

Searching for Dark Matter at Future e^+e^- Colliders

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- ◇ BG, M.Iglicki, K.Mekala and A. F.Zarnecki, "Searching for Dark Matter at Future e^+e^- Colliders", in progress,
 - ◇ D.Azevedo, M.Duch, BG, D.Huang, M.Iglicki, R.Santos, "One-loop contribution to dark matter-nucleon scattering in the pseudoscalar dark matter model", JHEP 1901 (2019) 138,
 - ◇ D.Azevedo, M.Duch, BG, D.Huang, M.Iglicki, R.Santos, "Testing scalar versus vector dark matter", Phys.Rev. D99 (2019) no.1, 015017,
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Pseudo-Goldstone Dark Matter (pGDM)

- J. Cline, T. Toma, “Pseudo-Goldstone dark matter confronts cosmic ray and collider anomalies”, arXiv:1906.02175,
- D. Karamitros, “Pseudo Nambu-Goldstone Dark Matter: Examples of Vanishing Direct Detection Cross Section”, Phys.Rev. D99 (2019) no.9, 095036,
- K. Kannike, M. Raidal, “Phase Transitions and Gravitational Wave Tests of Pseudo-Goldstone Dark Matter in the Softly Broken $U(1)$ Scalar Singlet Model”, Phys.Rev. D99 (2019) no.11, 115010,
- T. Alanne, M. Heikinheimo, V. Keus, N. Koivunen, K. Tuominen, “Direct and indirect probes of Goldstone dark matter”, Phys.Rev. D99 (2019) no.7, 075028,
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- K. Ghorbani, P. Hossein Ghorbani, “Leading Loop Effects in Pseudoscalar-Higgs Portal Dark Matter”, JHEP 1905 (2019) 096,
- K. Ishiwata, T. Toma, “Probing pseudo Nambu-Goldstone boson dark matter at loop level”, JHEP 1812 (2018) 089,
- D. Azevedo, M. Duch, BG, D. Huang, M. Iglicki, R. Santos, “One-loop contribution to dark matter-nucleon scattering in the pseudoscalar dark matter model”, JHEP 1901 (2019) 138,
- D. Azevedo, M. Duch, BG, D. Huang, M. Iglicki, R. Santos, “Testing scalar versus vector dark matter”, Phys.Rev. D99 (2019) no.1, 015017,
- C. Gross, O. Lebedev, and T. Toma, “Cancellation Mechanism for Dark-Matter–Nucleon Interaction”, Phys. Rev. Lett. 119 (2017), no. 19 191801.

$$S = \frac{1}{\sqrt{2}}(v_S + \phi + iA) \quad , \quad \text{and} \quad H = \begin{pmatrix} \pi^+ \\ \frac{1}{\sqrt{2}}(v + h + i\pi^0) \end{pmatrix}.$$

$$V_{\text{pGDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2 + (\mu^2 S^2 + \text{H.c.})$$

Positivity: $\lambda_H > 0$, $\lambda_S > 0$, $\kappa > -2\sqrt{\lambda_H \lambda_S}$

Symmetries:

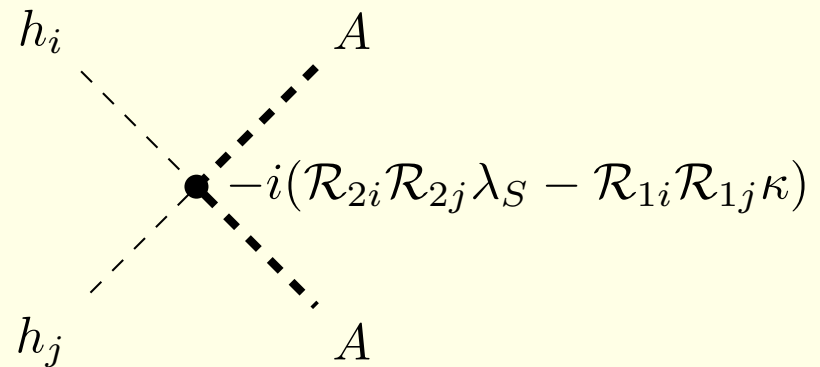
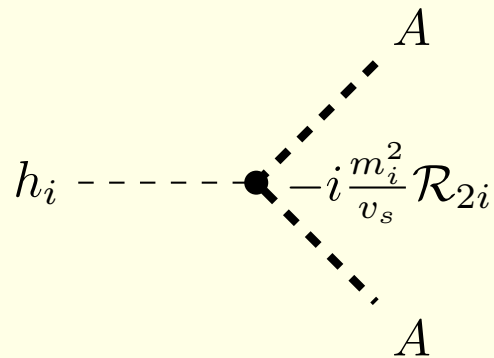
- $\mu^2 \neq 0$ breaks global $U(1)$ softly to residual $Z_2 : S \rightarrow -S$,
- rephase S such that $\text{Im } \mu^2 = 0$ (basis choice),
- $V \ni \mu^2(S^2 + S^{*2})$, then symmetry $S \xrightarrow{C} S^*$ ($\phi \rightarrow \phi$ and $A \rightarrow -A$) reveals itself,
- global minimum at $\langle S \rangle = \frac{v_S}{\sqrt{2}}$ with v_S being real, so C is unbroken and A is a stable DM candidate, $m_A^2 \propto \mu^2$, if $\mu^2 \rightarrow 0$ then A becomes a Goldstone boson of broken $U(1)$,

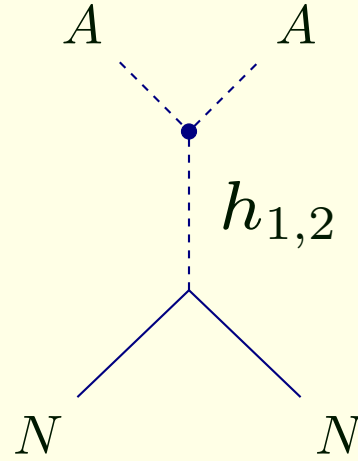
$$S = \frac{1}{\sqrt{2}}(v_S + \phi + iA) \quad , \quad \text{and} \quad H = \begin{pmatrix} \pi^+ \\ \frac{1}{\sqrt{2}}(v + h + i\pi^0) \end{pmatrix}.$$

$$\mathcal{M}^2 = \begin{pmatrix} 2\lambda_H v^2 & \kappa v v_S & 0 \\ \kappa v v_S & 2\lambda_S v_S^2 & 0 \\ 0 & 0 & -4\mu^2 \end{pmatrix}$$

$$\mathcal{M}_{\text{diag}}^2 = \begin{pmatrix} m_1^2 & 0 & 0 \\ 0 & m_2^2 & 0 \\ 0 & 0 & m_A^2 \end{pmatrix}, \quad R = \begin{pmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{pmatrix}, \quad \begin{pmatrix} h_1 \\ h_2 \end{pmatrix} = R^{-1} \begin{pmatrix} h \\ \phi \end{pmatrix},$$

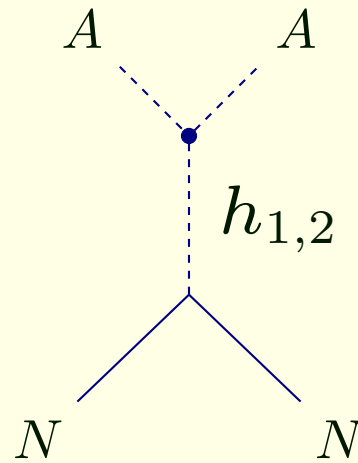
$$V_{\text{pGDM}} \ni \lambda_S |S|^4 + \kappa |S|^2 |H|^2 \ni \frac{A^2}{2} (2\lambda_S v_S \phi + \kappa v \phi) = \frac{A^2}{2v_S} (\sin \alpha \, \textcolor{red}{m}_1^2 h_1 + \cos \alpha \, \textcolor{red}{m}_2^2 h_2)$$





The DM direct detection signals are naturally suppressed in the pGDM model:

$$\begin{aligned}
 i\mathcal{M} &= -i \frac{\sin \alpha \cos \alpha f_N m_N}{vv_S} \left(\frac{m_1^2}{q^2 - m_1^2} - \frac{m_2^2}{q^2 - m_2^2} \right) \bar{u}_N(p_4) u_N(p_2) \\
 &\approx -i \frac{\sin \alpha \cos \alpha f_N m_N}{vv_S} \left(\frac{m_1^2 - m_2^2}{m_1^2 m_2^2} \right) q^2 \bar{u}_N(p_4) u_N(p_2) .
 \end{aligned}$$



The total cross section σ_{AN} :

$$\sigma_{AN}^{(\text{tree})} \propto \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 \times v_A^4,$$

where v_A is the A velocity in the lab frame. Since $v_A \sim 200$ km/s, the total DM nuclear recoil cross section σ_{AN} is greatly suppressed by the factor $v_A^4 \sim 10^{-13}$:

$$\sigma_{AN}^{(\text{tree})} \sim 10^{-70} \text{ cm}^2 \ll \sigma_{AN}^{(\text{XENON1T})} \sim 10^{-46} \text{ cm}^2$$



1-loop effects are leading

- if $q^2 \rightarrow 0$ then loop corrections are expected to be UV finite,
- if $m_A^2 \propto \mu^2 \rightarrow 0$ then loop corrections should vanish.

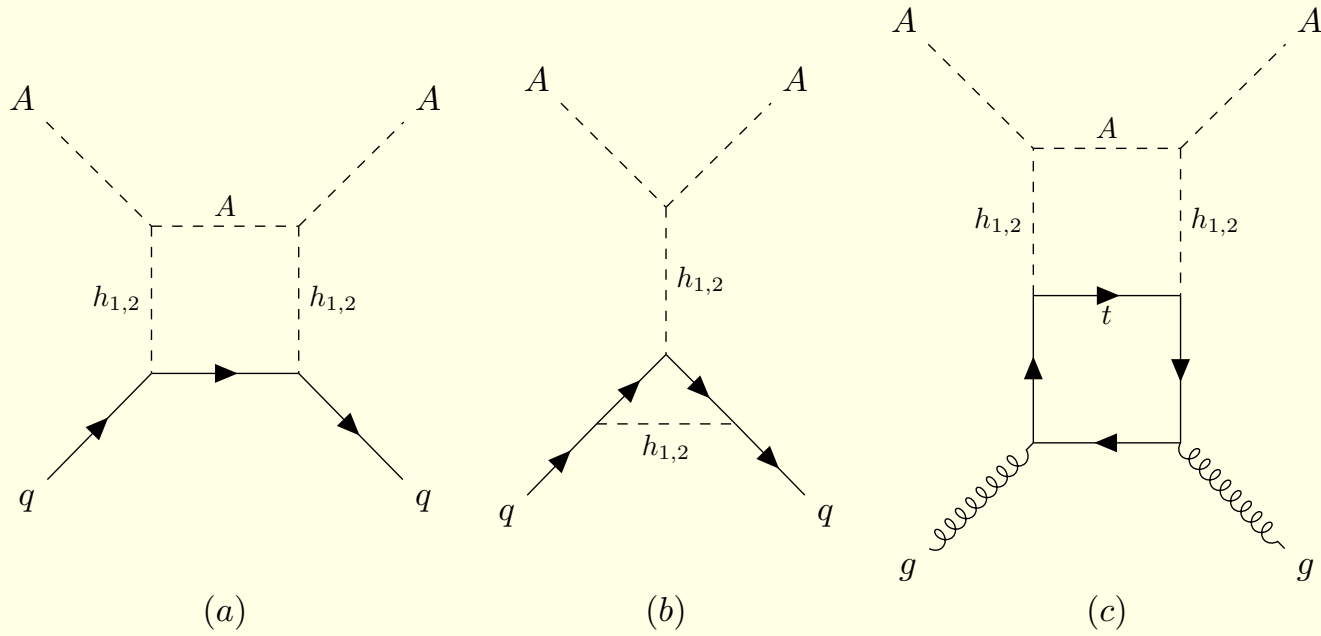


Figure 1: Examples of diagrams contributing to DM-nucleon scattering, which are discarded in our computation. Diagrams (a) and (b) represent the one-loop box and light-quark- $h_{1,2}$ vertex corrected diagrams which are ignored due to the multiple Yukawa coupling suppression, while the diagram (c) is an example of DM-gluon scattering with two Higgs lines inserted into the top-quark loop, which is assumed to be subdominant.

