

Open String Dynamics in Warped Backgrounds Gary Shiu University of Wisconsin

Marchesano, McGuirk, GS, 0812.2247
Chen, Nakayama, GS, 0905.4463
McGuirk, GS, Sumitomo, in progress

See also:

GS, Torroba, Underwood, Douglas, 0803.3068

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Compensators and Warping

$$\delta G^{\mu}_{\nu} = \delta^{\mu}_{\nu} u^{I} \delta_{I} \left\{ e^{2A} \left[-2\tilde{\nabla}^{2}A + 4(\widetilde{\nabla A})^{2} - \frac{1}{2}\tilde{R} \right] \right\} + e^{-2A} \left(\partial^{\mu}\partial_{\nu}u^{I} - \delta^{\mu}_{\nu}\Box u^{I} \right) \left(4\delta_{I}A - \frac{1}{2}\delta_{I}\tilde{g} \right) + \left(\partial^{\mu}\partial_{\nu}u^{I} - \delta^{\mu}_{\nu}\Box u^{I} \right) e^{2A}\tilde{\nabla}^{p} (B_{Ip} - \partial_{p}K_{I}) + e^{-2A} f^{K} \delta_{K} G^{(4)\mu}_{\nu} - \frac{1}{2} \left(\delta_{K} g^{\mu}_{\nu} - \delta^{\mu}_{\nu}\delta_{K} g^{\lambda}_{\lambda} \right) e^{2A}\tilde{\nabla}^{2} f^{K} ,$$
(A.14)

$$\delta G_m^{\mu} = \delta R_m^{\mu} = e^{-2A} \partial^{\mu} u^I \left\{ 2 \partial_m \delta_I A - 8 \partial_m A \delta_I A - \frac{1}{2} \partial_m \delta_I \tilde{g} + \partial_m A \delta_I \tilde{g} \right. \\ \left. - 2 \partial^{\tilde{p}} A \delta_I \tilde{g}_{mp} + \frac{1}{2} \tilde{\nabla}^p \delta_I \tilde{g}_{mp} \right. \\ \left. - \frac{1}{2} \tilde{\nabla}^p \left[e^{4A} \left(\tilde{\nabla}_p B_{Im} - \tilde{\nabla}_m B_{Ip} \right) \right] + 2 (\partial_m A B_{Ip} - \partial_p A B_{Im}) \tilde{\nabla}^p e^{4A} \right. \\ \left. + \frac{1}{2} e^{8A} B_{Im} \tilde{\nabla}^2 e^{-4A} - e^{4A} \tilde{R}_m^n B_{In} \right\},$$

$$(A.15)$$

$$\begin{split} \delta G_n^m = & u^I \delta_I \left\{ e^{2A} \left[\tilde{G}_n^m + 4 (\widetilde{\nabla A})^2 \delta_n^m - 8 \nabla_n A \widetilde{\nabla}^m A \right] \right\} - \frac{1}{2} e^{-2A} \Box u^I \tilde{g}^{mk} \delta_I \tilde{g}_{kn} \\ & + \delta_n^m e^{-2A} \Box u^I (-2\delta_I A + \frac{1}{2} \delta_I \tilde{g}) \\ \Box u^I \left(\frac{1}{2} e^{-2A} \left\{ \tilde{\nabla}^m \left[e^{4A} \left(B_{In} - \partial_n K_I \right) \right] + \tilde{\nabla}_n \left[e^{4A} \left(B_I^{\tilde{m}} - \partial^{\tilde{m}} K_I \right) \right] \right\} \\ & - \delta_n^m \tilde{\nabla}^p \left[e^{2A} \left(B_{Ip} - \partial_p K_I \right) \right] \right) \\ & + \frac{1}{2} \delta_K g_\mu^\mu \left\{ -\frac{1}{2} e^{-2A} \left[\tilde{\nabla}^m \left(e^{4A} \partial_n f^K \right) + \tilde{\nabla}_n \left(e^{4A} \partial^{\tilde{m}} f^K \right) \right] + \delta_n^m \tilde{\nabla}^p \left[e^{2A} \partial_p f^K \right] \right\} \\ & - \frac{1}{2} \delta_n^m f^K e^{-2A} \delta_K R^{(4)} \,. \end{split}$$

