

Physics of Condensed Matter I



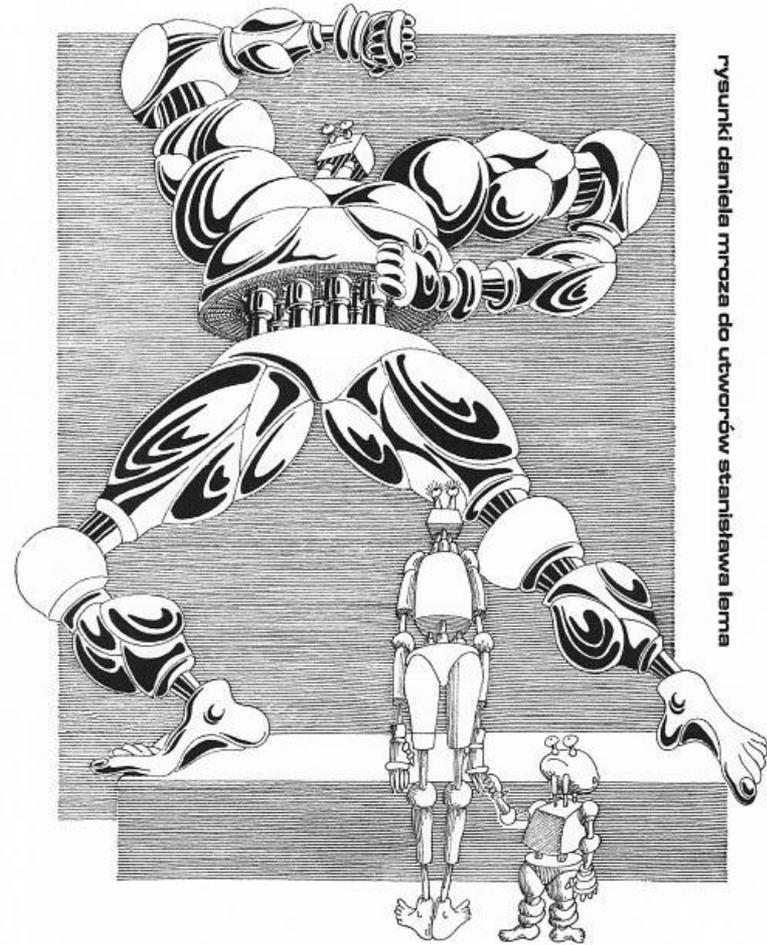
1100-4INZ`PC

Faculty of Physics UW

Jacek.Szczytko@fuw.edu.pl

Summary of the lecture

1. Quantum mechanics
2. Optical transitions
3. Lasers
4. Optics
5. Molecules
6. Properties of molecules
7. Crystals
8. Crystallography
9. Solid state
10. Electronic band structure
11. Effective mass approximation
12. Electrons and holes
13. Carriers
14. Transport
15. Optical properties of solids



rysunki daniela mroza do utworów stanisława lemna

Summary of the lecture

1. Quantum mechanics

Revision of quantum mechanics: quantum states, quantum numbers, quantum well. Perturbation theory in quantum mechanics. Electric and magnetic field.

2. Optical transitions

Spin. Spin-orbit interaction. Magnetic ordering. Time-dependent perturbation theory. Electric dipole transition, selection rules.

3. Lasers

Einstein coefficients. Photo absorption, spontaneous emission and induced emission. Population inversion for laser action.

4. Optics

Revision of optics. Maxwell equations. Polarizability. Refractive index. Lorentz model of refractive index. Optical properties of metals, plasma frequency.

5. Molecules

Chemical bonds (σ , π). H_2^+ molecule. Born-Oppenheimer approximation. Hartree-Fock method. Linear combination of atomic orbitals. Diatomic molecules.

Summary of the lecture

6. Properties of molecules

Molecular spectra. Electronic, vibrational and rotational transitions for diatomic molecule.

7. Crystals

Chemical bonds, hybridization (s,p,d). Symmetry of bonds. Condensation. Crystal structure. Primitive translation vectors. Bravais lattice.

8. Crystallography

Basic elements of crystallography. Miller indices. Scattering of X-rays. Atomic form factor. Reciprocal lattice.

9. Solid state

Drude model of metals, mobility, conductivity. One-electron model. Hartree-Fock approximation. Periodic potentials and Bloch's theorem. Empty lattice approximation. Brillouin zones.

10. Electronic band structure

Tight-binding approximation, Linear Combination of Atomic Orbitals (LCAO). Electronic band structure.

Summary of the lecture

11. Effective mass approximation

kp perturbation theory. Effective mass. Electrons and holes.

12. Electrons and holes

Quantum statistics. Fermi-Dirac distribution. Bose-Einstein distribution. Fermi energy in solid state. Density of states in 3D, 2D and 1D.

13. Carriers

Carriers concentration: intrinsic. Doping. Acceptors and donors. Hydrogen-like atom model. p-n junction.

14. Transport

Carriers in external electric field. Carriers in external electric and magnetic field. Conductivity and resistivity tensors. Hall effect.

15. Optical properties of solids

Absorption edge. Joint density of states. Excitons, trions in quantum wells. Quantum dots.

Assessment criteria:

Quiz (kartkówka) 20p.

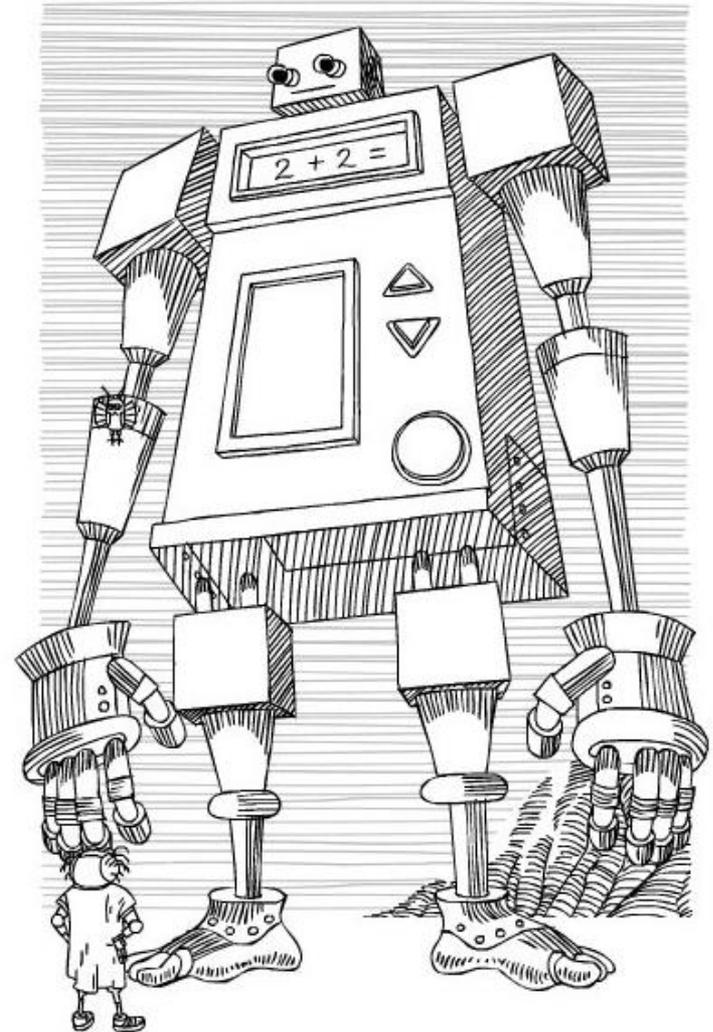
2 tests (3 exercises) $2 \times 30p = 60p$

Activity in exercises (dr Aneta Drabińska) 20p

passing grades: minimum 50p/100p

Exam: final test (40p) and oral exam

oral exam – minimum 70p/140p



Assessment criteria:

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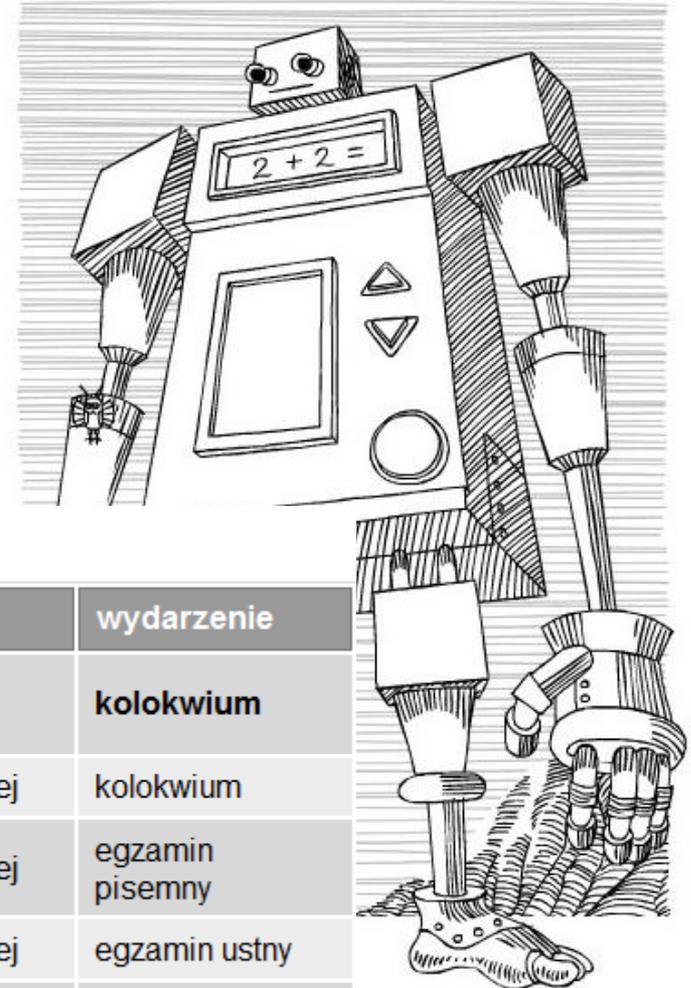
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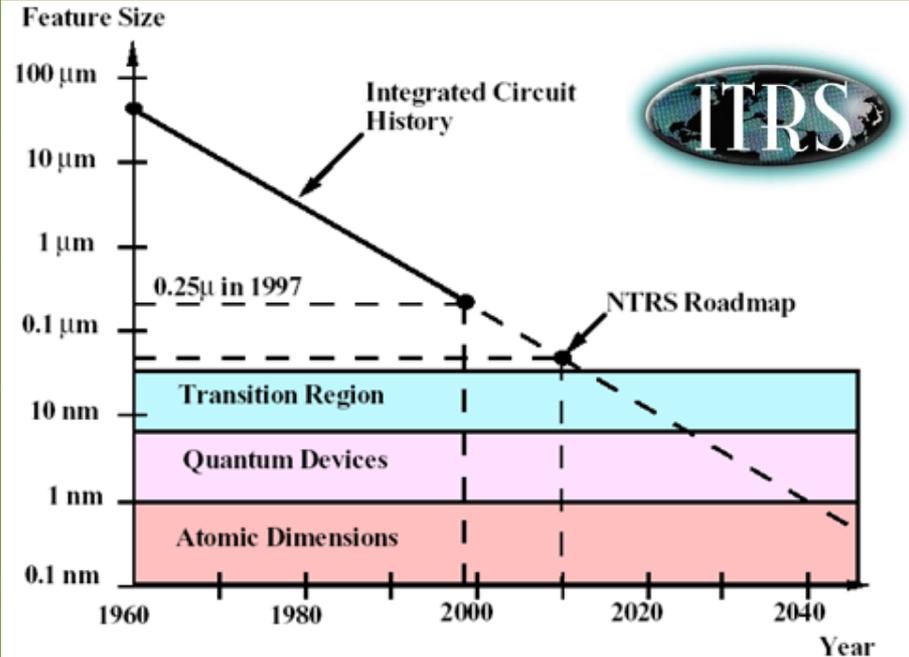
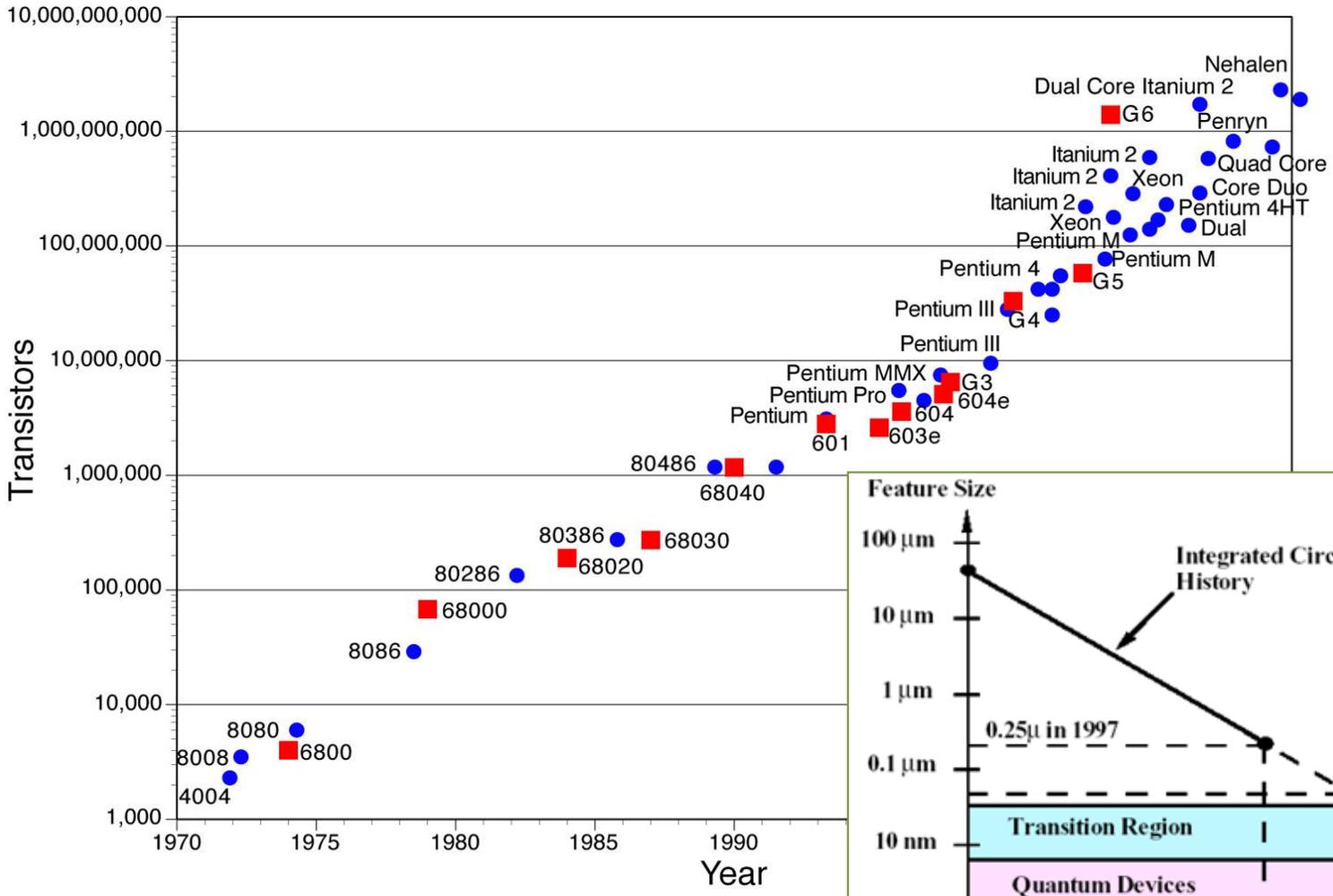
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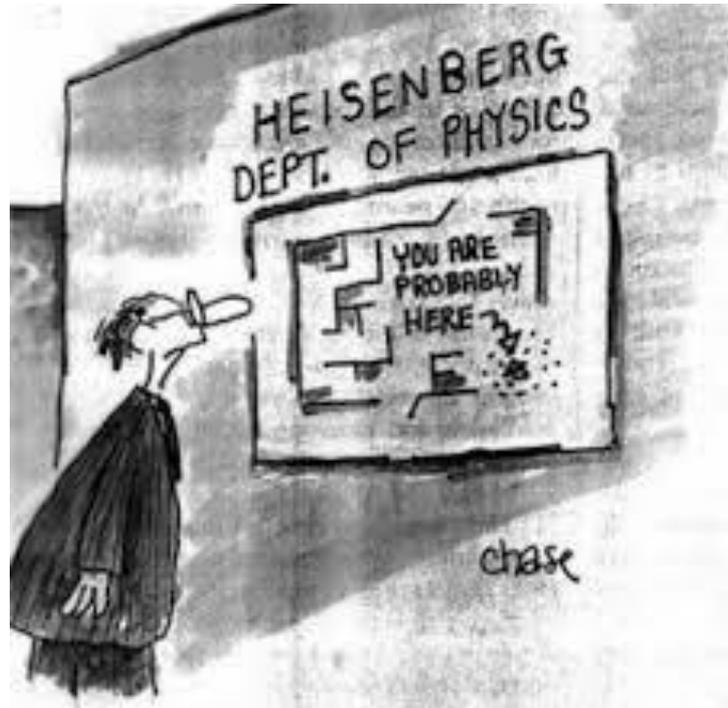
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2016.01.11	09:00 – 13:00	1.40	Fizyka materii skondensowanej	kolokwium
2016.01.29	14:00 – 17:00	Cyklotron A	Fizyka materii skondensowanej	egzamin pisemny
2016.02.01	09:00 – 13:00	1.37	Fizyka materii skondensowanej	egzamin ustny
2016.02.02	09:00 – 13:00	1.37	Fizyka materii skondensowanej	egzamin ustny



