec{k} + \Delta \vec{k}]} \times \frac{1}{\Delta V} \int d\vec{r} n(k, r) \\ &\quad \uparrow \\ &\quad \text{there is } \approx 1 \text{ such momentum per box} \\ &= \frac{l^3}{(2\pi)^3} \times 1 \times \frac{1}{l^3} \int d\vec{r} n(k, r) \\ &= \frac{1}{(2\pi)^3} \int d\vec{r} n(k, r) \end{aligned}$$

now take $T \rightarrow 0$ limit

$$\hookrightarrow n(k, r) \approx \Theta \left[\mu - \frac{\hbar^2 k^2}{2m} - U(\vec{r}) \right]$$





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$$\tilde{n}(k) \approx \frac{1}{(2\pi)^3} \int d^3\vec{r} \Theta \left[\frac{\hbar^2}{2m} (k_F^2 - k^2) - \frac{m}{2} \sum_i \omega_i^2 r_i^2 \right]$$

↑
anisotropic → a pain

$$\text{let } \omega_i r_i = q_i \quad \rightarrow \quad d^3\vec{r} = \frac{d^3\vec{q}}{\omega_x \omega_y \omega_z}$$

$$\tilde{n}(k) \approx \frac{1}{(2\pi \omega_{ho})^3} \int d^3\vec{q} \Theta \left[\frac{\hbar^2}{2m} (k_F^2 - k^2) - \frac{m}{2} |\vec{q}|^2 \right]$$

polar coordinates of \vec{q} : $|\vec{q}|, \theta_q, \phi_q$

$$= \frac{1}{(2\pi \omega_{ho})^3} \int_0^\infty |\vec{q}|^2 d|\vec{q}| \Theta \left[\frac{\hbar^2}{2m} (k_F^2 - k^2) - \frac{m}{2} |\vec{q}|^2 \right] \int_0^{2\pi} d\theta_q \int_0^\pi \sin\phi_q d\phi_q$$

$$= \frac{1}{2\pi^2 \omega_{ho}^3} \int_0^{q_{\max}} |\vec{q}|^2 d|\vec{q}| \quad \text{with } \frac{m}{2} q_{\max}^2 = \frac{\hbar^2}{2m} (k_F^2 - k^2)$$

$$\hookrightarrow q_{\max} = \sqrt{\frac{2\mu}{m} \left(1 - \left(\frac{k}{k_F} \right)^2 \right)}$$

$$= \frac{1}{6\pi^2 \omega_{ho}^3} q_{\max}^3$$

$$\tilde{n}(k) = \frac{2\sqrt{2} \mu^{3/2}}{6\pi^2 \omega_{ho}^3 m^{3/2}} \times \left[1 - \left(\frac{k}{k_F} \right)^2 \right]^{3/2} \quad \text{isotropic!}$$



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PHYSICAL REVIEW LETTERS

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Pauli Blocking of Collisions in a Quantum Degenerate Atomic Fermi Gas

B. DeMarco, S. B. Papp, and D. S. Jin*

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and Physics Department, University of Colorado, Boulder, Colorado 80309-0440
(Received 29 January 2001)*