$$\delta T^{\mu}_{\nu} = -\delta^{\mu}_{\nu} \frac{1}{4\kappa_{10}^2} \left\{ u^I \delta_I \left[e^{-6A} (\widetilde{\nabla \alpha})^2 \right] - 2e^{-6A} \Box u^I S_{Im} \partial^{\tilde{m}} \alpha - 2 \Box u^I K_I e^{-6A} (\widetilde{\nabla \alpha})^2 \right\},$$

$$(A.37)$$

$$\delta T^{\mu}_m = \frac{1}{2\kappa_{10}^2} \partial^{\mu} u^I e^{-6A} \left[\partial_m S_{Ip} - \partial_p S_{Im} + \partial_m \alpha B_{Ip} - \partial_p \alpha B_{Im} \right] \partial^{\tilde{p}} \alpha ,$$

$$(A.38)$$

$$\delta G_N^M = \kappa_{10}^2 \delta T_N^M$$

$$\delta T_n^m = -\frac{1}{2\kappa_{10}^2} u^I \delta_I \left\{ e^{-6A} \left[\partial_n \alpha \partial^{\tilde{m}} \alpha - \frac{1}{2} \delta_n^m (\widetilde{\nabla \alpha})^2 \right] \right\} + \frac{e^{-6A}}{2\kappa_{10}^2} \Box u^I \left\{ S_{In} \partial^{\tilde{m}} \alpha + \partial_n \alpha S_I^{\tilde{m}} - \delta_n^m S_{Ip} \partial^{\tilde{p}} \alpha + 2K_I \left[\partial_n \alpha \partial^{\tilde{m}} \alpha - \frac{1}{2} \delta_n^m (\widetilde{\nabla \alpha})^2 \right] \right\} .$$
(A.39)

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Inflation and UV Physics



sensitive to dimension 6, Planck suppressed corrections:

$$\delta V \sim \frac{V}{M_P^2} \phi^2 \quad \blacksquare \quad \forall \eta \equiv M_P^2 \frac{V''}{V} \sim \mathcal{O}(1)$$

such corrections may come from the Kahler potential which is not protected by holomorphy.

BSM and UV Physics

In gravity mediation, soft terms are generated by Planck suppressed operators:

$$m_0 \sim m_{1/2} \sim m_{3/2} \sim \frac{\langle F \rangle}{M_P}$$

Issue of FCNC can only be addressed with knowledge of UV physics.

Again, the Kahler potential comes into play.

Warping Corrections

In addition to g_s and α' , yet another correction:

For example: N D3-branes



Warp Factor:

$$Z \equiv e^{-4A} = 1 + \frac{g_s N \alpha'^2}{r^4}$$

Warping corrections can be important even for small

$$g_s$$
 & α'

Warped Closed Strings

Question:

What is the effect of warping in string models?

Warped Closed Strings

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One can quantify such effect in terms of a modified 4D EFT, including a "warped Kähler potential" K^w

Warped Closed Strings

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 Closed string/gravity sector: Giddings and Maharana'05 Burgess, Cámara, de Alwis, Giddings, Maharana, Quevedo'06
 Many subtle issues GS, Torroba, Underwood, Douglas'08 Douglas, Torroba'08
 Simple expressions for certain subsectors (universal Kähler modulus)

Frey, Torroba, Underwood, Douglas'08 Chen, Nakayama, GS '09

Warped Open Strings

Open string/gauge sector of the theory: K^w unexplored

Many immediate applications to particle physics & cosmology

D-brane Inflation







Chen, Nakayama, GS

D3-moduli

Warped Extra Dimensions



Analogous ideas for F-theory, magnetized/intersecting branes, twisted tori, etc, see talks of Heckman, Marchesano, Kobayashi, Bourjaily, Camara, Ohki,...

