

E.g. Luminance [lx]

noon in shade	weak	20 - 40	Webb's law
	medium	40 - 80	
	strong	60 - 100	(feeling) \propto
Screen in cinema		80 - 250	log (source)
Paper in sunshine		100 000	

E.g. Electric bulb

514

7 W

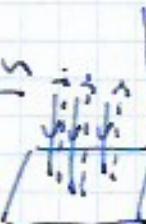
470 nm \leftrightarrow  in tri. set

warm white

25000 h

Absorption, Emission, Kirchhoff's Law

Df. absorption is

$$A(\omega, \tau) = \frac{\text{absorbed energy at } \omega}{\text{total energy at } \omega}$$


$$[A] = [1]$$

$\tau = A(\omega, \tau, \text{chemistry, colour, phase, etc.})$

$$A(\omega, \tau) = \text{constant. } \forall \omega - \text{gray body}$$

$$A(\omega, \tau) = 1 - \text{black body}$$

$$A(\omega, \tau) = 0 - \text{perfect mirror}$$

Df. emission

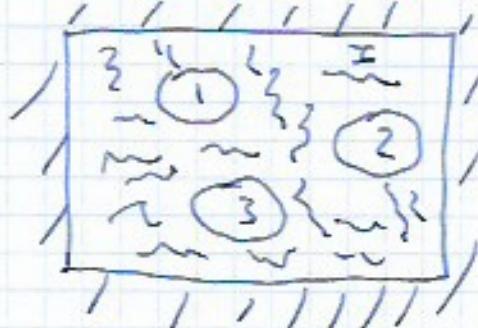
$$E(\omega, \tau) = \text{emitted energy at } \omega \text{ per time}$$

$$[E] = [W = \frac{J}{s}]$$

18.59 - Kirchhoff's Law

$$\frac{E(\omega, T)}{A(\omega, T)} = \varepsilon(\omega, T) \quad - \text{universal function of } \omega \text{ and } T \text{ for all bodies.}$$

Proof. \Rightarrow bodies in thermal equilibrium with radiation inside an isolated mirror box.



\Rightarrow Each body emits and absorbs the same

amount of energy ΔS and Δt
(if not in thermal equilibrium: impairs)

$$E_1(\omega, T) \Delta S, \Delta t_1 = A_1(\omega, T) \Delta S, \Delta t + I(\omega, T)$$

\uparrow light intensity
(the same for all
bodies)

$$\Rightarrow \frac{E_1(\omega, T)}{A_1(\omega, T)} = \frac{E_2(\omega, T)}{A_2(\omega, T)} = \dots = I(\omega, T) = \varepsilon(\omega, T)$$

□

For a black body ~~radiation~~ $A_{bb}(\omega, T) = 1$

$$\frac{E_{bb}(\omega, T)}{A_{bb}(\omega, T)} = \varepsilon(\omega, T) = I(\omega, T) = \frac{c}{4} u(\omega, T)$$

$$\Rightarrow \boxed{\varepsilon(\omega, T) = \frac{c}{4} u(\omega, T) \sim \text{Planck distribution for all bodies.}}$$

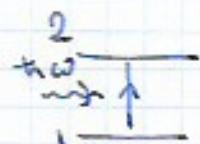
\uparrow
Planck
distribution

\Rightarrow A body more radiates if it more absorbs.

E.g. black coffee wool faster than water at the same T.

Einstein coefficients

1) Absorption



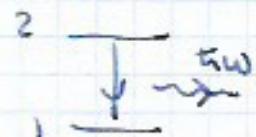
transition rate

$$\left(\frac{dN}{dt} \right)_{ab} = N_1 B_{12} \mu(\omega)$$

↑
occupation
of initial
state

↑ density
energy
proportionality
coefficient

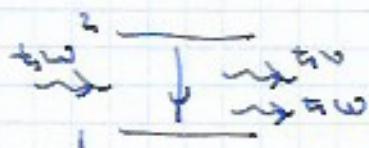
2) Spontaneous emission



$$\left(\frac{dN}{dt} \right)_{sp-em} = A N_2$$

↑ coefficient ↑ occupation of
initial state

3) Stimulated emission



$$\left(\frac{dN}{dt} \right)_{st-em} = N_2 B_{21} \mu(\omega)$$

Detailed balance condition

$$\left(\frac{dN}{dt} \right)_{ab} = \left(\frac{dN}{dt} \right)_{sp-em} + \left(\frac{dN}{dt} \right)_{st-em}$$

