

Open String Dynamics in Warped Backgrounds Gary Sbiu
University of Wisconsin

## Based on:

- Marchesano, McGuirk, GS, 0812.2247
- Chen, Nakayama, GS, 0905.4463
- McGuirk, GS, Sumitomo, in progress

See also:

- GS, Torroba, Underwood, Douglas, 0803.3068


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## Compensators and Warping

$\begin{aligned} \delta G_{\nu}^{\mu}= & \delta_{\nu}^{\mu} u^{I} \delta_{I}\left\{e^{2 A}\left[-2 \tilde{\nabla}^{2} A+4(\widetilde{\nabla A})^{2}-\frac{1}{2} \tilde{R}\right]\right\}+e^{-2 A}\left(\partial^{\mu} \partial_{\nu} u^{I}-\delta_{\nu}^{\mu} \square u^{I}\right)\left(4 \delta_{I} A-\frac{1}{2} \delta_{I} \tilde{g}\right) \\ & +\left(\partial^{\mu} \partial_{\nu} u^{I}-\delta_{\nu}^{\mu} \square u^{I}\right) e^{2 A} \tilde{\nabla}^{p}\left(B_{I p}-\partial_{p} K_{I}\right) \\ & +e^{-2 A} f^{K} \delta_{K} G_{\nu}^{(4) \mu}-\frac{1}{2}\left(\delta_{K} g_{\nu}^{\mu}-\delta_{\nu}^{\mu} \delta_{K} g_{\lambda}^{\lambda}\right) e^{2 A} \tilde{\nabla}^{2} f^{K},\end{aligned}$

$$
\begin{align*}
\delta G_{m}^{\mu}=\delta R_{m}^{\mu}= & e^{-2 A} \partial^{\mu} u^{I}\left\{2 \partial_{m} \delta_{I} A-8 \partial_{m} A \delta_{I} A-\frac{1}{2} \partial_{m} \delta_{I} \tilde{g}+\partial_{m} A \delta_{I} \tilde{g}\right. \\
& -2 \partial^{\tilde{p}} A \delta_{I} \tilde{g}_{m p}+\frac{1}{2} \tilde{\nabla}^{p} \delta_{I} \tilde{g}_{m p} \\
& -\frac{1}{2} \tilde{\nabla}^{p}\left[e^{4 A}\left(\tilde{\nabla}_{p} B_{I m}-\tilde{\nabla}_{m} B_{I p}\right)\right]+2\left(\partial_{m} A B_{I p}-\partial_{p} A B_{I m}\right) \tilde{\nabla}^{p} e^{4 A} \\
& \left.+\frac{1}{2} e^{8 A} B_{I m} \tilde{\nabla}^{2} e^{-4 A}-e^{4 A} \tilde{R}_{m}^{n} B_{I n}\right\} \tag{A.15}
\end{align*}
$$

$$
\begin{aligned}
\delta G_{n}^{m}= & u^{I} \delta_{I}\left\{e^{2 A}\left[\tilde{G}_{n}^{m}+4(\widetilde{\nabla A})^{2} \delta_{n}^{m}-8 \nabla_{n} A \tilde{\nabla}^{m} A\right]\right\}-\frac{1}{2} e^{-2 A} \square u^{I} \tilde{g}^{m k} \delta_{I} \tilde{g}_{k n} \\
& +\delta_{n}^{m} e^{-2 A} \square u^{I}\left(-2 \delta_{I} A+\frac{1}{2} \delta_{I} \tilde{g}\right) \\
& \square u^{I}\left(\frac{1}{2} e^{-2 A}\left\{\tilde{\nabla}^{m}\left[e^{4 A}\left(B_{I n}-\partial_{n} K_{I}\right)\right]+\tilde{\nabla}_{n}\left[e^{4 A}\left(B_{I}^{\tilde{m}}-\partial^{\tilde{m}} K_{I}\right)\right]\right\}\right. \\
& \left.-\delta_{n}^{m} \tilde{\nabla}^{p}\left[e^{2 A}\left(B_{I p}-\partial_{p} K_{I}\right)\right]\right) \\
& +\frac{1}{2} \delta_{K} g_{\mu}^{\mu}\left\{-\frac{1}{2} e^{-2 A}\left[\tilde{\nabla}^{m}\left(e^{4 A} \partial_{n} f^{K}\right)+\tilde{\nabla}_{n}\left(e^{4 A} \partial^{\tilde{m}} f^{K}\right)\right]+\delta_{n}^{m} \tilde{\nabla}^{p}\left[e^{2 A} \partial_{p} f^{K}\right]\right\} \\
& -\frac{1}{2} \delta_{n}^{m} f^{K} e^{-2 A} \delta_{K} R^{(4)}
\end{aligned}
$$