Warped Open Strings



Type IIB warped background:

$$ds_{10}^2 = \Delta^{-1/2} \eta_{\mu\nu} dx^{\mu} dx^{\nu} + \Delta^{1/2} e^{\Phi} \hat{g}_{mn} dy^m dy^n$$

Consistency requires:

$$F_5 = (1 + *_{10})F_5^{\text{int}} \qquad F_5^{\text{int}} = \hat{*}_6 d(\Delta e^{\Phi})$$

Warped Open Strings



Compute the open string wavefunctions in a warped background
 Deduce the open string Kähler potential

Warmup: Warped Flat Space

Introduce a probe D7-brane in this background, see how its internal fluctuations are affected by the presence of Z and F5

 D_3

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Our results will be generalized later to warped Calabi-Yau and backgrounds with other SUGRA and/or worldvolume fluxes

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Q: How do the D7 wavefunctions couple to F_5 ?

• Bosonic action $S_{D7}^{bos} = S_{D7}^{DBI} + S_{D7}^{CS}$

Fermionic action

Martucci, Rosseel, Van den Bleeken, Van Proeyen'05

$$S_{\mathrm{D7}}^{\mathrm{fer}} = \tau_{\mathrm{D7}} \int d^{8}\xi \, e^{\Phi} \sqrt{|\det G|} \, \bar{\Theta} P_{-}^{\mathrm{D7}} \left(\Gamma^{\alpha} \mathcal{D}_{\alpha} + \frac{1}{2} \mathcal{O} \right) \Theta_{\text{see also Graña 'O2}}$$

Marolf, Martucci, Silva 'O3

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$$\Theta = \begin{pmatrix} \theta_{1} \\ \theta_{2} \end{pmatrix} \quad \text{10D MW bispinor (type IIB superspace)}$$

 $P_{\pm}^{D7} = \frac{1}{2} (\mathbb{I} \mp \Gamma_{8D} \otimes \sigma_2)$ halves the dof's down to $\mathcal{N}=1$ 8D SYM

 $\begin{array}{lll} \delta_{\epsilon}\psi_{M} &=& \mathcal{D}_{M}\epsilon & \mbox{type IIB gravitino variation} \\ \delta_{\epsilon}\lambda &=& \mathcal{O}\epsilon & \mbox{type IIB dilatino variation} \end{array} (contain F_{p}) \end{array}$

Contains the coupling of fermions to RR fluxes

 $S_{\rm D7}^{\rm bos} = S_{\rm D7}^{\rm DBI} + S_{\rm D7}^{\rm CS}$ Bosonic action

Fermionic action

Martucci, Rosseel, Van den Bleeken, Van Proeyen'05

$$\begin{split} S_{\mathrm{D7}}^{\,\mathrm{fer}} &= \tau_{\mathrm{D7}} \int d^8 \xi \, e^{\Phi} \sqrt{|\det G|} \, \bar{\Theta} \, P_{-}^{\mathrm{D7}} \left(\Gamma^{\alpha} \mathcal{D}_{\alpha} + \frac{1}{2} \mathcal{O} \right) \underbrace{\Theta}_{\substack{\text{see also Graña '02} \\ \mathcal{M}arolf, \, \mathcal{M}artucci, \, \mathcal{S}ilva '03}}_{\mathcal{M}arolf, \, \mathcal{M}artucci, \, \mathcal{S}ilva '03} \\ \Theta &= \begin{pmatrix} \theta_1 \\ \theta_2 \end{pmatrix} \quad \text{10D MW bispinor (type IIB superspace)} \\ P_{\pm}^{\mathrm{D7}} &= \frac{1}{2} \left(\mathbb{I} \mp \Gamma_{8D} \otimes \sigma_2 \right) \quad \text{halves the dof's down to } \mathcal{N}=1 \text{ 8D SYM} \end{split}$$