NANO era



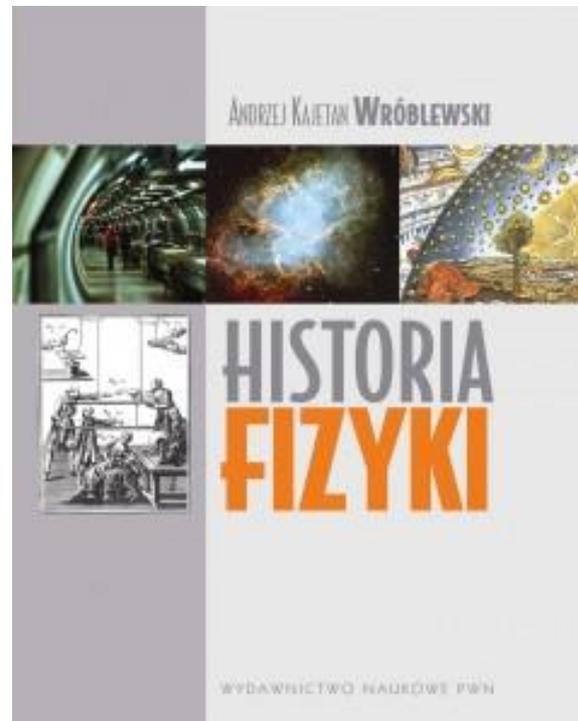
Quantum mechanics



The history

History of physical sciences 1100-HPS-OG, 1101-3N`HPS

Prof. dr hab Andrzej Kajetan Wróblewski



Historia fizyki. Od czasów najdawniejszych do współczesności
Wydawnictwo Naukowe PWN

The Math and the Nature?

Scientific method:

The conversation with the Nature must be hold in the language of mathematics, otherwise nature does not answer our questions.

prof. **Michał Heller**

Dialog z przyrodą musi być prowadzony w języku matematyki, w przeciwnym razie przyroda nie odpowiada na nasze pytania.

prof. **Michał Heller**



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The history



Sir Isaac Newton (4 January 1643 – 31 March 1727)



Michael Faraday, FRS (September 22, 1791 – August 25, 1867)

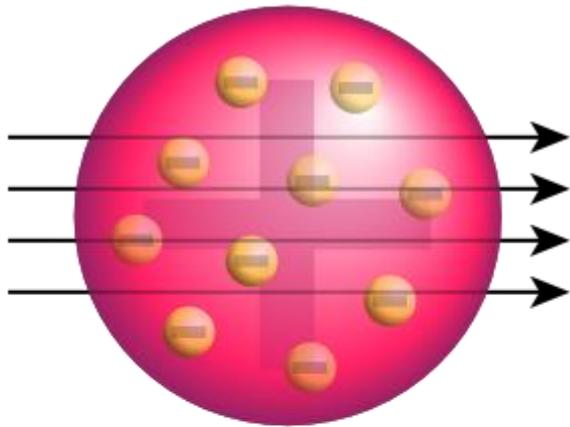


James Clerk Maxwell (13 June 1831 – 5 November 1879)

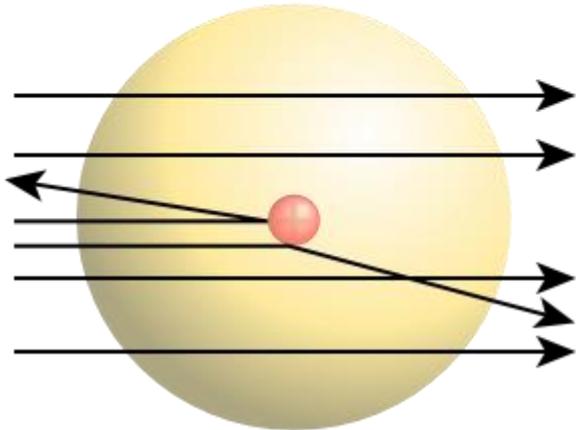
<http://www-personal.umich.edu/~jbouj/money1.htm>

The history

- In the XIX century: the matter is granular, the energy (mostly e-m) is a wave



The plum pudding model (J.J. Thomson 1904)



1909 gold foil experiment of Hans Geiger and Ernest Marsden.

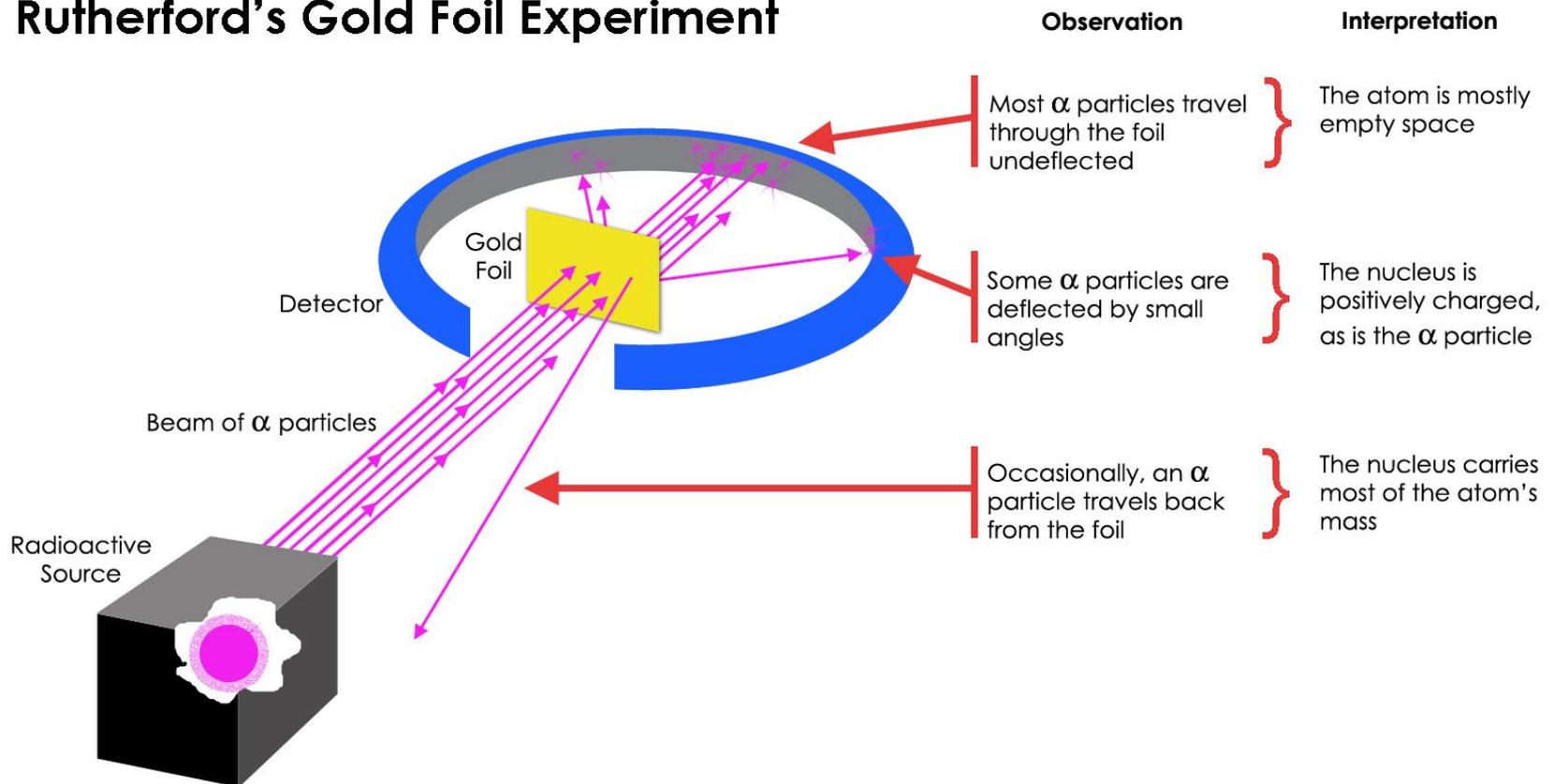
Interpreted by Ernest Rutherford in 1911

<http://rutherford.pl/informacje/doswiadczenie-rutherforda/>

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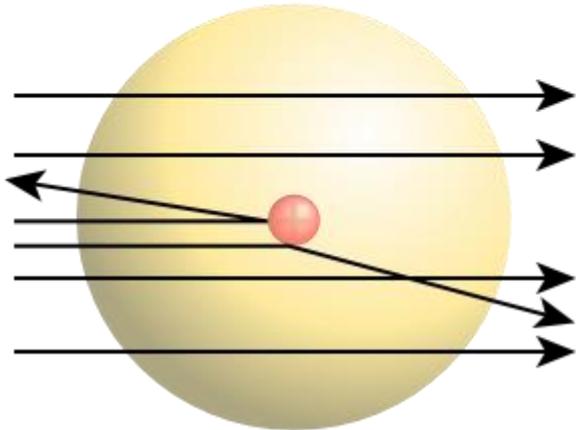
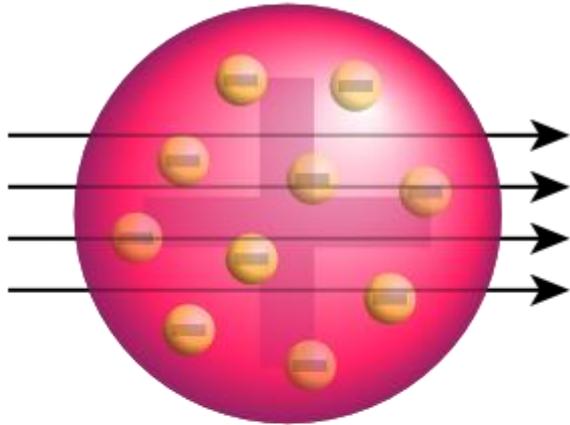
Rutherford's Gold Foil Experiment



<http://socratic.org/questions/what-were-the-aims-of-rutherford-experiment>

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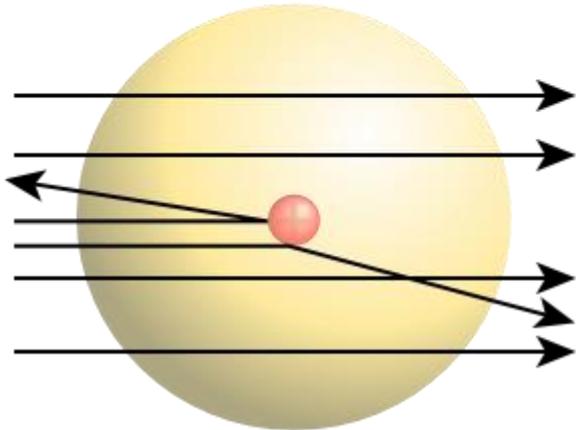
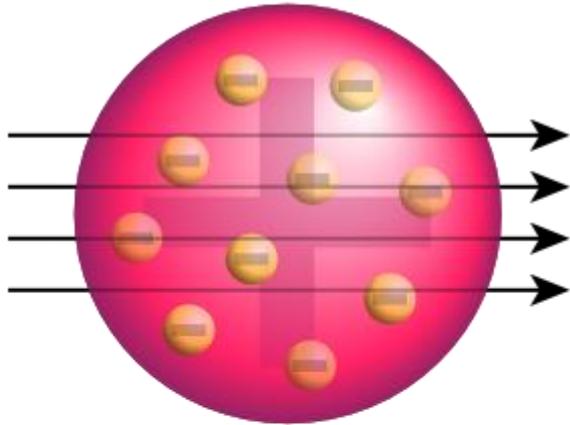
Sir Ernest Rutherford (1871 -1937)

