## Leptogenesis and Grand Unification

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- introduction
- a quick review of leptogenesis
- successful leptogenesis in SO(10) unification
- a predictive SO(10) leptogenesis scenario

- P. Hosteins, S. L. and C. Savoy, hep-ph/0606078
- A. Abada, P. Hosteins, F.-X. Josse-Michaux and S. L., 0808.2058 [hep-ph]
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## Introduction

The baryon asymmetry of the universe (BAU)

 $\frac{n_B - n_{\bar{B}}}{n_{\gamma}} \simeq \frac{n_B}{n_{\gamma}} = (6.21 \pm 0.16) \times 10^{-10}$  (WMAP 5y) must be explained by some dynamical mechanism  $\Rightarrow$  baryogenesis (I) **B** violation

Sakharov's conditions: (2) C and CP violation (3) departure from thermal equilibrium

(1) and (2) are present in the SM – B+L anomaly  $\Rightarrow$  transitions between vacua with different (B+L) possible at  $T \gtrsim M_{weak}$ , where nonperturbative (B+L)-violating processes (sphalerons) are in equilibrium

Electroweak baryogenesis fails in the SM because (3) is not satisfied [also CP violation is too weak]  $\Rightarrow$  need either new physics at Mweak, or generate a (B-L) asymmetry at  $T > T_{EW}$ 

Leptogenesis (generation of a L asymmetry above TEW, which is then converted into a B asymmetry by sphalerons) belongs to the second class

Attractive mechanism since (in its simplest versions) connects neutrino masses to the BAU

A lot of work has been done in the past decade:

- refinement of the calculation of the generated baryon asymmetry (finite T effects, spectator processes, lepton flavour effects...)

- alternative scenarios to the standard one, including low-scale scenarios

- attempts to relate leptogenesis to measurable parameters, in particular to low-energy CP violation (no direct connection in general)

<u>This talk</u>: possibility of realizing successful (and possibly predictive) lepogenesis in SO(10) GUTs

### A quick review of (standard) leptogenesis

Generate a B-L asymmetry through the out-of-equilibrium decays of the heavy Majorana neutrinos responsible for neutrino mass (Fukugita, Yanagida)

 $N_i^c \equiv C\bar{N}_i^T = N_i$  (Majorana)  $\Rightarrow$  decays both into I<sup>+</sup> and I<sup>-</sup>

Γ

$$\frac{N_{i}}{V_{ix}^{*}} - H \qquad \frac{N_{i}}{V_{ix}} - H^{*} \\
N_{i} \rightarrow L_{x} H \qquad N_{i} \rightarrow \overline{L}_{x} H^{*} \\
tree(N_{i} \rightarrow LH) = \Gamma_{tree}(N_{i} \rightarrow \overline{L}H^{*}) = \frac{M_{i}}{16\pi}(YY^{\dagger})_{ii}$$

CP asymmetry due to interference between tree and 1-loop diagrams:



 $\Rightarrow \Gamma(N_i \to LH) \neq \Gamma(N_i \to LH^*)$ 

When M<sub>1</sub> << M<sub>2</sub>, M<sub>3</sub>, L-violating processes involving N<sub>1</sub> tend to erase the asymmetry generated from N<sub>2</sub> and N<sub>3</sub> decays, and it is often assumed that the final baryon asymmetry is dominated by the CP asymmetry in N<sub>1</sub> decays:

$$\epsilon_{N_1} \equiv \frac{\Gamma(N_1 \to LH) - \Gamma(N_1 \to \bar{L}H^{\star})}{\Gamma(N_1 \to LH) + \Gamma(N_1 \to \bar{L}H^{\star})} \simeq \frac{3}{16\pi} \sum_k \frac{\operatorname{Im}[(YY^{\dagger})_{k1}^2]}{(YY^{\dagger})_{11}} \frac{M_k}{M_1}$$

Covi, Roulet, Vissani Buchmüller, Plümacher

Generated lepton asymmetry:  $Y_L \equiv \frac{n_L - n_{\bar{L}}}{s} = 0.42 \frac{\eta \epsilon_{N_1}}{g_{\star}}$ 

g\* = total number of relativistic d.o.f. (g\* = 106.75 in the SM)

 $\eta$  = efficiency factor that takes into account the initial population of N<sub>1</sub>, the out-of-equilibrium condition for their decays, and the dilution of the lepton asymmetry by L-violating processes ( $LH \rightarrow N_1$ ,  $LH \rightleftharpoons \bar{L}H^* \cdots$ ) Conversion into a baryon asymmetry:  $at T > M_{weak}$ , the sphalerons (which violate B+L, but preserve B-L) are in thermal equilibrium

 $\Rightarrow$  Y<sub>L</sub> partially converted into Y<sub>B</sub>:

$$\langle Y_B \rangle_T = C \langle Y_{B-L} \rangle_T$$
  $C = \frac{8N_f + 4N_H}{22N_f + 13N_H} = \frac{28}{79}$  (SM)  
hence  $Y_B = -0.42 C \frac{\eta \epsilon_{N_1}}{g_{\star}} = -1.4 \times 10^{-3} \eta \epsilon_{N_1}$  (SM)

Can leptogenesis explain the observed baryon asymmetry?