$$N_1 B_{12} \mu(\omega) = N_2 A + N_2 B_{21} \mu(\omega)$$

In thermal equilibrium

$$N_1 = \text{const. } e^{-\beta E_1}, \quad N_2 = \text{const. } e^{-\beta E_2} \Rightarrow \frac{N_1}{N_2} = e^{-\beta(E_2-E_1)} = e^{\beta \Delta E}$$

$$\mu(\omega) = \frac{N_2 A}{N_1 B_{12} - N_2 B_{21}} = \frac{A}{B_{21}} \cdot \frac{1}{e^{\beta \Delta E} - 1}$$

In the limit $T \rightarrow \infty$

$$B_{12} = B_{21} = B \Rightarrow \mu(\omega) = \frac{A}{B} \cdot \frac{1}{e^{\beta \Delta E} - 1}$$

absorption = stimulated emission

Comparing with Planck distribution

$$\boxed{\frac{A}{B} = \frac{\omega^3 k}{h^2 c^3}}$$

A and B are Einstein
coefficients

Relation to absorption and emission

$$\frac{E(\omega, \tau)}{A(\omega, \tau)} = \frac{c}{4} \mu(\omega, \tau) = \frac{c}{4} \frac{A}{B} \frac{1}{e^{\beta \hbar \omega / k} - 1}$$

bivariate analysis

$$[A] : \left[\frac{1}{s} \right] = \left[\left(\frac{dn}{dt} \right)_{\text{em}} \right] = A n_2 \Rightarrow [A] = \left[\frac{1}{s} \right]$$

$$[B] : \left[\frac{1}{s} \right] = \left[\left(\frac{dn}{dt} \right)_{\text{abs-em}} \right] = [B N_2 \mu(\omega)] =$$

$$= [B N_2 \underbrace{\frac{\hbar \omega^3}{\pi^2 c^3}}_{\text{abs}} \frac{1}{e^{\beta \hbar \omega / k} - 1}]$$

$$[B] = \left[\frac{1}{s} \cdot \frac{\omega^3 / s^3}{2 \cdot s \cdot \frac{\hbar \omega^3}{\pi^2 c^3}} = \frac{\omega^3}{s} \underbrace{\frac{1}{4 \pi^2 c^3} \frac{s^3}{\hbar \omega^3}}_{\text{abs}} = \frac{m}{k_B T} \right]$$

$$\Rightarrow \boxed{\left[\frac{A}{B} \right] = \left[\frac{k_B T}{m \cdot s} \right]}$$

On the other hand:

$$\frac{E(\omega, \tau)}{A(\omega, \tau)} : \frac{A}{B} \frac{c}{4} \frac{1}{e^{\beta \hbar \omega / k} - 1}$$

$$\left[\frac{E}{A} \right] = \left[\frac{\frac{1}{s}}{\frac{1}{s} \cdot \frac{\hbar \omega^3}{\pi^2 c^3}} \right] = \left[\frac{k_B \cdot \frac{m}{s^2}}{\frac{1}{s} \cdot \frac{\hbar \omega^3}{\pi^2 c^3}} = \frac{k_B}{s^2} \right]$$

$$\Rightarrow \boxed{\left[\frac{A}{B} \right] = \left[\frac{k_B}{s^2} \frac{1}{m \cdot s} \right] = \left[\frac{k_B}{m \cdot s} \right]}$$