Atoms	→ 10^{-10} m
Nucleus	→ 10^{-14} m
Proton, neutron	→ 10^{-15} m

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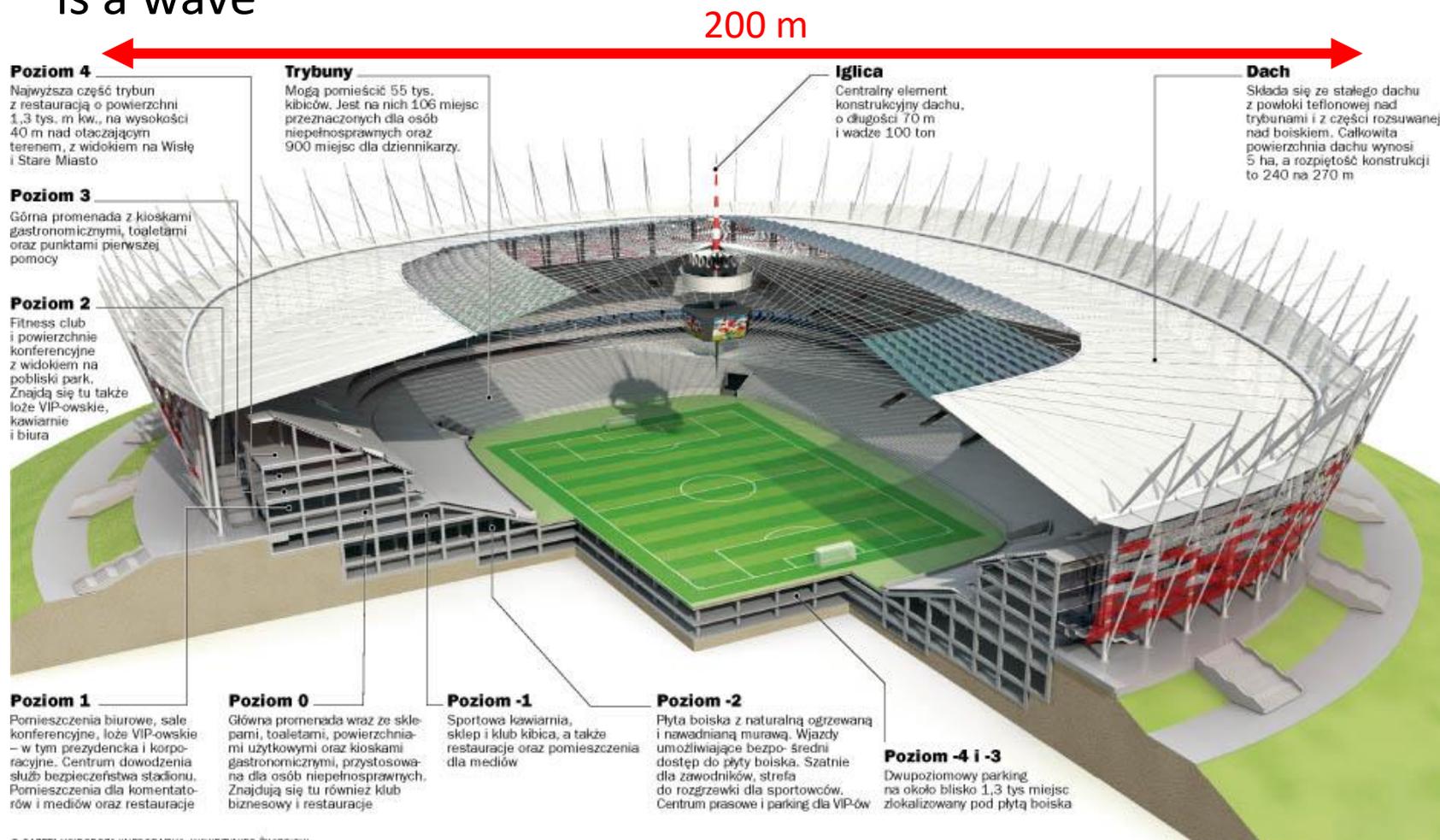
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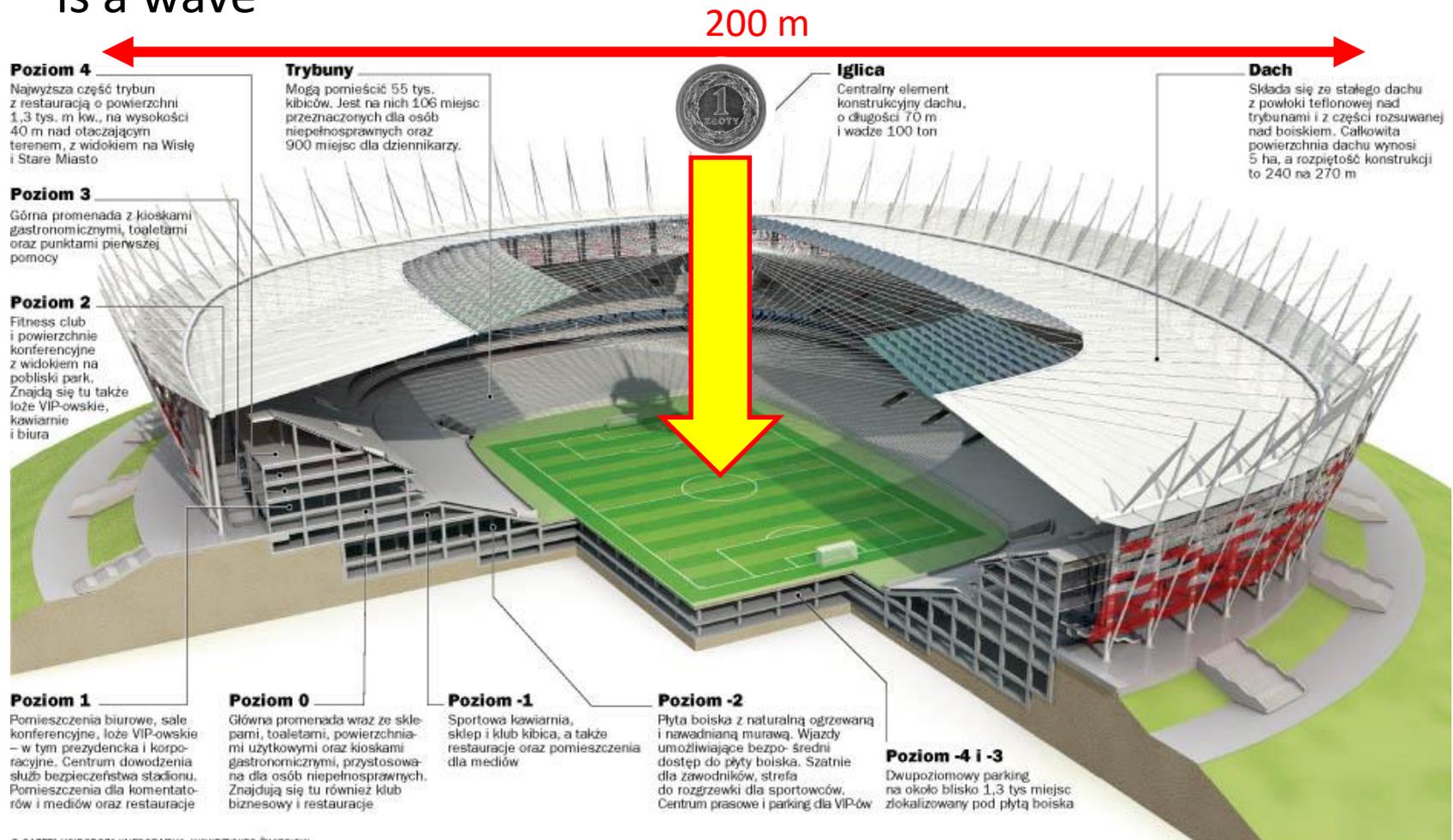
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The history

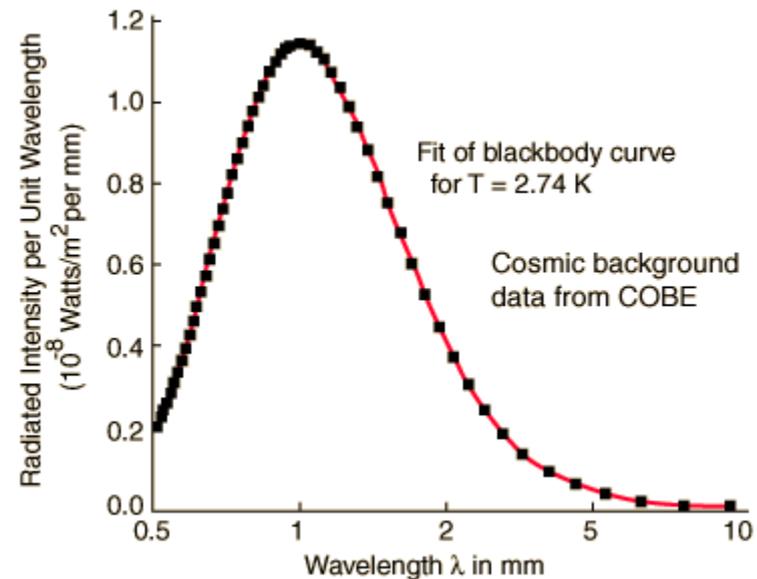
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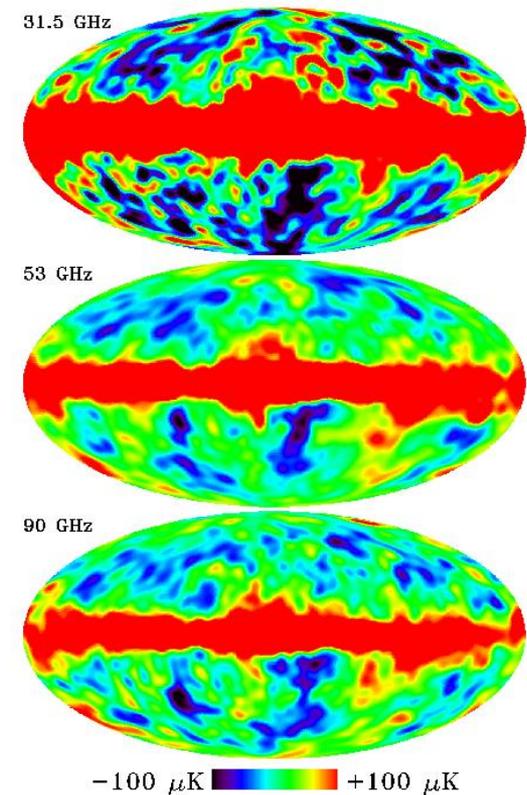
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- In the XIX century: the matter is granular, the energy (mostly e-m) is a wave
- Problems **NOT** solved
 - Black body radiation
 - Photoelectric effect
 - Origin of spectral lines of atoms



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 - Black body radiation spectrum
 - Photoelectric effect
 - Origin of spectral lines of atoms



The history



Albert Einstein, 5 Israeli Lira (1968)



Erwin Schrödinger, 1000 Austrian Schilling (1983)



Niels Bohr, 500 Danish Kroner



Lord Ernest Rutherford, 100 New Zealand Dollars

The history

- In the XIX century: the matter is granular, the energy (mostly e-m) is a wave
- In the XX century: the matter is (also) a wave and the energy is (also) granular (corpuscular)

The history

- In the XX century: the matter is (also) a wave and the energy is (also) granular (corpuscular)

- **Solved** problems:

- Black body radiation spectrum (Planck 1900, Nobel 1918)
- Photoelectric effect (Einstein 1905, Nobel 1922)
- Origin of spectral lines of atoms (Bohr 1913, Nobel 1922)

$$p = h / \lambda$$

- *Photons* – *energy*: $E = h \nu$ ($h = 6.626 \times 10^{-32} \text{ J s} = 4.136 \times 10^{-15} \text{ eV s}$)

- *momentum*: $p = E / c = h / \lambda$

Count Dooku's Geonosian solar sailer



light mill - Crookes radiometer

Trochę historii

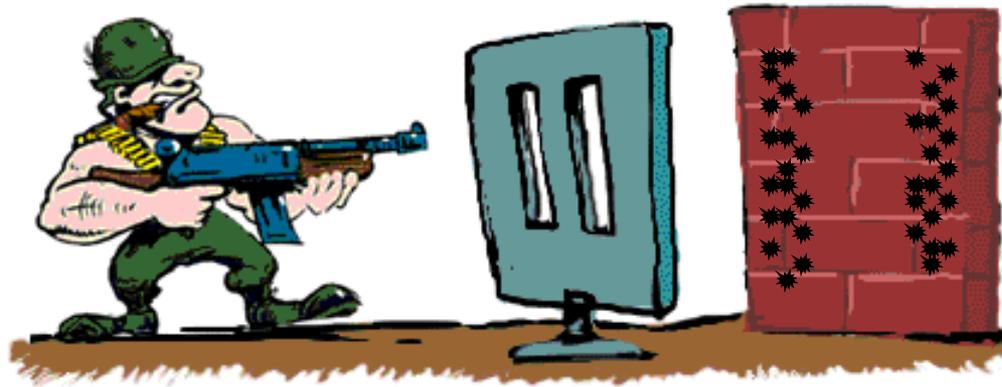
In the XX century: the matter is (also) a wave and the energy is (also) granular (corpuscular)

Matter waves – De Broglie 1924 (Nobel 1929), experiments G.P. Thomson L.H. Germer and C.J. Davisson (Nobel 1937)

