

I MICROCANONICAL ENSEMBLE

§1. FOUNDATIONS OF STATISTICAL PHYSICS

Goal: Description of phenomena in
macroscopic world

- electrons in metals and insulators
- superconductivity and superfluidity
- magnetism
- different phases of water (ice, liquid, vapor)
- other phase transitions
- gases
- liquids
- amorphous systems, alloys
- polymers, membranes
- biological systems (DNA, amino acids, ...)
- ...

I Common feature:

huge number of components / particles



$$\sim 10^{19} - 10^{24} \frac{1}{\text{cm}^3}$$

Amount of matter in SI unit system is given in mol (from molecule) [20/05/2019]

1 mol of matter contains exactly

$$N_A = 6,02214076 \cdot 10^{23}$$

structural elements (need to be pointed out)

Avogadro number exact value by definition

$$N_A = 6,02214076 \cdot 10^{23} \text{ mol}^{-1}$$

[A day of mol is celebrated on Oct. 23
10/23 6⁰² am ÷ 6⁰² pm]

N_A - is approximately a number of ^{12}C carbon isotope atoms in 0,012 kg of carbon (old definition of mol).

www.gum.gov.pl

Examples

(i) 18 g of water (18 ml) - 1 mol H_2O

(ii) 22,4 l of the air - 1 mol of O_2 and N_2 together

(iii) 7,8 bil people on the Earth

(or 7,8 billion) $7,8 \cdot 10^9$ - $1,3 \cdot 10^{-14}$ mol of people

(iv) visible universe $3 \cdot 10^{23}$ stars - 1 mol of stars

□ (3)

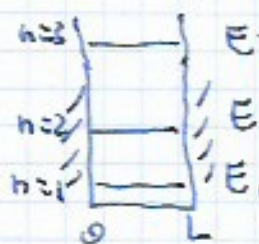
ii) Common feature:

Quantum systems \rightarrow discrete energy levels

Example

1) infinite quantum well

$$\hat{H} = \frac{\hat{p}^2}{2m}$$




$$E_n = \frac{\hbar^2 k_n^2}{2mL^2} n^2$$

$$n = 1, 2, 3, \dots$$

$$\psi_n(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi}{L}x\right)$$

2) harmonic oscillator

$$\hat{H} = \underbrace{\frac{\hat{p}^2}{2m}}_{E_{kin}} + \underbrace{\frac{k}{2}x^2}_{E_{pot}}$$



$$\omega = \frac{k}{2}$$

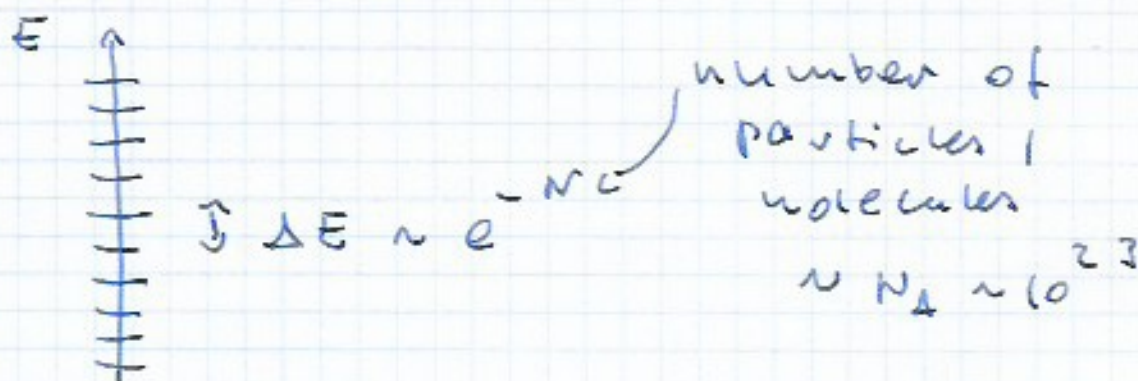
$$E_n = \hbar\omega \left(n + \frac{1}{2}\right)$$

$$n = 0, 1, 2, \dots \quad \frac{m\omega^2}{2} \frac{x^2}{2}$$

$$\psi_n(x) = N_n \psi_n(x) e^{-\dots}$$

□

In macroscopic systems distances between discrete levels are exponentially small



Hence, an arbitrary small ^(weak) perturbation (any interaction with environment) yields transitions from E to $E \pm \Delta E$ almost continuously in time.