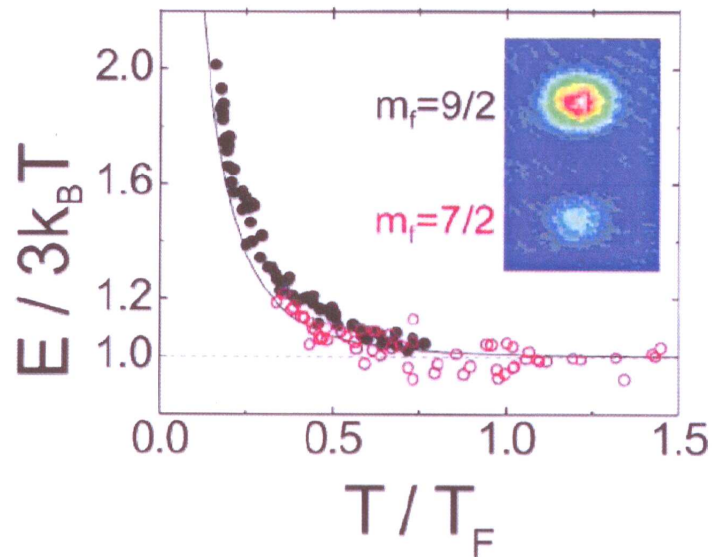


FIG. 3. (Color online) Evidence for quantum degeneracy effects in trapped Fermi gases. The average energy per particle, extracted from absorption images, is shown for two-spin mixtures. In the quantum degenerate regime, the data agree well with the ideal Fermi gas prediction (solid line). The horizontal dashed line corresponds to the result of a classical gas. From De Marco *et al.*, 2001.



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Lecture 2 : BCS THEORY - SUPERFLUIDITY

J. Bardeen L. Cooper J. Schrieffer

Phys. Rev. 108 1175 (1957)

originally for superconductivity = superfluidity of e^-

INTRO :

How is a superfluid (SF) made in the Bose gas?

OFF-DIAGONAL LONG RANGE ORDER (ODLRO) is needed

where $\langle \hat{f} \rangle \neq 0$ for some non-Hermitian \hat{f}

- have BEC which provides ODLRO via the
order parameter $\langle \hat{\psi}(x) \rangle = \psi(x)$

- kills a boson

- coherently so $\langle \hat{\psi} \rangle \neq 0$

because there is a well defined phase

condensate wavefunction

In Helium SF in He^4 at 2.18 K

He^3 at ~ 1 mK

How to make a SF with fermions?



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suggestion : same way as with bosons.

Q: how to kill a boson when you have fermions?

A: kill 2 fermions.

let $\hat{\Pi}(x)$ be a boson [= fermion pair] annihilator

$$\hat{\Pi}(x) = \hat{\psi}(x) \hat{\psi}(x)$$

if we had ODLRO in $\hat{\Pi}$: $\langle \hat{\Pi}(x) \rangle = \Pi(x) \neq 0$
it should lead to superfluidity just like
(or at least similarly to) with $\psi(x)$ for bosons.

small problem: due to fermion properties

$$\hat{\psi}(x) \hat{\psi}(x) = -\hat{\psi}(x) \hat{\psi}(x) \quad (\text{exchange } x \leftrightarrow x)$$

$$\rightarrow \hat{\psi}(x) \hat{\psi}(x) = 0 \quad \ddot{\vdots}$$

(theoretically) easy fix : two components. \uparrow and \downarrow

(cooper) pair : $\hat{\psi}_{\uparrow}(x) \hat{\psi}_{\downarrow}(x)$ consists of one \uparrow and one \downarrow

$$\text{ODLRO : } \langle \hat{\psi}_{\uparrow}(x) \hat{\psi}_{\downarrow}(x) \rangle = \Delta(x)$$

macroscopic wavefunction of cooper pairs

called "the gap"



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BCS HAMILTONIAN : TRUE \hat{H}

consider $\hat{K} = \hat{H} - \mu \sum_{\sigma} \hat{N}_{\sigma}$ the "Kamiltonian"

$$\hat{K} = \sum_{\sigma} \int d^3r \left\{ \hat{\psi}_{\sigma}^{\dagger}(r) H_0 \hat{\psi}_{\sigma}(r) + \frac{g}{2} \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}(r) \hat{\psi}_{\sigma}(r) \right\}$$

$\sigma = \pm \frac{1}{2}$

Contact interactions
attractive: $g < 0$

$$H_0 = \frac{-\hbar^2 \nabla^2}{2m} - \mu$$

[uniform gas] equal μ for both components σ
& same number of ^{atoms} ~~spins~~ in each spin

EFFECTIVE \hat{H} (\hat{K})

Want an "effective" mean field theory that describes single fermions in a mean field produced by the rest

↳ want only pairs of operators $\hat{\psi}, \hat{\psi}^{\dagger}$

Postulate:

$$\hat{K}_{\text{eff}} = \int d^3r \left\{ \sum_{\sigma} \left[\hat{\psi}_{\sigma}^{\dagger}(r) H_0 \hat{\psi}_{\sigma}(r) + W(r) \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}(r) \right] \right. \\ \left. + \frac{1}{2} F(r) \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}(r) + \text{h.c.} \right. \\ \left. + \frac{1}{2} \Delta(r) \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}^{\dagger}(r) + \text{h.c.} \right\}$$

with $\Delta(r)$
 $W(r)$
 $F(r)$
to be determined



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How to choose Δ, W, F ??