$$\lambda = h / p$$

$$p = h / \lambda$$

classically



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quantum

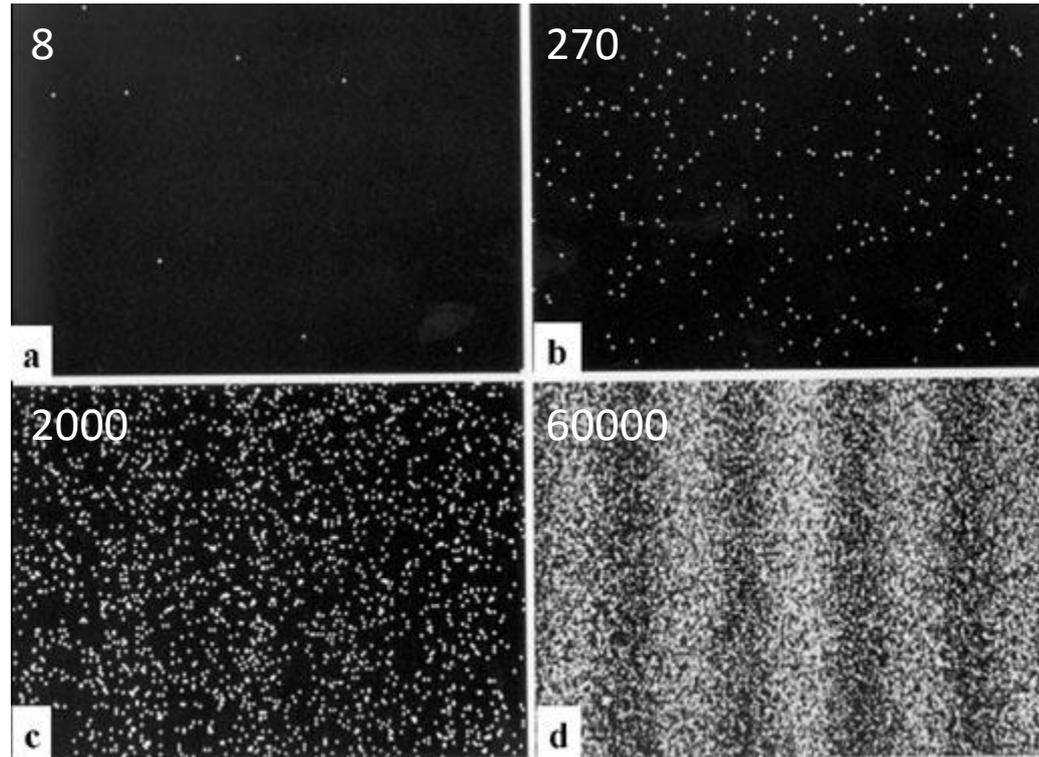


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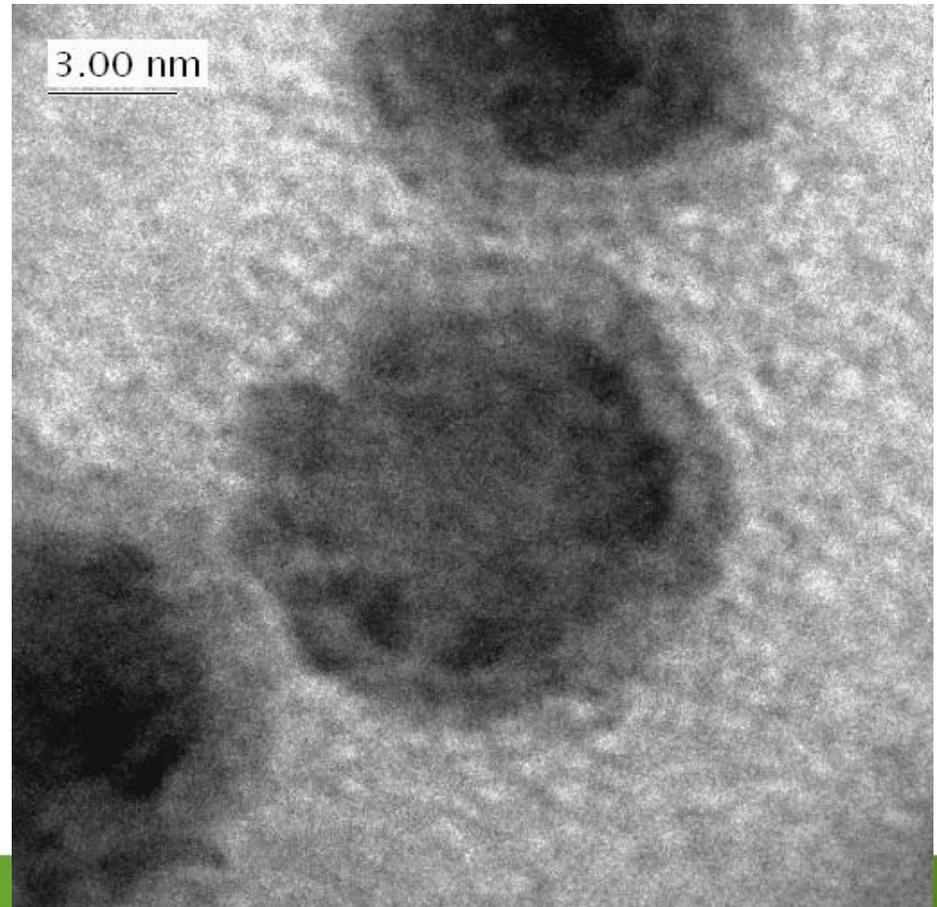
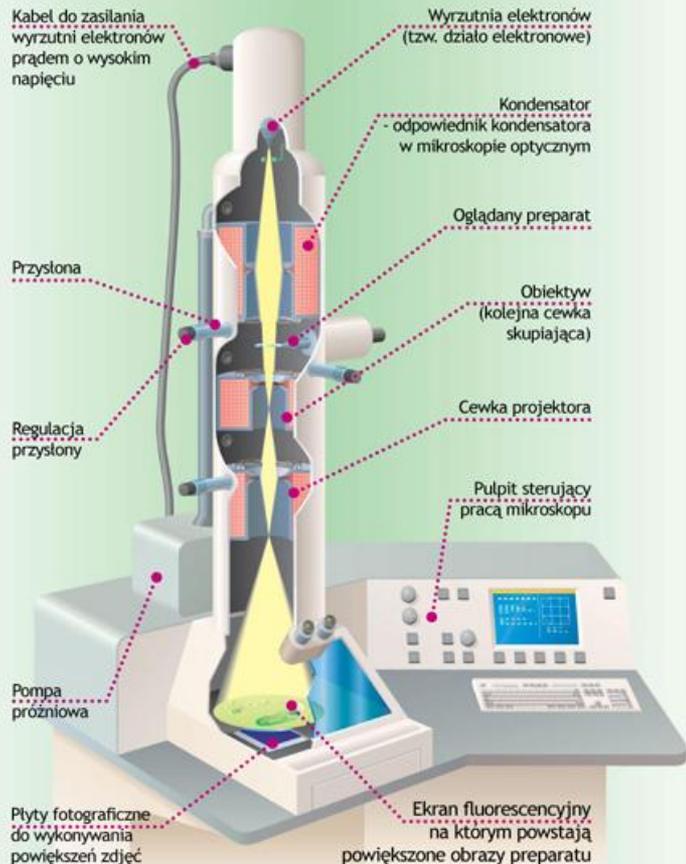
Single-electron events build up over a 20 minute exposure to form an interference pattern in this double-slit experiment by Akira Tonomura and co-workers. (a) 8 electrons; (b) 270 electrons; (c) 2000 electrons; (d) 60,000. A video of this experiment will soon be available on the web www.hqrd.hitachi.co.jp/em/doubleslit.html.

Trochę historii

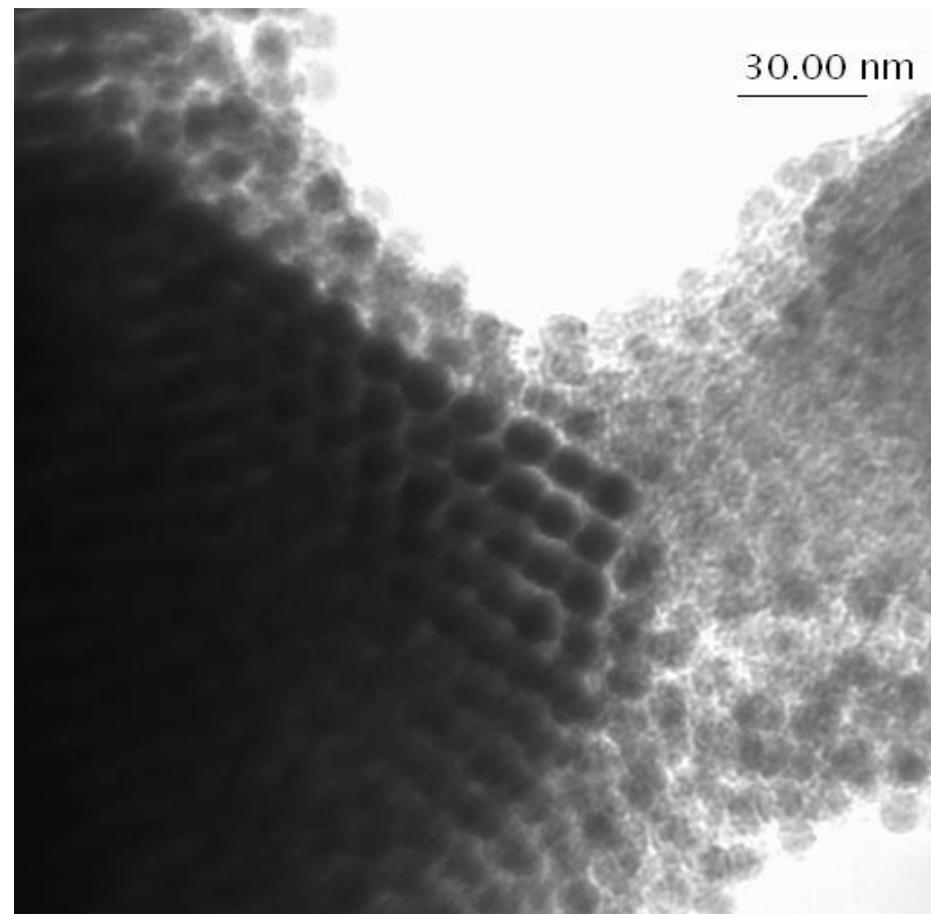
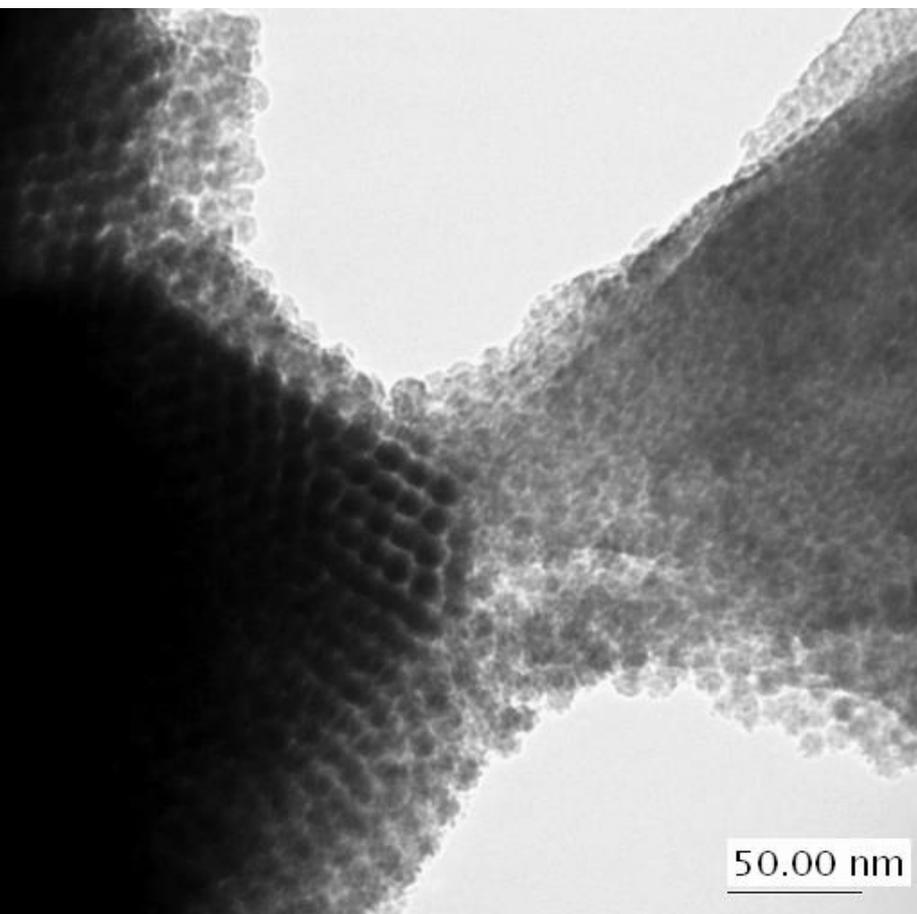
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Magnetic QDs



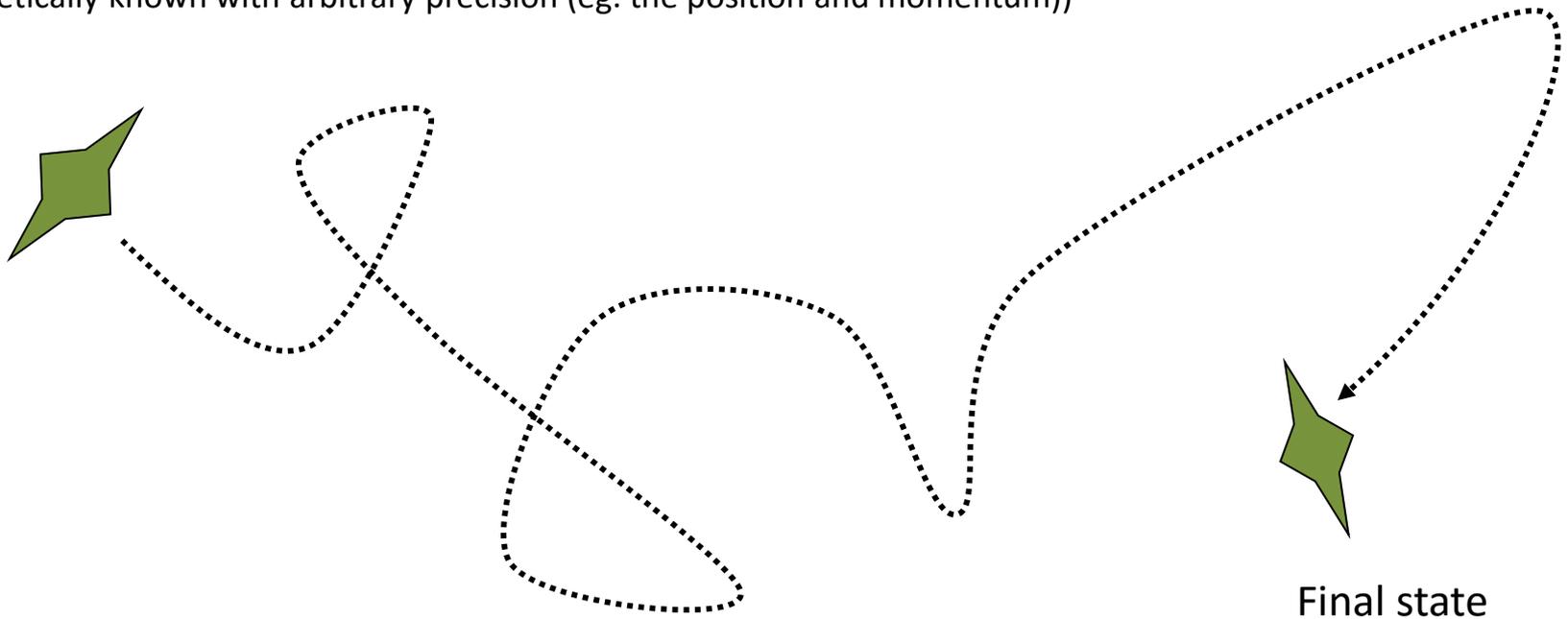
Synteza: Paweł Majewski, TEM: Jolanta Borysiuk

Classical and quantum universe

Classical mechanics:

Initial conditions

(theoretically known with arbitrary precision (eg. the position and momentum))



Evolution (so-called. *state space*)
(Lagrange, Hamilton)

Classical and quantum universe

Classical mechanics:

Initial conditions

(theoretically known with arbitrary precision (eg. the position and momentum))

$$\vec{r}(t=0) = [x_0, y_0, z_0]$$

$$\vec{v} = [v_x, 0, v_z]$$

$$\vec{r}(t) = \left[x_0 + v_x t, y_0, z_0 + v_z t - \frac{gt^2}{2} \right]$$



Final state

Evolution (so-called. *state space*)
(Lagrange, Hamilton)

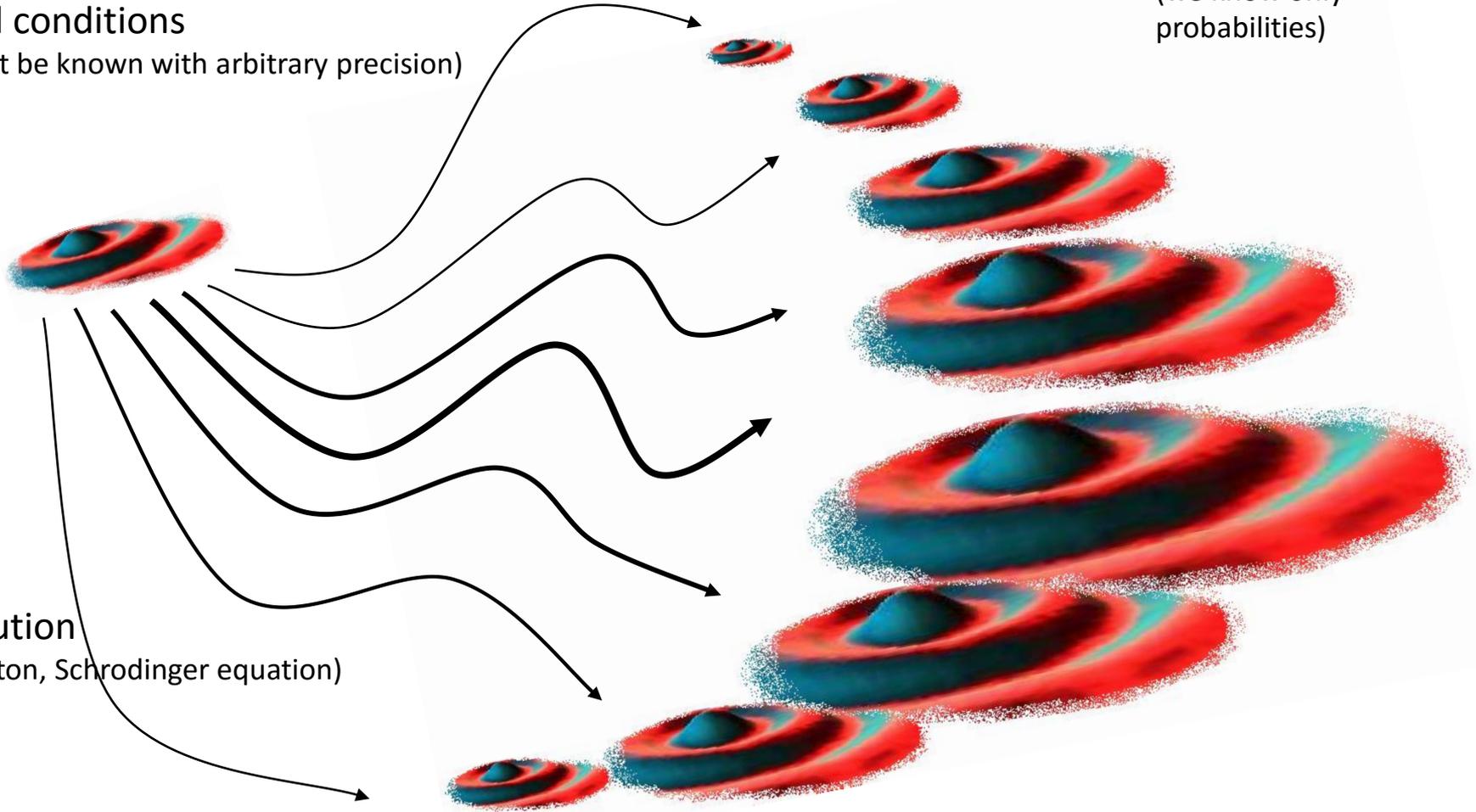
Classical and quantum universe

Quantum mechanics:

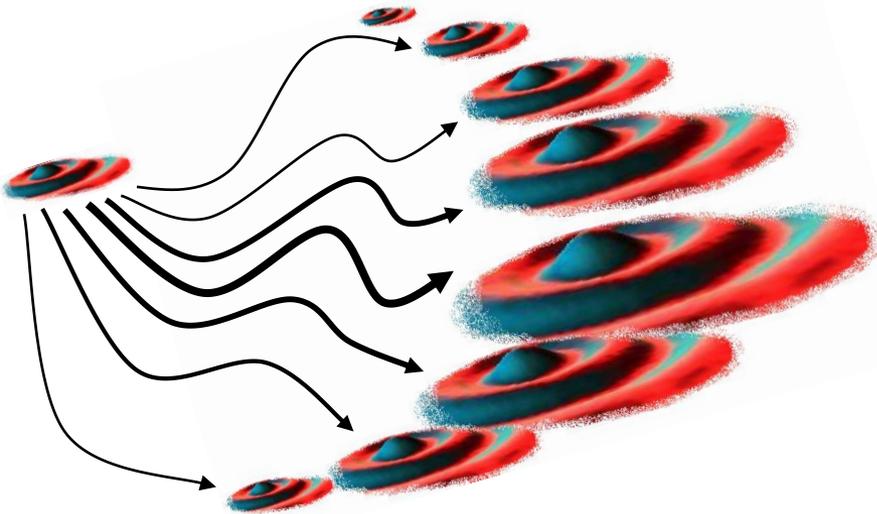
Initial conditions
(cannot be known with arbitrary precision)

Final states
(we know only probabilities)

Evolution
(Hamilton, Schrodinger equation)



Classical and quantum universe



Quantum state must be defined in the whole space
→ Wave function

$$\Psi_n(\vec{r}, t)$$

n – quantum numbers

Remark 1: QUANTUM NUMBERS determine wave function:

Remark 2: the wave function is defined in the WHOLE SPACE, in that sense its evolution describes all of the possible trajectories (in space and time) of particles :

Remark 3: a LINEAR COMBINATION of wave functions is also a wave function (the principle of superposition)

Remark 4: evolution of the wave function is DETERMINISTIC. However, by measurement a superposition of several eigenstates reduces to a single eigenstate (*wave function collapse*)

Remark 5: quantum particles are INDISTINGUISHABLE

Vector space of wave functions (Hilbert space), operators, eigen states and eigen functions, etc.