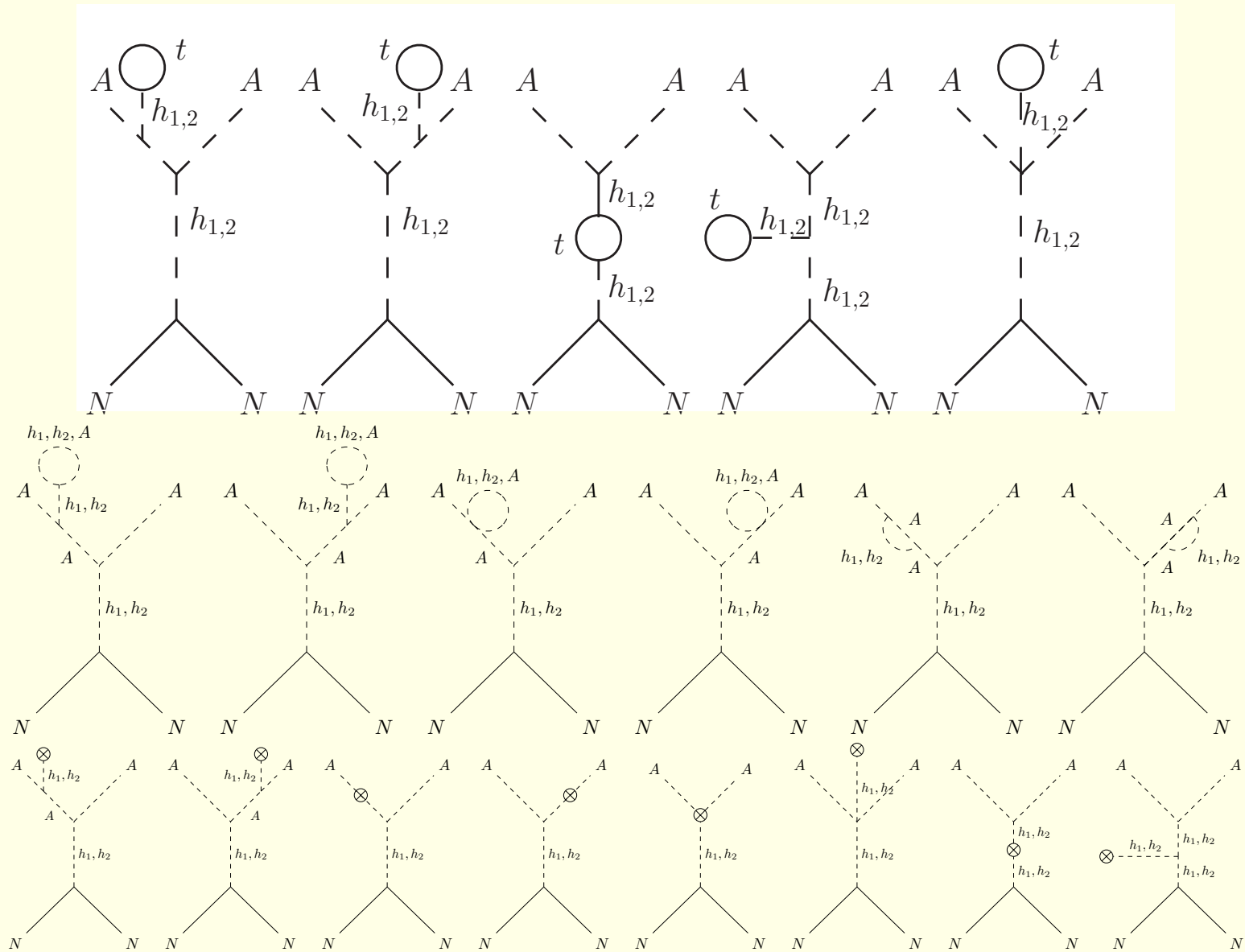


Figure 2: 1-loop diagrams that do not contribute to A -nucleon scattering.

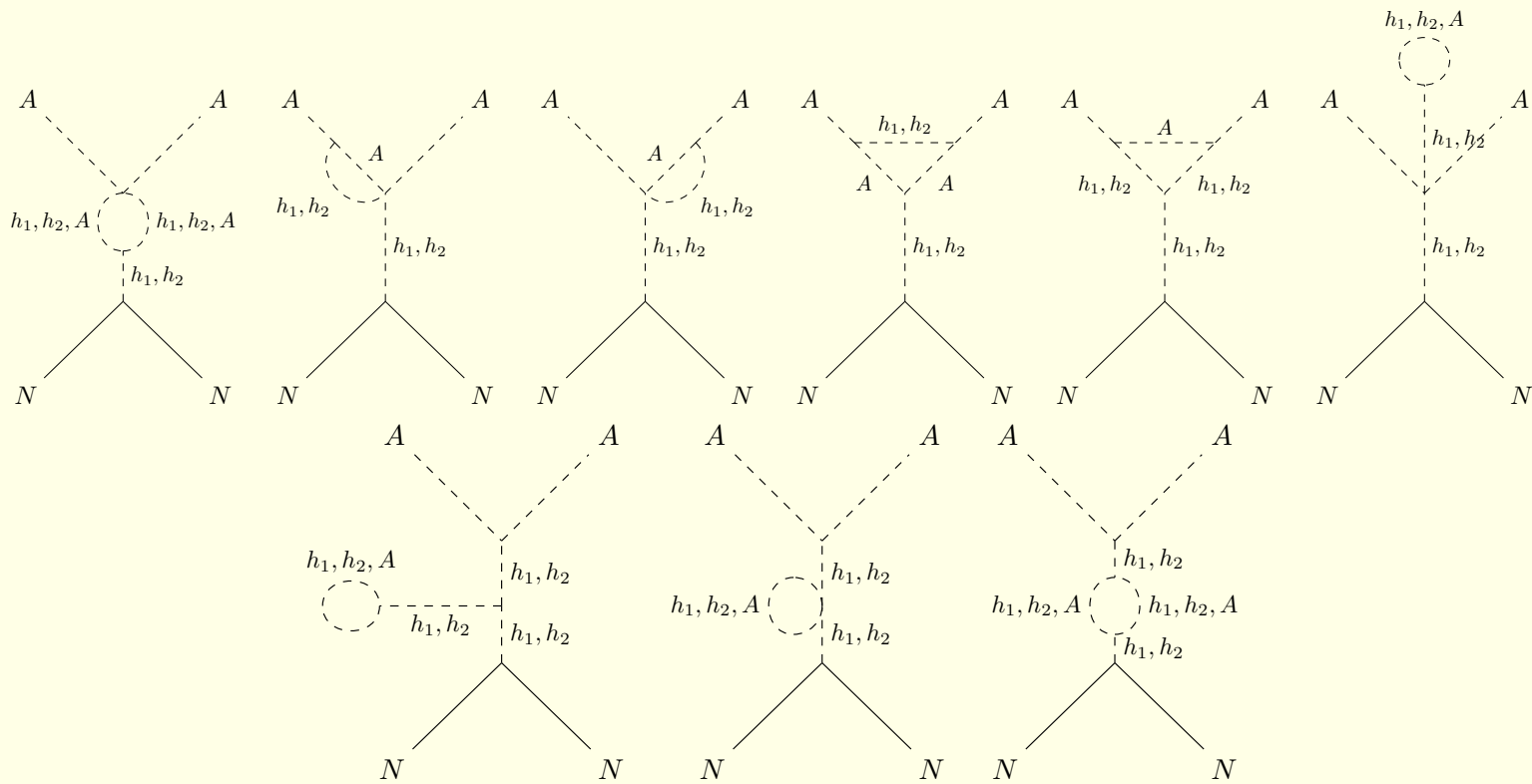


Figure 3: 1-loop diagrams contributing to A -nucleon scattering.

$$\sigma_{AN}^{(1)} = \frac{f_N^2}{\pi v^2} \frac{m_N^2 \mu_{AN}^2}{m_A^2} \mathcal{F}^2,$$

where the one-loop function $\mathcal{F}$ is defined as

$$\mathcal{F} = \frac{V_{AA1}^{(1)} c_\alpha}{m_1^2} - \frac{V_{AA2}^{(1)} s_\alpha}{m_2^2}$$

with $V_{AA1,AA2}^{(1)}$ as one-loop corrections to the vertices $h_1 A^2$ and $h_2 A^2$.

$$\sigma_{\text{AN}} = \frac{\mu^2 m_{\text{DM}}^2}{\pi} \cdot \frac{m_N^2}{v^2} \frac{f_N^2}{m_1^4 m_2^4} \cdot \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 \cdot \left[\frac{\mathcal{A}}{64\pi^2 v v_S^2} \right]^2 ,$$

$$\begin{aligned} \mathcal{A} = & a_1 \cdot C_2(0, m_{\text{DM}}^2, m_{\text{DM}}^2, m_1^2, m_2^2, m_{\text{DM}}^2) + \\ & a_2 \cdot D_3(0, 0, m_{\text{DM}}^2, m_{\text{DM}}^2, 0, m_{\text{DM}}^2, m_1^2, m_1^2, m_2^2, m_{\text{DM}}^2) + \\ & a_3 \cdot D_3(0, 0, m_{\text{DM}}^2, m_{\text{DM}}^2, 0, m_{\text{DM}}^2, m_1^2, m_2^2, m_2^2, m_{\text{DM}}^2) \end{aligned}$$

with

$$\begin{aligned} a_1 &= 2(m_1^2 \sin^2 \alpha + m_2^2 \cos^2 \alpha) [2v(m_1^2 \sin^2 \alpha + m_2^2 \cos^2 \alpha) - (m_1^2 - m_2^2) \sin 2\alpha v_S] \\ a_2 &= -2m_1^4 \sin \alpha [(m_1^2 + 5m_2^2)v_S \cos \alpha - (m_1^2 - m_2^2)(v_S \cos 3\alpha + 4v \sin^3 \alpha)] \\ a_3 &= 2m_2^4 \cos \alpha [(5m_1^2 + m_2^2)v_S \sin \alpha - (m_1^2 - m_2^2)(v_S \sin 3\alpha + 4v \cos^3 \alpha)] \end{aligned}$$

Comments:

- The one loop amplitude $\mathcal{F}$ is UV finite in the limit of zero momentum transfer $q^2 \rightarrow 0$,
- $\mathcal{F} \rightarrow 0$ for $m_A \rightarrow 0$.

$$S = \frac{1}{\sqrt{2}}(v_S + \phi)e^{iA/v_S},$$

- A is odd under the Z_2 symmetry transformation $S \leftrightarrow S^*$, it is DM candidate.
- The only terms that contain A are the kinetic and the $U(1)$ symmetry softly-breaking terms:

$$\begin{aligned}\mathcal{L}_A &= \partial^\mu S^* \partial_\mu S - \frac{M^3}{\sqrt{2}}(S + S^*) - \mu^2(S^2 + S^{*2}) \\ &\supset \frac{1}{2}\partial^\mu A \partial_\mu A + \frac{1}{2}\left(4\mu^2 + \frac{M^3}{v_s}\right)A^2 + \frac{\phi}{v_s}\partial^\mu A \partial_\mu A + \left(4\mu^2 + \frac{M^3}{2v_s}\right)\frac{\phi}{v_s}A^2,\end{aligned}$$

so $m_A^2 = -4\mu^2 - M^3/v_s$.