1) consider the free energy based on \hat{K}_{ef}

$$F_{ef} = \langle \hat{K}_{ef} \rangle_{ef} - TS \quad \text{using the basis that diagonalizes } \hat{K}_{ef}$$

because of this, F_{ef} jest stacjonarna

$$\delta F_{ef} = 0 \quad \text{with respect to eigenstates.}$$

2) consider the free energy based on \hat{K} (full) : F
calculated in the \hat{K}_{ef} basis. $F = \langle \hat{K} \rangle_{ef} - TS$

To have $\delta F = 0$ also we want

$$\delta F = \delta F_{ef}$$

[then \hat{K}_{ef} is chosen so that its eigenstates make
the full free energy minimized as well.]





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what is δE_{ef} ?

$$\delta E_{ef} = \int d^3r \sum_{\sigma} \left\{ \delta \langle \hat{\psi}_{\sigma}^{\dagger}(r) \mu_0 \hat{\psi}_{\sigma}(r) \rangle + W(r) \delta \langle \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}(r) \rangle \right. \\ \left. + \frac{1}{2} F(r) \delta \langle \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}(r) \rangle + \frac{1}{2} F(r)^* \delta \langle \hat{\psi}_{-\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}(r) \rangle \right. \\ \left. + \frac{1}{2} \Delta(r) \delta \langle \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}^{\dagger}(r) \rangle + \frac{1}{2} \Delta(r)^* \delta \langle \hat{\psi}_{-\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}^{\dagger}(r) \rangle \right\} \\ - T \delta S$$

what is δF ?

e.g. $\langle F \rangle$ involves $\langle \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}^{\dagger}(r) \hat{\psi}_{-\sigma}(r) \hat{\psi}_{\sigma}(r) \rangle$

using the "generalised" Wick theorem

[applies to statistical ensembles of the form $e^{-\hat{K}_{ef}/k_B T}$ with quadratic \hat{K}_{ef}]

$$\rightarrow = \langle \psi_{\sigma}^{\dagger}(r) \psi_{-\sigma}^{\dagger}(r) \rangle \langle \bar{\psi}_{-\sigma}(r) \bar{\psi}_{\sigma}(r) \rangle + \langle \psi_{\sigma}^{\dagger}(r) \psi_{\sigma}(r) \rangle \langle \psi_{-\sigma}^{\dagger}(r) \psi_{-\sigma}(r) \rangle \\ - \langle \psi_{\sigma}^{\dagger}(r) \psi_{-\sigma}(r) \rangle \langle \psi_{-\sigma}^{\dagger}(r) \psi_{\sigma}(r) \rangle$$



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hence...

$$\begin{aligned} \delta F = \int d^3r \sum_{\sigma} \left\{ \delta \langle \psi_{\sigma}^{\dagger} H_0 \psi_{\sigma} \rangle + \frac{g}{2} \langle \psi_{\sigma}^{\dagger} \psi_{-\sigma}^{\dagger} \rangle \delta \langle \psi_{-\sigma} \psi_{\sigma} \rangle \right. \\ \left. + \frac{g}{2} \delta \langle \psi_{\sigma}^{\dagger} \psi_{-\sigma}^{\dagger} \rangle \langle \psi_{-\sigma} \psi_{\sigma} \rangle + \frac{g}{2} \langle \psi_{\sigma}^{\dagger} \psi_{\sigma} \rangle \delta \langle \psi_{-\sigma}^{\dagger} \psi_{-\sigma} \rangle + \frac{g}{2} \delta \langle \psi_{\sigma}^{\dagger} \psi_{\sigma} \rangle \langle \psi_{-\sigma}^{\dagger} \psi_{-\sigma} \rangle \right. \\ \left. - \frac{g}{2} \delta \langle \psi_{\sigma}^{\dagger} \psi_{-\sigma} \rangle \langle \psi_{-\sigma}^{\dagger} \psi_{\sigma} \rangle - \frac{g}{2} \langle \psi_{\sigma}^{\dagger} \psi_{-\sigma} \rangle \delta \langle \psi_{-\sigma}^{\dagger} \psi_{\sigma} \rangle \right\} - T \delta S \end{aligned}$$