See also: Ten theorems about quantum mechanics (3)
 measurements, D.G. van Kampen,
 Physical A113, p7-113 (1988)

Microscopic description of many body-system (mechanical)

Quantum mechanics

$$\Psi = \Psi(t, \underbrace{\vec{r}_1, \dots, \vec{r}_N}_{\sim 10^{23} \text{ variables}}) \quad \text{complex wave function}$$

Unitary evolution $i\hbar \frac{\partial \Psi}{\partial t} = \hat{H} \Psi$ - Schrödinger eq.

$$|\Psi(t, \vec{r}_1, \dots, \vec{r}_N)|^2 d^3N r = \text{probability to find } N \text{ particles at } \vec{r}_1, \vec{r}_2, \dots, \vec{r}_N \text{ at time } t.$$

this probability changes in time. The knowledge of the wave function seems to be impractical.

Classical mechanics

phase space

$$(\vec{q}, \vec{p}) = (\underbrace{\vec{q}_1, \dots, \vec{q}_N}_{6N \text{ degrees of freedom}}, \underbrace{\vec{p}_1, \dots, \vec{p}_N}_{\sim 10^{23}})$$

position \vec{q}_i

momentum

$$\vec{p}_i = m \frac{d\vec{q}_i}{dt}$$

Newton's equation

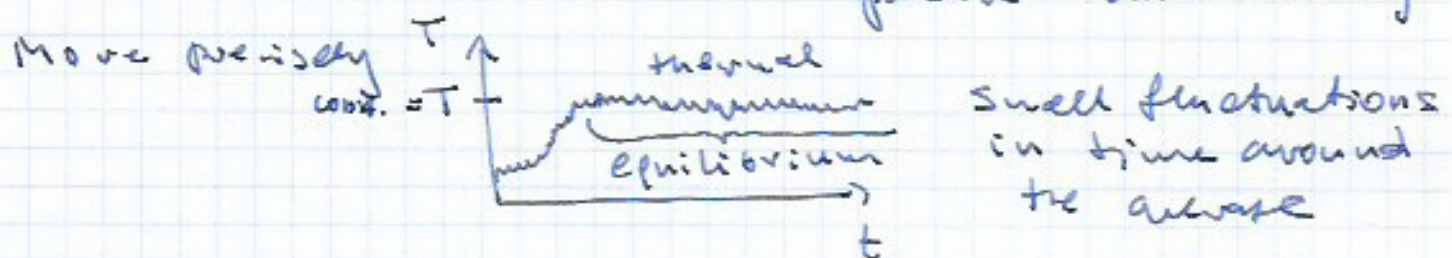
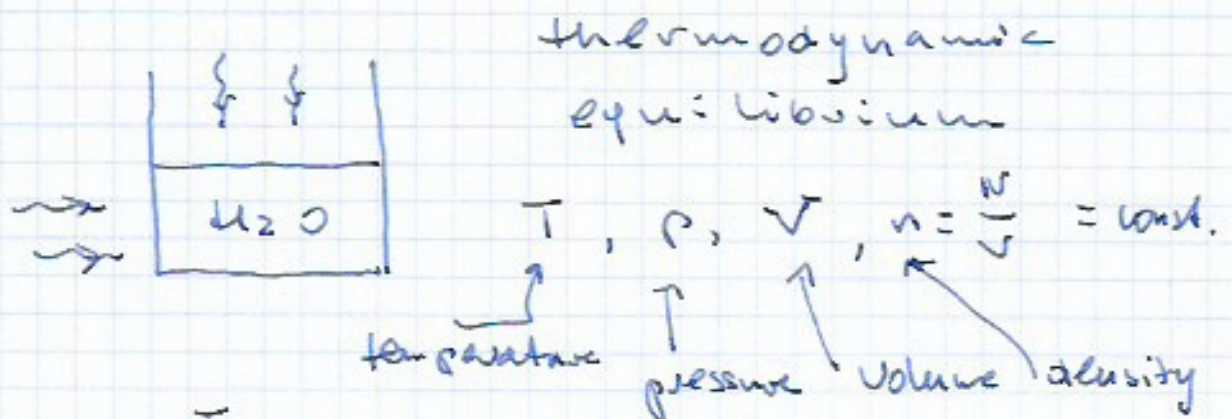
$$\frac{d\vec{p}_i}{dt} = \vec{F}_i + \sum_{j \neq i} \vec{F}_{ij}$$

$$i = 1, \dots, N \sim 10^{23}$$

Even if we know $(\vec{q}_i(t), \vec{p}_i(t)) \forall t$ this seems to be impractical.

Important observation

though $\Psi = \Psi(t)$, $(\bar{r}(t), \bar{p}(t))$ are changing in time due to chaotic motion of particles and an interaction with the environment some macroscopic properties of the system do not change



Thermodynamics (hydrodynamics) deals with a macroscopic description of many-body systems providing relationships between macroscopic quantities determined in experiments

- (p, v) - (pressure, volume)
- (\vec{B}, \vec{M}) - (magnetic induction, magnetization)
- (\vec{E}, \vec{P}) - (electric field intensity, polarization)
- (γ, L) - (surface tension, length)
- \vdots

Statistical physics wants to derive thermodynamic relations from microscopic equations of motion.