- Repeatedly integrating by parts and adopting free equations of motion for A and h_i , one finds the pseudo-Goldstone-Higgs vertices as follows

$$\mathcal{L}_A \supset \frac{1}{2}(\partial^\mu A \partial_\mu A - m_A^2 A^2) - \frac{R_{2i}}{2v_s} \left(m_i^2 + \frac{M^3}{v_S} \right) h_i A^2$$

The Vector Dark Matter (VDM) model

- T. Hambye, “Hidden vector dark matter”, JHEP 0901 (2009) 028,
- O. Lebedev, H. M. Lee, and Y. Mambrini, “Vector Higgs-portal dark matter and the invisible Higgs”, Phys.Lett. B707 (2012) 570,
- Y. Farzan and A. R. Akbarieh, “VDM: A model for Vector Dark Matter”, JCAP 1210 (2012) 026,
- S. Baek, P. Ko, W.-I. Park, and E. Senaha, “Higgs Portal Vector Dark Matter : Revisited”, JHEP 1305 (2013) 036,
- Ch. Gross, O. Lebedev, Y. Mambrini, “Non-Abelian gauge fields as dark matter”, arXiv:1505.07480,
- . . .

The model:

- extra $U(1)_X$ gauge symmetry (A_X^μ), **DM candidate:** A_X^μ ,
- a complex scalar field S , whose vev generates a mass for the $U(1)$'s vector field, $S = (0, \mathbf{1}, \mathbf{1}, 1)$ under $U(1)_Y \times SU(2)_L \times SU(3)_c \times U(1)_X$.
- SM fields neutral under $U(1)_X$,
- in order to ensure stability of the new vector boson a $\mathbb{Z}_2$ symmetry is assumed to forbid $U(1)$ -kinetic mixing between $U(1)_X$ and $U(1)_Y$. The extra gauge boson A_μ and the scalar S transform under $\mathbb{Z}_2$ (dark charge conjugation) as follows

$$A_X^\mu \xrightarrow{C} -A_X^\mu, \quad S \xrightarrow{C} S^*$$

The scalar potential

$$V_{\text{VDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2.$$

The vector bosons masses:

$$m_W = \frac{1}{2}gv, \quad m_Z = \frac{1}{2}\sqrt{g^2 + g'^2}v \quad \text{and} \quad m_X = g_X v_S,$$

whit

$$S = \frac{1}{\sqrt{2}}(v_S + \phi + i\pi_S) \quad , \quad \text{and} \quad H = \begin{pmatrix} \pi^+ \\ \frac{1}{\sqrt{2}}(v + h + i\pi^0) \end{pmatrix}.$$

Positivity of the potential implies

$$\lambda_H > 0, \quad \lambda_S > 0, \quad \kappa > -2\sqrt{\lambda_H \lambda_S}.$$

The minimization conditions for scalar fields

$$(2\lambda_H v^2 + \kappa v_S^2 - 2\mu_H^2)v = 0 \quad \text{and} \quad (\kappa v^2 + 2\lambda_S v_S^2 - 2\mu_S^2)v_S = 0$$

For $\kappa^2 < 4\lambda_H\lambda_S$ the global minimum is

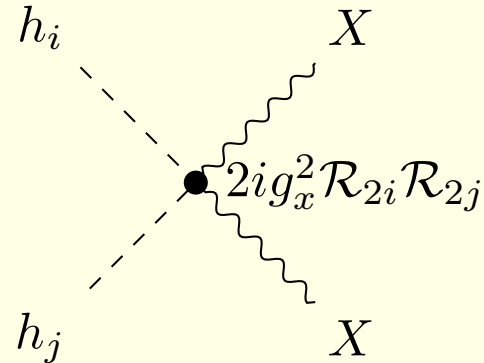
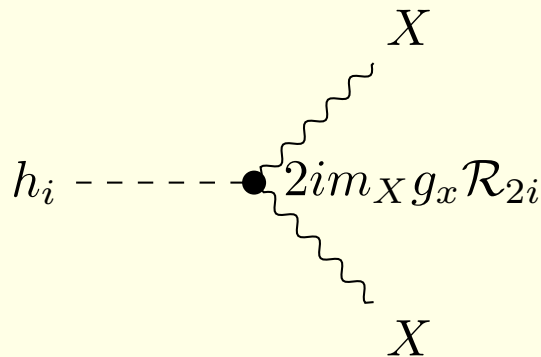
$$v^2 = \frac{4\lambda_S\mu_H^2 - 2\kappa\mu_S^2}{4\lambda_H\lambda_S - \kappa^2} \quad \text{and} \quad v_S^2 = \frac{4\lambda_H\mu_S^2 - 2\kappa\mu_H^2}{4\lambda_H\lambda_S - \kappa^2}$$

The mass squared matrix $\mathcal{M}^2$ for the fluctuations (h, ϕ) and their eigenvalues read

$$\mathcal{M}^2 = \begin{pmatrix} 2\lambda_H v^2 & \kappa v v_S \\ \kappa v v_S & 2\lambda_S v_S^2 \end{pmatrix}$$

$$m_{\pm}^2 = \lambda_H v^2 + \lambda_S v_S^2 \pm \sqrt{\lambda_S^2 v_S^4 - 2\lambda_H\lambda_S v^2 v_S^2 + \lambda_H^2 v^4 + \kappa^2 v^2 v_S^4}$$

$$\mathcal{M}_{\text{diag}}^2 = \begin{pmatrix} m_1^2 & 0 \\ 0 & m_2^2 \end{pmatrix}, \quad R = \begin{pmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{pmatrix}, \quad \begin{pmatrix} h_1 \\ h_2 \end{pmatrix} = R^{-1} \begin{pmatrix} h \\ \phi \end{pmatrix},$$



Fermion Dark Matter (FDM)

- DM: χ - a left-handed Dirac fermion,
- spin 0 mediator: S - a real field.

$$\mathbb{Z}_4 : S \rightarrow -S, \chi \rightarrow i\chi$$

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + i\bar{\chi}\not{\partial}\chi + \frac{1}{2}\partial^\mu S \partial_\mu S - \frac{y_x}{2}(\bar{\chi}^c\chi + \bar{\chi}\chi^c)S - V(H, S) ,$$

$$V_{\text{FDM}}(H, S) = -\mu_H^2|H|^2 + \lambda_H|H|^4 - \frac{\mu_S^2}{2}S^2 + \frac{\lambda_S}{4}S^4 + \frac{\kappa}{2}|H|^2S^2 ,$$

where $\chi^c \equiv -i\gamma_2\chi^*$ and

$$S = v_S + \phi , \quad H = \begin{pmatrix} \pi^+ \\ \frac{v+h+i\pi^0}{\sqrt{2}} \end{pmatrix}$$

After SSB relevant parts of the Lagrangian take the following form:

$$i\bar{\chi}\not{\partial}\chi + \frac{1}{2}\partial^\mu S \partial_\mu S - \frac{y_x}{2}(\bar{\chi}^c\chi + \bar{\chi}\chi^c)S \rightarrow \frac{i}{2}\bar{\psi}\not{\partial}\psi + \frac{1}{2}\partial^\mu\phi \partial_\mu\phi - \frac{y_x v_S}{2}\bar{\psi}\psi - \frac{y_x}{2}\bar{\psi}\psi\phi$$

where $\psi = \psi^c \equiv \chi + \chi^c$ is a Majorana mass eigenstate with $m_\psi = y_x v_S$.

$$\begin{pmatrix} h_1 \\ h_2 \end{pmatrix} = \mathcal{R}^{-1} \begin{pmatrix} h \\ \phi \end{pmatrix}, \mathcal{R} = \begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix}, \alpha \in \left[-\frac{\pi}{4}, \frac{\pi}{4}\right], \tan 2\alpha = \frac{\kappa v v_S}{\lambda_H v^2 - \lambda_S v_S^2}.$$

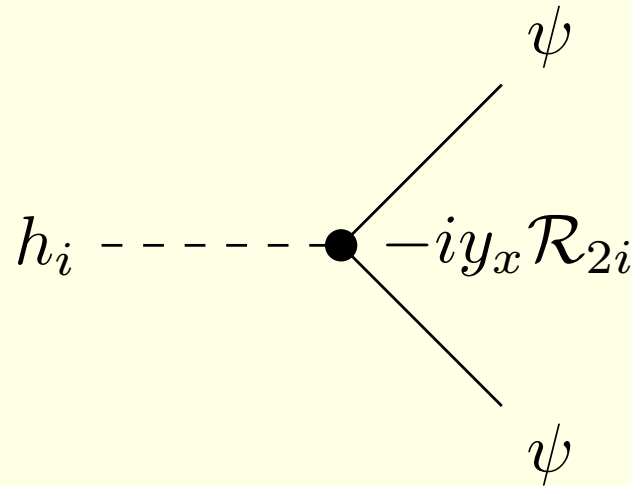


Figure 4: The vertex relevant for the FDM model.

Parameters

$$V_{\text{pGDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2 + (\mu^2 S^2 + \text{H.c.})$$

$$V_{\text{VDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$

$$V_{\text{FDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \frac{\mu_S^2}{2} S^2 + \frac{\lambda_S}{4} S^4 + \frac{\kappa}{2} |H|^2 S^2$$

$$S = v_S + \phi (+iA) , \quad H = \begin{pmatrix} \pi^+ \\ \frac{v+h+i\pi^0}{\sqrt{2}} \end{pmatrix}$$

$$\begin{pmatrix} h_1 \\ h_2 \end{pmatrix} = \mathcal{R}^{-1} \begin{pmatrix} h \\ \phi \end{pmatrix} , \mathcal{R} = \begin{bmatrix} \cos \alpha & -\sin \alpha \\ \sin \alpha & \cos \alpha \end{bmatrix}$$

It is convenient to use the same input parameters for all the models:

$$m_2, \sin \alpha, m_{DM} \equiv (m_A, m_X, m_\psi) \quad \text{and} \quad v_S$$

- **CEPC:** $\sqrt{s} = 240 \text{ GeV} \implies m_{DM} \lesssim 72.5 \text{ GeV}$
 $52.5 \text{ GeV} < m_{DM} < 72.5 \text{ GeV}, 85 \text{ GeV} < m_2 < 165 \text{ GeV}$
- **ILC:** $\sqrt{s} = 250 \text{ GeV} \implies m_{DM} \lesssim 77.5 \text{ GeV}$

Vacuum stability

$$V_{\text{VDM}}(H, S) = -\mu_H^2 |H|^2 + \lambda_H |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$

$$\lambda_H(Q) > 0, \quad \lambda_S(Q) > 0, \quad \kappa(Q) + 2\sqrt{\lambda_H(Q)\lambda_S(Q)} > 0$$