 κ -symmetry $\Theta \rightarrow \Theta + P^{D7}\kappa$

Convenient choices:

$$\Theta = \begin{pmatrix} \theta \\ 0 \end{pmatrix} \qquad \text{or}$$

$$P_{-}^{\mathrm{D7}}\Theta = 0$$





If one now decomposes the 10D MW spinor as

 $\theta = \chi + B^* \chi^*$ $\chi = \theta_{4D} \otimes \theta_{6D}$ B: Majorana matrix

and performs a KK reduction, the 4D mass eigenstate eq. is

$$\left[\partial_{4}^{\text{int}} - \frac{1}{8} \left(\partial_{4}^{\text{int}} \ln Z\right) \left(1 + 2\Gamma_{\text{Extra}}\right)\right] \theta_{6D}^{0} = 0$$

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and so the 4D zero modes are

$$\begin{aligned} \theta_{6D}^0 &= Z^{-1/8} \eta_- & \text{for} & \Gamma_{\text{Extra}} \eta_- = -\eta_- \\ \theta_{6D}^0 &= Z^{3/8} \eta_+ & \text{for} & \Gamma_{\text{Extra}} \eta_+ = \eta_+ \end{aligned}$$

in contrast to $\theta_{6D}^0 = Z^{1/8}\eta$, the result in the absence of F₅

Acharya, Benini, Valandro'06

Upon dimensional red., such fermion zero modes

$$\theta_{6D}^0 = Z^{-1/8} \eta_-$$
$$\theta_{6D}^0 = Z^{3/8} \eta_+$$

imply the following kinetic terms

$$S_{\mathrm{D7}}^{\mathrm{fer}} = \tau_{\mathrm{D7}} e^{\Phi_0} \int_{\mathbb{R}^{1,3}} d^4 x \,\bar{\theta}_{4D} \partial_{\mathbb{R}^{1,3}} \theta_{4D} \int d^4 y \,\eta_-^{\dagger} \eta_-$$

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that indeed match the kinetic terms of the bosonic zero modes (e.g., $f_{D7} \sim \int Z + i C_4$). This allows to identify them in terms of their bosonic superpartners.

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$$\begin{aligned} \theta_{6D}^{0} &= Z^{-1/8} \eta_{-} & \text{Wilsonini} \quad A_{a} \\ \theta_{6D}^{0} &= Z^{3/8} \eta_{+} & \text{gaugino} + \text{modulino} \quad A_{\mu} , \end{aligned}$$

$$\begin{aligned} \text{imply the following kinetic terms} \\ S_{\text{D7}}^{\text{fer}} &= \tau_{\text{D7}} e^{\Phi_{0}} \int_{\mathbb{R}^{1,3}} d^{4}x \, \bar{\theta}_{4D} \partial_{\mathbb{R}^{1,3}} \theta_{4D} \int d^{4}y \, \eta_{-}^{\dagger} \eta_{-} \end{aligned} \\ \begin{aligned} \text{vsual Volume} \\ S_{\text{D7}}^{\text{fer}} &= \tau_{\text{D7}} e^{\Phi_{0}} \int_{\mathbb{R}^{1,3}} d^{4}x \, \bar{\theta}_{4D} \partial_{\mathbb{R}^{1,3}} \theta_{4D} \int d^{4}y \, \eta_{-}^{\dagger} \eta_{-} \end{aligned} \\ \end{aligned}$$

that indeed match the kinetic terms of the bosonic zero modes (e.g., $f_{D7} \sim \int Z + i C_4$). This allows to identify them in terms of their bosonic superpartners.

A subtlety

Our strategy to compute the zero modes appears to be:

$$S_{\mathrm{D7}}^{\mathrm{fer}} = \int d^8 \xi \,\bar{\theta} D^w \theta \longrightarrow D^w \theta^0 = 0$$

• For a 10D MW spinor, θ and $\overline{\theta}$ cannot be varied independently.