$$
\begin{gather*}
\delta T_{\nu}^{\mu}=-\delta_{\nu}^{\mu} \frac{1}{4 \kappa_{10}^{2}}\left\{u^{I} \delta_{I}\left[e^{-6 A}(\widetilde{\nabla \alpha})^{2}\right]-2 e^{-6 A} \square u^{I} S_{I m} \partial^{\tilde{m}} \alpha-2 \square u^{I} K_{I} e^{-6 A}(\widetilde{\nabla \alpha})^{2}\right\}  \tag{A.38}\\
\delta T_{m}^{\mu}=\frac{1}{2 \kappa_{10}^{2}} \partial^{\mu} u^{I} e^{-6 A}\left[\partial_{m} S_{I p}-\partial_{p} S_{I m}+\partial_{m} \alpha B_{I p}-\partial_{p} \alpha B_{I m}\right] \partial^{\tilde{p}} \alpha, \tag{A.37}
\end{gather*}
$$

$$
\begin{aligned}
\delta T_{n}^{m} & =-\frac{1}{2 \kappa_{10}^{2}} u^{I} \delta_{I}\left\{e^{-6 A}\left[\partial_{n} \alpha \partial^{\tilde{m}} \alpha-\frac{1}{2} \delta_{n}^{m}(\widetilde{\nabla \alpha})^{2}\right]\right\} \\
& +\frac{e^{-6 A}}{2 \kappa_{10}^{2}} \square u^{I}\left\{S_{I n} \partial^{\tilde{m}} \alpha+\partial_{n} \alpha S_{I}^{\tilde{m}}-\delta_{n}^{m} S_{I p} \partial^{\tilde{p}} \alpha+2 K_{I}\left[\partial_{n} \alpha \partial^{\tilde{m}} \alpha-\frac{1}{2} \delta_{n}^{m}(\widetilde{\nabla \alpha})^{2}\right]\right\}
\end{aligned}
$$

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## The Ubiquitous Throat



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## Cosmology

Inflation, sequestered DM,
see Langlois,
McAllister's talks


## String Theory

Moduli stabilization
AdS/CFT, ...

## The Ubiquitous Throat



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## Inflation and UV Physics


sensitive to dimension 6, Planck suppressed corrections:

$$
\delta V \sim \frac{V}{M_{P}^{2}} \phi^{2} \quad \rightleftarrows \quad \eta \equiv M_{P}^{2} \frac{V^{\prime \prime}}{V} \sim \mathcal{O}(1)
$$

such corrections may come from the Kahler potential which is not protected by holomorphy.

## BSM and UV Physics

In gravity mediation, soft terms are generated by Planck suppressed operators:

$$
m_{0} \sim m_{1 / 2} \sim m_{3 / 2} \sim \frac{<F>}{M_{P}}
$$

Issue of FCNC can only be addressed with knowledge of UV physics.

Again, the Kahler potential comes into play.

## Warping Corrections

In addition to $g_{s}$ and $\alpha^{\prime}$, yet another correction:
For example: N D3-branes

$$
\begin{aligned}
& \text { Warp Factor: } \\
& \qquad Z \equiv e^{-4 A}=1+\frac{g_{s} N \alpha^{\prime 2}}{r^{4}}
\end{aligned}
$$

Warping corrections can be important even for small

$$
g_{s} \quad \& \quad \alpha^{\prime}
$$

## Warped Closed Strings

## Question:

What is the effect of warping in string models?

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* One can quantify such effect in terms of a modified 4D EFT, including a "warped Kähler potential" Kw


## Warped Closed Strings

## Question:

> What is the effect of warping in string models?
$\%$ One can quantify such effect in terms of a modified 4D EFT, including a "warped Kähler potential" K ${ }^{\text {w }}$
$\%$ Closed string/gravity sector:
Giddings and Maharana'05
$\uparrow$ Many subtle issues
Burgess, Cámara, de Alwis, Giddings, Maharana, Quevedo'06 GS, Torroba, Underwood, Douglas'08

Douglas, Torroba'08
$\downarrow$ Simple expressions for certain subsectors (universal Kähler modulus)

## Warped Open Strings

$\because$ Open string/gauge sector of the theory: $\mathrm{K}^{\mathrm{w}}$ unexplored
$\because$ Many immediate applications to particle physics \& cosmology

D-brane Inflation


D3-moduli
Chen, Nakayama, GS

Warped Extra Dimensions


Marchesano, McGuirk, GS

## Warped Extra Dimensions



Analogous ideas for F-theory, magnetized/intersecting branes, twisted tori, etc, see talks of Heckman, Marchesano, Kobayashi, Bourjaily, Camara, Ohki,...