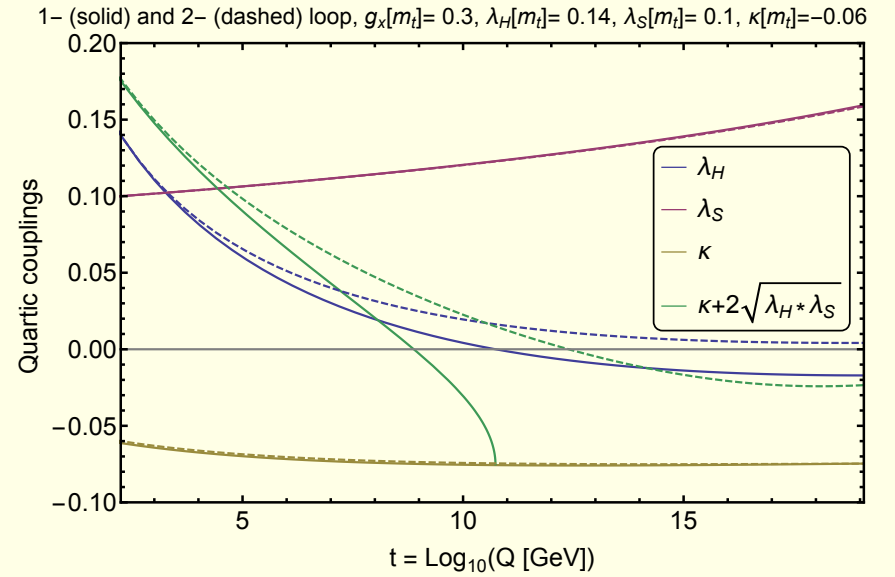
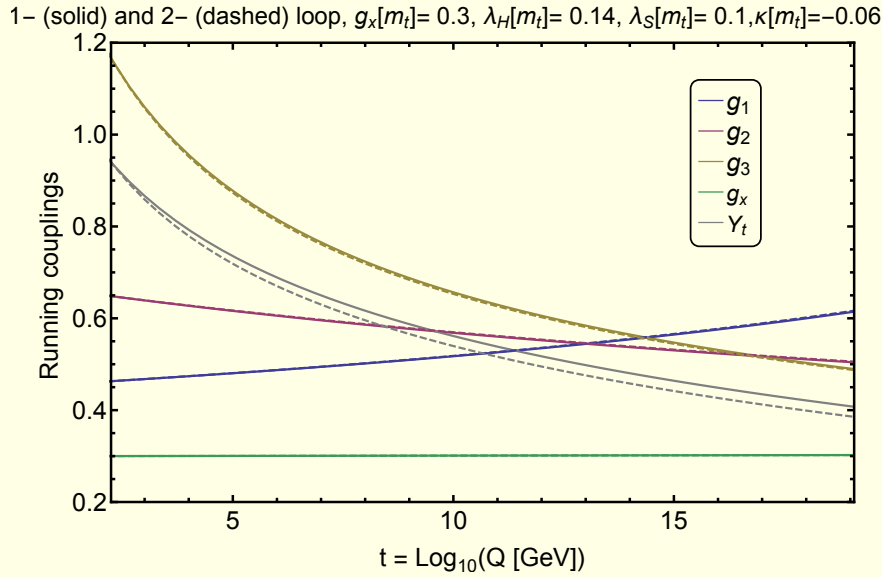


Figure 5: Running of various parameters at 1- and 2-loop, in solid and dashed lines respectively. For this choice of parameters $\lambda_H(Q) > 0$ at 2-loop (right panel blue) but not at 1-loop. $\lambda_S(Q)$ is always positive (right panel red), running of $\kappa(Q)$ is very limited, however the third positivity condition $\kappa(Q) + 2\sqrt{\lambda_H(Q)\lambda_S(Q)} > 0$ is violated at higher scales even at 2-loops (right panel green).

The mass of the Higgs boson is known experimentally therefore within *the SM* the initial condition for running of $\lambda_H(Q)$ is fixed

$$\lambda_H(m_t) = M_{h_1}^2/(2v^2) = \lambda_{SM} = 0.13$$

For VDM this is not necessarily the case:

$$M_{h_1}^2 = \lambda_H v^2 + \lambda_S v_S^2 \pm \sqrt{\lambda_S^2 v_S^4 - 2\lambda_H \lambda_S v^2 v_S^2 + \lambda_H^2 v^4 + \kappa^2 v^2 v_S^4}.$$

VDM:

- Larger initial values of λ_H such that $\lambda_H(m_t) > \lambda_{SM}$ are allowed delaying the instability (by shifting up the scale at which $\lambda_H(Q) < 0$).
- Even if the initial λ_H is smaller than its SM value, $\lambda_H(m_t) < \lambda_{SM}$, still there is a chance to lift the instability scale if appropriate initial value of the portal coupling $\kappa(m_t)$ is chosen.

$$\beta_{\lambda_H}^{(1)} = \beta_{\lambda_H}^{SM(1)} + \kappa^2$$

Constraints

1. Perturbativity: $y_x < 4\pi$, $g_X < 4\pi$ ($v_S > \frac{m_X}{4\pi}$), $\kappa < 4\pi$.
2. The mixing angle: $0 < \sin \alpha < 0.3$ (1501.02234, 1604.04552).

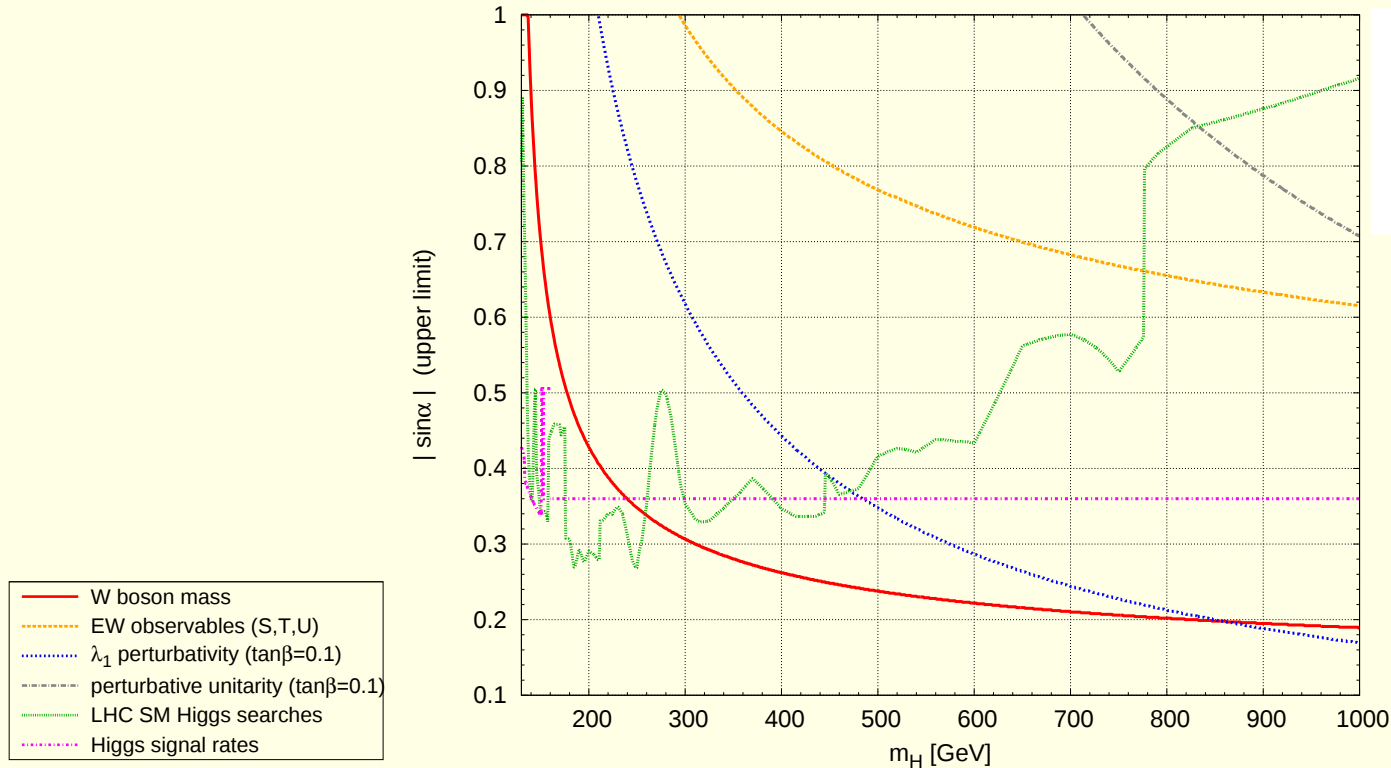


Figure 6: Upper limit for $|\sin \alpha|$. (1611.03007)

3. Higgs invisible decays (1809.05937): $\text{BR}(h_1 \rightarrow \text{inv}) = \text{BR}(h_1 \rightarrow DM DM) < 19\%$.

4. The Planck data (1807.06209): $h^2 \Omega = 0.12 \pm 0.0012$. The thermally averaged cross section ($DM DM \rightarrow \bar{f} f$) reads:

$$\langle \sigma v \rangle = \frac{n_c m_{\text{DM}} m_f^2}{3 \pi v^2} \cdot \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 \cdot \frac{(m_{\text{DM}}^2 - m_f^2)^{3/2}}{(4m_{\text{DM}}^2 - m_1^2)^2 (4m_{\text{DM}}^2 - m_2^2)^2} \cdot$$

$$\times \begin{cases} 12 & (\text{pGDM}) \\ 1 & (\text{VDM}) \\ \frac{9}{4} \left(\frac{m_{\text{DM}}}{T}\right)^{-1} + [\text{higher orders in } \left(\frac{m_{\text{DM}}}{T}\right)^{-1}] & (\text{FDM}) \end{cases}$$

$$\langle \sigma v \rangle = \sigma_0 x^{-n} \implies h^2 \Omega \propto \frac{(n+1) x_f^{n+1}}{\sigma_0} \quad (x \equiv \frac{m_{\text{DM}}}{T})$$

$$\Downarrow$$

$$\frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 =$$

$$= 2.1 \cdot 10^{-5} \text{ GeV}^{-2} \frac{(4m_{\text{DM}}^2 - m_1^2)^2 (4m_{\text{DM}}^2 - m_2^2)^2}{m_{\text{DM}} (m_{\text{DM}}^2 - m_b^2)^{3/2}} \cdot \begin{cases} \frac{1}{12} & (\text{pGDM}) \\ 1 & (\text{VDM}) \\ \frac{4m_{\text{DM}}}{9T_f} & (\text{FDM}) \end{cases}$$