 $\Rightarrow$  must compare YB computed from leptogenesis with observed value

-  $\eta$  essentially depends on M<sub>1</sub> and on  $\tilde{m}_1 \equiv (YY^{\dagger})_{11}v^2/M_1$ , which controls the out-of-equ. decay condition / strength of washout processes:

 $\Gamma_{N_1} < H(T = M_1) \iff \tilde{m}_1 < \tilde{m}_1^* = 2.2 \times 10^{-3} \,\mathrm{eV}$ 

-  $\mathcal{E}_{N_1}$  depends on the Ni masses and couplings, but is bounded by a simple function of M<sub>1</sub>, m<sub>1</sub> and  $\tilde{m}_1$  [case  $M_1 \ll M_2, M_3$ ]:

$$|\epsilon_{N_1}| \leq \frac{3}{16\pi} \frac{M_1(m_3 - m_1)}{v^2} f\left(\frac{m_1}{\tilde{m}_1}\right) \qquad 0 \leq f\left(\frac{m_1}{\tilde{m}_1}\right) \leq 1 \quad \begin{array}{l} \text{Davidson, Ibarra} \\ \text{Hambye et al.} \end{array}$$

Isocontours of  $\eta$  in the  $\tilde{m}_1^{\eta}$  ( $\tilde{m}_1^{\eta}$ ,  $M_1$ ) plane:



weak washout regime:  $\tilde{m}_1 \ll \tilde{m}_1^{\star} \Rightarrow \eta \sim \tilde{m}_1/\tilde{m}_1^{\star}$ strong washout regime:  $\tilde{m}_1 \gg \tilde{m}_1^{\star} \Rightarrow \eta \sim (\tilde{m}_1^{\star}/\tilde{m}_1)^{1.16}$ (for  $M_1 \ll 10^{14} \text{ GeV}$  and zero initial Ni abundance) Region in the  $(\tilde{m}_1, M_1)$  plane where leptogenesis can reproduce the observed baryon asymmetry:



 $\Rightarrow M_1 \ge (0.5 - 2.5) \times 10^9 \,\text{GeV}$  depending on the initial conditions

Case  $M_1 \approx M_2$ : if  $|M_1 - M_2| \sim \Gamma_2$ , the self-energy part of  $\epsilon_{N_1}$  has a resonant behaviour, and  $M_1 \ll 10^9 \text{ GeV}$  is compatible with successful leptogenesis ("resonant leptogenesis")

Covi, Roulet, Vissani Pilaftsis Flavour effects in leptogenesis

Barbieri, Creminelli, Strumia, Tetradis Endoh et al. - Nardi et al. - Abada et al. Blanchet, Di Bari, Raffelt - Pascoli, Petcov, Riotto - ...

"one-flavour approximation": leptogenesis described in terms of a single direction in flavour space, the lepton  $\mathcal{L}_1 \propto \sum_{\alpha} Y_{1\alpha} L_{\alpha}$  to which N1 couples  $\Rightarrow$  valid as long as the charged lepton Yukawas  $\lambda_{\alpha}$  are out of equilibrium

At  $T \lesssim 10^{12} \,\text{GeV}$ ,  $\lambda_{\tau}$  is in equilibrium and destroys the coherence of  $\mathcal{L}_1$  $\Rightarrow$  2 relevant flavours:  $L_{\tau}$  and a combination of Le and  $L_{\mu}$ 

At  $T \lesssim 10^9 \, {
m GeV}$ ,  $\lambda_{\tau}$  and  $\lambda_{\mu}$  are in equilibrium  $\Rightarrow$  must distinguish between Le , L<sub>µ</sub> and L<sub>τ</sub>