equating $\delta F = \delta F_{eff}$ get the "optimized" choice:

$$\Delta(r) = g \langle \hat{\psi}_{-\sigma}(r) \hat{\psi}_{\sigma}(r) \rangle \leftarrow \text{"Gap equation"}$$

$$W(r) = g \langle \hat{\psi}_{\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}(r) \rangle$$

$$F(r) = -g \langle \hat{\psi}_{-\sigma}^{\dagger}(r) \hat{\psi}_{\sigma}(r) \rangle$$

NOTES: i) if $F \neq 0$ then Spin is not conserved in system
(not what we have in mind)

↳ Assume $F=0$

ii) $\Delta(r)$ can be identified with the pairing wavefunction suggested earlier.

iii) $W(r)$ is usually omitted \rightarrow it is not very important for superfluidity. (see below)



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The full description then involves \hat{K}_{eff} and the Gape equation

they must be solved self-consistently.

UNIFORM GAS

Better in k -space

$$\hat{\psi}_{\sigma}(\vec{r}) = \frac{1}{(2\pi)^{3/2}} \int d^3\vec{k} e^{i\vec{k}\cdot\vec{r}} \hat{\psi}_{\sigma}(\vec{k})$$

$$\Delta(r) = \text{const.} = \Delta$$

obtain:

$$\hat{K}_{\text{eff}} = \int d^3\vec{k} \left\{ \varepsilon(\vec{k}) \hat{\psi}_{\sigma}^{+\dagger}(\vec{k}) \hat{\psi}_{\sigma}(\vec{k}) + \frac{\Delta}{2} \hat{\psi}_{\sigma}^{+\dagger}(\vec{k}) \hat{\psi}_{\sigma}^{+\dagger}(-\vec{k}) + \frac{\Delta^*}{2} \hat{\psi}_{\sigma}(\vec{k}) \hat{\psi}_{\sigma}(-\vec{k}) \right\}$$

with $\varepsilon(\vec{k}) = \frac{\hbar^2 k^2}{2m} - \mu + W$

W is seen as a harmless correction to μ

Note the pairing of k & $-k$.



Kurs letni: inżynieria kwantowa – Zimne fermiony

in creation operators : $\hat{\psi}_{\sigma}(\mathbf{r}) \sim \frac{\hat{a}_{\sigma, \mathbf{k}}}{\sqrt{(\Delta \mathbf{k})^3}}$

$$\hat{K}_{\text{eff}} = \sum_{\mathbf{k}} \left\{ \varepsilon(\mathbf{k}) \sum_{\sigma} a_{\sigma \mathbf{k}}^{\dagger} a_{\sigma \mathbf{k}} + \Delta a_{\uparrow \mathbf{k}}^{\dagger} a_{\downarrow -\mathbf{k}}^{\dagger} + \Delta^* a_{\downarrow -\mathbf{k}} a_{\uparrow \mathbf{k}} \right\}$$

$$= \sum_{\mathbf{k}} \left[a_{\uparrow \mathbf{k}}^{\dagger}, a_{\downarrow -\mathbf{k}} \right] \Omega \begin{bmatrix} a_{\uparrow \mathbf{k}}^{\dagger} \\ a_{\downarrow, -\mathbf{k}}^{\dagger} \end{bmatrix} + C$$

$$C = \sum_{\mathbf{k}} \varepsilon(\mathbf{k})$$

$$\Omega = \begin{bmatrix} \varepsilon_{\mathbf{k}} & \Delta \\ \Delta^* & -\varepsilon_{\mathbf{k}} \end{bmatrix}$$

and Gap equation is

$$\Delta = g \sum_{\mathbf{k}} \langle a_{\downarrow -\mathbf{k}} a_{\uparrow \mathbf{k}} \rangle$$