## Warped Open Strings



D7-branes wrap $S_{4} \subset X_{6}$

Type IIB warped background:

$$
d s_{10}^{2}=\Delta^{-1 / 2} \eta_{\mu \nu} d x^{\mu} d x^{\nu}+\Delta^{1 / 2} e^{\Phi} \hat{g}_{m n} d y^{m} d y^{n}
$$

Consistency requires:

$$
F_{5}=\left(1+*_{10}\right) F_{5}^{\text {int }} \quad F_{5}^{\text {int }}=\hat{*}_{6} d\left(\Delta e^{\Phi}\right)
$$

## Warped Open Strings



- Compute the open string wavefunctions in a warped background
$\checkmark$ Deduce the open string Kähler potential


## Warmup: Warped Flat Space

\% Introduce a probe D7-brane in this background, see how its internal fluctuations are affected by the presence of $Z$ and $F_{5}$

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Our results will be generalized later to warped Calabi-Yau and backgrounds with other SUGRA and/or worldvolume fluxes


## Warmup: Warped Flat Space

\& Introduce a probe D7-brane in this background, see how its internal fluctuations are affected by the presence of $Z$ and $F_{5}$

Our results will be generalized later to warped Calabi-Yau and backgrounds with other SUGRA and/or worldvolume fluxes


Q: How do the D7 wavefunctions couple to $\mathrm{F}_{5}$ ?

## D7-brane action

$\&$ Bosonic action $\quad S_{\mathrm{D} 7}^{\mathrm{bos}}=S_{\mathrm{D} 7}^{\mathrm{DBI}}+S_{\mathrm{D} 7}^{\mathrm{CS}}$
\&Fermionic action

Martucci, Rasseel, Van den Bleeken, Van Proeyen'05

$$
S_{\mathrm{D} 7}^{\mathrm{fer}}=\tau_{\mathrm{D} 7} \int d^{8} \xi e^{\Phi} \sqrt{|\operatorname{det} G|} \bar{\Theta} P_{-}^{\mathrm{D} 7}\left(\Gamma^{\alpha} \mathcal{D}_{\alpha}+\frac{1}{2} \mathcal{O}\right) \Theta
$$

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& S_{\mathrm{DT}}^{\text {fer }}=\tau_{\mathrm{D} 7} \int d^{8} \xi e^{\Phi} \sqrt{|\operatorname{det} G|} \bar{\Theta} P_{-}^{\mathrm{D} 7}\left(\Gamma^{\alpha} \mathcal{D}_{\alpha}+\frac{1}{2} \mathcal{O}\right) \Theta \\
& \text { Maralf, Martucci, Sikua'03 } \\
& \Theta=\binom{\theta_{1}}{\theta_{2}} \text { 10D MW bispinor (type IIB superspace) } \\
& P_{ \pm}^{\text {DT }}=\frac{1}{2}\left(\mathbb{I} \mp \Gamma_{8 D} \otimes \sigma_{2}\right) \quad \text { halves the dof's down to } \mathbb{N}=18 \mathrm{D} \text { SYM } \\
& \delta_{\epsilon} \psi_{M}=\mathcal{D}_{M} \epsilon \quad \text { type IIB gravitino variation } \\
& \delta_{\epsilon} \lambda=\mathcal{O} \epsilon \quad \text { type IIB dilatino variation }
\end{aligned}
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& \kappa \text {-symmetry } \quad \Theta \rightarrow \Theta+P_{-}^{D 7} \kappa \\
& \text { Convenient } \\
& \text { choices: } \\
& \Theta=\binom{\theta}{0} \quad \text { or } \quad P_{-}^{\mathrm{D} 7} \Theta=0
\end{aligned}
$$

## D7-brane action

\% In warped flat space:

$$
\begin{aligned}
\mathcal{O} & =0 \\
\mathcal{D}_{M} & =\nabla_{M}+\frac{1}{8} F_{5}^{\mathrm{int}} \Gamma_{M} i \sigma_{2}
\end{aligned}
$$