Classical and quantum universe

Remark 1: QUANTUM NUMBERS determine wave function::

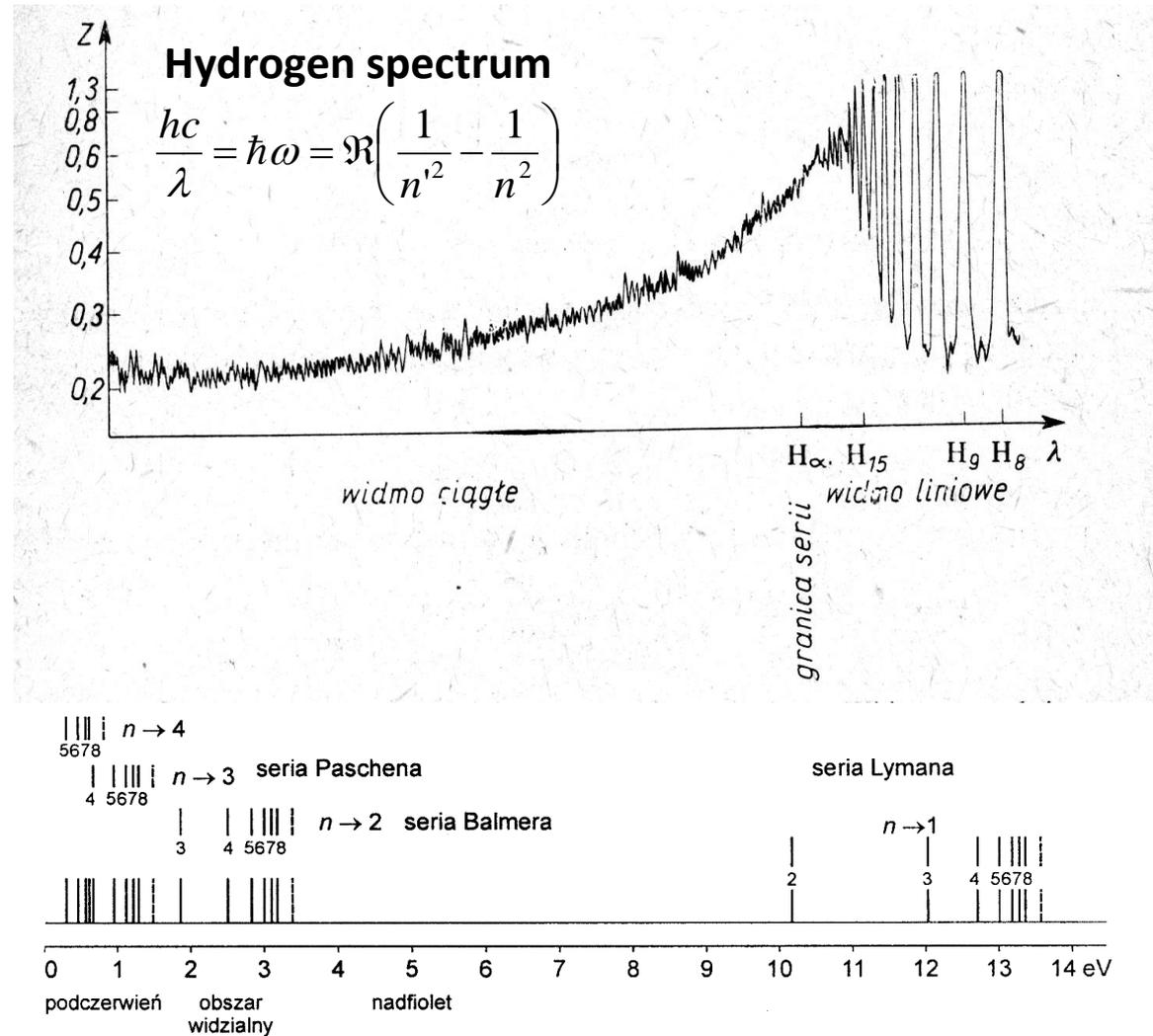
$$\Psi_n(\vec{r}, t)$$



n – quantum numbers

Exaples of „quantum”

mass, charge, energy ("energy levels" in an atom in the crystal), momentum (orbital, spin), the projection of angular momentum, minimum light intensity of the energy $E = h \nu$, polarization of light, etc ..

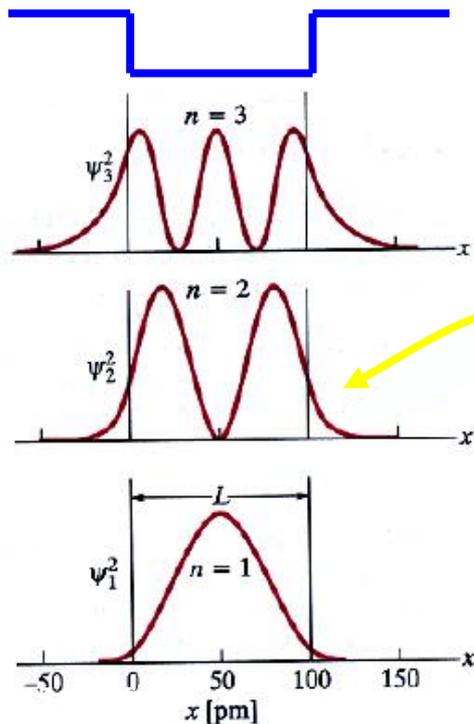


Classical and quantum universe

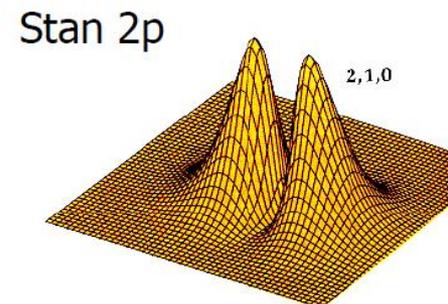
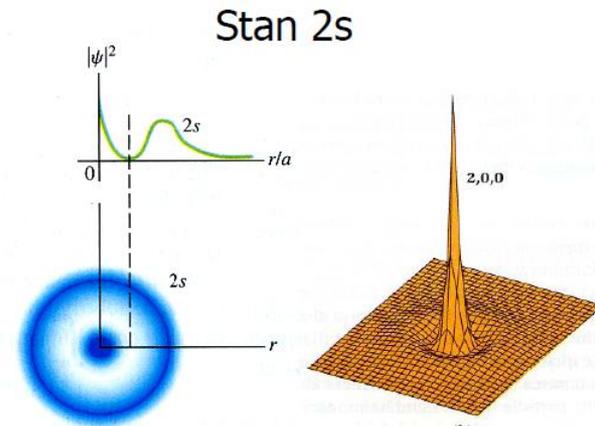
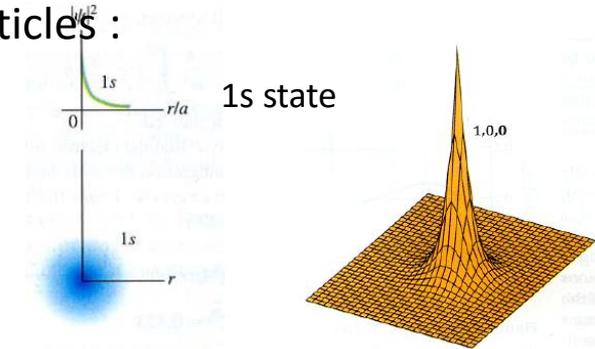
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$$|\Psi_n(\vec{r}, t)|^2 \text{ probability density}$$

$$\hat{H} \Psi(\vec{r}) = [E_{kin}(\vec{p}) + E_{pot}(\vec{r})]\Psi(\vec{r}) = E \Psi(\vec{r})$$

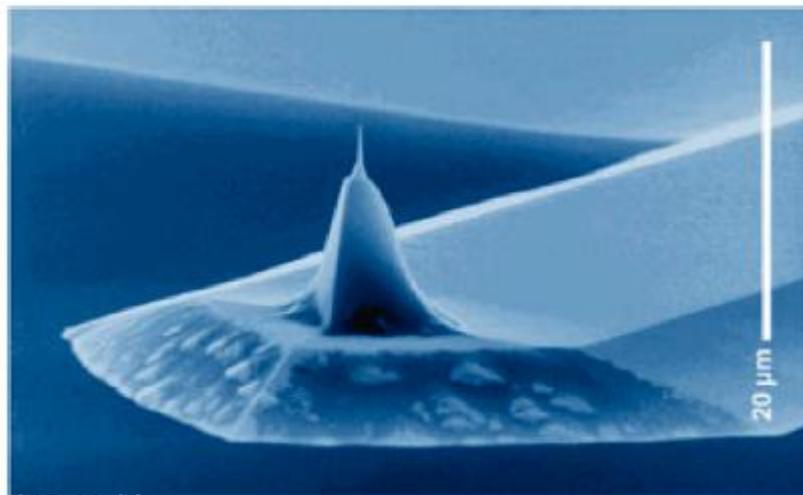
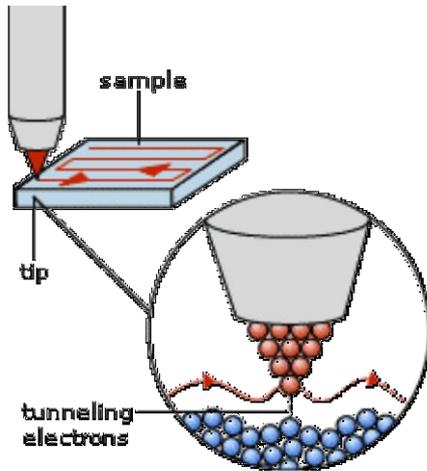