Relevant parameters for the discussion of flavour effects:

$$\epsilon_{N_1}^{\alpha} \equiv \frac{\Gamma(N_1 \to L_{\alpha}H) - \Gamma(N_1 \to \bar{L}_{\alpha}H^{\star})}{\Gamma(N_1 \to L_{\alpha}H) + \Gamma(N_1 \to \bar{L}_{\alpha}H^{\star})} \qquad \tilde{m}_1^{\alpha} \equiv \frac{|Y_{1\alpha}|^2 v^2}{M_1}$$

qualitatively  $Y_B \approx \sum_{\alpha} \epsilon^{\alpha}_{N_1} \eta(\tilde{m}^{\alpha}_1) \Rightarrow$  can deviate from the one-flavour approximation if e.g.  $\epsilon^{\tau}_{N_1} \gg \epsilon^e_{N_1}, \epsilon^{\mu}_{N_1}$  and  $\tilde{m}^{\tau}_1 \ll \tilde{m}^e_1, \tilde{m}^{\mu}_1$ 

## Leptogenesis in SO(10) models

Right-handed neutrinos are suggestive of SO(10) unification:

(i) 16 =  $(Q, \bar{u}, \bar{d}, L, \bar{e}) \oplus \bar{N}$ 

(ii) B-L is a generator of SO(10)  $\Rightarrow$  the mass scale of the NR is associated with the breaking of the gauge group  $\Rightarrow$  MR >> Mweak natural

However, successful leptogenesis is not so easy to achieve in SO(10) models with a type I seesaw mechanism:

 $M_D \propto M_u \Rightarrow$  very hierarchical right-handed neutrino masses  $\Rightarrow M_1 << 10^8 \ GeV \ , \ below \ the \ Davidson-Ibarra \ bound$ 

Ways out:

- flavour-dependent N2 leptogenesis [Vives]: N2 decays generate an asymmetry in a lepton flavour that is only mildly washed out by N1
- large corrections to  $M_D = M_u$
- other versions of the seesaw mechanism: type II (heavy scalar SU(2)L triplet exchange), type I + II (left-right symmetric seesaw mechanism)

### SO(10) models with a left-right symmetric seesaw

Type I+II seesaw mechanism:

 $\Delta L = SU(2)L$  triplet with couplings fLij to lepton doublets



$$M_{\nu} = f_L v_L - \frac{v^2}{v_R} Y^T f_R^{-1} Y \equiv M_{\nu}^{II} + M_{\nu}^{I}$$

**Right-handed neutrino mass matrix:**  $M_R = f_R v_R$ 

 $v_R \equiv \langle \Delta R \rangle$  scale of B-L breaking

 $\Delta R = SU(2)R$  triplet with couplings fRij to right-handed neutrinos

vL is small since it is an induced vev:  $v_L \equiv \langle \Delta_L \rangle \sim v^2 v_R / M_{\Delta_L}^2$ 

In a broad class of theories with underlying left-right symmetry (such as SO(10) with a  $\overline{126}_H$ ), one has  $Y = Y^T$  and  $f_L = f_R \equiv f$ 

→ left-right symmetric seesaw mechanism

The SU(2) triplet also contributes to leptogenesis. If  $M_1 \le M_{\Delta L}$ , it mainly affects leptogenesis by contributing to the CP asymmetry in N1 decays:

The total CP asymmetry is (for  $M_1 \ll M_2, M_3, M_{\Delta_L}$ ):

$$\epsilon_{N_1} = \epsilon_{N_1}^I + \epsilon_{N_1}^{II} \simeq \frac{3}{8\pi} \frac{\sum_{k,l} \operatorname{Im} \left[Y_{1k} Y_{1l} \left(M_{\nu}\right)_{kl}^{\star}\right]}{(YY^{\dagger})_{11}} \frac{M_1}{v^2}$$

Since the triplet is heavy, the dilution of the generated lepton asymmetry is mainly due to N1-related processes and depends on the effective mass parameter  $\tilde{m}_1 \equiv (YY^{\dagger})_{11}v^2/M_1$  as in the type I case

In order to study leptogenesis, need to reconstruct the fij (which determine both the triplet couplings and the RHN mass matrix) as a function of the Yij (predicted by the theory) and of the light neutrino parameters (in principe accessible to experiment)

### Reconstruction of the fij couplings

Assuming that Y is known in the basis of charged lepton mass eigenstates, the LR symmetric seesaw formula  $v^2$ 

$$M_{\nu} = f_L v_L - \frac{v^2}{v_R} Y f_R^{-1} Y$$

admits  $2^3 = 8$  solutions for f (for 3 generations) [Akhmedov, Frigerio]

 $\rightarrow$  new possibilities for leptogenesis with respect to the type I case

One can obtain useful analytical expressions from a reconstruction procedure using complex orthogonal matrices [Hosteins, S.L., Savoy]

Application:SO(10) models with two 10's and a  $\overline{126}$  in the Higgs sector $W \ni Y_{ij}^{(1)} 16_i 16_j 10_1 + Y_{ij}^{(2)} 16_i 16_j 10_2 + f_{ij} 16_i 16_j \overline{126}$  $Y^{(1)}, Y^{(2)}$  symmetric $\overline{126} \ni \Delta_L, \Delta_R$  with  $f_L = f_R = f$ 