\% The D7-brane sees the warped metric

$$
d s_{\mathrm{D} 7}^{2}=Z^{-1 / 2} \eta_{\mu \nu} d x^{\mu} d x^{\nu}+Z^{1 / 2} \sum_{a, b=4}^{7} \delta_{a b} d y^{a} d y^{b}
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$\therefore \kappa$-fixing $\Theta=\binom{\theta}{0}$, the 8 D Dirac action is given by

$$
\begin{aligned}
S_{\mathrm{D} 7}^{\mathrm{fer}} & =\tau_{\mathrm{D} 7} e^{\Phi_{0}} \int_{\mathbb{R}^{1,3}} d^{4} x \int d^{4} y \bar{\theta} \not D^{w} \theta \\
\not D^{w} & =\sum_{\mu} \Gamma^{\mu} \mathcal{D}_{\mu}+\sum_{a} \Gamma^{a} \mathcal{D}_{a}+\frac{1}{2} \mathcal{O} \\
& =\not \ddot{y}_{4}^{\mathrm{ext}}+\not \ddot{y}_{4}^{\mathrm{int}}-\frac{1}{8}\left(\not \ddot{y}_{4}^{\mathrm{int}} \ln Z\right)\left(1+2 \Gamma_{\text {Extra }}\right)
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$$

## D7-brane zero modes

If one now decomposes the 10D MW spinor as

$$
\theta=\chi+B^{*} \chi^{*} \quad \chi=\theta_{4 D} \otimes \theta_{6 D} \quad \text { B: Majorana matrix }
$$

and performs a KK reduction, the 4D mass eigenstate eq. is

$$
\left[\not \ddot{\partial}_{4}^{\text {int }}-\frac{1}{8}\left(\ddot{\partial}_{4}^{\text {int }} \ln Z\right)\left(1+2 \Gamma_{\text {Extra }}\right)\right] \theta_{6 D}^{0}=0
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$$

and so the 4D zero modes are

$$
\begin{array}{cll}
\theta_{6 D}^{0}=Z^{-1 / 8} \eta_{-} & \text {for } & \Gamma_{\text {Extra }} \eta_{-}=-\eta_{-} \\
\theta_{6 D}^{0}=Z^{3 / 8} \eta_{+} & \text {for } & \Gamma_{\text {Extra }} \eta_{+}=\eta_{+}
\end{array}
$$

in contrast to $\theta_{6 D}^{0}=Z^{1 / 8} \eta$, the result in the absence of $F_{5}$

## D7-brane zero modes

※Upon dimensional red., such fermion zero modes

$$
\begin{aligned}
& \theta_{6 D}^{0}=Z^{-1 / 8} \eta_{-} \\
& \theta_{6 D}^{0}=Z^{3 / 8} \eta_{+}
\end{aligned}
$$

imply the following kinetic terms

$$
\begin{aligned}
S_{\mathrm{D} 7}^{\mathrm{fer}} & =\tau_{\mathrm{D} 7} e^{\Phi_{0}} \int_{\mathbb{R}^{1,3}} d^{4} x \bar{\theta}_{4 D}{\not \mathbb{R}^{1,3}} \theta_{4 D} \int d^{4} y \eta_{-}^{\dagger} \eta_{-} \\
S_{\mathrm{D} 7}^{\mathrm{fer}} & =\tau_{\mathrm{D} 7} e^{\Phi_{0}} \int_{\mathbb{R}^{1,3}} d^{4} x \bar{\theta}_{4 D} \partial_{\mathbb{R}^{1,3}} \theta_{4 D} \int d^{4} y Z \eta_{+}^{\dagger} \eta_{+}
\end{aligned}
$$

that indeed match the kinetic terms of the bosonic zero modes (e.g., $f_{D 7} \sim \int Z+i C_{4}$ ). This allows to identify them in terms of their bosonic superpartners.

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## D7-brane zero modes

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\theta_{6 D}^{0}=Z^{-1 / 8} \eta_{-} & \text {Wísonini } & A_{a} \\
\theta_{6 D}^{0}=Z^{3 / 8} \eta_{+} & \text {gaugino }+ \text { modulino } & A_{\mu},
\end{array}
$$

imply the following kinetic terms


$$
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\end{aligned}
$$

that indeed match the kinetic terms of the bosonic zero modes (e.g., fD7 $\sim \int Z+i C_{4}$ ). This allows to identify them in terms of their bosonic superpartners.

## A subtlety

٪ Our strategy to compute the zero modes appears to be:

$$
S_{\mathrm{DT}}^{\mathrm{fer}}=\int d^{8} \xi \bar{\theta} D^{w} \theta \quad \rightarrow \quad D^{w} \theta^{0}=0
$$

$\because$ For a 10D MW spinor, $\theta$ and $\bar{\theta}$ cannot be varied independently.
$\%$ For example:

$$
\tau_{\mathrm{D} 7} \int d^{8} \xi \bar{\theta} \Gamma^{\alpha} \partial_{\alpha} \theta \quad \text { and } \quad \tau_{D 7} \int d^{8} \xi \bar{\theta} \Gamma^{\alpha}\left(\partial_{\alpha}-\partial_{\alpha} \ln f\right) \theta
$$

both give $\Gamma^{\alpha} \partial_{\alpha} \theta=0$ since $\bar{\theta} \Gamma^{a_{1} \ldots a_{n}} \theta \neq 0$ only if $n=3,7$
$\because$ Naively, the warp factor dependence drops out in eom.