\hat{K}_{eff} can be diagonalised into

$$\hat{K}_{\text{eff}} = \sum_{\mathbf{k}} E_{\mathbf{k}} \left[b_{\uparrow \mathbf{k}}^{\dagger} b_{\uparrow \mathbf{k}} + b_{\downarrow -\mathbf{k}}^{\dagger} b_{\downarrow -\mathbf{k}} \right] + \tilde{C}$$



Kurs letni: inżynieria kwantowa – Zimne fermiony

with $\tilde{c} = \sum_{\mathbf{k}} (\xi_{\mathbf{k}} c_{\mathbf{k}} - E_{\mathbf{k}})$

$$E_{\mathbf{k}} = \sqrt{\xi_{\mathbf{k}}^2 + \Delta^2} \quad \text{from eigenvalues of } \Omega$$

by the Bogoliubov transformation

$$\begin{cases} a_{\uparrow\mathbf{k}} = b_{\uparrow\mathbf{k}} u_{\mathbf{k}} + b_{\downarrow-\mathbf{k}}^{\dagger} v_{\mathbf{k}} \\ a_{\downarrow-\mathbf{k}} = b_{\uparrow\mathbf{k}} v_{\mathbf{k}} + b_{\downarrow-\mathbf{k}}^{\dagger} u_{\mathbf{k}} \end{cases} \quad \text{with fermi operators } \hat{b}_{\sigma\mathbf{k}}$$

~~eigenvectors~~ commutation relations on a, a^{\dagger}

give ~~$u_{\mathbf{k}} v_{\mathbf{m}}^{\dagger} - v_{\mathbf{k}} u_{\mathbf{m}}^{\dagger} = 0$~~

$$u_{\mathbf{k}} u_{\mathbf{m}}^{\dagger} + v_{\mathbf{k}} v_{\mathbf{m}}^{\dagger} = \delta_{\mathbf{k}\mathbf{m}}$$

squaring eigenvalue eqn. for $u_{\mathbf{k}}$ obtain

$$(E_{\mathbf{k}} u_{\mathbf{k}})^2 = (\xi_{\mathbf{k}} u_{\mathbf{k}} + \Delta v_{\mathbf{k}})^2$$

$$(\xi_{\mathbf{k}}^2 + \Delta^2) u_{\mathbf{k}}^2 = \xi_{\mathbf{k}}^2 u_{\mathbf{k}}^2 + \Delta^2 v_{\mathbf{k}}^2 + 2\xi_{\mathbf{k}} \Delta u_{\mathbf{k}} v_{\mathbf{k}}$$

↳ obtain $u_{\mathbf{k}} v_{\mathbf{k}} = \frac{\Delta}{2E_{\mathbf{k}}}$

$$u_{\mathbf{k}}^2 = \frac{1}{2} \left(1 + \frac{\xi_{\mathbf{k}}}{E_{\mathbf{k}}} \right)$$

$$v_{\mathbf{k}}^2 = \frac{1}{2} \left(1 - \frac{\xi_{\mathbf{k}}}{E_{\mathbf{k}}} \right)$$



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this leads to

$$\Delta = -g \sum_{\mathbf{k}} \frac{\Delta}{2 \sqrt{\epsilon_{\mathbf{k}}^2 + \Delta^2}} \quad \text{Gap equation}$$

using the fact that we are in
the ground state and $\hat{b}_{\sigma\mathbf{k}} |GS\rangle = 0$

this can be solved to give

$$\Delta \approx \left(\frac{2}{e}\right)^{2/3} E_F \exp\left[-\frac{\pi^{1/3}}{2|a| (6n)^{1/3}}\right]$$