For example:

 $au_{\mathrm{D7}} \int d^8 \xi \, \bar{\theta} \Gamma^{\alpha} \partial_{\alpha} \theta \quad \text{and} \quad au_{\mathrm{D7}} \int d^8 \xi \, \bar{\theta} \Gamma^{\alpha} \big(\partial_{\alpha} - \partial_{\alpha} \ln f \big) \theta$ both give $\Gamma^{\alpha} \partial_{\alpha} \theta = 0$ since $\overline{\theta} \Gamma^{a_1 \dots a_n} \theta \neq 0$ only if n = 3, 7

Naively, the warp factor dependence drops out in eom.

But a careful analysis gives:

$$\delta S_{\mathrm{D7}}^{\mathrm{fer}} = \tau_{\mathrm{D7}} e^{\Phi_0} \int d^8 \xi \,\overline{\delta\theta} \, D^w \theta + \bar{\theta} \, D^w \delta\theta = 2\tau_{\mathrm{D7}} e^{\Phi_0} \int d^8 \xi \,\overline{\delta\theta} \, D^w \theta$$

A subtlety

Implicitly a choice of gauge is made in the D-brane fermionic action (choice of supercoord. system)

$$P_{-}^{\mathrm{D7}}\left(\Gamma^{\alpha}\mathcal{D}_{\alpha}^{E}+\frac{1}{2}\mathcal{O}^{E}\right)\Theta=0$$

The gauge choice should be consistent with the gauge choices in the bosonic sector. One can check this by dimensionally reducing the SUSY variations

ex: gauge boson 8D 4D

$$\delta_{\epsilon}A_{\alpha=\mu} = \overline{\epsilon}\Gamma_{\alpha=\mu}\theta \rightarrow \delta_{\epsilon}A_{\mu} = \overline{\epsilon}\lambda$$

 $\uparrow \uparrow \uparrow \uparrow$
 $Z^{0} = Z^{-\frac{1}{8} - \frac{1}{4} + \frac{3}{8}}$

see Bandos and Sorokin'06

Recap

In general, the open string wavefunctions have an internal profile of the form

$$\psi^{\text{int}} = Z^p \eta, \qquad \eta = \text{const.}$$

✤ Their kinetic terms group into 4D 𝒴=1 multiplets $\int_{\mathbb{R}^{1,3}} d^4x \, \bar{\phi} D\phi \int d^4y Z^q$

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Their kinetic terms group into 4D N=1 multiplets

$$\int_{\mathbb{R}^{1,3}} d^4x \,\bar{\phi} D\phi \int d^4y Z^{\mathbf{q}}$$

$\mathrm{D7}$		
4D Field	p	q
gauge boson/modulus	0	1
gaugino/modulino	3/8	T
Wilson line	0	0
Wilsonino	-1/8	U

\mathbf{RS}		$\mathrm{D7}$			
4D Field	p	q	4D Field	p	q
gauge boson	0	1 / 1	gauge boson/modulus	0	1
gaugino	3/8	1/4	gaugino/modulino	3/8	1
matter scalar	(3-2c)/8	(1 a)/9	Wilson line	0	0
matter fermion	(2-c)/4	(1-c)/2	Wilsonino	-1/8	U

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- ✤ 5-form fluxes
- **3**-form fluxes

D7-brane worldvolume flux

Generalizations

The same is obtained if, instead of warped flat space, one considers a warped Calabi-Yau and a BPS D7-brane

 $\psi^{\text{int}} = Z^p \eta, \qquad \eta = \text{const.} \rightarrow \eta = \text{cov.const.}$



Generalizations

The same is obtained if, instead of warped flat space, one considers a warped Calabi-Yau and a BPS D7-brane

 $\psi^{\text{int}} = Z^p \eta, \quad \eta = \text{const.} \quad \to \quad \eta = \text{cov.const.}$

In addition one may also consider type IIB backgrounds with G₃ fluxes, as well as with varying dilaton.