5. Indirect-detection (1611.03184) limits.

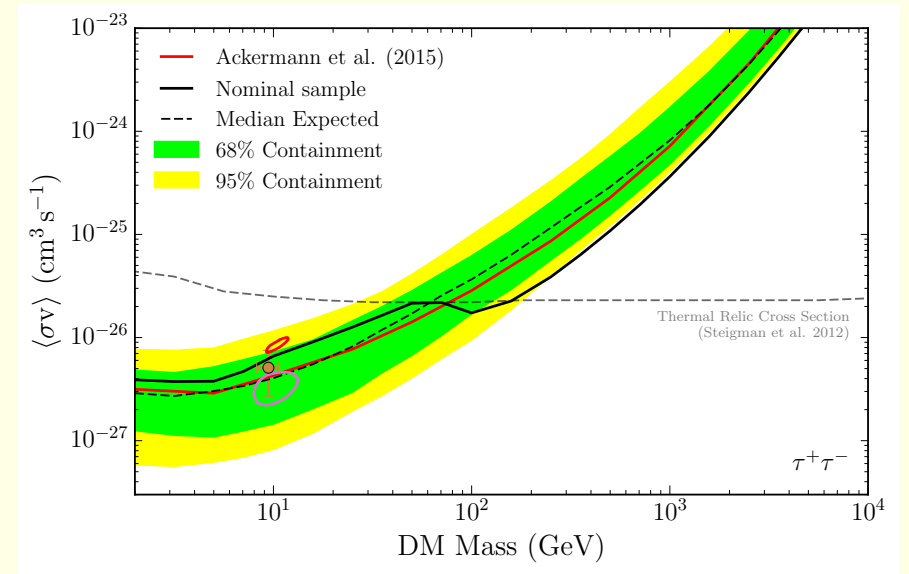
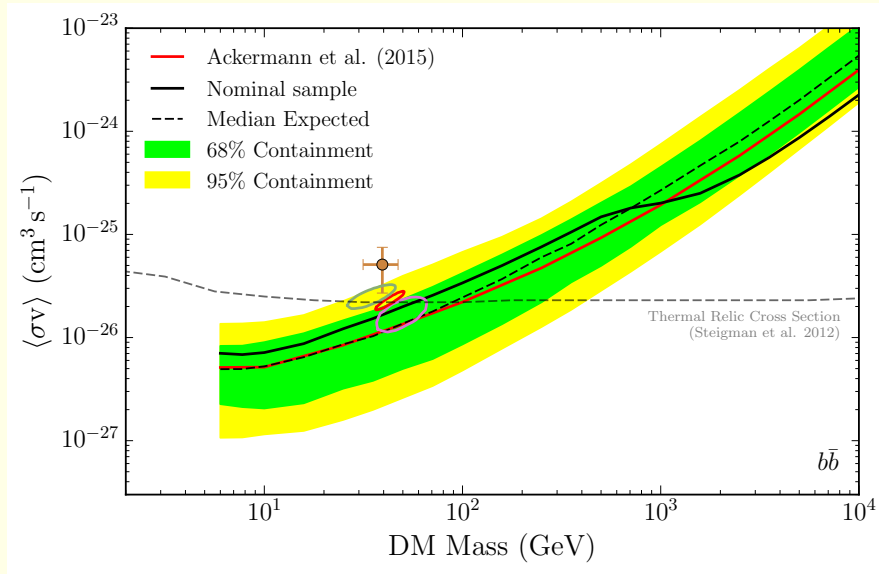


Figure 7: Relic-density and indirect-detection limits. (1611.03184)

- For pGDM and VDM: $m_{\text{pGDM}}, m_{\text{VDM}} \gtrsim 30 \text{ GeV}$, since otherwise the proper value of annihilation cross section is forbidden.
- For FDM due to the T_0/T_f factor, current annihilation cross-section corresponding to the correct value of relic density is orders of magnitude times smaller than the one for the pGDM and VDM, hence it satisfies the ID limit.

6. Direct-detection (1805.12562) constraint the spin-independent nucleon-scattering cross section.

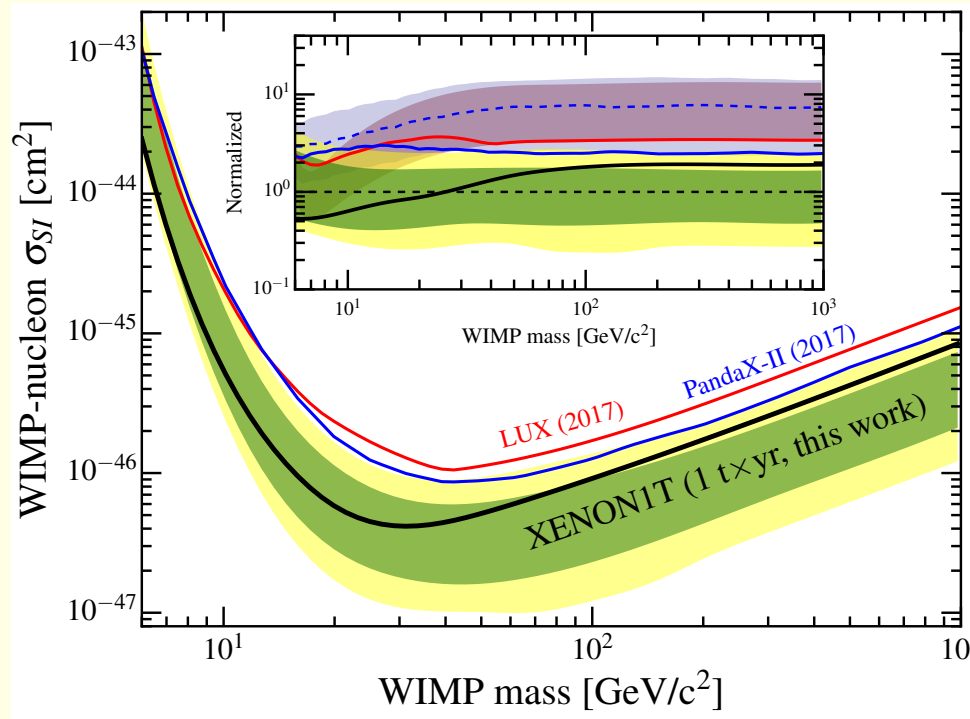


Figure 8: Direct-detection limits. (1805.12562)

For VDM and FDM:

$$\sigma_{\text{SI}} \simeq \frac{\mu^2 m_{\text{DM}}^2}{\pi} \cdot \frac{m_N^2}{v^2} \frac{f_N^2}{m_1^4 m_2^4} \cdot \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 ,$$

For pGDM 1-loop calculations are needed (1810.06105).

$$\sigma_{\text{SI}} = \frac{\mu^2 m_{\text{DM}}^2}{\pi} \cdot \frac{m_N^2}{v^2} \frac{f_N^2}{m_1^4 m_2^4} \cdot \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 \cdot \left[\frac{\mathcal{A}}{64\pi^2 v v_S^2} \right]^2 ,$$

$$\mathcal{A} = a_1 \cdot C_2(0, m_{\text{DM}}^2, m_{\text{DM}}^2, m_1^2, m_2^2, m_{\text{DM}}^2) +$$

$$a_2 \cdot D_3(0, 0, m_{\text{DM}}^2, m_{\text{DM}}^2, 0, m_{\text{DM}}^2, m_1^2, m_1^2, m_2^2, m_{\text{DM}}^2) +$$

$$a_3 \cdot D_3(0, 0, m_{\text{DM}}^2, m_{\text{DM}}^2, 0, m_{\text{DM}}^2, m_1^2, m_2^2, m_2^2, m_{\text{DM}}^2)$$

with

$$a_1 = 2(m_1^2 \sin^2 \alpha + m_2^2 \cos^2 \alpha) [2v(m_1^2 \sin^2 \alpha + m_2^2 \cos^2 \alpha) - (m_1^2 - m_2^2) \sin 2\alpha v_S]$$

$$a_2 = -2m_1^4 \sin \alpha [(m_1^2 + 5m_2^2)v_S \cos \alpha - (m_1^2 - m_2^2)(v_S \cos 3\alpha + 4v \sin^3 \alpha)]$$

$$a_3 = 2m_2^4 \cos \alpha [(5m_1^2 + m_2^2)v_S \sin \alpha - (m_1^2 - m_2^2)(v_S \sin 3\alpha + 4v \cos^3 \alpha)]$$

$$\left[\frac{\mathcal{A}}{64\pi^2 v v_S^2} \right]^2 \lesssim 10^{-5}$$

In considered range of parameters, in the case of pGDM, the DD upper bound on the value of $\frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2$ is always higher than the value corresponding to the correct relic density (for $\sin \alpha$ that maximize the cross section for given m_2 and m_{DM}).

$$\sigma_{\text{SI}} = \frac{\mu^2 m_{\text{DM}}^2}{\pi} \cdot \frac{m_N^2}{v^2} \frac{f_N^2}{m_1^4 m_2^4} \cdot \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 \cdot \left[\frac{\mathcal{A}}{64\pi^2 v v_S^2} \right]^2 ,$$

The XENON1T limit for $m_{\text{DM}} \gtrsim 100 \text{ GeV}$ can be parametrized as:

$$\frac{\sigma_{\text{SI}}^{\text{max}}}{1 \text{ cm}^2} = \frac{m_{\text{DM}}}{1 \text{ GeV}} \cdot 10^{-48.05} . \quad (1)$$

Hence, the strictest possible DD limit reads

$$\begin{aligned} \frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 &< \frac{m_2^4}{m_{\text{DM}}} \frac{v^2}{m_N^2} \frac{m_1^4}{f_N^2} \frac{\pi}{\mu^2} \frac{1 \text{ cm}^2}{1 \text{ GeV}} \cdot 10^{-48.05} \cdot \begin{cases} \left[\frac{\mathcal{A}}{64\pi^2 v v_S^2} \right]^{-2} & (\text{pGDM}) \\ 1 & (\text{FDM}, \end{cases} \\ &= \frac{m_2^4}{m_{\text{DM}}} \cdot 1.5 \cdot 10^{-6} \text{ GeV}^{-1} \cdot \begin{cases} 10^5 & (\text{pGDM}) \\ 1 & (\text{FDM}, \text{VDM}) \end{cases} . \end{aligned} \quad (2)$$

Dark matter at e^+e^- colliders

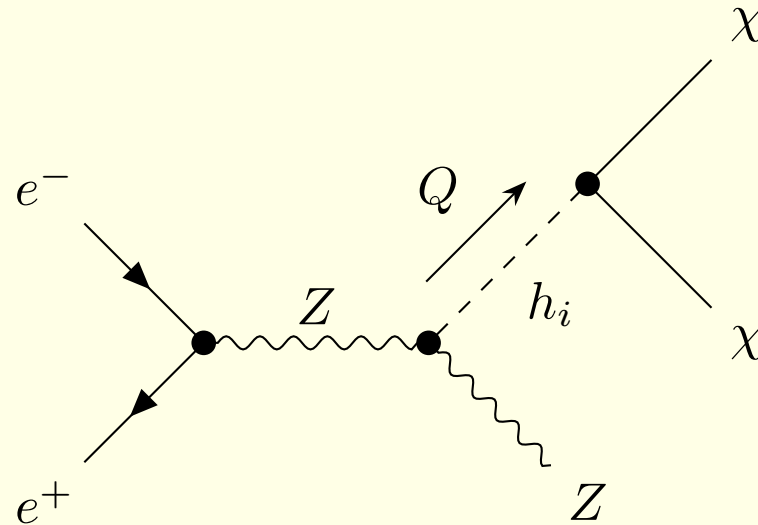


Figure 9: Feynman diagram for $e^+e^- \rightarrow Z\chi\bar{\chi}$, χ denotes the dark particle ($\chi = A, X, \psi$).