* But a careful analysis gives:

$$
\delta S_{\mathrm{D} 7}^{\mathrm{fer}}=\tau_{\mathrm{D} 7} e^{\Phi_{0}} \int d^{8} \xi \overline{\delta \theta} D^{w} \theta+\bar{\theta} D^{\omega} \delta \theta=2 \tau_{\mathrm{D} 7} e^{\Phi_{0}} \int d^{8} \xi \overline{\delta \theta} D^{D^{w}} \theta
$$

## A subtlety

\% Implicitly a choice of gauge is made in the D-brane fermionic action (choice of supercoord. system)

$$
P_{-}^{\mathrm{DT}}\left(\Gamma^{\alpha} \mathcal{D}_{\alpha}^{E}+\frac{1}{2} \mathcal{O}^{E}\right) \Theta=0
$$

\%The gauge choice should be consistent with the gauge choices in the bosonic sector. One can check this by dimensionally reducing the SUSY variations
ex: qauge boson $\delta_{\delta_{\epsilon} A_{\alpha=\mu}=}$


## Recap

$\%$ In general, the open string wavefunctions have an internal profile of the form

$$
\psi^{\text {int }}=Z^{p} \eta, \quad \eta=\text { const. }
$$

*Their kinetic terms group into 4D $\mathbb{N}=1$ multiplets

$$
\int_{\mathbb{R}^{1}, 3} d^{4} x \bar{\phi} D \phi \int d^{4} y Z^{q}
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$$

| D7 |  |  |
| :--- | :---: | :---: |
| 4D Field | $p$ | $q$ |
| gauge boson/modulus | 0 | 1 |
| gaugino/modulino | $3 / 8$ |  |
| Wilson line | 0 | 0 |
| Wilsonino | $-1 / 8$ |  |

## Comparison to RS

| RS |  |  | D7 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4D Field | $p$ | $q$ | 4D Field | $p$ | $q$ |
| gauge boson gaugino | $\begin{gathered} \hline 0 \\ 3 / 8 \end{gathered}$ | $1 / 4$ | gauge boson/modulus gaugino/modulino | $\begin{gathered} 0 \\ 3 / 8 \end{gathered}$ | 1 |
| matter scalar matter fermion | $\begin{gathered} (3-2 c) / 8 \\ (2-c) / 4 \end{gathered}$ | $(1-c) / 2$ | Wilson line Wilsonino | $\begin{gathered} 0 \\ -1 / 8 \end{gathered}$ | 0 |

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| :---: | :---: | :---: | :---: | :---: | :---: |
| 4D Field | $p$ | $q$ | 4D Field | $p$ | $q$ |
| gauge boson gaugino | $\begin{gathered} \hline 0 \\ 3 / 8 \\ \hline \end{gathered}$ | $1 / 4$ | gauge boson/modulus gaugino/modulino | $\begin{gathered} \hline 0 \\ 3 / 8 \end{gathered}$ | 1 |
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| :---: | :---: | :---: | :---: | :---: | :---: |
| 4D Field | $p$ | $q$ | 4D Field | $p$ | $q$ |
| gauge boson gaugino | $\begin{gathered} \hline 0 \\ 3 / 8 \\ \hline \end{gathered}$ | $1 / 4$ | gauge boson/modulus gaugino/modulino | $\begin{gathered} \hline 0 \\ 3 / 8 \end{gathered}$ | 1 |
| matter scalar matter fermion | $\begin{gathered} (3-2 c) / 8 \\ (2-c) / 4 \end{gathered}$ | $(1-c) / 2$ | Wilson line Wilsonino | $\begin{gathered} 0 \\ -1 / 8 \end{gathered}$ | 0 |



## Comparison to RS

| RS |  |  | D7 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4D Field | $p$ | $q$ | 4D Field | $p$ | $q$ |
| gauge boson gaugino | $\begin{gathered} \hline 0 \\ 3 / 8 \\ \hline \end{gathered}$ | $1 / 4$ | gauge boson/modulus gaugino/modulino | $\begin{gathered} \hline 0 \\ 3 / 8 \end{gathered}$ | 1 |
| matter scalar matter fermion | $\begin{gathered} (3-2 c) / 8 \\ (2-c) / 4 \end{gathered}$ | $(1-c) / 2$ | Wilson line Wilsonino | $\begin{gathered} 0 \\ -1 / 8 \end{gathered}$ | 0 |



## Generalizations

$\because$ The same is obtained if, instead of warped flat space, one considers a warped Calabi-Yau and a BPS D7-brane

$$
\psi^{\text {int }}=Z^{p} \eta, \quad \eta=\text { const. } \quad \rightarrow \quad \eta=\text { cov.const. }
$$



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$$
\psi^{\text {int }}=Z^{p} \eta, \quad \eta=\text { const. } \quad \rightarrow \quad \eta=\text { cov.const. }
$$

\% In addition one may also consider type IIB backgrounds with $\mathrm{G}_{3}$ fluxes, as well as with varying dilaton.