If the doublets in the  $\overline{126}$  do not get vevs, this leads to:

$$Yv_u \equiv M_D = M_u \qquad \qquad M_d = M_e$$

(in general, Y and  $M_V$  contain physical high-energy phases)



<u>Inputs</u>: normal hierarchy with  $m_1 = 10^{-3}$  eV,  $\theta_{13} = 0$ , all PMNS and high-energy phases vanish –  $V^2 = 0.1$  VL VR

Among the 8 solutions, 3 different patterns emerge for leptogenesis:

- 2 solutions with a rising  $M_1 \Rightarrow$  large  $\epsilon_{N_1}$  for large  $v_R$
- 2 solutions with  $M_1 \sim {
  m few}\, 10^9\,{
  m GeV}$

- 4 solutions with  $M_1 \sim 10^5 \,\text{GeV} \Rightarrow \epsilon_{N_1}$  too small, but  $M_2 \sim \text{few } 10^9 \,\text{GeV}$ or rises with  $v_R \Rightarrow$  the observed baryon asymmetry could be generated from N<sub>2</sub> decays (relevance of flavour effects)

In all cases, the washout tends to be important and a numerical resolution of the Boltzmann equations is required to tell whether some of these solutions lead to successful leptogenesis

#### Relevant ingredients:

- contribution of N<sub>2</sub>
- lepton flavour effects (independent evolution of the asymmetries in the e,  $\mu$  and  $\tau$  flavours)
- corrections to Md = Me from appropriate SO(10) operators (affects the reconstruted RHN spectra)

### Computation of the baryon asymmetry

Solve the Boltzmann equations with flavour effects and decays of  $N_1$  and  $N_2$  Relevant quantities:

- <u>flavour-dependent CP asymmetries:</u>

$$\epsilon_{N_i}^{\alpha} \equiv \frac{\Gamma(N_i \to L_{\alpha}H) - \Gamma(N_i \to \bar{L}_{\alpha}H^{\star})}{\Gamma(N_i \to L_{\alpha}H) + \Gamma(N_i \to \bar{L}_{\alpha}H^{\star})}$$

- <u>wash-out processes</u>:  $\Delta L$  and N<sub>3</sub> very heavy  $\Rightarrow$  associated wash-out processes suppressed. Furthermore, we neglect  $\Delta L=2$  processes since we deal with masses M<sub>1</sub> and M<sub>2</sub> < 10<sup>12</sup> GeV

 $\Rightarrow$  only inverse decays and  $\Delta L=1$  scatterings associated with N<sub>1</sub> and N<sub>2</sub> enter the Boltzmann equations. The relevant washout parameters are:

$$\tilde{m}_i^{\alpha} \equiv \frac{|Y_{i\alpha}|^2 v^2}{M_i}$$

Both the  $\epsilon_{N_i}^{\alpha}$  and the  $\tilde{m}_i^{\alpha}$  depend on the Mi and on the Yia, hence on the reconstructed fij couplings



<u>Inputs</u>: normal hierarchy with  $m_1 = 10^{-3} \text{ eV}$ ,  $\theta_{13} = 0$ ,  $\delta = 0$ , different choices of Majorana and high-energy phases  $-V^2 = 0.1 \text{ VL VR} - \text{Tin} = 10^{11} \text{ GeV}$ 



Successful leptogenesis possible for  $v_R \gtrsim 10^{13} \, \text{GeV}$ 



Successful leptogenesis possible for  $v_R \sim (10^{13} - 10^{14}) \,\mathrm{GeV}$ 

The corrections to  $M_d = M_e$  play a crucial role here (not enough baryon asymmetry produced for  $M_d = M_e$ ).



In spite of a huge enhancement by flavour effects, the baryon asymmetry generated from N<sub>2</sub> decays fails to reproduce the observed value (we did not find successful parameters – confirmed by Di Bari, Riotto in the type I case). Case ++- marginally successful, however not for TRH <  $10^{10}$  GeV

### Impact of flavour effects (case Md = Me)



Figure 2: The final baryon asymmetry as a function of  $v_R$  for the four reference solutions, in the one-flavour approximation (dashed black line) and with flavour effects taken into account (solid red line). The GUT-scale mass relation  $M_d = M_e$  is assumed. Inputs: hierarchical light neutrino masses with  $m_1 = 10^{-3}$  eV,  $\theta_{13} = 0$  and no CP violation in the PMNS mixing matrix;  $\Phi_2^u = \pi/4$  and all other high-energy phases are set to zero;  $\beta/\alpha = 0.1$ . The Boltzmann equations are evolved starting from  $T_{in} = 10^{11}$  GeV. The thick horizontal line corresponds to the WMAP constraint.

