

Huygens' principle and anomalously small radiation tails

Piotr Bizoń

Jagiellonian University, Kraków

Based on joint work with Tadek Chmaj and Andrzej Rostworowski

Outline:

- Introduction and motivation
- Late-time tails of linear and nonlinear waves
- Anomalous tails and huygensian systems

Myron Mathisson and Huygens' principle

A wave equation is said to satisfy **Huygens' principle** if:

- the solution at a point P depends only on the initial data at the intersection of the past light cone of P with the Cauchy hypersurface or, equivalently,
- the solution vanishes at all points which cannot be reached from the initial data by a null geodesic (there are no tails).

Hadamard's conjecture: The only Huygensian linear second-order hyperbolic equation of the form

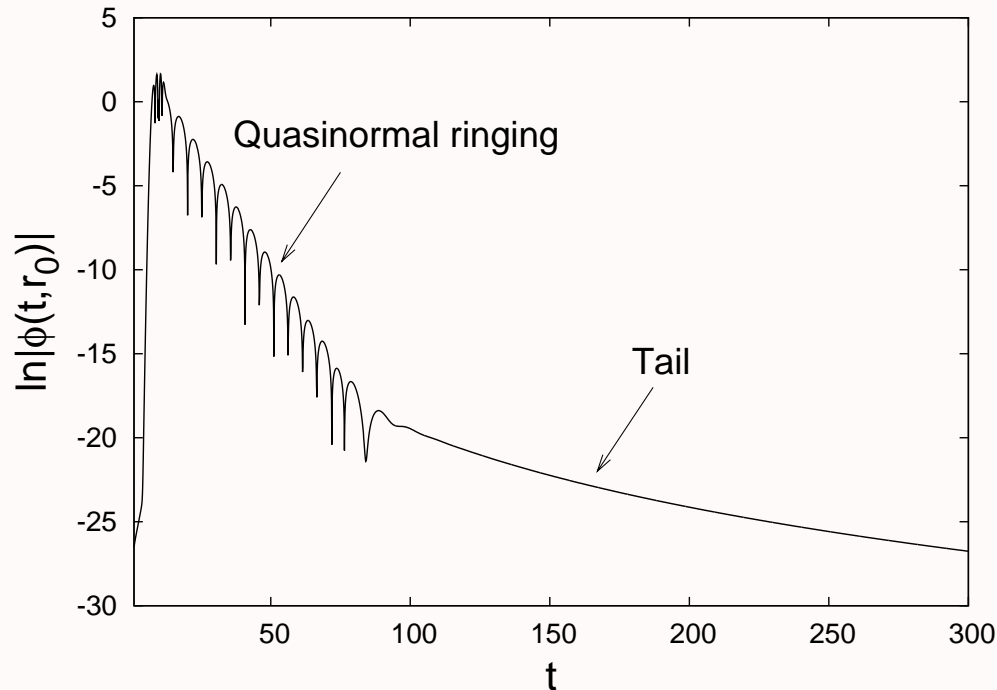
$$g^{\mu\nu}(x)\nabla_{\mu}\nabla_{\nu}\phi + A^{\mu}(x)\nabla_{\mu}\phi + B(x)\phi = 0$$

is the ordinary wave equation $\square\phi = 0$ in even dimensions ≥ 4 (modulo trivial transformations).

Mathisson (1939) proved this conjecture in the case of four dimensional Minkowski spacetime.

Our motivation

- Goal: understanding of relaxation to equilibrium for nonlinear wave equations
- Mechanism of relaxation: dissipation by dispersion (only for spatially unbounded domains)
- Example: formation of the [skyrmion](#)



Model and assumptions

$$\square\phi + V(x)\phi + N(t, x, \phi, \nabla\phi) = 0, \quad \square = \partial_t^2 - \Delta, \quad (t, x) \in \mathbb{R}^{1+d}$$

- Spherically symmetric smooth initial data with compact support
- Spatial dimension $d = \text{odd} \geq 3$
- Why $d > 3$?

(i)
$$\left(\partial_t^2 - \partial_r^2 - \frac{d-1}{r} \partial_r \right) \phi = \frac{1}{r^l} \left(\partial_t^2 - \partial_r^2 - \frac{2}{r} \partial_r + \frac{l(l+1)}{r^2} \right) (r^l \phi)$$
 where $l = (d-3)/2$.