Two important concepts

(i) Microstate ^(mechanical state) - a state of a system at time t given by $\Psi(t, \vec{r}_1, \dots, \vec{r}_N)$ or $(\vec{p}(t), \vec{r}(t))$ or any other quantities describing the system as much accurately as possible.

(ii) Macrostate (thermodynamical state) - a state of a system described by few relevant macroscopic (thermodynamic) variables:

either (U, V, N, \dots) - extensive variables
(proportional to number of particles)
↑ internal energy ↑ volume ↑ number of particles

or (T, p, μ, \dots) - intensive variables
(independent from a number of particles)
↑ temperature ↑ pressure ↑ chemical potential

or their combinations.

Example: two-level system

binary alloy, spins $1/2$, atoms on optical lattices,
ions

consider N spins $1/2$

Microstate: one of 2^N combinations of
spin arrangements

$\left\{ \begin{array}{cccccccccccc} \uparrow & \downarrow & \downarrow & \downarrow & \uparrow & \uparrow & \downarrow & \downarrow & \uparrow & \uparrow & \downarrow & \dots & \dots & \dots & \uparrow & \downarrow \\ 1 & 2 & 3 & 4 & 5 & 6 & 7 & 8 & 9 & 10 & 11 & \dots & \dots & \dots & N-1 & N \end{array} \right\}$

Macrostate: macroscopic variable $M = N_{\uparrow} - N_{\downarrow}$
a state with a fixed M

e.g. $N=3$, $2^3 = 8$ - microstates

$\{\uparrow\uparrow\uparrow\}$	$M = 3$	} four possible macrostates
$\{\downarrow\uparrow\uparrow\}, \{\uparrow\downarrow\downarrow\}, \{\uparrow\downarrow\uparrow\}$	$M = 1$	
$\{\uparrow\downarrow\downarrow\}, \{\downarrow\downarrow\uparrow\}, \{\downarrow\uparrow\downarrow\}$	$M = -1$	
$\{\downarrow\downarrow\downarrow\}$	$M = -3$	

Two macrostates $n = \pm 1$ are realized by
three microstates, ~~two~~ correspondingly.

More during tutorials. □

§ 2. BASIC POSTULATE OF STATISTICAL PHYSICS

We have seen that ~~for~~ a given macrostate can be realized by many different microstates.

Which one does the nature choose?

Ludwig Boltzmann - hypothesis of molecular chaos:

For a closed / isolated system in thermal equilibrium all possible microstates realizing a given macrostate are equally probable.

Two views:

elements of ignorance { 1) If we cannot decide in which equivalent microstate the system exists then because of our ignorance we have to average over all possible microstates with the same probabilities.

self-averaging hypothesis { 2) Measurement means to observe a system for some time. In this time a system can change microstates many times ($\sim 10^{22}$). The observed quantity is an average over those microstates.

$\{ \uparrow \uparrow \downarrow \} \quad \{ \uparrow \downarrow \uparrow \} \quad \{ \uparrow \uparrow \downarrow \downarrow \} \quad \{ \downarrow \uparrow \uparrow \} \quad \{ \uparrow \downarrow \uparrow \}$