where $a =$ scattering length

$$g = \frac{4\pi\hbar^2 a}{m}$$

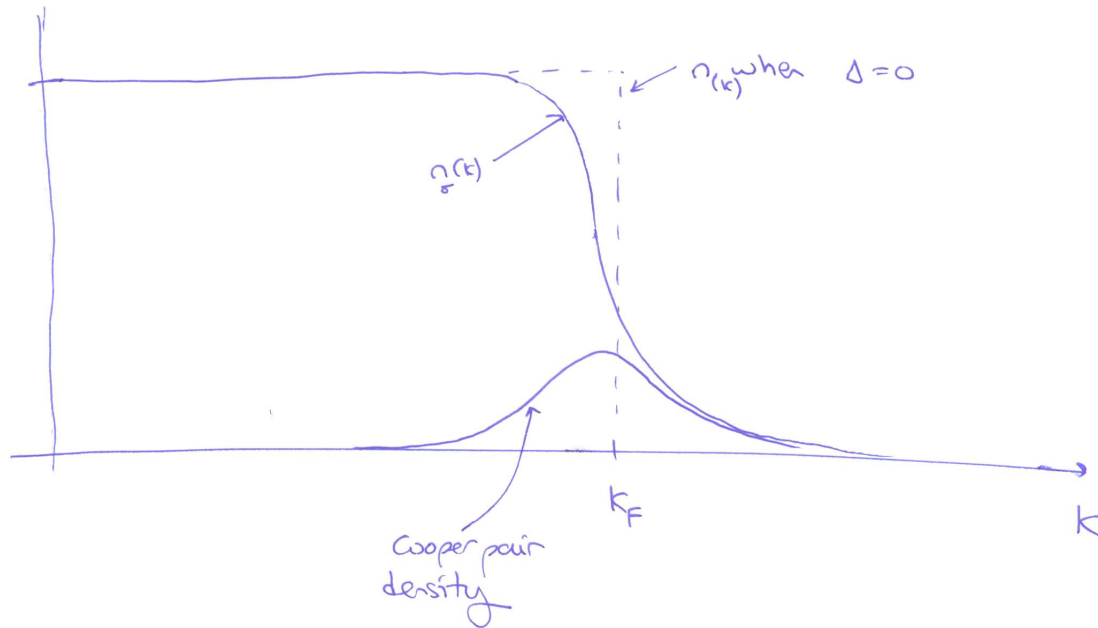
also $T_c \approx 0.28 E_F \exp[\dots]$



Kurs letni: inżynieria kwantowa – Zimne fermiony

Fermion density $n_{\sigma}(k) = \langle a_{\sigma k}^{\dagger} a_{\sigma k} \rangle = v_{k}^2$

Cooper pair density $|\langle a_{\sigma k} a_{\sigma-k} \rangle|^2 = u_{k}^2 v_{k}^2$



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ARTICLES

Vortices and superfluidity in a strongly interacting Fermi gas

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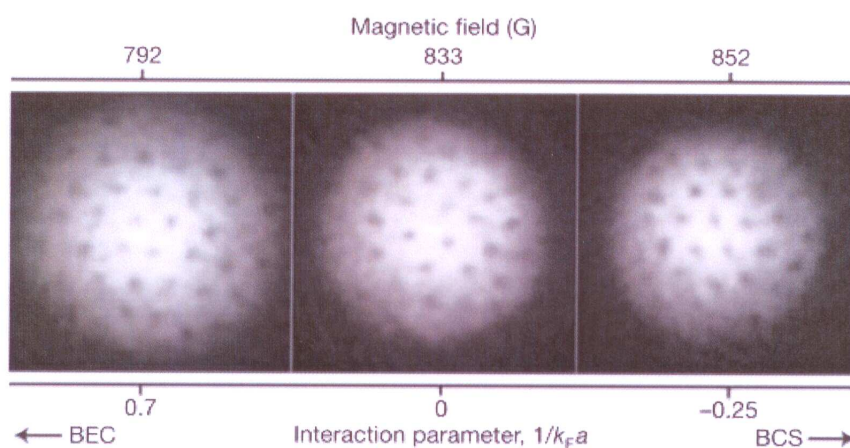


Figure 3 | Optimized vortex lattices in the BEC-BCS crossover. After a vortex lattice was created at 812 G, the field was ramped in 100 ms to 792 G (BEC-side), 833 G (resonance) and 853 G (BCS-side), where the cloud was held for 50 ms. After 2 ms of ballistic expansion, the magnetic field was ramped to 735 G for imaging (see text for details). The field of view of each image is $880 \mu\text{m} \times 880 \mu\text{m}$.



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