### Flavour-dependent N2 leptogenesis

At  $T \ll M_2$ , the evolution of the flavour asymmetries  $Y_{\Delta_{\alpha}}$  ( $\Delta_{\alpha} \equiv B/3 - L_{\alpha}$ ) generated in N2 decays is governed by the Boltzmann equation:

$$\frac{dY_{\Delta_{\alpha}}}{dz} \simeq -2 |A_{\alpha\alpha}| \kappa_{1\alpha} W_1(z) Y_{\Delta_{\alpha}}(z)$$

where  $z \equiv M_1/T$ ,  $\kappa_{1\alpha} \equiv \tilde{m}_1^{\alpha}/\tilde{m}_1^{\star}$ ,  $|A_{\alpha\alpha}| \approx 1$ 

and  $W_1(z)$  = rate of N<sub>1</sub> inverse decays (main washout processes)

$$(Y_{\Delta_{\alpha}})_{\text{final}} \simeq (Y_{\Delta_{\alpha}})_{N_{2}} e^{-2|A_{\alpha\alpha}|\kappa_{1\alpha}\int_{z_{in}}^{\infty} dz \ W_{1}(z)}$$
$$\simeq (Y_{\Delta_{\alpha}})_{N_{2}} e^{-\frac{3\pi}{4}|A_{\alpha\alpha}|\kappa_{1\alpha}}$$

(using  $\int_0^\infty dz \ W_1(z) = 3\pi/8$ ). If e.g.  $\kappa_{1e} \ll 1 \ll \kappa_{1\mu}, \kappa_{1\tau}$ , the asymmetry in the electron flavour is almost unaffected, while the asymmetries in the muon and tau flavours are exponentially diluted. This results in a large final B-L asymmetry. By contrast, in the one-flavour approximation, the asymmetry generated in N2 decays is completely washed out:

$$Y_{B-L} \simeq e^{-\frac{3\pi}{4}\sum_{\alpha}\kappa_{1\alpha}} (Y_{B-L})_{N_2} \ll (Y_{B-L})_{N_2}$$

#### Supersymmetric thermal leptogenesis $\Rightarrow$ gravitino problem

If impose  $T_{RH} < 10^{10}$  GeV, only 4 solutions survive (generically) No successful realization of "N2 leptogenesis"



Figure 10: Regions of the  $(v_R, T_{in})$  parameter space where  $|Y_B| > Y_B^{WMAP}$  for solutions (+, +, +), (+, -, +) and (+, +, -), and where  $|Y_B| > 0.1 Y_B^{WMAP}$  for solution (-, -, -). These regions are delimited by the thick black contour in the (+, +, +) case, the dashed red contour for (+, -, +), the long-dashed blue contour for (+, +, -), and the thin black contour for (-, -, -). Inputs: set 1 of the Appendix for  $U_m$  and the high-energy phases; other input parameters as in Fig. 2.

#### Impact of corrections to MD = Mu

**plots:**  $m_{D_2} = (0.1 - 10) m_c(M_{GUT})$ 

- successful flavour-dependent N2 leptogenesis (however, solution -- fails if impose  $T_{RH} < 10^{10} \text{ GeV}$ )

- solution + - + successful for vR as large as MGUT



Figure 9: The final baryon asymmetry as a function of  $v_R$  for different values of  $y_2$ , from  $y_2/y_c(M_{GUT}) = 0.1$  (yellow/light grey) to  $y_2/y_c(M_{GUT}) = 10$  (blue/dark grey). The reference case  $y_2 = y_c$  is plotted in black. Left panel: solution (-, -, -), set 1 for  $U_m$  and the high-energy phases; right panel: solution (+, -, +), set 4 for  $U_m$  and the high-energy phases. The other input parameters are as in Fig. 2.