- (ii) Some geometric wave equations in $3+1$ are equivalent to scalar wave equations in $d+1$ for $d > 3$. Example: Wave maps in $3+1$ dimensions.

$$\left(\partial_t^2 - \partial_r^2 - \frac{2}{r} \partial_r \right) \psi + \frac{\sin(2\psi)}{r^2} = 0 \text{ after substitution } \psi = r\phi \text{ becomes}$$
$$\left(\partial_t^2 - \partial_r^2 - \frac{4}{r} \partial_r \right) \phi + \frac{4}{3} \phi^3 + \text{higher order terms} = 0$$

Tools (elementary)

- Solution of $\square\phi = 0$ with smooth radial data ($d = 2l + 3$):

$$\phi(t, r) = \frac{1}{r^{2l+1}} \sum_{k=0}^l \frac{2^{k-l}(2l-k)!}{k!(l-k)!} r^k \left(a^{(k)}(t-r) - (-1)^k a^{(k)}(t+r) \right) \quad (1)$$

- Solution of $\square\phi = F(t, r)$ with zero data (Duhamel formula):

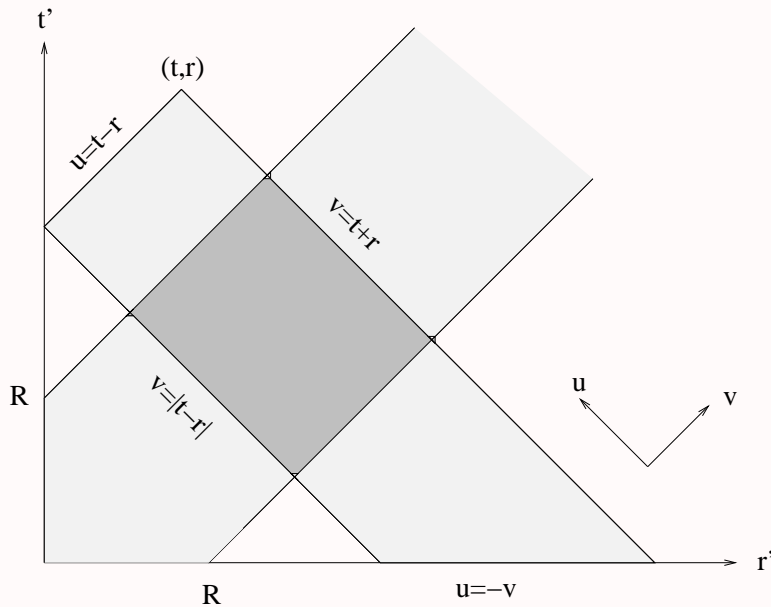
$$\phi(t, r) = \frac{1}{2r^{l+1}} \int_0^t d\tau \int_{|t-r-\tau|}^{t+r-\tau} K(t, r; \tau, \rho) F(\tau, \rho) d\rho,$$

where $K(t, r; \tau, \rho) = \rho^{l+1} P_l(\mu)$ and $\mu = \frac{r^2 + \rho^2 - (t - \tau)^2}{2r\rho}$.

In terms of $u = \tau - \rho$ and $v = \tau + \rho$ this becomes

$$\phi(t, r) = \frac{1}{4r^{l+1}} \int_{|t-r|}^{t+r} dv \int_{-v}^{t-r} K(t, r; u, v) F(u, v) du \quad (2)$$

Simplification due to Huygens



$$F(u, v) = 0 \text{ for } |u| > R$$

$$\phi(t, r) = \frac{1}{4r^{l+1}} \int_{-\infty}^{\infty} du \int_{t-r}^{t+r} K(t, r; u, v) F(u, v) dv$$

Linear tails

$$\square\phi + \lambda V\phi = 0, \quad (\phi(0, r), \partial_t\phi(0, r)) = (f(r), g(r))$$

Perturbation series:

$$\phi = \phi_0 + \lambda\phi_1 + \lambda^2\phi_2 + \dots$$

Iteration:

$$\square\phi_0 = 0, \quad \square\phi_1 = -V\phi_0, \quad \square\phi_2 = -V\phi_1, \text{ etc.}$$