## Magnetized D7-branes

※ Finally, one can consider internally magnetized D7-branes, a necessary ingredient for 4D chirality in CY/F-theory models


6D Chiral fermion

## Magnetized D7-branes

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## Magnetized D7-branes

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$\begin{array}{ll}e^{-\Phi_{0} / 2} \mathcal{F}=B_{i} d \operatorname{vol}_{\left(\mathbf{T}^{2}\right)_{i}}+B_{j} d \operatorname{vol}_{\left(\mathbf{T}^{2}\right)_{j}} & B_{i}=e^{-\Phi_{0} / 2} Z^{-1 / 2} b_{i} \\ \text { \& Result: } & \end{array} \quad$ BPS $\Longleftrightarrow b_{i}=-b_{j}$.

$$
\begin{aligned}
& \theta_{6 D}^{0}=\frac{Z^{-1 / 8}}{1+i B_{i} \Gamma_{\mathrm{T}_{i}^{2}}^{2}} \eta_{-} \quad \begin{array}{c}
\text { Wísoniní } \\
\theta_{6 D}^{0}=Z^{3 / 8} \eta_{+}
\end{array} \quad \text { gaugino }+ \text { modulino }
\end{aligned} \longrightarrow \int d^{4} y \eta_{-}^{\dagger} \eta_{-}
$$

$$
\longrightarrow \int d^{4} y\left|Z^{1 / 2}+i e^{\Phi_{0} / 2} b\right|^{2} \eta_{+}^{\dagger} \eta_{+}
$$

## Warped EFT

$\%$ Is all this compatible with the closed string results?
\% Let us consider a D7-brane wrapping a 4-cycle $S_{4}$ in a warped Calabi-Yau, and with $\mathcal{F}=0$
\% Gauge kinetic function:

$$
f_{\mathrm{DT}}=\left(8 \pi^{3} k^{2}\right)^{-1} \int_{\mathcal{S}_{4}} \frac{\mathrm{dvol} l_{\mathcal{S}_{4}}}{\sqrt{\hat{g}_{\mathcal{S}_{4}}}}\left(Z \sqrt{\hat{g}_{\mathcal{S}_{4}}}+i C_{4}^{\mathrm{int}}\right) \quad k=2 \pi \alpha^{\prime}
$$

$\rightarrow$ Can be understood as a holomorphic function

## Warped EFT

\% Is all this compatible with the closed string results?
※Geometric moduli $\zeta^{a} \quad a=1, \ldots, h^{(0,2)}\left(\mathcal{S}_{4}\right)$
Unwarped kin. terms:

$$
\begin{gathered}
i \tau_{\mathrm{D} 7} \int_{\mathbb{R}^{1}, 3} \mathrm{e}^{\Phi} \mathcal{L}_{A \bar{B}} \mathrm{~d} \zeta^{A} \wedge *_{4} \mathrm{~d} \overline{\zeta^{\bar{B}}} \\
\mathcal{L}_{A \bar{B}}=\frac{\int_{\mathcal{S}_{4}} m_{A} \wedge m_{\bar{B}}}{\int_{X_{6}} \Omega^{\mathrm{CY}} \wedge \bar{\Omega}^{\mathrm{CY}}}
\end{gathered}
$$

Couple to the dilaton as

$$
S=t-\kappa_{A}^{2} \tau_{\mathrm{D} T} \mathcal{L}_{A \bar{B}} \zeta^{A} \bar{\zeta}^{\bar{B}} \Rightarrow \mathcal{K} \ni \ln \left[-i(S-\bar{S})-2 i \kappa_{4}^{2} \tau_{\mathrm{D} T} \mathcal{L}_{A \bar{B}} \zeta^{A} \bar{\zeta}^{\bar{B}}\right]
$$