### SO(10) models with type II seesaw mechanism

Much more difficult: RHNs belong to the matter representation (16), hence are always around and couple to lepton doublets

Way out: "non-standard" embedding of the SM fermions into SO(10) representations

 $\begin{array}{rcl} 16_i & = & 10_i \ \oplus & . \ \oplus \ 1_i \\ 10_i & = & . \ \oplus \ \overline{5}_i^{10} \end{array}$ 

 $(5_i^{10}, \bar{5}_i^{16})$  form a vector-like pair of matter fields

Motivation: E6?  $27_i = 16_i \oplus 10_i \oplus 1_i$ 

How to achieve this?  $W = \frac{1}{2} y_{ij} 16_i 16_j 10 + h_{ij} 16_i 10_j 16$ SU(5) singlet in the I6:  $v_1^{16} \neq 0 \Rightarrow$  GUT-scale masses for  $(5_i^{10}, \overline{5}_i^{16})$  $5_i^{10} \equiv (L_i^c, D_i)$  heavy anti-lepton doublets and quark singlets SM matter fields:  $10_i^{16} = (Q_i, u_i^c, e_i^c), \quad \overline{5}_i^{10} = (L_i, d_i^c), \quad 1_i^{16} = \nu_i^c$ 

$$W = \frac{1}{2} y_{ij} 16_i 16_j 10 + h_{ij} 16_i 10_j 16$$

Quark and lepton masses:  $M_u = y v_u^{10}$   $M_d = M_e^T = h v_d^{16}$ 

No neutrino Dirac couplings at tree level: RHNs couple to heavy leptons

The heavy leptons (quarks) have hierarchical masses proportional to down-type fermion masses:  $M_i = h_i v_1^{16} = m_{e_i} v_1^{16} / v_d^{16}$ 

Assumed:

- matter parity

- no mass term 10; 10; no 54 vev  $\Rightarrow$  no mixing  $\overline{5}_i^{10}/\overline{5}_i^{16}$ 

### Leptogenesis

Requires a CP asymmetry in triplet decays. In standard triplet leptogenesis, the fij 's are not enough: need a second set of (flavour) couplings, otherwise  $\epsilon_{\Delta} \propto \text{Im}[\text{Tr}(ff^*ff^*)] = 0$ 

 $\Rightarrow$  introduce e.g. a second triplet with couplings f'ij to leptons

 $\Rightarrow$  loose predictivity: no direct connection between leptogenesis and neutrino masses (usual problem of leptogenesis: see e.g. Davidson et al, Petcov et al.)

However, in our scenario the states in the loop are heavy:



(the self-energy diagram does not contribute to the asymmetry)



 $\epsilon_{\Delta} \propto \sum_{kl} c_{kl} \,\theta(M_{\Delta} - M_k - M_l) \,\mathrm{Im}[f_{kl}(f^*ff^*)_{kl}]$ 

The  $\tilde{L}_i^c$  are heavy with hierarchical masses:  $(M_1, M_2, M_3) \sim (2 \times 10^{11}, 4 \times 10^{13}, 7 \times 10^{14}) \text{ GeV} \left(\frac{\tan \beta}{10}\right) \left(\frac{v_1^{16}}{10^{16} \text{ GeV}}\right)$ If e.g.  $M_3 > M_\Delta$ , the trace is incomplete and  $\epsilon_\Delta \neq 0$ 

Assuming  $M_1 \ll M_{\Delta} < M_1 + M_2$  and  $M_S = M_T = M_{24} \gg M_{\Delta}$ , one obtains:  $\epsilon_{\Delta} \simeq \frac{1}{10\pi} \frac{M_{\Delta}}{M_{24}} \frac{\lambda_L^4}{\lambda_L^2 + \lambda_{L_1^c}^2 + \lambda_{H_u}^2 + \lambda_{H_d}^2} \frac{\text{Im}[M_{11}(M^*MM^*)_{11}]}{(\sum_i m_i^2)^2}$ where  $\lambda_L^2 \equiv \sum_{i,i=1}^3 |f_{ij}|^2$ ,  $\lambda_{L_1^c}^2 \equiv |f_{11}|^2$ ,  $\lambda_{H_{u,d}}^2 \equiv |\sigma \alpha_{u,d}^2|^2$ 

 $\Delta_s \to \tilde{L}_1^c \tilde{L}_1^c$ : opposite CP asymmetry  $(-\epsilon_\Delta) / \Delta_s \to \overline{\tilde{H}}_d \overline{\tilde{H}}_d, H_u H_u$ : no CP asymmetry

### Dependence on the light neutrino parameters

$$\frac{\operatorname{Im}[M_{11}(M^*MM^*)_{11}]}{\overline{m}^4} = -\frac{1}{\overline{m}^4} \left\{ c_{13}^4 c_{12}^2 s_{12}^2 \sin(2\rho) m_1 m_2 \Delta m_{21}^2 \right. \\ \left. + c_{13}^2 s_{13}^2 c_{12}^2 \sin 2(\rho - \sigma) m_1 m_3 \Delta m_{31}^2 - c_{13}^2 s_{13}^2 s_{12}^2 \sin(2\sigma) m_2 m_3 \Delta m_{32}^2 \right\} \\ \left. U_{ei} = \left( c_{13} c_{12} e^{i\rho}, c_{13} s_{12}, s_{13} e^{i\sigma} \right) \right\}$$