Trapped particle penetrates the potential barrier

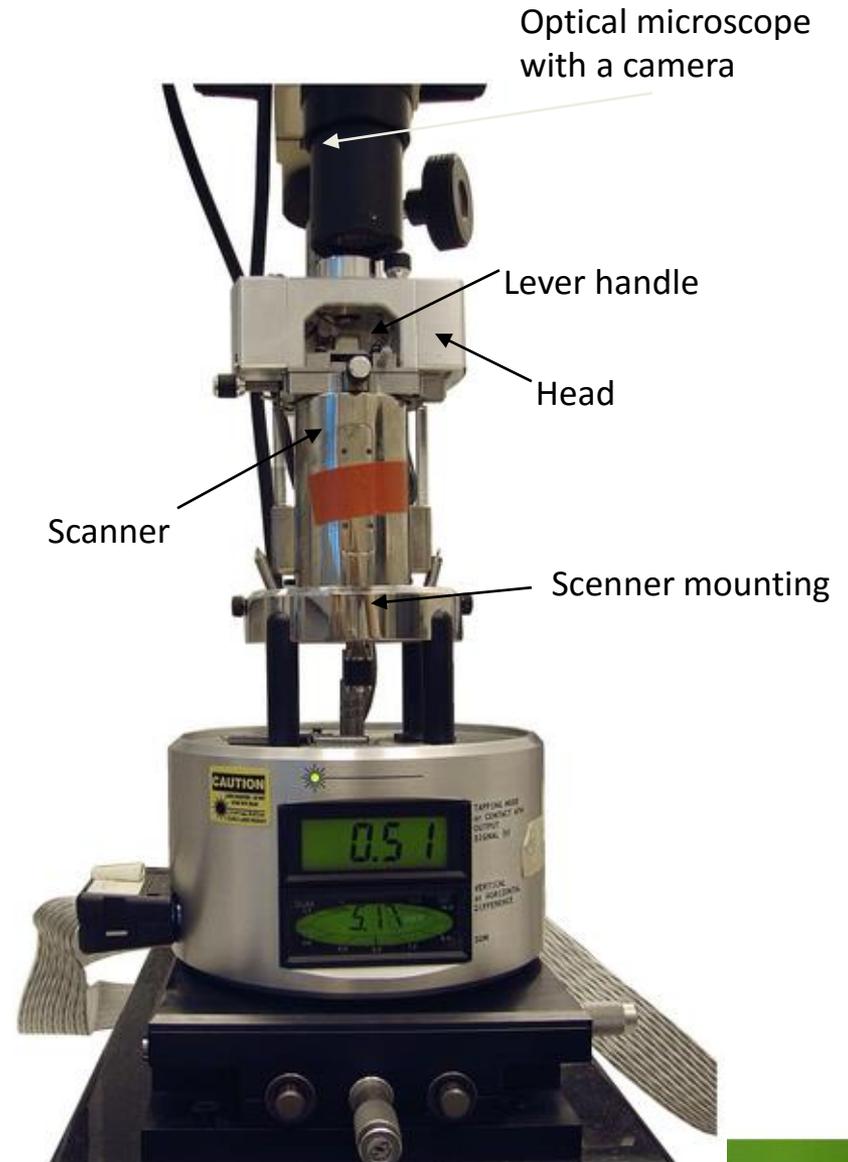


Classical and quantum universe

STM – Scanning Tunnelling Microscope
Nobel 1986 Gerd Binnig, Heinrich Rohrer

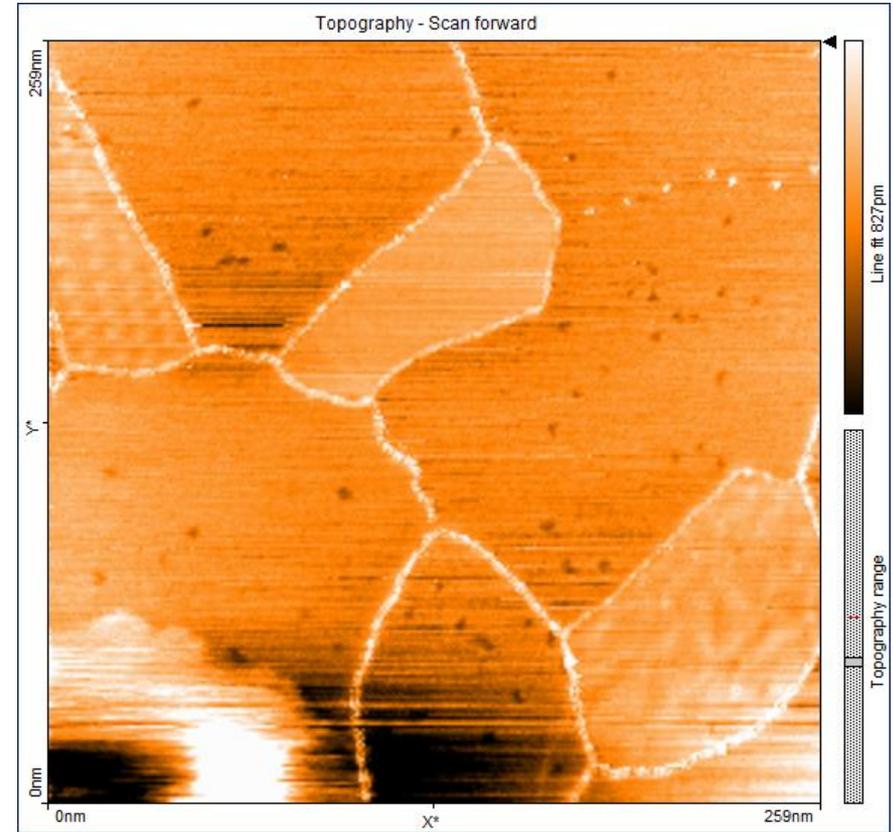
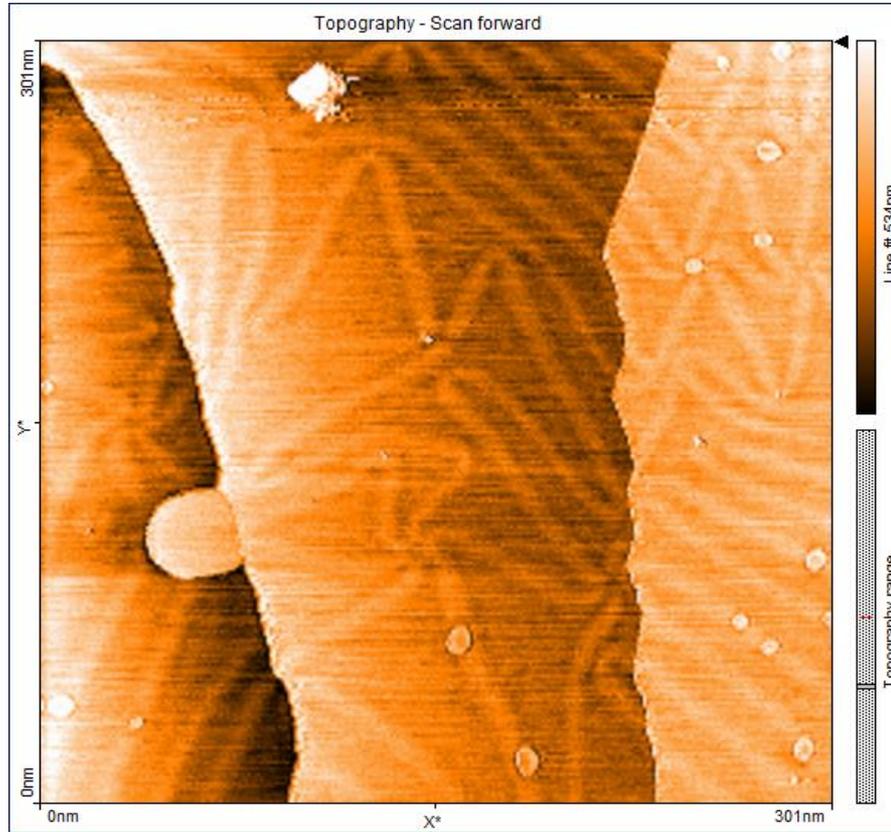


www.nanosensors.com



Classical and quantum universe

Polish graphene



Classical and quantum universe

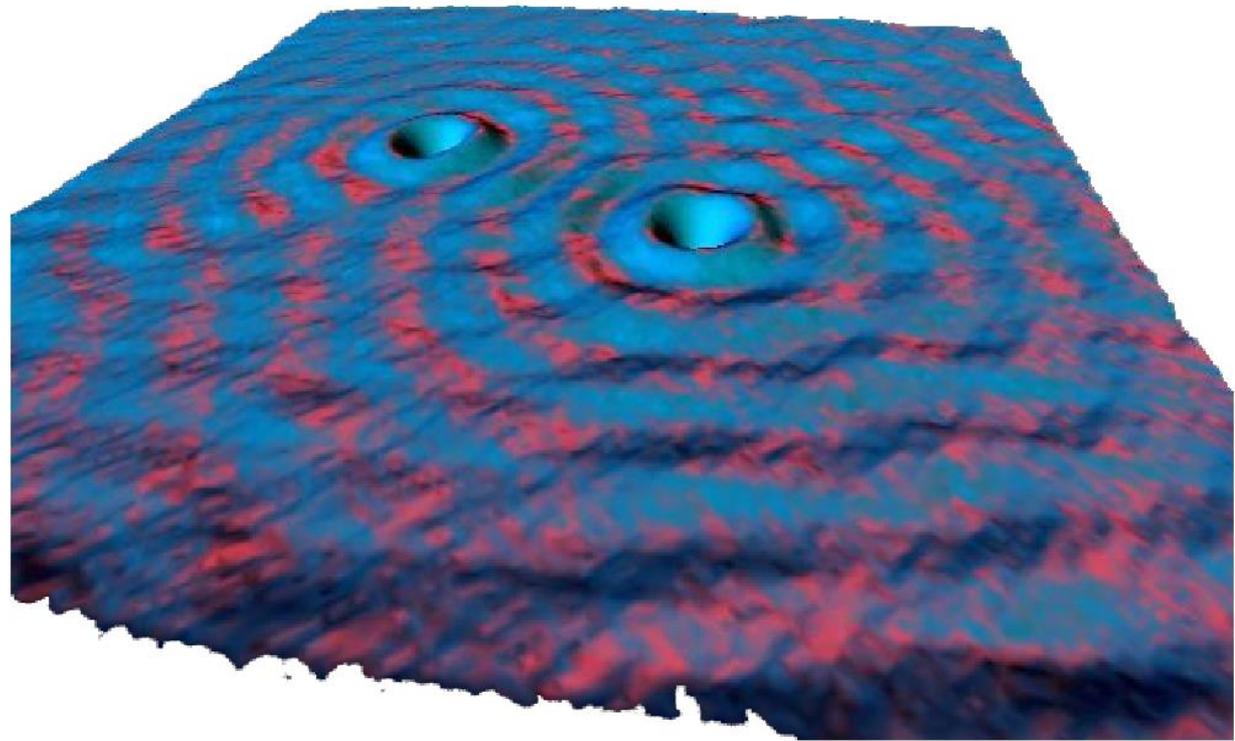
Remark 3: a LINEAR COMBINATION of wave functions is also a wave function (the principle of superposition)

$$\Psi = A\Psi_A + B\Psi_B$$

~~$$|\Psi|^2 \neq A^2|\Psi_A|^2 + B^2|\Psi_B|^2$$~~

$$|\Psi|^2 = A^2|\Psi_A|^2 + B^2|\Psi_B|^2 + 2AB\Psi_A\Psi_B$$

interference term



A quantum waves

In this scanning tunneling microscope (STM) image, electron density waves are seen to be breaking around two atom-sized defects on the surface of a copper crystal. The resultant standing waves result from the interference of the electron waves scattering from the defects. Courtesy, Don Eigler, IBM.

Classical and quantum universe

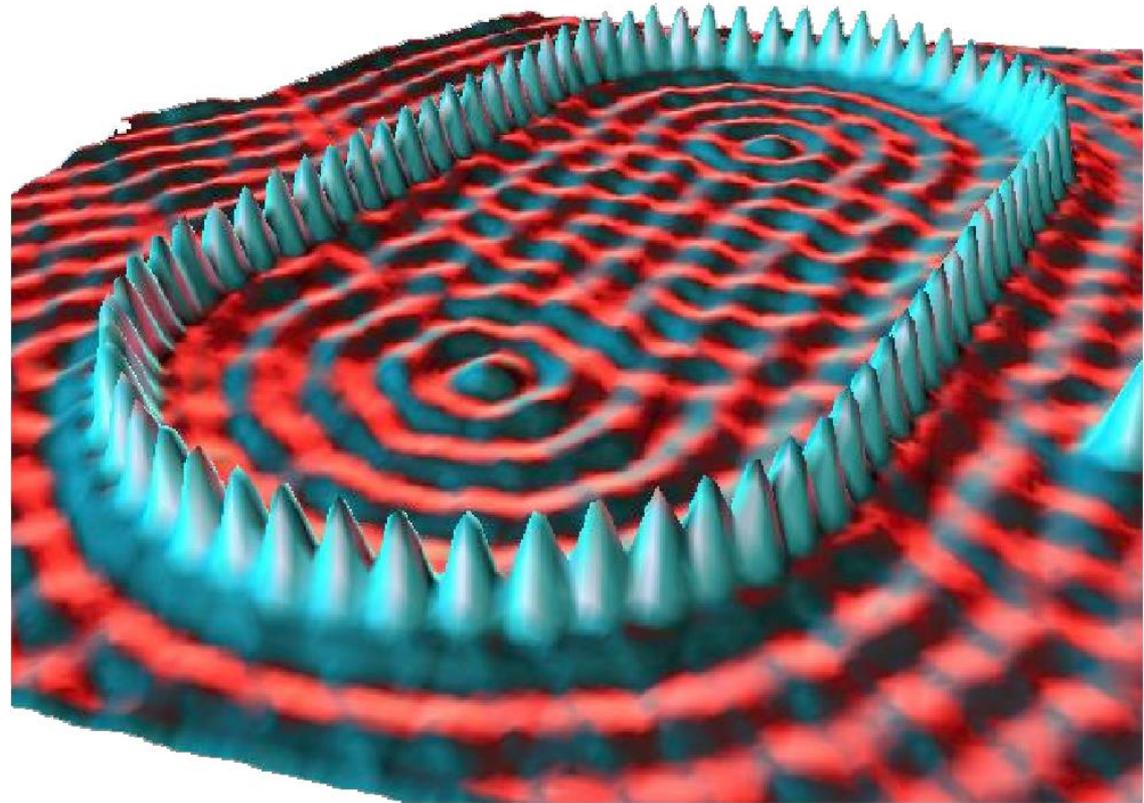
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interference term



A quantum corral

Scanning tunnelling microscope (STM) picture of a stadium-shaped "quantum corral" made by positioning iron atoms on a copper surface. This structure was designed for studying what happens when surface electron waves in a confined region. Courtesy, Don Eigler, IBM.

Classical and quantum universe

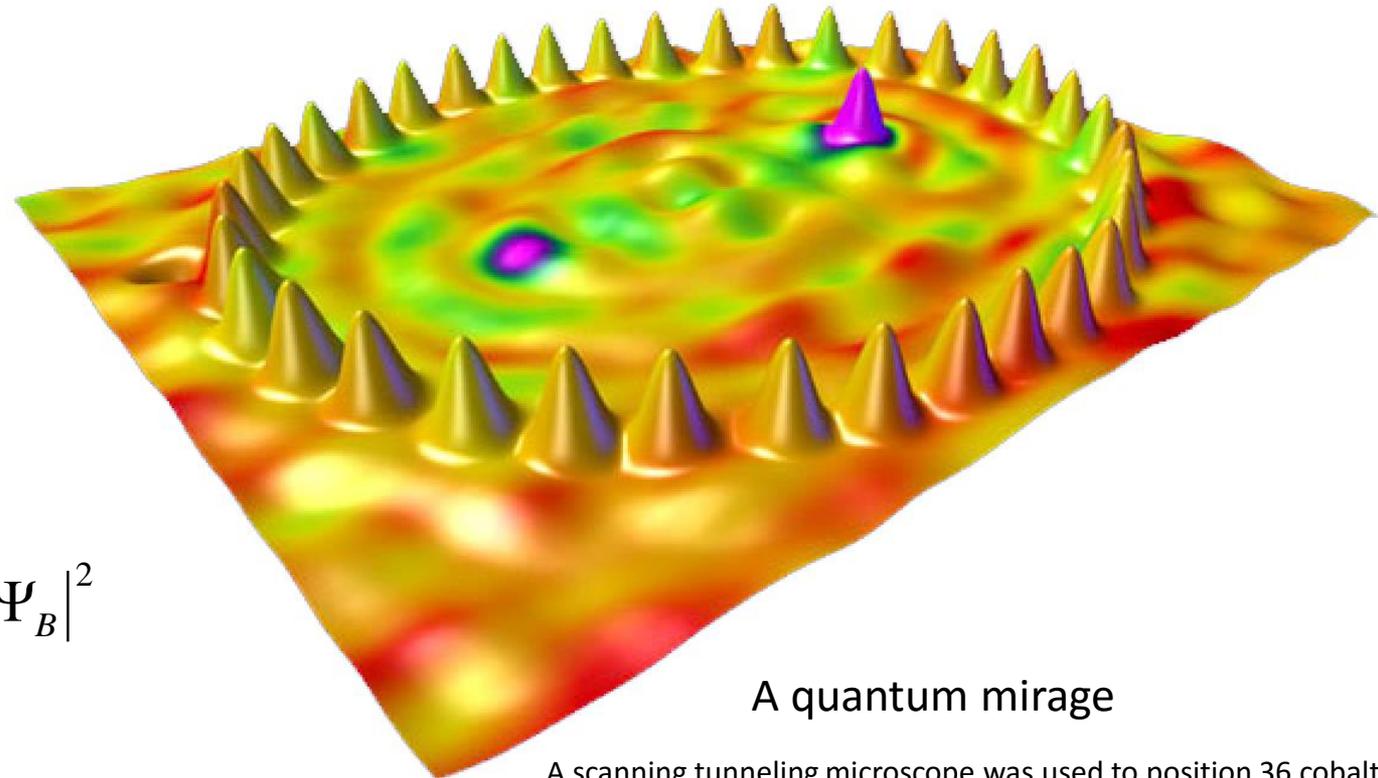
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interference term



A quantum mirage

A scanning tunneling microscope was used to position 36 cobalt atoms in an elliptical structure known as a "quantum corral." Electron waves moving in the copper substrate interact both with a magnetic cobalt atom carefully positioned at one of the foci of the ellipse and apparently with a "mirage" of another cobalt atom (that isn't really there) at the other focus. (Courtesy of IBM.) reported by: Manoharan et al., in [Nature](#), 3 February 2000