Assuming that $V(r) \sim r^{-\alpha}$ ($\alpha > 2$) for $r \rightarrow \infty$ we get (at fixed r and $t \rightarrow \infty$)

$$\phi_1(t, r) = \frac{C(l, \alpha)}{t^{\alpha+2l}} \left[A + \mathcal{O}\left(\frac{1}{t}\right) \right]$$

$$C(l, \alpha) = -\frac{2^{\alpha+2l-1}}{(2l+1)!!} \left(\frac{\alpha-3}{2}\right)^l \left(\frac{\alpha}{2}\right)^{\bar{l}} \quad \text{and} \quad A = \int_{-\infty}^{+\infty} a(u) du$$

This was first derived by Ching et al. (1995). Notation:

$$\begin{aligned} x^{\underline{0}} &:= 1, & x^{\underline{k}} &:= x \cdot (x-1) \cdot \dots \cdot (x-(k-1)), & k > 0, \\ x^{\bar{0}} &:= 1, & x^{\bar{k}} &:= x \cdot (x+1) \cdot \dots \cdot (x+(k-1)), & k > 0. \end{aligned}$$

How good is the first order approximation?

- All higher-order terms $\phi_n(t, r)$ decay in the same manner as $\phi_1(t, r)$
- Convergence of the perturbation series? (proved for $d = 3$, Szpak)
- The perturbation series is asymptotic to the solution (for any $d \geq 3$)

$$\phi(t, r) - \lambda\phi_1(t, r) \sim \mathcal{O}(\lambda^2)t^{-(\alpha+2l)} \quad (3)$$

Numerical evidence.

Example:

$$V = \lambda \frac{\tanh^{\alpha+2} r}{r^\alpha}$$

$$\lambda = 0.1$$

$\alpha = 4$		Theory	Numerics
$d = 3$	Exponent	4	4.00002
$(l = 0)$	Amplitude	-0.3545	-0.3320
$d = 5$	Exponent	6	5.9999
$(l = 1)$	Amplitude	-0.2363	-0.2318
$d = 7$	Exponent	8	7.9999
$(l = 2)$	Amplitude	0.1418	0.1404

Nonlinear tails

$$\square \phi - \phi^p = 0, \quad (\phi(0, r), \partial_t \phi(0, r)) = (\varepsilon f(r), \varepsilon g(r))$$

Perturbation series:

$$\phi = \varepsilon \phi_0 + \varepsilon^2 \phi_1 + \varepsilon^3 \phi_2 + \dots$$

Iteration:

$$\square \phi_0 = 0, \quad \square \phi_p = \phi_0^p, \text{ etc.}$$

At fixed r and $t \rightarrow \infty$

$$\phi_p(t, r) = \frac{\tilde{C}(l, p)}{t^{(l+1)p-1}} \left[\tilde{A} + \mathcal{O}\left(\frac{1}{t}\right) \right]$$

$$\tilde{C}(l, p) = (-1)^l \frac{2^{(l+1)(p+1)-1}}{(2l+1)!!} [(l+1)(p-1) - 2]^l \quad \text{and} \quad \tilde{A} = \int_{-\infty}^{+\infty} [a^{(l)}(u)]^p du$$

The perturbation series is convergent for $d = 3$ ($l = 0$) (Szpak) and, at least, asymptotic for $d > 3$ ($l > 0$).

Competition between linear and nonlinear tails

$$\text{linear tail} \sim t^{-(\alpha+2l)} \quad \text{vs.} \quad \text{nonlinear tail} \sim t^{1-(l+1)p}$$

Thus

$$\phi(t, r) \sim t^{-\gamma}, \quad \gamma = \min\{\alpha + 2l, (l + 1)p - 1\}$$

Example: relaxation of the skyrmion

$$(w_F \dot{F})' - (w_F F')' + \sin(2F) + \alpha^2 \sin(2F) \left(\frac{\sin^2 F}{r^2} + F'^2 - \dot{F}^2 \right) = 0,$$

where $w_F = r^2 + 2\alpha^2 \sin^2 F$. Asymptotically $F(t, r) \rightarrow S(r) = \text{skyrmion}$.
The perturbation $\phi(t, r) = r w_S^{-1/2} [F(t, r) - S(r)]$ satisfies

$$\left(\partial_t^2 - \partial_r^2 - \frac{4}{r} \partial_r \right) \phi + V \phi + \frac{4}{3} \phi^3 + \text{higher order terms} = 0, \quad V(r) \sim r^{-6}.$$