- P. Ko, H. Yokoya, “Search for Higgs portal DM at the ILC”, JHEP 1608 (2016) 109,
- T. Kamon, P. Ko, J. Li “Characterizing Higgs portal dark matter models at the ILC”, Eur.Phys.J. C77 (2017) no.9, 652

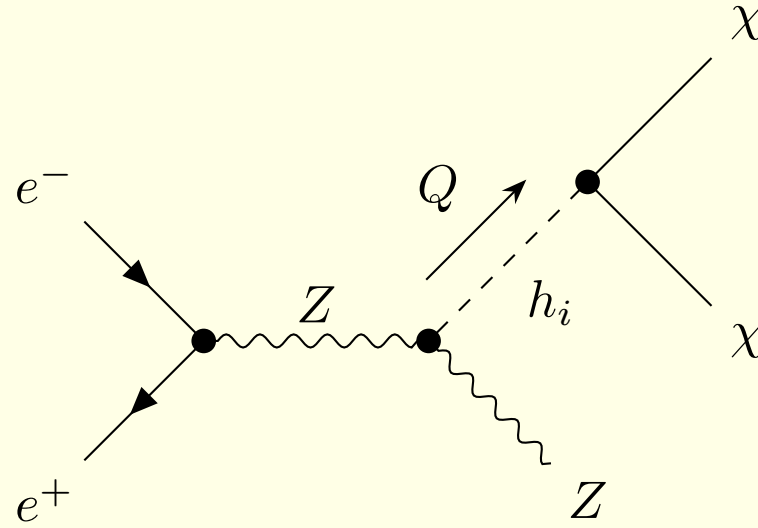


Figure 10: Feynman diagram for $e^+e^- \rightarrow Z\chi\bar{\chi}$, χ denotes the dark particle ($\chi = A, X, \psi$).

$$\frac{d\sigma}{dQ^2}(Q^2) = \frac{f(s, \sqrt{Q^2})}{32\pi^2} \frac{(Q^2)^2 \cdot \sin^2 \alpha \cos^2 \alpha \cdot [(m_1^2 - m_2^2)^2 + (m_1\Gamma_1 - m_2\Gamma_2)^2]}{[(Q^2 - m_1^2)^2 + (m_1\Gamma_1)^2] [(Q^2 - m_2^2)^2 + (m_2\Gamma_2)^2]} \times$$

$$\times \sqrt{1 - 4\frac{m_{\text{DM}}^2}{Q^2}} \cdot \frac{1}{v_S^2} \cdot \begin{cases} 1 & (\text{pGDM}) \\ 1 - 4\frac{m_{\text{DM}}^2}{Q^2} + 12\left(\frac{m_{\text{DM}}^2}{Q^2}\right)^2 & (\text{VDM}) \\ 2\left[\frac{m_{\text{DM}}^2}{Q^2} - 4\left(\frac{m_{\text{DM}}^2}{Q^2}\right)^2\right] & (\text{FDM}) \end{cases}$$

$$f(s, \sqrt{Q^2}) \equiv \frac{g_V^2 + g_A^2}{24\pi} \left(\frac{g^2}{\cos \theta_W^2} \frac{1}{s - m_Z^2} \right)^2 \frac{\lambda^{1/2}(s, Q^2, m_Z^2) [12 s m_Z^2 + \lambda(s, Q^2, m_Z^2)]}{8s^2}$$

$$Q^2 = Q^2(s, E_Z) \equiv s - 2E_Z\sqrt{s} + m_Z^2$$

$$E_Z(Q^2 = m_i^2) = E_i \equiv \frac{s - m_i^2 + m_Z^2}{2\sqrt{s}}.$$

$$E_{\max} = \frac{s - 4m_{DM}^2 + m_Z^2}{2\sqrt{s}},$$

$$\sqrt{s} = 1.5 \text{ TeV}, \quad m_2 = 700 \text{ GeV}, \quad v_S = 5.54 \text{ TeV}$$

— two-pole case: $m_{DM} = 60 \text{ GeV}$, $\sin \alpha = 0.01$

— one-pole case: $m_{DM} = 200 \text{ GeV}$, $\sin \alpha = 0.05$

— no-pole case: $m_{DM} = 500 \text{ GeV}$, $\sin \alpha = 0.3$

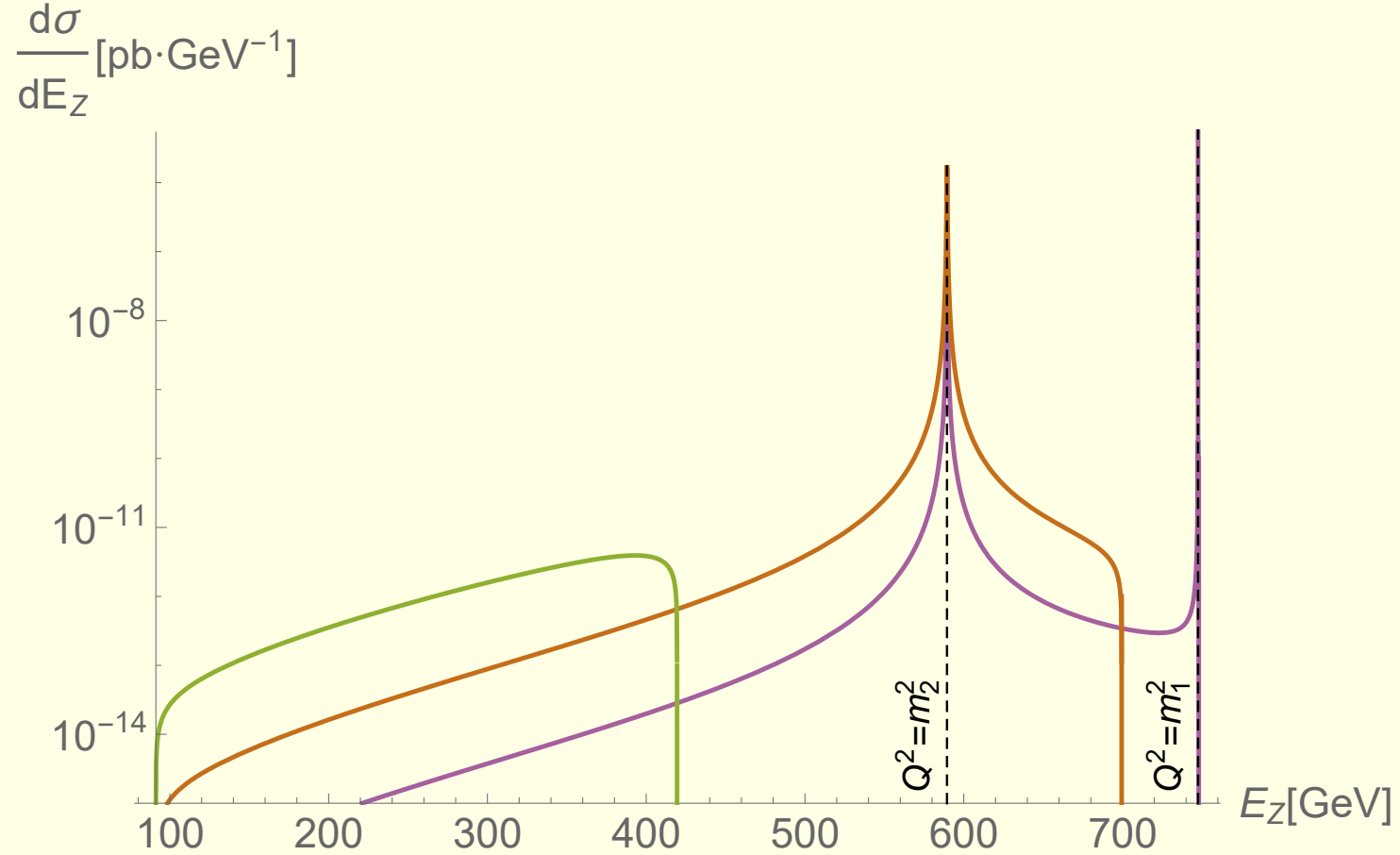


Figure 11: $\frac{d\sigma}{dE_Z}$ for the pGDM model.

$$E_Z(Q^2 = m_i^2) = E_i \equiv \frac{s - m_i^2 + m_Z^2}{2\sqrt{s}}, \quad E_{\max} = \frac{s - 4m_{DM}^2 + m_Z^2}{2\sqrt{s}}$$

Background

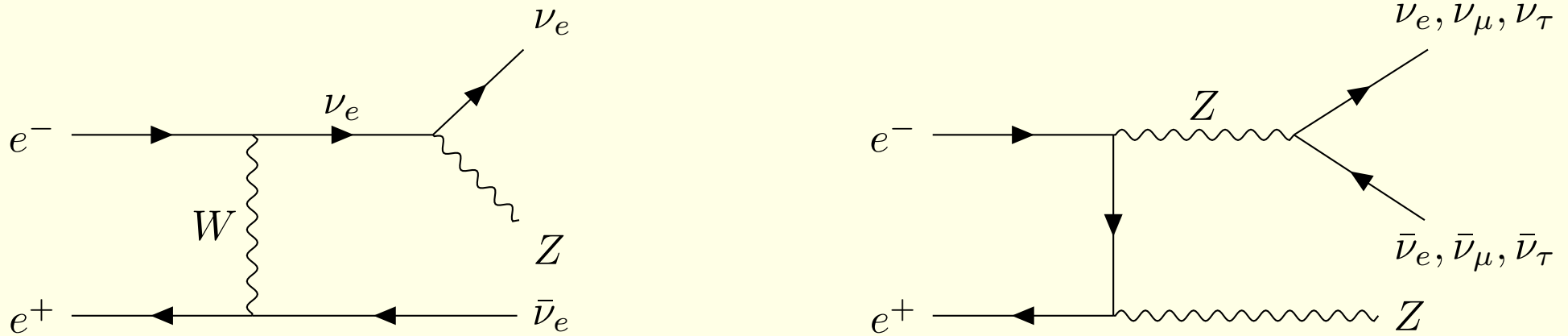


Figure 12: Exemplary diagrams of the Standard Model background processes. Neutrinos contribute to missing energy and can therefore mimic dark particles. The background cross-section could be reduced by polarizing the initial e^+ and e^- beams.

Strategy

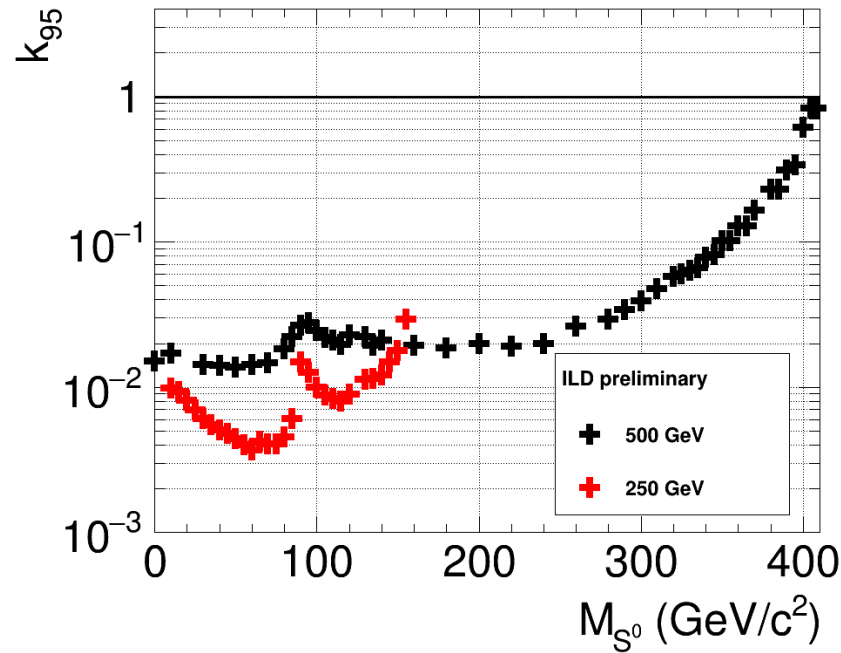
- Fix $\sqrt{s}$ e.g. at 240 GeV for the CEPC, or $\sqrt{s} = 250$ GeV for the ILC,
- Parameters: m_2 , m_{DM} , $\sin \alpha$ and v_S ,
- For given (m_2, m_{DM}) and $\sin \alpha$, the value of v_S is derived by solving the relic-density condition given: so v_S is no longer independent:

$$\frac{\sin^2 \alpha \cos^2 \alpha}{v_S^2} (m_1^2 - m_2^2)^2 =$$
$$= 2.1 \cdot 10^{-5} \text{ GeV}^{-2} \frac{(m_1^2 - 4m_{\text{DM}}^2)^2 (m_2^2 - 4m_{\text{DM}}^2)^2}{m_{\text{DM}}(m_{\text{DM}}^2 - m_b^2)^{3/2}} \cdot \begin{cases} \frac{1}{12} & (\text{pGDM}) \\ 1 & (\text{VDM}) \\ \frac{4m_{\text{DM}}}{9T_f} & (\text{FDM}) \end{cases}$$

- For each point (m_2, m_{DM}) of the plot, we choose such value of $\sin \alpha \leq 0.3$ that maximizes the total cross section.