$P_{\uparrow \uparrow \downarrow}$

$P_{\uparrow \downarrow \uparrow}$

$P_{\uparrow \uparrow \downarrow \downarrow}$

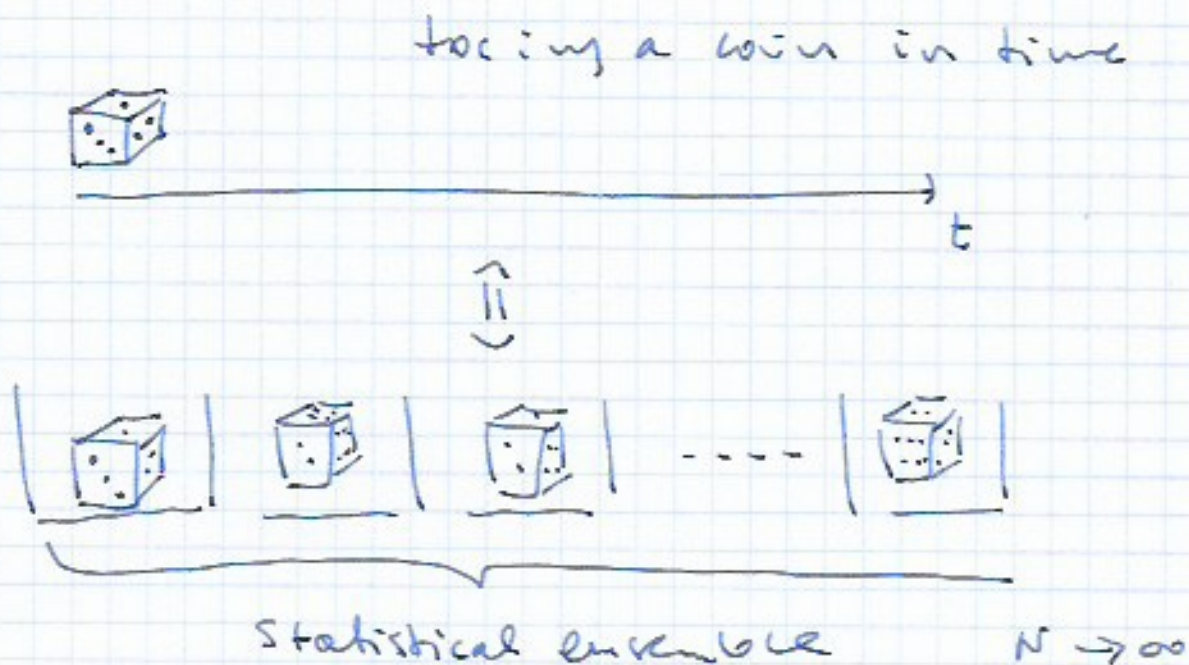
$P_{\downarrow \uparrow \uparrow}$

$P_{\uparrow \downarrow \uparrow}$

time \rightarrow

We need to know these probability distributions (9)

According to ergodic hypothesis of Boltzmann we can consider an ensemble of systems with many (infinitely many) elements with given probabilities.



Instead of averaging in time a process of observation we average over many samples from a statistical ensemble of this system (no more time is involved)

Statistical physics \leftrightarrow thermodynamics

Description and understanding of a macroscopic system is statistical in its origin.

How to find corresponding probabilities?

From Boltzmann hypothesis all microstates realizing a given macrostate are equally probable a priori.

This statement is sufficient to find all statistical properties of the system.

How it works?

Ω - number of all microstates (partition function)

y_i - a variable characterizing a given macrostate

Ω_i - number of microstates with the same y_i

From the postulates of equal probabilities a priori we get

$i = 1, \dots, n$

$$P_i = P(y = y_i) = \frac{\Omega_i}{\Omega}$$

Consistency proof \uparrow random variable

$$P_i = \sum_{\substack{\text{all} \\ \text{microstates} \\ \text{with } y=y_i}} \frac{1}{\Omega} = \frac{\Omega_i}{\Omega} \quad \square$$

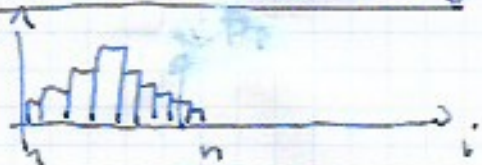
Average $\langle y \rangle = \sum_{i=1}^n y_i P_i = \frac{1}{\Omega} \sum_{i=1}^n y_i \Omega_i$

n-th moment $\mu_n = \langle y^n \rangle = \sum_{i=1}^n y_i^n P_i = \frac{1}{\Omega} \sum_{i=1}^n y_i^n \Omega_i$

Variance $\sigma_y^2 = \langle (y - \langle y \rangle)^2 \rangle = \langle y^2 \rangle - \langle y \rangle^2 = \mu_2 - \mu_1^2$

probability distribution of random variable y

$$P(y) = \frac{\Omega_i}{\Omega} = \frac{\Omega(y=y_i)}{\Omega}$$



In case of continuous random variable y

$$P(y) = \frac{\Omega(y)}{\Omega}$$



$$\mu_n = \langle y^n \rangle = \int_{-\infty}^{\infty} dy y^n P(y) = \frac{1}{\Omega} \int_{-\infty}^{\infty} dy y^n \Omega(y)$$

e.g. average $\langle y \rangle = \frac{1}{\Omega} \int_{-\infty}^{\infty} y \Omega(y) dy$

normalization

$$1 = \langle 1 \rangle = \frac{1}{\Omega} \underbrace{\int_{-\infty}^{\infty} dy \Omega(y)}_{\Omega} = \frac{\Omega}{\Omega} = 1$$

Warning: in statistical physics with many degrees of freedom the integral measure has to be modified properly

$$d_{3N} \mathbf{p} \, d_{3N} \mathbf{r} \rightarrow \frac{d_{3N} \mathbf{p}}{h^{3N}} \frac{d_{3N} \mathbf{r}}{N!}$$

↑
Planck
constant.

see later on.