We have $d = 5$ ($l = 1$), $\alpha = 6$, $p = 3$, thus

$$\text{linear tail} \sim t^{-8}, \quad \text{nonlinear tail} \sim t^{-5}$$

Anomalous tails

- Linear case ($V(r) = \lambda r^{-\alpha}$ for $r > R$)

$$C(l, \alpha) \propto \left(\frac{\alpha - 3}{2}\right)^l = 0 \quad \text{if } \alpha = \text{odd integer} \leq 2l + 1$$

No tail in the first order! \Leftrightarrow The system $\square\phi_0 = 0$, $\square\phi_1 = -V\phi_0$ is Huygensian
Solving the second order equation $\square\phi_2 = -V\phi_1$ we get

$$\phi(t, r) \approx \lambda^2 \phi_2(t, r) = \lambda^2 \frac{D(l, \alpha)}{t^{2(\alpha+l-1)}} \left[A + \mathcal{O}\left(\frac{1}{t}\right) \right]$$

- Nonlinear case ($\square\phi = \phi^p$):

$$C(l, p) \propto [(l+1)(p-1) - 2]^l = 0 \quad \text{if } p = 2 \quad \text{and} \quad l \geq 1$$

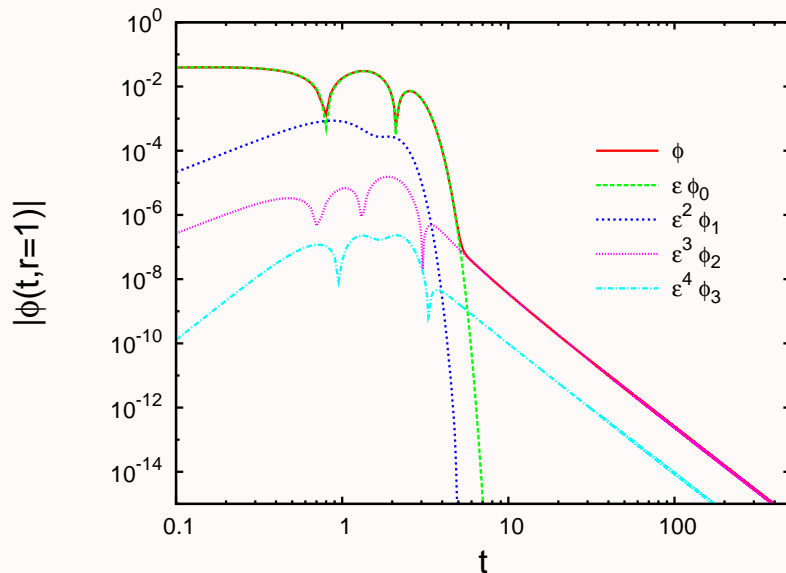
No tail in the first order! \Leftrightarrow The system $\square\phi_0 = 0$, $\square\phi_1 = \phi_0^2$ is Huygensian
Solving the second order equation $\square\phi_2 = 2\phi_0\phi_1$ we get

$$\phi(t, r) \approx \varepsilon^3 \phi_2(t, r) \sim \varepsilon^3 \frac{c(l)}{t^{3l+1}}$$

Examples of anomalous tails

- Spherically-symmetric Yang-Mills in $3 + 1$ dimensions

$$\left(\partial_t^2 - \partial_r^2 - \frac{4}{r}\partial_r\right)\phi + 3\phi^2 + r^2\phi^3 = 0$$



Near timelike infinity

$$\phi(t, r) \approx \varepsilon^3 \phi_2(t, r) \sim \varepsilon^3 c t^{-4}$$

$$c = -8 \int_{-\infty}^{+\infty} a(u) a'(u)^2 du$$

Tails on Schwarzschild background

- Thoroughly studied in $d = 3$. Price's law: $\phi(t, r) \sim t^{-3}$
- Almost unexplored for $d > 3$

$$\partial_t^2 \psi - \partial_x^2 \psi + U(x)\psi = 0, \quad dr/dx = 1 - 1/r^{d-2}$$

For $x \rightarrow \infty$

$$U(x) \sim \frac{l(l+1)}{x^2} + \frac{a}{x^d} + \frac{b}{x^{2d-2}}, \quad l = j + \frac{d-3}{2}$$

Conjecture: For odd $d > 3$ the tail behaves as $t^{-\gamma}$, where

$$\gamma = \alpha + 2l = 2d - 2 + 2\left(j + \frac{d-3}{2}\right) = 2j + 3d - 5$$

In higher dimensions black holes become bald very fast