Extra scalars at 500 ILC $\rightarrow$ Exclusion Limits

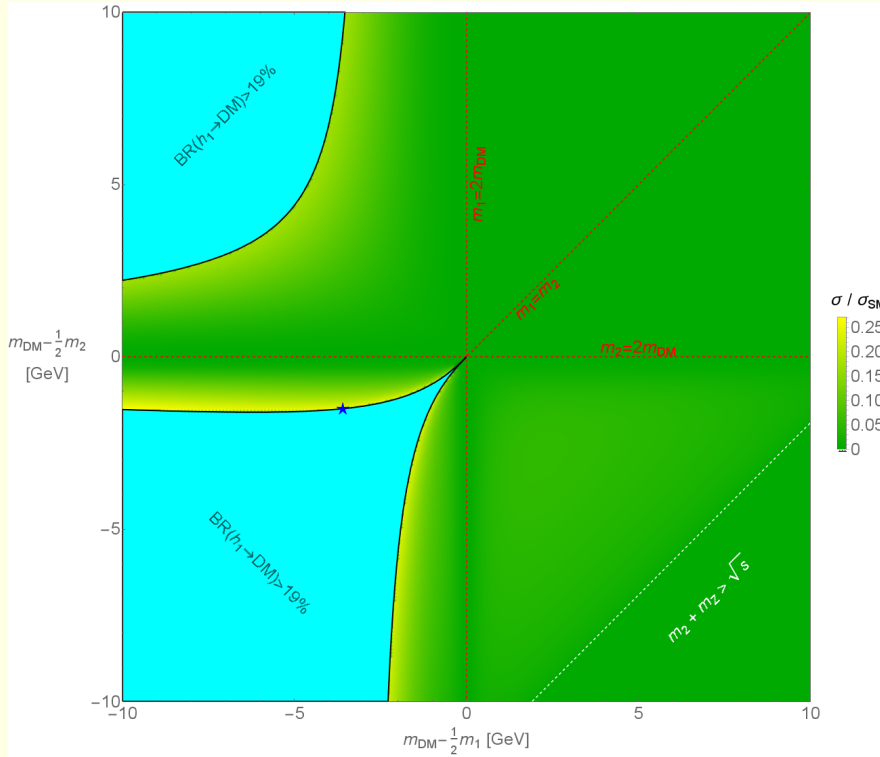
Preliminary results for 500 GeV.



Yan Wang | Searching for new extra scalars at the ILC | October 25, 2018 | 17/18



Figure 13: Expected sensitivity for the measurement of the cross-section for $e^+e^- \rightarrow DM DM Z$ at the ILC, from Yan Wang (DESY, IHEP) at LCWS 2018, Arlington, October 25, 2018. $\kappa \equiv \sigma(e^+e^- \rightarrow \dots + Z)/\sigma_{SM}(e^+e^- \rightarrow h_{SM}Z)$



Pseudo Goldstone dark matter (pGDM)

$$m_2 = 120.8 \text{ GeV} , \quad m_{\text{DM}} = 58.9 \text{ GeV} ,$$

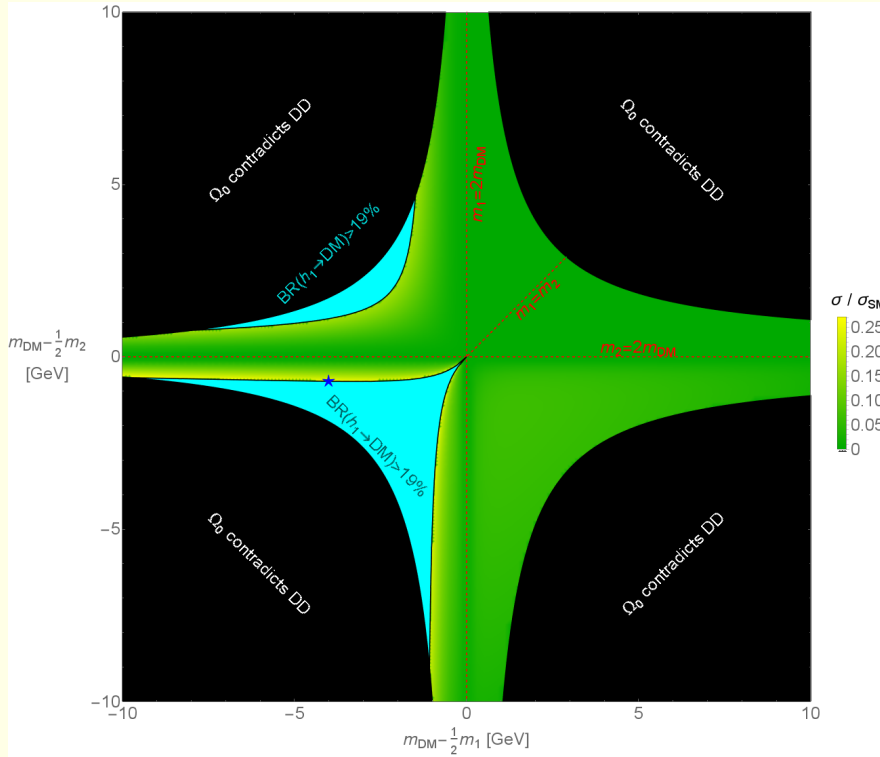
$$\sin \alpha = 0.30 , \quad v_S = 646 \text{ GeV} ,$$

$$\Gamma_1 = 7.4 \cdot 10^{-3} \text{ GeV} , \quad \Gamma_2 = 9.8 \cdot 10^{-3} \text{ GeV} ,$$

$$\text{BR}(h_1 \rightarrow \text{DM}) = 19\% , \quad \text{BR}(h_2 \rightarrow \text{DM}) = 95\% ,$$

$$\frac{\sigma}{\sigma_{SM}} = 0.26, \quad \sigma = 62 \text{ fb}$$

$\sigma(e^+e^- \rightarrow AA Z)/\sigma_{SM}(e^+e^- \rightarrow h_{SM}Z)$ for the pGDM model. **Cyan:** forbidden region, $\text{BR}(h_1 \rightarrow \text{DM}) > 19\%$. **Gray:** small cross section, h_2 cannot be produced on-shell due to required energy exceeding $\sqrt{s}$. The **star** denotes the proposed benchmark point: $m_2 = 120.8 \text{ GeV}$, $m_{\text{DM}} = 58.9 \text{ GeV}$.



Vector dark matter (VDM)

$$m_2 = 118.4 \text{ GeV} , \quad m_{\text{DM}} = 58.5 \text{ GeV} ,$$

$$\sin \alpha = 0.30 , \quad v_S = 561 \text{ GeV} ,$$

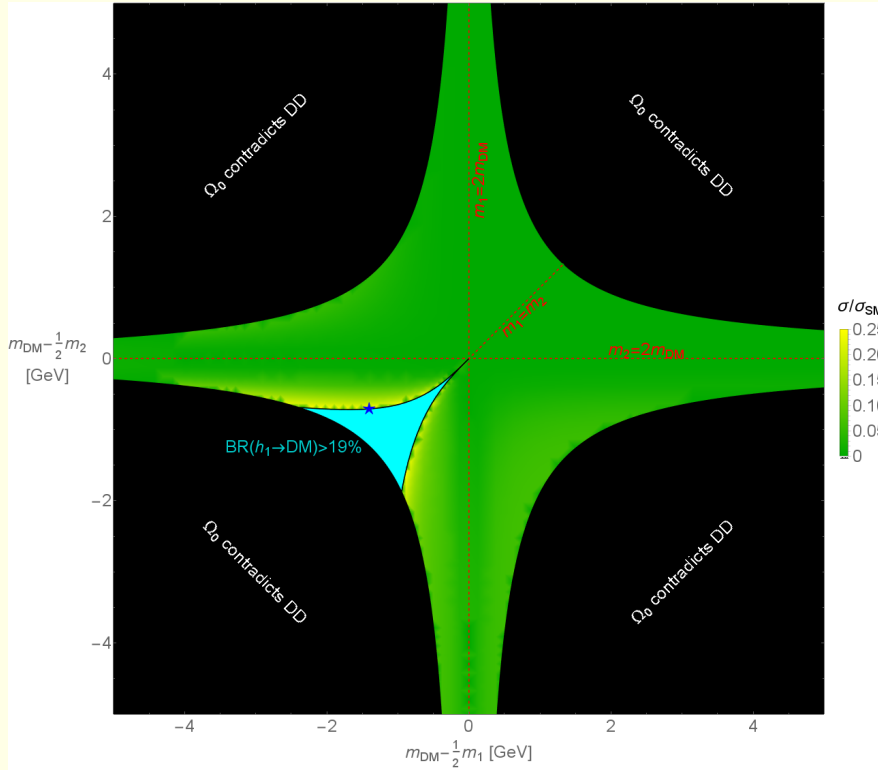
$$\Gamma_1 = 7.4 \cdot 10^{-3} \text{ GeV} , \quad \Gamma_2 = 6.4 \cdot 10^{-3} \text{ GeV} ,$$

$$\text{BR}(h_1 \rightarrow \text{DM}) = 18\% , \quad \text{BR}(h_2 \rightarrow \text{DM}) = 92\% ,$$

$$\frac{\sigma}{\sigma_{SM}} = 0.25, \quad \sigma = 61 \text{ fb}$$

$\sigma(e^+e^- \rightarrow XX Z)/\sigma_{SM}(e^+e^- \rightarrow h_{SM}Z)$ for the VDM model. **Black:** forbidden region, correct relic density cannot be obtained within the direct detection limits.

Cyan: forbidden region, $\text{BR}(h_1 \rightarrow \text{DM}) > 19\%$. The **star** denotes the proposed benchmark point: $m_2 = 118.4 \text{ GeV}$, $m_{\text{DM}} = 58.5 \text{ GeV}$.