 $\rightarrow \epsilon_{\Delta}$  does not depend on high-scale flavour parameters - only on the light neutrino parameters and on  $\lambda_L, \lambda_{H_u}, \lambda_{H_d}, M_{\Delta}/M_{24}$ 

 $\rightarrow$  the CP violation needed for leptogenesis is provided by the CP-violating phases of the PMNS matrix (the Majorana phases to which neutrinoless double beta decay is sensitive in the case  $M_1 < M_\Delta < M_1 + M_2$ )

 $\rightarrow \epsilon_{\Delta}$  can be large ( $\lambda_L^2$  is bounded by perturbativity):

$$\epsilon_{\Delta}^{\max} \simeq 2.2 \times 10^{-4} \lambda_L^2 \qquad (\text{maximum } \theta_{13}) ,$$
  
$$\simeq 3.4 \times 10^{-5} \lambda_L^2 \qquad (\text{vanishing } \theta_{13}) ,$$



Isocontours of the CP asymmetry in units of  $\lambda_L^2$ in the  $(\sin^2 \theta_{13}, m_{\text{lightest}})$  plane, maximized with respect to the CP-violating phases and to  $M_{\Delta}/M_{24}$ 

 $\frac{n_B}{s} = 7.62 \times 10^{-3} \, \eta \, \epsilon_{\Delta} \text{ agrees with the WMAP value } (8.82 \pm 0.23) \times 10^{-11}$ 

if  $\eta \epsilon_{\Delta} \approx 10^{-8} \Rightarrow$  the efficiency factor can be as small as  $10^{-5} - 10^{-4}$ in the region where the CP asymmetry is maximal

This regime must be studied numerically. There is also a large efficiency regime that can be discussed analytically, namely the regime where

 $K_{L_1^c} \ll 1$ ,  $K_L, K_{H_u} \gtrsim 1$  and  $M_{24} \gg M_\Delta$ with  $K_a \equiv \Gamma(\Delta_s \to aa)/H(M_\Delta)$   $(a = \tilde{L}_1^c, \bar{L}, H_u)$ 

Even though triplet decays are in equilibrium, a lepton asymmetry is generated thanks to  $K_{L_1^c} \ll 1$  [Hambye, Raidal, Strumia]

Unfortunately, the condition  $K_{L_1^c} \ll 1$  corresponds to a suppressed neutrinoless double beta decay rate, hence to a suppressed CP asymmetry



Normal hierarchy with  $m_1 \ll m_2$ ,  $\sin^2 \theta_{13} = 0.05$  and  $\sin 2\sigma = 1$  $\tan \beta = 10$ ,  $\lambda_{H_d} = 0$  and  $M_{\Delta}/M_{24} = 0.1$ 

#### We find that successful leptogenesis is possible for $M_{\Delta} \gtrsim 10^{12} \, {\rm GeV}$

This scale is problematic in view of the gravitino problem, which requires  $T_{RH} \lesssim (10^9 - 10^{10}) \,\text{GeV}$  in the most favourable cases (unstable gravitino with  $m_{3/2} \gtrsim 10 \,\text{TeV}$  or gravitino LSP with harmless NLSP for BBN)

#### <u>Ways out:</u>

- very light gravitino (< 16 eV required by WMAP)</li>
- very heavy gravitino (>> 100 TeV)
- non-thermal production of the triplets ( $T_{RH} \ll M_{\Delta}$ )
- non-supersymmetric scenario with a real 54

## Conclusions

• Ways to realize successful leptogenesis in GUTs:

- SO(10) models with a left-right symmetric seesaw mechanism: successful leptogenesis even with MD = Mu
- SO(10) models with SM fermions split among 16 and 10 matter multiplets and type II seesaw ⇒ predictive leptogenesis
- Work in progress [L. Calibbi, M. Frigerio, S.L., A. Romanino]
  - build complete SO(10) models with SM fermions in 16 and 10 matter multiplets
  - study flavour violating effects in these models

# Back-up slides

### Reconstruction of the heavy neutrino mass spectrum

The starting point is the left-right symmetric seesaw formula:

$$M_{\nu} = f v_L - \frac{v^2}{v_R} Y f^{-1} Y$$

with f,Y complex and symmetric. The goal is to reconstruct f assuming that Y is known in the basis of charged lepton mass eigenstates

Akhmedov and Frigerio (hep-ph/0509299) showed that there are  $2^n$  solutions for n generations, connected 2 by 2 by a "seesaw duality":

$$f \longrightarrow \hat{f} \equiv \frac{M_{\nu}}{v_L} - f$$

and provided explicit expressions for the fij up to n=3

In hep-ph/0606078, we proposed an alternative reconstruction procedure which employs complex orthogonal matrices