Magnetized D7-branes

Finally, one can consider internally magnetized D7-branes, a necessary ingredient for 4D chirality in CY/F-theory models



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$$\begin{array}{rcl} & & & & & & & & \\ \hline & & & & & \\ e^{-\Phi_0/2}\mathcal{F} & = & B_i \, d\mathrm{vol}_{(\mathbf{T}^2)_i} + B_j \, d\mathrm{vol}_{(\mathbf{T}^2)_j} & & & \\ B_i & = e^{-\Phi_0/2} Z^{-1/2} b_i \\ & & & \\ BPS & \Leftrightarrow b_i = -b_j \end{array} \\ & & & \\ \theta^0_{6D} & = \frac{Z^{-1/8}}{1+iB_i\Gamma_{\mathbf{T}_i^2}} \eta_- & & \\ \theta^0_{6D} & = Z^{3/8} \eta_+ & & gaugino + modulino \\ & & \longrightarrow \int d^4 y \, |Z^{1/2} + i e^{\Phi_0/2} b|^2 \eta_+^\dagger \eta_+ \end{array}$$

Is all this compatible with the closed string results?

• Let us consider a D7-brane wrapping a 4-cycle S_4 in a warped Calabi-Yau, and with $\mathcal{F} = 0$

Gauge kinetic function:

$$f_{\rm D7} = \left(8\pi^3 k^2\right)^{-1} \int_{\mathcal{S}_4} \frac{\hat{\rm dvol}_{\mathcal{S}_4}}{\sqrt{\hat{g}}_{\mathcal{S}_4}} \left(Z\sqrt{\hat{g}}_{\mathcal{S}_4} + iC_4^{\rm int}\right) \qquad k = 2\pi\alpha'$$

→ Can be understood as a holomorphic function

Baumann, Dymarsky, Klebanov, Maldacena, McAllister, Murugan'06

Is all this compatible with the closed string results?

Geometric moduli
$$\zeta^{a}$$
 $a = 1, ..., h^{(0,2)}(S_{4})$
Unwarped kin. terms:
 $i\tau_{D7} \int_{\mathbb{R}^{1,3}} e^{\Phi} \mathcal{L}_{A\bar{B}} d\zeta^{A} \wedge *_{4} d\bar{\zeta}^{\bar{B}}$
 $\mathcal{L}_{A\bar{B}} = \frac{\int_{S_{4}} m_{A} \wedge m_{\bar{B}}}{\int_{X_{6}} \Omega^{CY} \wedge \bar{\Omega}^{CY}}$

Couple to the dilaton as

•

$$S = t - \kappa_4^2 \tau_{\mathrm{D7}} \mathcal{L}_{A\bar{B}} \zeta^A \bar{\zeta}^{\bar{B}} \implies \mathcal{K} \ni \ln\left[-i\left(S - \bar{S}\right) - 2i\kappa_4^2 \tau_{\mathrm{D7}} \mathcal{L}_{A\bar{B}} \zeta^A \bar{\zeta}^{\bar{B}}\right]$$

see, e.g., Jockers and Louis'04

Is all this compatible with the closed string results?

Suggest a coupling

 $S^{\mathbf{W}} = t - \kappa_4^2 \tau_{\mathrm{D7}} \mathcal{L}_{A\bar{B}}^{\mathbf{W}} \zeta^A \bar{\zeta}^{\bar{B}} \implies \mathcal{K} \ni \ln\left[-i\left(S^{\mathbf{W}} - \bar{S}^{\mathbf{W}}\right) - 2i\kappa_4^2 \tau_{\mathrm{D7}} \mathcal{L}_{A\bar{B}}^{\mathbf{W}} \zeta^A \bar{\zeta}^{\bar{B}}\right]$

compatible with Shiu, Torroba, Underwood, Douglas'08

Is all this compatible with the closed string results?