Fermion dark matter (FDM)

$$m_2 = 123.6 \text{ GeV} , \quad m_{\text{DM}} = 61.1 \text{ GeV} ,$$

$$\sin \alpha = 0.30 , \quad v_S = 76 \text{ GeV} ,$$

$$\Gamma_1 = 7.4 \cdot 10^{-3} \text{ GeV} , \quad \Gamma_2 = 5.9 \cdot 10^{-3} \text{ GeV} ,$$

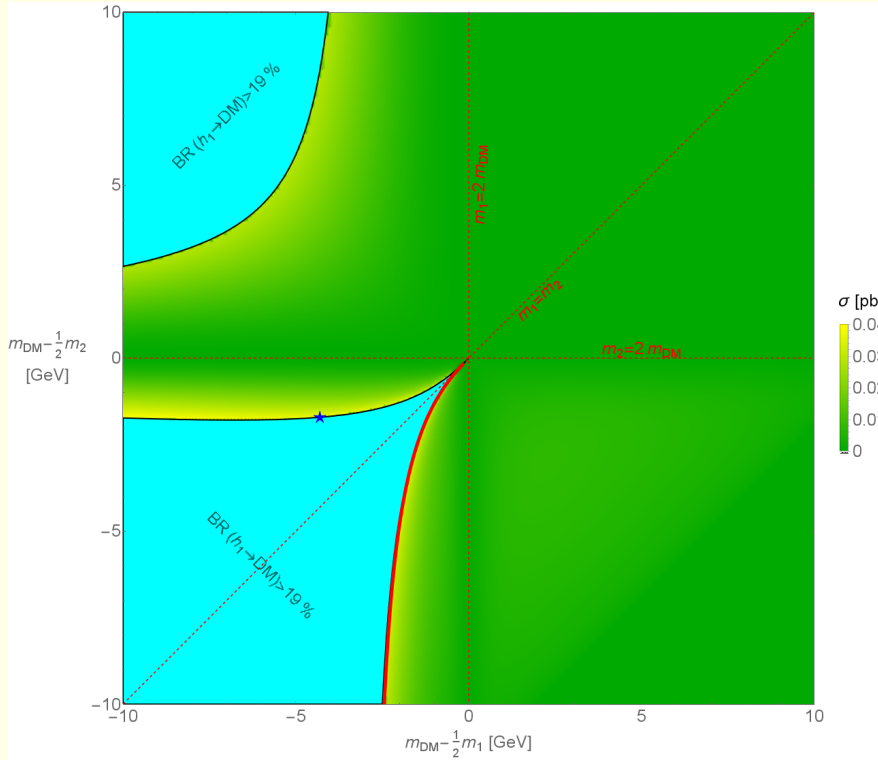
$$\text{BR}(h_1 \rightarrow \text{DM}) = 18\% , \quad \text{BR}(h_2 \rightarrow \text{DM}) = 91\% ,$$

$$\frac{\sigma}{\sigma_{SM}} = 0.24, \quad \sigma = 59 \text{ fb}$$

$\sigma(e^+e^- \rightarrow \psi\psi Z)/\sigma_{SM}(e^+e^- \rightarrow h_{SM}Z)$ in the FDM model. **Black:** forbidden region, correct relic density cannot be obtained within the direct detection limits. **Cyan:** forbidden region, $\text{BR}(h_1 \rightarrow \text{DM}) > 19\%$. The **star** denotes the proposed benchmark point: $m_2 = 123.6 \text{ GeV}$, $m_{\text{DM}} = 61.1 \text{ GeV}$.

Summary

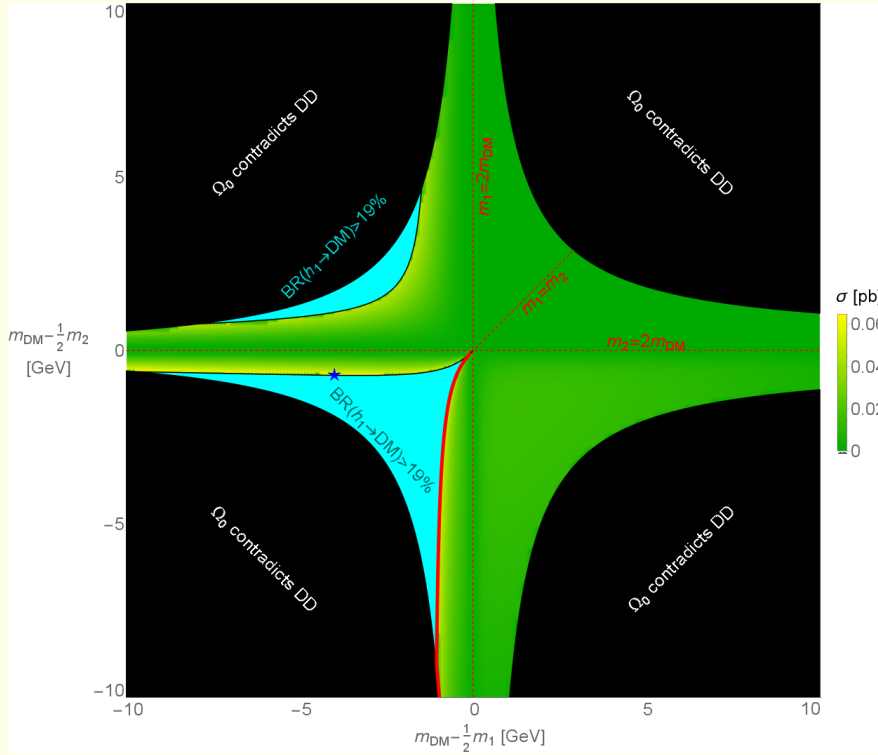
1. The following models were discussed and compared:
 - pGDM with complex scalar field S and $U(1)$ global symmetry softly broken by $\mu^2(S^2 + S^{*2})$,
 - VDM with gauged $U(1)_X$ and complex S ,
 - FDM with chiral χ and real S .
2. Vacuum stability in both models.
3. Direct detection efficiently suppressed in the pGDM model, $\sigma_{\text{SI}} \propto v_A^4$, as a consequence of A being a pseudo-Goldstone boson, 1-loop calculations were performed and adopted. For $q^2 = 0$ the 1-loop results are UV finite and vanish in the limit $m_A = m_{DM} \rightarrow 0$.
4. In some regions of (m_2, m_{DM}) space e^+e^- colliders might be useful to disentangle the models.



Pseudo Goldstone dark matter (pGDM)

$$\begin{aligned}
 m_2 &= 120.8 \text{ GeV} , & m_{\text{DM}} &= 58.9 \text{ GeV} , \\
 \sin \alpha &= 0.30 , & v_S &= 646 \text{ GeV} , \\
 \Gamma_1 &= 7.4 \cdot 10^{-3} \text{ GeV} , & \Gamma_2 &= 9.8 \cdot 10^{-3} \text{ GeV} , \\
 \text{BR}(h_1 \rightarrow \text{DM}) &= 19\% , & \text{BR}(h_2 \rightarrow \text{DM}) &= 95\% , \\
 \sigma &= 62 \text{ fb}
 \end{aligned}$$

$\sigma(e^+e^- \rightarrow AA Z)$ for the pGDM model. **Cyan:** forbidden region, $\text{BR}(h_1 \rightarrow \text{DM}) > 19\%$. **Gray:** small cross section, h_2 cannot be produced on-shell due to required energy exceeding $\sqrt{s}$. The **star** denotes the proposed benchmark point: $m_2 = 120.8 \text{ GeV}$, $m_{\text{DM}} = 58.9 \text{ GeV}$.

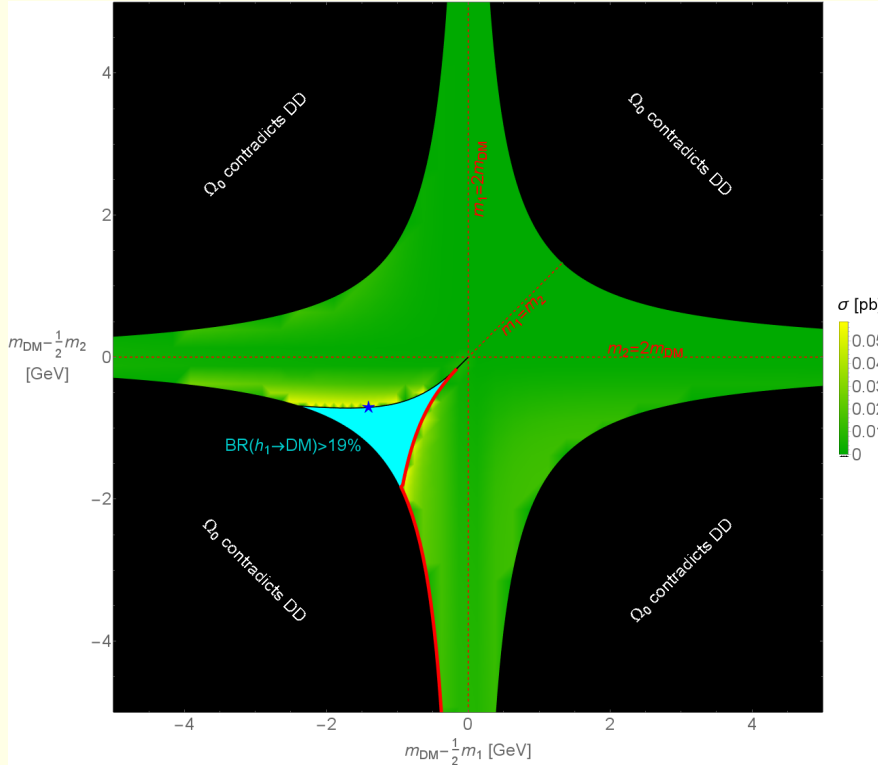


Vector dark matter (VDM)

$$\begin{aligned}
 m_2 &= 118.4 \text{ GeV} , & m_{\text{DM}} &= 58.5 \text{ GeV} , \\
 \sin \alpha &= 0.30 , & v_S &= 561 \text{ GeV} , \\
 \Gamma_1 &= 7.4 \cdot 10^{-3} \text{ GeV} , & \Gamma_2 &= 6.4 \cdot 10^{-3} \text{ GeV} , \\
 \text{BR}(h_1 \rightarrow \text{DM}) &= 18\% , & \text{BR}(h_2 \rightarrow \text{DM}) &= 92\% , \\
 \sigma &= 61 \text{ fb}
 \end{aligned}$$

$\sigma(e^+e^- \rightarrow XX Z)$ for the VDM model. **Black:** forbidden region, correct relic density cannot be obtained within the direct detection limits. **Cyan:** forbidden region, $\text{BR}(h_1 \rightarrow \text{DM}) > 19\%$. The **star** denotes the proposed benchmark point: $m_2 = 118.4 \text{ GeV}$, $m_{\text{DM}} = 58.5 \text{ GeV}$.

Backup slides



Fermion dark matter (FDM)

$$\begin{aligned}
 m_2 &= 123.6 \text{ GeV} , & m_{\text{DM}} &= 61.1 \text{ GeV} , \\
 \sin \alpha &= 0.30 , & v_S &= 76 \text{ GeV} , \\
 \Gamma_1 &= 7.4 \cdot 10^{-3} \text{ GeV} , & \Gamma_2 &= 5.9 \cdot 10^{-3} \text{ GeV} , \\
 \text{BR}(h_1 \rightarrow \text{DM}) &= 18\% , & \text{BR}(h_2 \rightarrow \text{DM}) &= 91\% , \\
 \sigma &= 59 \text{ fb}
 \end{aligned}$$

$\sigma(e^+e^- \rightarrow \psi\psi Z)$ in the FDM model. **Black:** forbidden region, correct relic density cannot be obtained within the direct detection limits. **Cyan:** forbidden region, $\text{BR}(h_1 \rightarrow \text{DM}) > 19\%$. The **star** denotes the proposed benchmark point: $m_2 = 123.6 \text{ GeV}$, $m_{\text{DM}} = 61.1 \text{ GeV}$.