First rewrite the LR symmetric seesaw formula  $M_{\nu} = \alpha f - \beta Y f^{-1} Y$  as

$$Z = \alpha X - \beta X^{-1}$$
  
with  $\alpha \equiv v_L$ ,  $\beta \equiv v^2/v_R$  and  
$$Z \equiv N_Y^{-1} M_\nu (N_Y^{-1})^T \qquad X \equiv N_Y^{-1} f(N_Y^{-1})^T$$

where NY is such that  $Y = N_Y N_Y^T$  (Y invertible)

Z complex symmetric  $\Rightarrow$  can be diagonalized by a complex orthogonal matrix Oz if its eigenvalues zi are all distinct:

$$Z = O_Z \operatorname{Diag}(z_1, z_2, z_3) O_Z^T , \qquad O_Z O_Z^T = \mathbf{1}$$

Then X can be diagonalized by the same orthogonal matrix as Z, and its eigenvalues are the solutions of:

$$z_i = \alpha x_i - \beta x_i^{-1}$$
  $(i = 1, 2, 3)$ 

2 solutions  $x_i^+$ ,  $x_i^-$  for each i  $\Rightarrow$  2<sup>3</sup> = 8 solutions for X, hence for f:

$$f = N_Y O_Z \begin{pmatrix} x_1 & 0 & 0 \\ 0 & x_2 & 0 \\ 0 & 0 & x_3 \end{pmatrix} O_Z^T N_Y^T , \qquad x_i = x_i^{\pm}$$

The corresponding right-handed neutrino masses  $M_i = f_i v_R$  are obtained by diagonalizing f with a unitary matrix:

$$f = U_f \begin{pmatrix} f_1 & 0 & 0 \\ 0 & f_2 & 0 \\ 0 & 0 & f_3 \end{pmatrix} U_f^T , \qquad U_f U_f^{\dagger} = \mathbf{1}$$

and the couplings of the NR mass eigenstates are  $U_f^{\dagger}Y$ 

Note: diagonalization of a complex symmetric matrix by a complex orthogonal matrix

I) the eigenvalues of Z are the roots of  $Det(Z - z\mathbf{1}) = 0$ 

2) the eigenvectors associated with zi are the solutions of  $Z.\vec{v} = z_i \vec{v}$ 

It is always possible to find solutions of the latter equation, but in case of multiple solutions, it is not always possible to find an orthonormal basis of the eigenspace. The problem arises when one non-trivial solution has a zero norm in the SO(3, C) sense, i.e.  $\vec{v}.\vec{v} = 0$ ; then Z cannot be diagonalized.

If all eigenvalues of Z are distinct, the eigenvectors automatically satisfy  $\vec{v}.\vec{v} \neq 0$ , hence Z is diagonalizable (it can be written as  $O_Z \text{Diag}(z_1, z_2, z_3) O_Z^T$ )



Figure 5: Regions of the  $(m_1, v_R)$  parameter space where  $|Y_B| > Y_B^{WMAP}$  for solutions (+, +, +), (+, -, +) and (+, +, -), and where  $|Y_B| > 0.1 Y_B^{WMAP}$  for solution (-, -, -). These regions are delimited by the thick black contour in the (+, +, +) case, the dashed red contour for (+, -, +), the long-dashed blue contour for (+, +, -), and the thin black contour for (-, -, -). Inputs: set 1 of the Appendix for  $U_m$  and the high-energy phases; other input parameters as in Fig. 2.



Figure 6: The final baryon asymmetry as a function of  $v_R$  in the four reference solutions, for  $\delta_{PMNS} = 0$  and different values of  $\theta_{13}$ :  $\theta_{13} = 0^{\circ}$  (black), 2° (purple), 5° (blue), 9° (red) and 13° (green / light grey). Inputs: set 1 of the Appendix for  $U_m$  and the high-energy phases;  $T_{in} = 7 \times 10^9$  GeV for (+, +, +) and (+, -, +), while  $T_{in} = 5 \times 10^{10}$  GeV for (+, +, -) and (-, -, -); other input parameters as in Fig. 2.



Figure 7: Contour lines of the ratio  $|Y_B|/Y_B^{WMAP}$  in the four reference solutions, as a function of  $\theta_{13}$  and  $\delta_{PMNS}$ . The input parameters are the same as in Fig. 6, and the B-L breaking scale has been fixed at  $v_R = 5 \times 10^{13}$  GeV for (+, +, +) and (+, -, +), and at  $v_R = 6 \times 10^{13}$  GeV for (+, +, -) and (-, -, -).