## Warped EFT

\% Is all this compatible with the closed string results?
© Geometric moduli $\zeta^{a} \quad a=1, \ldots, h^{(0,2)}\left(\mathcal{S}_{4}\right)$
Warped kin. terms:

$$
\begin{aligned}
& i \tau_{\mathrm{D} 7} \int_{\mathbb{R}^{1,3}} \mathrm{e}^{\Phi} \mathcal{L}_{A \bar{B}}^{\mathrm{w}} \mathrm{~d} \zeta^{A} \wedge *_{4} \mathrm{~d} \bar{\zeta}^{\bar{B}} \\
& \mathcal{L}_{A \bar{B}} \rightarrow \mathcal{L}_{A \bar{B}}^{\mathrm{w}}=\frac{\int_{\mathcal{S}_{4}} Z m_{A} \wedge m_{\bar{B}}}{\int_{X_{6}} Z \Omega^{\mathrm{CY}} \wedge \bar{\Omega}^{\mathrm{CY}}}
\end{aligned}
$$



Suggest a coupling

$$
S^{\mathbf{w}}=t-\kappa_{4}^{2} \tau_{\mathrm{D} 7} \mathcal{L}_{A \bar{B}}^{\mathbf{w}} \zeta^{A} \bar{\zeta}^{\bar{B}} \Rightarrow \mathcal{K} \ni \ln \left[-i\left(S^{\mathbf{w}}-\bar{S}^{\mathbf{w}}\right)-2 i \kappa_{4}^{2} \tau_{\mathrm{D} 7} \mathcal{L}_{A \bar{B}}^{\mathbf{w}} \zeta^{A} \bar{\zeta}^{\bar{B}}\right]
$$

## Warped EFT

$\%$ Is all this compatible with the closed string results?
$\%$ Wilson line moduli $\quad W^{I} \quad I=1, \ldots, h^{(0,1)}\left(\mathcal{S}_{4}\right)$
Unwarped kin. terms:

$$
i \frac{2 \tau_{\mathrm{D} 7} k^{2}}{\mathcal{V}} \int_{\mathbb{R}^{1}, 3} \mathcal{C}_{\alpha}^{I \bar{J}^{\alpha}} v^{\alpha} \mathrm{d} w_{I} \wedge *_{4} \mathrm{~d} \bar{w}_{\bar{J}} \quad \mathcal{C}_{\alpha}^{I \bar{J}}=\int_{\mathcal{S}_{4}} P\left[\omega_{\alpha}\right] \wedge W^{I} \wedge \bar{W}^{\bar{J}} .
$$

Couple to Kähler moduli as

$$
\begin{array}{ll}
T_{\alpha}+\bar{T}_{\alpha}=\frac{3}{2} \mathcal{K}_{\alpha}+6 i \kappa_{4}^{2} \tau_{\mathrm{D} 7} k^{2} \mathcal{C}_{\alpha}^{I \bar{J}} w_{I} \bar{w}_{\bar{J}} & \mathcal{K}_{\alpha}=\mathcal{I}_{\alpha \beta \gamma} v^{\beta} v^{\gamma} \\
& \mathcal{V}=\frac{1}{6} \mathcal{I}_{\alpha \beta \gamma} v^{\alpha} v^{\beta} v^{\gamma}
\end{array}
$$

## Warped EFT

$\%$ Is all this compatible with the closed string results?
$\therefore$ Wilson line moduli $\quad W^{I} \quad I=1, \ldots, h^{(0,1)}\left(\mathcal{S}_{4}\right)$
Warped kin. terms:

$$
i \frac{2 \tau_{\mathrm{D} 7} k^{2}}{\mathcal{V}_{\mathrm{w}}} \int_{\mathbb{R}^{1}, 3} \mathcal{C}_{\alpha}^{I \bar{J}^{\alpha}} v^{\alpha} \mathrm{d} w_{I} \wedge *_{4} \mathrm{~d} \bar{w}_{\bar{J}} \quad \mathcal{C}_{\alpha}^{I \bar{J}}=\int_{\mathcal{S}_{4}} P\left[\omega_{\alpha}\right] \wedge W^{I} \wedge \bar{W}^{\bar{J}} .
$$

Suggest the following def. for "warped Kähler modulus"

$$
\begin{aligned}
& T_{\alpha}^{\mathbf{w}}+\bar{T}_{\alpha}^{\mathbf{w}}= \frac{3}{2} \mathcal{I}_{\alpha \beta \gamma}^{\mathbf{w}} \gamma^{\beta} v^{\gamma}+6 i \kappa_{4}^{2} \tau_{\mathrm{D} 7} k^{2} \mathcal{C}_{\alpha}^{I \bar{J}} w_{I} \bar{w}_{\bar{J}} \\
& \mathcal{I}_{\alpha \beta \gamma}^{\mathbf{w}}=\int_{X^{6}} Z \omega_{\alpha} \wedge \omega_{\beta} \wedge \omega_{\gamma} \Rightarrow \quad V_{\mathbf{w}}=\frac{1}{6} \mathcal{I}_{\alpha \beta \gamma}^{\mathbf{w}} v^{\alpha} v^{\beta} v^{\gamma}
\end{aligned}
$$