• Wilson line moduli W^I $I = 1, \ldots, h^{(0,1)}(\mathcal{S}_4)$

Unwarped kin. terms:

$$i\frac{2\tau_{\mathrm{D7}}k^{2}}{\mathcal{V}}\int_{\mathbb{R}^{1,3}}\mathcal{C}_{\alpha}^{I\bar{J}}v^{\alpha}\mathrm{d}w_{I}\wedge\ast_{4}\mathrm{d}\overline{w}_{\bar{J}}\qquad\qquad\mathcal{C}_{\alpha}^{I\bar{J}}=\int_{\mathcal{S}_{4}}P\left[\omega_{\alpha}\right]\wedge W^{I}\wedge\overline{W}^{\bar{J}}$$
$$J^{\mathrm{CY}}=v^{\alpha}\omega_{\alpha}$$

Couple to Kähler moduli as

$$T_{\alpha} + \overline{T}_{\alpha} = \frac{3}{2} \mathcal{K}_{\alpha} + 6i\kappa_{4}^{2} \tau_{\mathrm{D7}} k^{2} \mathcal{C}_{\alpha}^{I\bar{J}} w_{I} \overline{w}_{\bar{J}} \qquad \qquad \mathcal{K}_{\alpha} = \mathcal{I}_{\alpha\beta\gamma} v^{\beta} v^{\gamma}$$
$$\mathcal{V} = \frac{1}{6} \mathcal{I}_{\alpha\beta\gamma} v^{\alpha} v^{\beta} v^{\gamma}$$

see again Jockers and Louis'04

Is all this compatible with the closed string results?

• Wilson line moduli W^I $I = 1, \ldots, h^{(0,1)}(\mathcal{S}_4)$

Warped kin. terms:

$$i\frac{2\tau_{\mathrm{D7}}k^{2}}{\mathcal{V}_{\mathbf{W}}}\int_{\mathbb{R}^{1,3}}\mathcal{C}_{\alpha}^{I\bar{J}}v^{\alpha}\mathrm{d}w_{I}\wedge\ast_{4}\mathrm{d}\overline{w}_{\bar{J}}\qquad\qquad\mathcal{C}_{\alpha}^{I\bar{J}}=\int_{\mathcal{S}_{4}}P\left[\omega_{\alpha}\right]\wedge W^{I}\wedge\overline{W}^{\bar{J}}$$
$$J^{\mathrm{CY}}=v^{\alpha}\omega_{\alpha}$$

Suggest the following def. for "warped Kähler modulus"

$$T^{\mathbf{w}}_{\alpha} + \overline{T}^{\mathbf{w}}_{\alpha} = \frac{3}{2} \mathcal{I}^{\mathbf{w}}_{\alpha\beta\gamma} v^{\beta} v^{\gamma} + 6i\kappa_{4}^{2} \tau_{\mathrm{D7}} k^{2} \mathcal{C}^{I\bar{J}}_{\alpha} w_{I} \overline{w}_{\bar{J}}$$
$$\mathcal{I}^{\mathbf{w}}_{\alpha\beta\gamma} = \int_{X^{6}} \mathbf{Z} \,\omega_{\alpha} \wedge \omega_{\beta} \wedge \omega_{\gamma} \quad \Rightarrow \quad \mathcal{V}_{\mathbf{w}} = \frac{1}{6} \mathcal{I}^{\mathbf{w}}_{\alpha\beta\gamma} v^{\alpha} v^{\beta} v^{\gamma}$$

Is all this compatible with the closed string results?