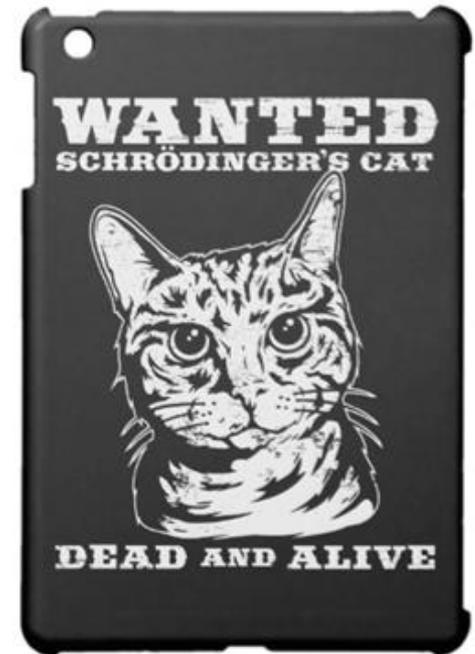
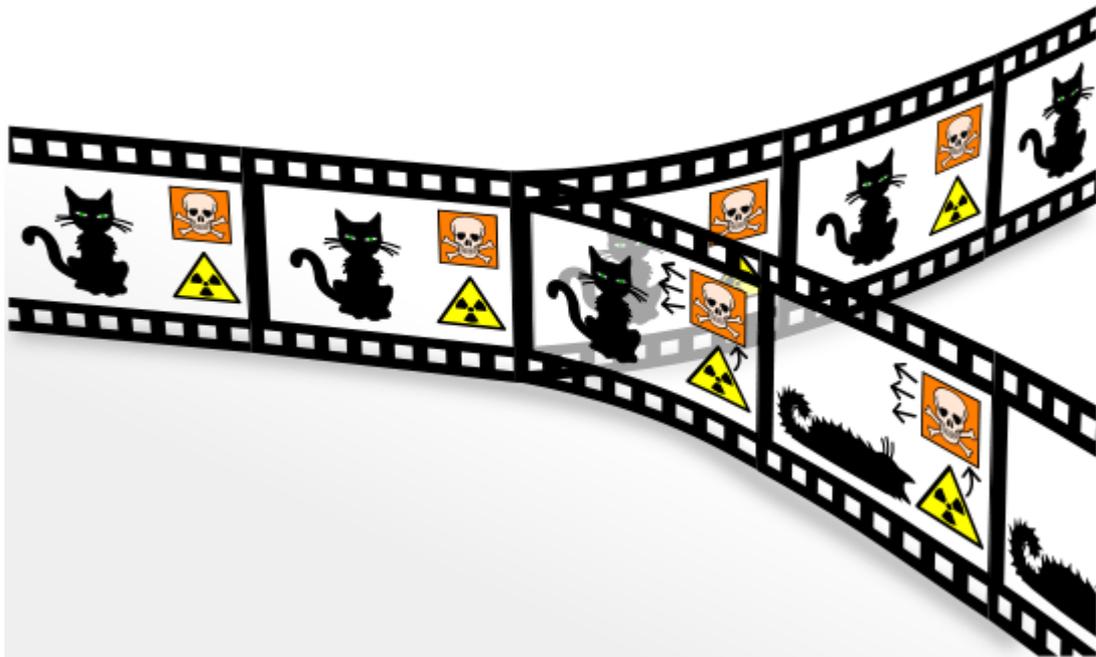
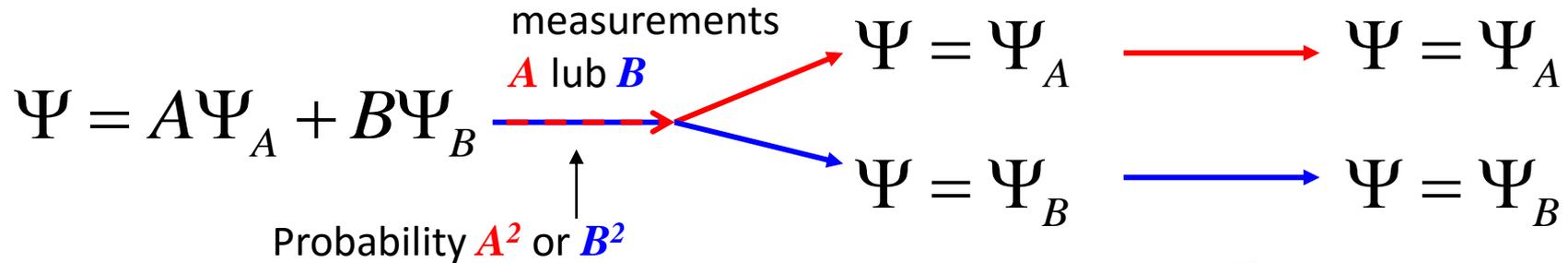
Classical and quantum universe

Remark 4: evolution of the wave function is DETERMINISTIC. However, by measurement a superposition of several eigenstates reduces to a single eigenstate (*wave function collapse*)



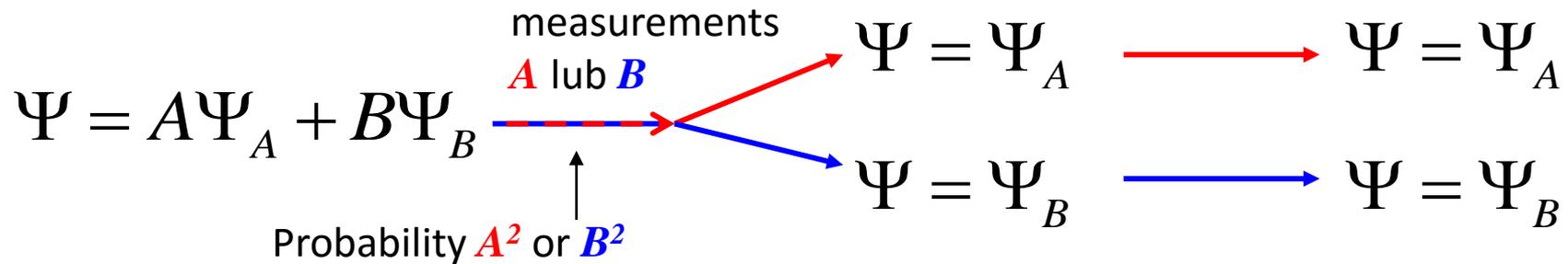
Świat klasyczny i kwantowy

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Classical and quantum universe

Remark 4: evolution of the wave function is DETERMINISTIC. However, by measurement a superposition of several eigenstates reduces to a single eigenstate (*wave function collapse*)



Sometimes the ORDER of a measurement is important :

$$[\hat{p}, \hat{x}] = -i\hbar$$

measurement = observables = operators = commutation rules

Heisenberg's uncertainty principle (position and momentum, energy and time, projections of momentum and momentum, etc.).

$$\Delta p \Delta x \geq \frac{1}{2} \hbar$$

$$\hat{U}\hat{W}\Psi \neq \hat{W}\hat{U}\Psi \quad \text{gdy} \quad \hat{U}\hat{W} - \hat{W}\hat{U} = [\hat{U}, \hat{W}] \neq 0$$

Superposition of states, entanglement

Synthesis of quantum dots

Heisenberg's uncertainty principle



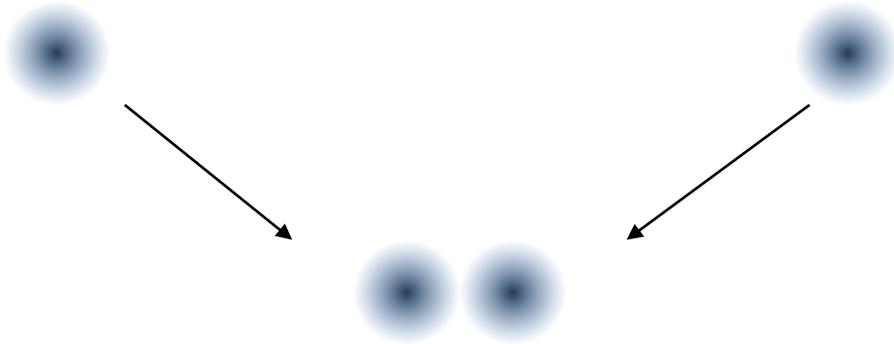
Classical and quantum universe

Remark 5: quantum particles are INDISTINGUISHABLE



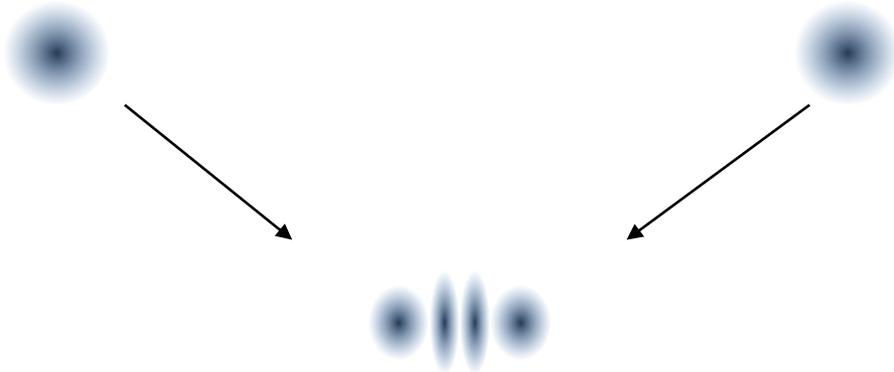
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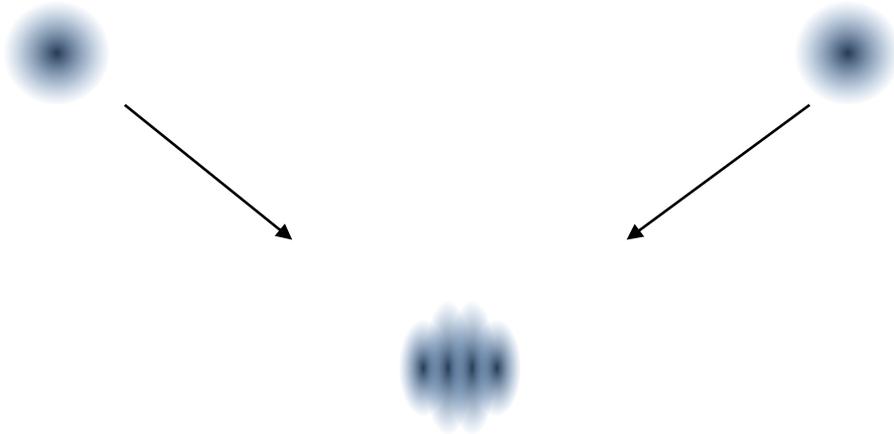
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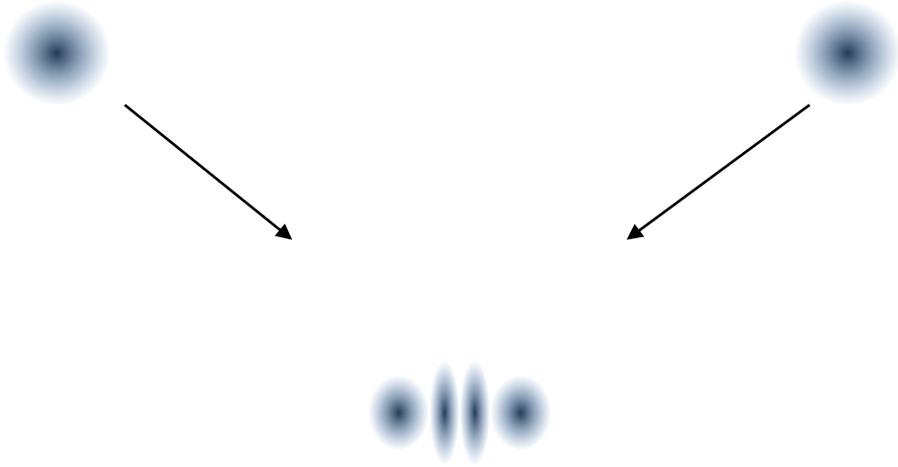
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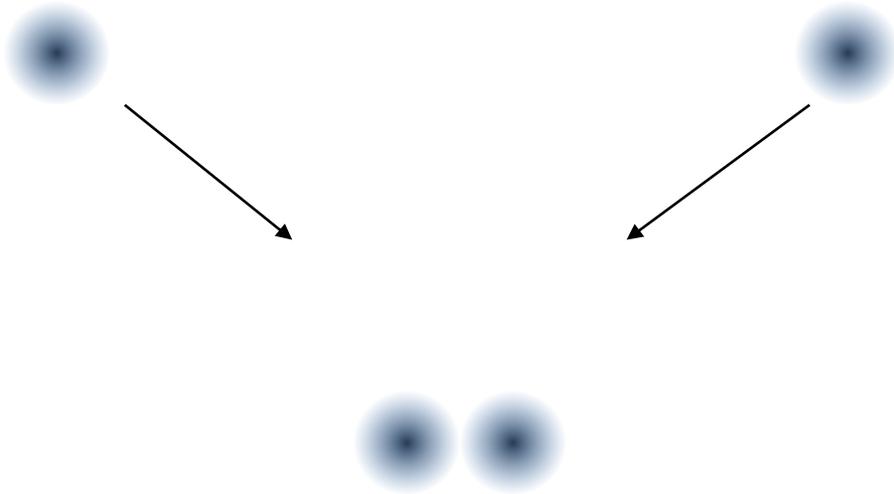
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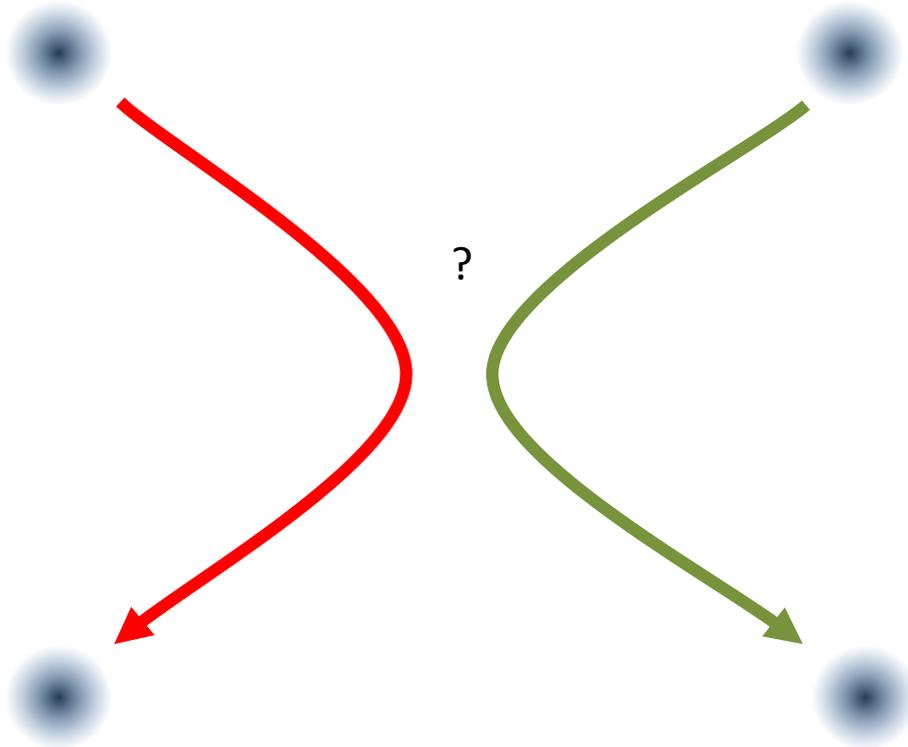
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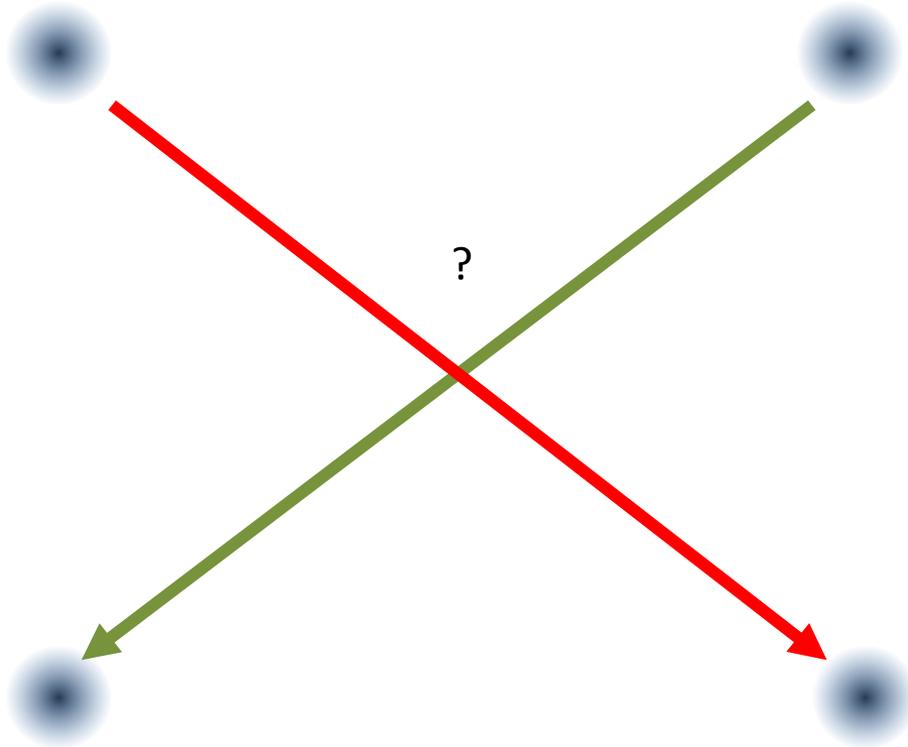
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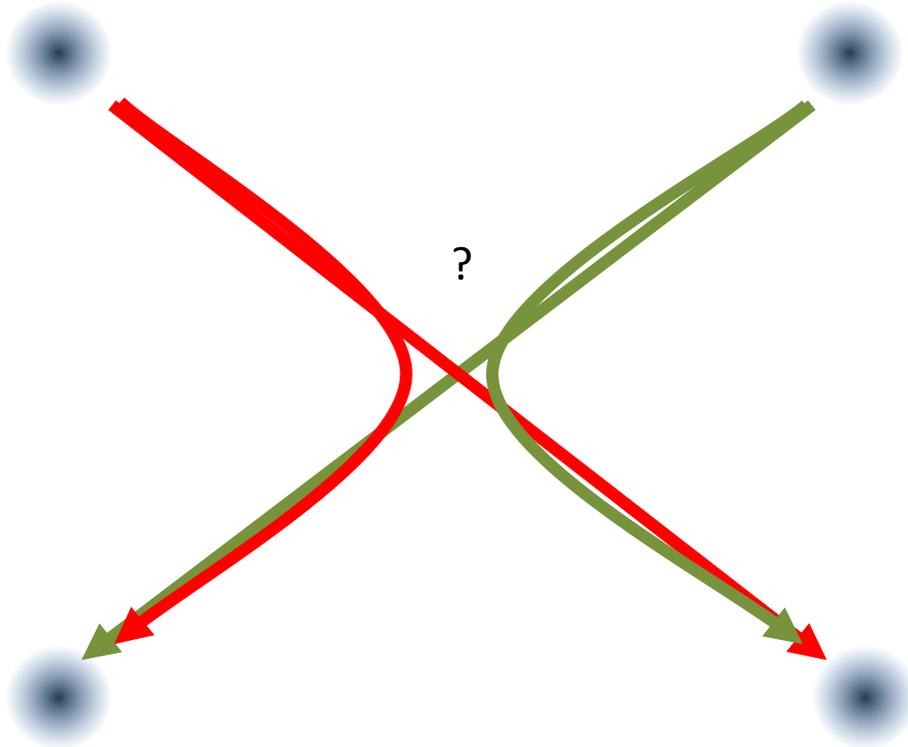
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Classical and quantum universe