## Warped EFT

$\%$ Is all this compatible with the closed string results?
\% In addition, for a single Kähler modulus $\wedge$ we have that

$$
K=-3 \ln \left[T_{\Lambda}^{\mathrm{w}}+\bar{T}_{\Lambda}^{\mathrm{w}}\right] \simeq-3 \ln \frac{\mathcal{V}_{\mathrm{w}}}{v^{\Lambda}}
$$

the fluctuation of such modulus is $\quad Z(x, y)=Z_{0}(y)+c(x)$
$\Rightarrow \mathcal{V}_{\mathrm{w}}(x)=\mathcal{V}_{\mathrm{w}}^{0}+c(x) \mathcal{V}_{\mathrm{CY}} \Rightarrow K \simeq-3 \ln \left(c+\frac{\mathcal{V}_{\mathrm{w}}^{0}}{\mathcal{V}_{\mathrm{CY}}}\right)-3 \ln \frac{\mathcal{V}_{\mathrm{CY}}}{v^{\Lambda}}$

## Holographic SUSY Breaking

McGuirk, GS, Sumitomo, in progress


## Via gauge/gravity duality, analyze strong dynamics from a weak coupling gravity dual !

However, using the backreacted $\overline{\mathrm{D} 3}$ background valid in large r region:
DeWolfe, Kachru, Mullugan
one finds the leading gravity computation gives vanishing gaugino mass.

## Holographic SUSY Breaking

Deformed Conifold: $\quad \sum_{i=1}^{4} z_{i}^{2}=\epsilon^{2} \quad$ McGuirk, GS, Sumitomo, in progress
R-symmetry: $\quad z_{i} \rightarrow e^{-i \alpha} z_{i} \quad$ exact as $\epsilon \rightarrow 0$ is broken to $\mathbb{Z}_{2}$ only in the IR .

Backreacted solution in the IR sources $(0,3)+(3,0)$ besides the $(1,2)+(2,1)$ fluxes already present in the $U V$.

Using gaugino wavefunction $Z^{3 / 8} \eta_{+} \quad$ Marchesano, McGuirk, GS
Gaugino mass: $\operatorname{Tr}\left(\lambda^{2}\right)\left(G^{3}\right)^{*}$ c.f. Camara, lbanez, Uranga LI $\lambda)(G \overline{123})$ See also: Grana, Grimm, Jockers, Louis;
Lust, Reffert, Stieberger
Gravitino mass: $\quad m_{3 / 2} \sim \int \Omega \wedge G_{3}$

## Open+Closed String Fluctuations

Chen, Nakayama, GS

$$
\begin{aligned}
\mathrm{P}(g)_{\mu \nu} & =e^{2 A(Y, u)+2 \Omega(u)}\left\{\tilde{g}_{\mu \nu}(x)+2 \partial_{\mu} \partial_{\nu} u^{I}(x) \mathbf{K}_{I}(Y)+2 \mathbf{B}_{i I}(Y) \partial_{\mu} u^{I}(x) \partial_{\nu} Y^{i}\right\} \\
& +e^{-2 A(Y, u)} \tilde{g}_{i j}(Y, u) \partial_{\mu} Y^{i} \partial_{\nu} Y^{j} \cdot \text { suggests a convenient gauge B=0 }
\end{aligned}
$$

Hamiltonian constraints:

$$
\begin{aligned}
& D^{M}\left(h^{-1 / 2} \pi_{M \alpha}\right)=0, \\
& D^{M}\left(h^{-1 / 2} \pi_{M i}\right)+\kappa_{10}^{2} \delta^{(6)}(y-Y) \frac{P_{i}}{\sqrt{h}}=0 .
\end{aligned}
$$

Combined Kahler potential:

$$
\kappa_{4}^{2} \mathcal{K}(\rho, Y)=-3 \log \left[\rho+\bar{\rho}-\gamma k(Y, \bar{Y})+2 \frac{V_{W}^{0}}{V_{C Y}}\right], \quad \gamma=\frac{T_{3} \kappa_{4}^{2}}{3}
$$

where

$$
\rho=\left(c+\frac{\gamma}{2} k(Y, \bar{Y})\right)+i \chi
$$

## Summary

\% Effective action for warped compactifications is much needed in drawing precise predictions of such models.
$\because$ Many subtleties in deriving warped Kahler potential. Inclusion of open string moduli essential for several applications to warped string models of particle physics and cosmology.
$\%$ Computed open string wavefunctions in warped backgrounds, and extracted the open string wEFT. Results agree with closed string computation.
\%Combined Kahler potential involving both D3 and universal Kahler modulus.

## THANKS

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