• In addition, for a single Kähler modulus Λ we have that

$$K = -3\ln\left[T_{\Lambda}^{\rm w} + \bar{T}_{\Lambda}^{\rm w}\right] \simeq -3\ln\frac{\mathcal{V}_{\rm w}}{v^{\Lambda}}$$

the fluctuation of such modulus is $Z(x,y) = Z_0(y) + c(x)$

Giddings and Maharana'05

$$\Rightarrow \mathcal{V}_{w}(x) = \mathcal{V}_{w}^{0} + c(x)\mathcal{V}_{CY} \Rightarrow K \simeq -3\ln\left(c + \frac{\mathcal{V}_{w}^{0}}{\mathcal{V}_{CY}}\right) - 3\ln\frac{\mathcal{V}_{CY}}{v^{\Lambda}}$$

reproduces Frey, Torroba, Underwood, Douglas'08 Chen, Nakayama, Shiu, 09

Holographic SUSY Breaking

McGuirk, GS, Sumitomo, in progress



Via gauge/gravity duality, analyze strong dynamics from a weak coupling gravity dual !

However, using the backreacted D3 background valid in large r region: DeWolfe, Kachru, Mullugan

one finds the leading gravity computation gives vanishing gaugino mass.

Benini, Dymarsky, Franco, Kachru, Simic, Verlinde

Holographic SUSY Breaking

Deformed Conifold:

A McGuirk, GS, Sumitomo, in progress
$$\sum_{i=1}^{4} z_i^2 = \epsilon^2$$

R-symmetry: $z_i \to e^{-i\alpha} z_i$ exact as $\epsilon \to 0$

is broken to \mathbb{Z}_2 only in the IR.

Backreacted solution in the IR sources (0,3)+(3,0) besides the (1,2)+(2,1) fluxes already present in the UV.

Using gaugino wavefunction $Z^{3/8}\eta_+$

Marchesano, McGuirk, GS

Gaugino mass: $\operatorname{Tr}(\lambda^2) \left(G_{\overline{123}}^3\right)^*$ c.f. Camara, Ibanez, Uranga See also: Grana, Grimm, Jockers, Louis; Lust, Reffert, Stieberger

Gravitino mass:

$$m_{3/2} \sim \int \Omega \wedge G_3$$

Open+Closed String Fluctuations

Chen, Nakayama, GS

$$P(g)_{\mu\nu} = e^{2A(Y,u)+2\Omega(u)} \left\{ \tilde{g}_{\mu\nu}(x) + 2\partial_{\mu}\partial_{\nu}u^{I}(x)\mathbf{K}_{I}(Y) + 2\mathbf{B}_{iI}(Y)\partial_{\mu}u^{I}(x)\partial_{\nu}Y^{i} \right\} \\ + e^{-2A(Y,u)}\tilde{g}_{ij}(Y,u)\partial_{\mu}Y^{i}\partial_{\nu}Y^{j} .$$
 suggests a convenient gauge B=0

Hamiltonian constraints:

$$D^{M} \left(h^{-1/2} \pi_{M\alpha} \right) = 0,$$

$$D^{M} \left(h^{-1/2} \pi_{Mi} \right) + \kappa_{10}^{2} \delta^{(6)} (y - Y) \frac{P_{i}}{\sqrt{h}} = 0.$$

Combined Kahler potential:

$$\begin{split} \kappa_4^2 \mathcal{K}(\rho,Y) &= -3\log\left[\rho + \bar{\rho} - \gamma k(Y,\overline{Y}) + 2\frac{V_W^0}{V_{CY}}\right], \quad \gamma = \frac{T_3\kappa_4^2}{3}, \\ \text{where} \qquad \rho = \begin{pmatrix} c + \frac{\gamma}{2}k(Y,\overline{Y}) \end{pmatrix} + i\chi \,. \\ & \swarrow \\ & \swarrow \\ \text{"breathing mode"} \qquad \text{axion} \end{split}$$

Summary

- Effective action for warped compactifications is much needed in drawing precise predictions of such models.
- Many subtleties in deriving warped Kahler potential. Inclusion of open string moduli essential for several applications to warped string models of particle physics and cosmology.
- Computed open string wavefunctions in warped backgrounds, and extracted the open string wEFT. Results agree with closed string computation.
- Combined Kahler potential involving both D3 and universal Kahler modulus.

THANKS

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