Remark 5: quantum particles are INDISTINGUISHABLE



The wave function of the whole system
STATISTICS:

World fermions (e, p, n)

World bosons (photon, W boson)

Pauli exclusion principle, Fermi-Dirac statistics, Bose-Einstein; exchange interaction, ferromagnetism

Conclusions (certain)

The mathematical description by **eigenstates** (*orthogonal*)

$$\{|A\rangle, |B\rangle\}$$

Two atomic levels $\{|g\rangle, |e\rangle\}$ np. $g = 1s, e = 2s$

Electron spin $\{|\uparrow\rangle, |\downarrow\rangle\}$

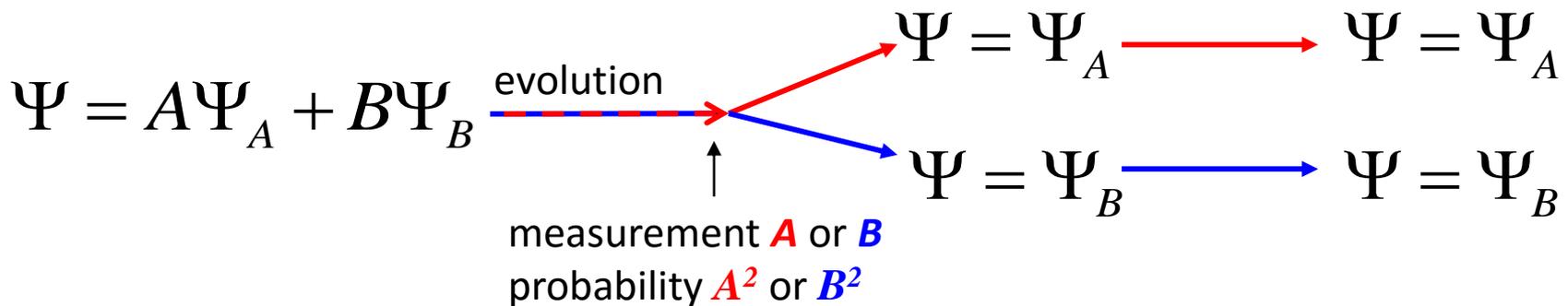
$$\{|\curvearrowright\rangle, |\curvearrowleft\rangle\}$$

Photon of two mutually orthogonal polarization states

$$\{|\rightarrow\rangle, |\uparrow\rangle\}$$

If the state of the particles describes two states A and B (superposition of states), the particle cannot be observed in both of them at the same time (so-called. orthogonality of states)!

$$\Psi = A\Psi_A + B\Psi_B$$

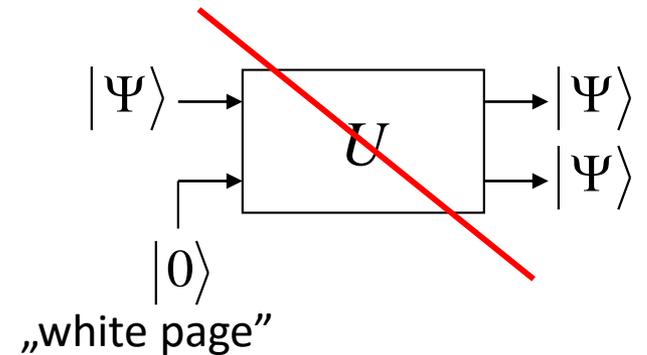


No cloning theorem

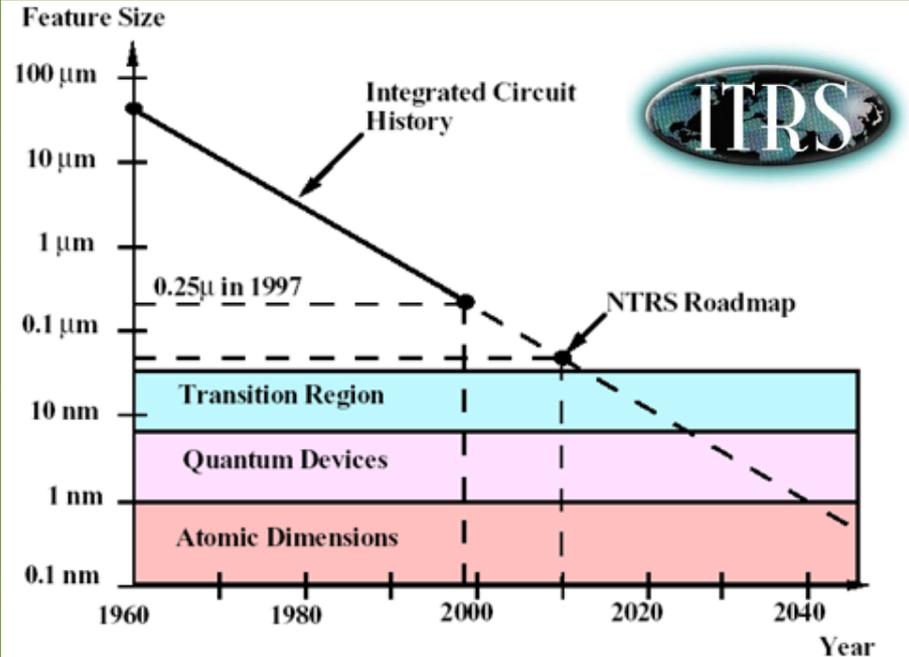
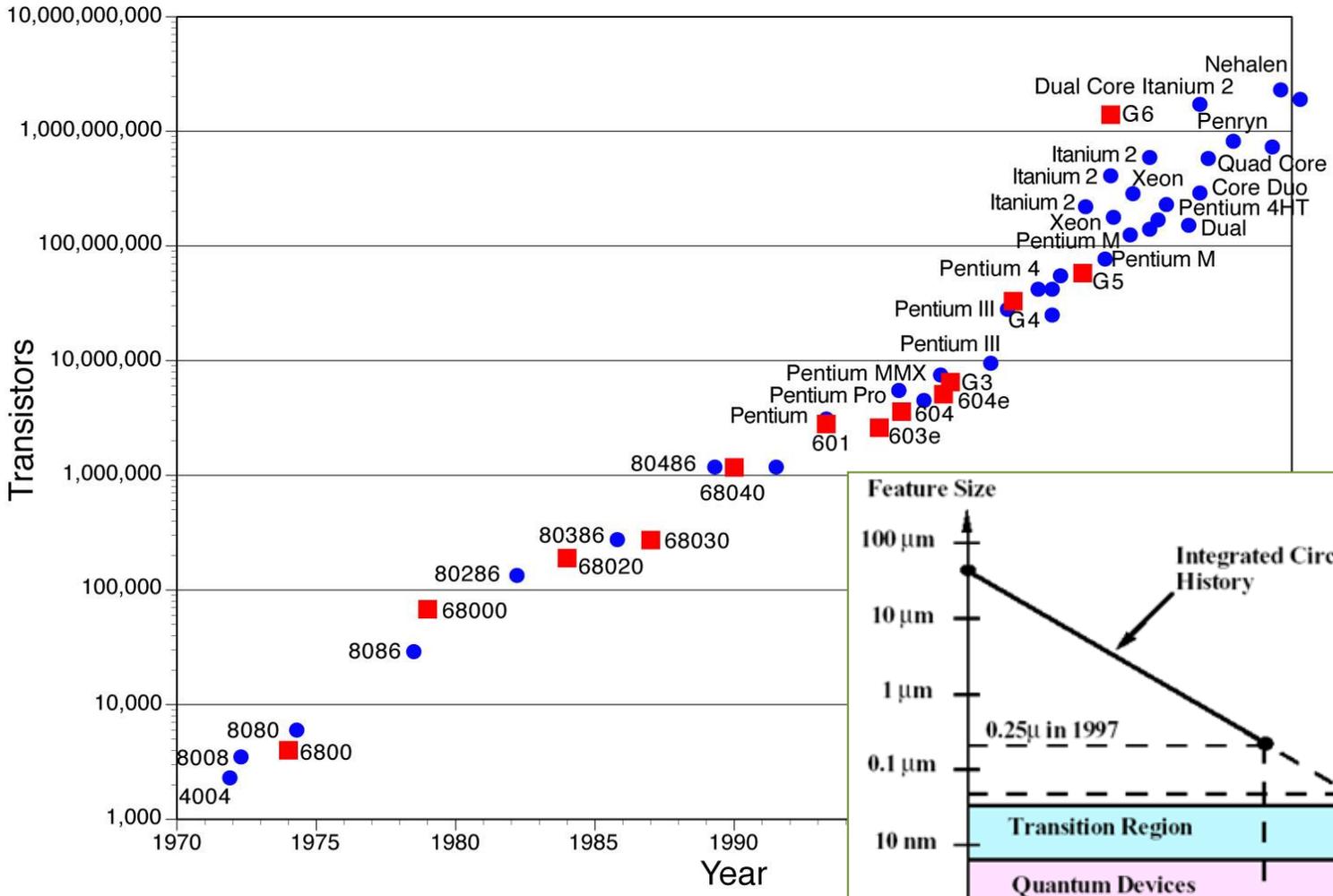


Źródło: Lucas Film

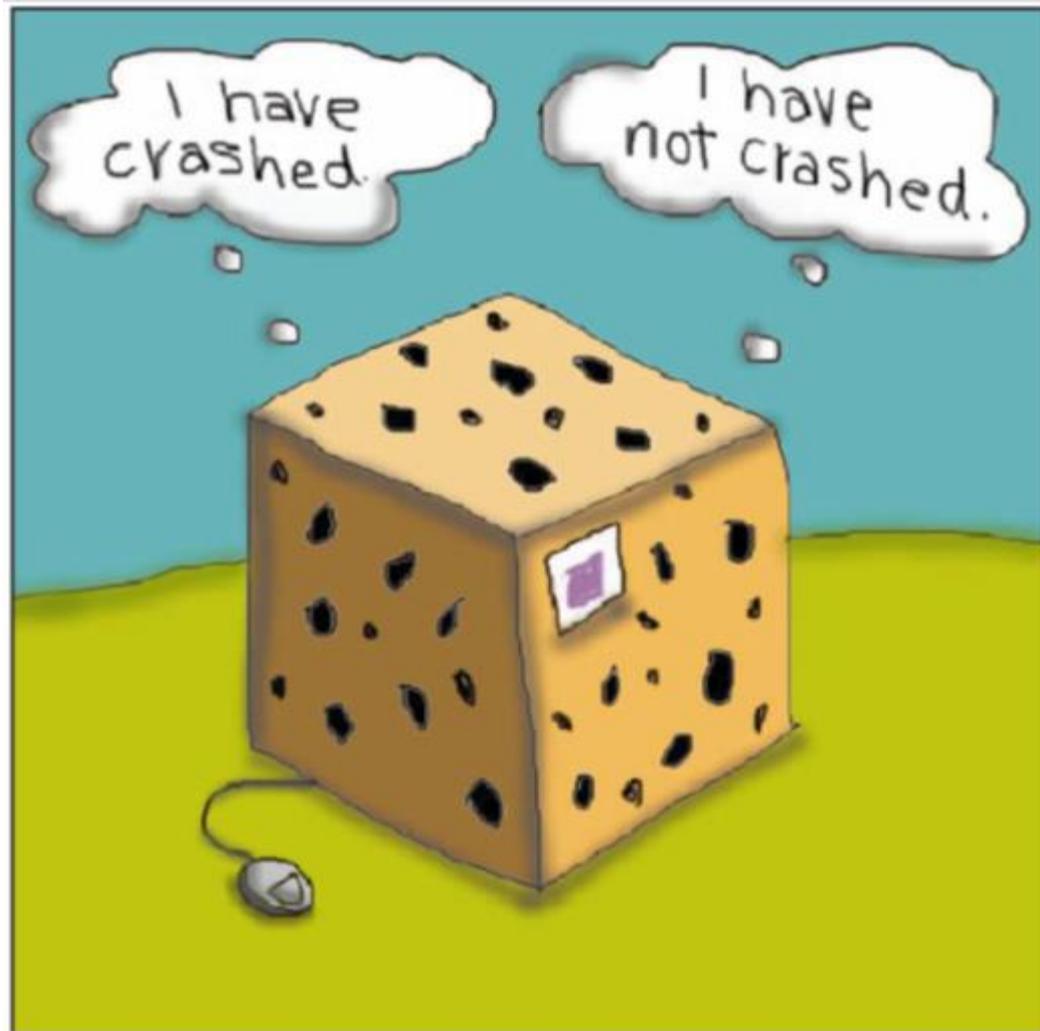
$$|\Psi\rangle \rightarrow |\Psi\rangle, |\Psi\rangle, |\Psi\rangle, |\Psi\rangle, |\Psi\rangle, |\Psi\rangle, |\Psi\rangle \dots$$



NANO era



QWindows



Schrödinger's computer.

—Sally O. Lee