□

§ 3 SPECIFIC HEAT OF SOLIDS - PHONONS

Df. heat capacity

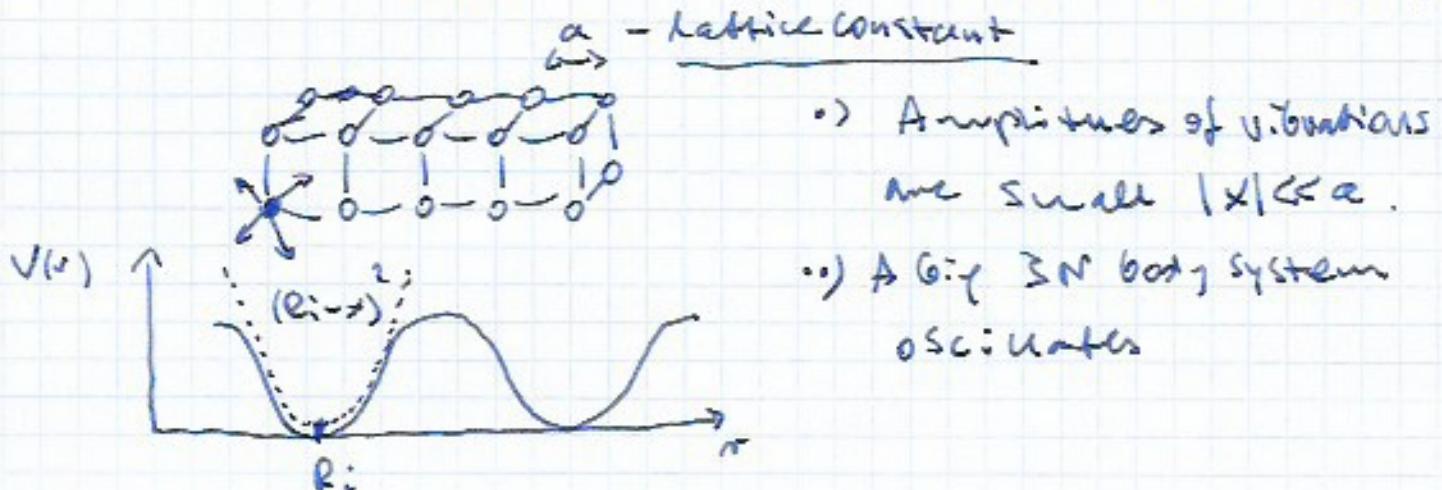
$$C_x = \left(\frac{\delta Q}{\delta T} \right)_x \quad \left[\frac{J}{K} \right]$$

at x - constant during the process, extensive quantity.

heat capacities per unit mass or per mole are called specific heats $\left[\frac{J}{kg \cdot K} \right]$ and are intensive quantities.

They measure how the system responds to the uptake of heat from its surroundings - thermodynamic response function

Solid crystals - ions vibrate around equilibrium positions, lattice sites.
(not move around the volume like in gas or liquid)



⇒ Amplitudes of vibrations are small $|x| \ll a$.

⇒ A 64g SiN body system oscillates

...) In this small amplitudes limit we have a set of 3N (coupled) harmonic oscillators

Energy of a harmonic crystal

$$\mathcal{E}(q, p) = \sum_{n=1}^{3N} \left(\frac{p_n^2}{2m_n} + \frac{1}{2} m_n \omega_n^2 q_n^2 \right)$$

m_n, ω_n - constants in harmonic approximations

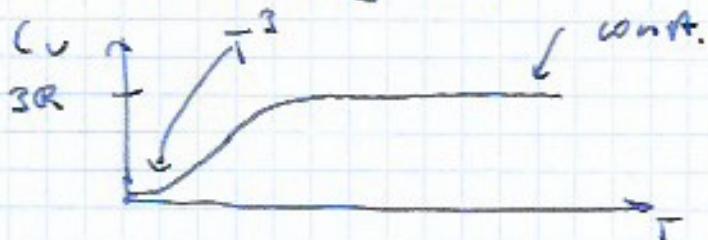
q_n, p_n - generalized coordinates

History:

→ Classical statistical physics predicted that

$$\begin{cases} U = \langle \mathcal{E} \rangle = 3Nk_B T & \text{for one mode} \\ C_V = \left(\frac{\partial U}{\partial T} \right)_V = 3Nk_B \approx 3R \approx \text{const.} & \\ & [\text{Dulong-Petit law (1812)}] \end{cases}$$

→ in second half of 19th c. it was shown experimentally that



••• Einstein model (1907) - used Planck distribution to describe crystal vibrations

if $\omega_n = \omega$ - each atom is the same oscillator

$$m_n = m$$

$$U = 3N_A \langle \varepsilon \rangle = 3N_A \left[\frac{\hbar \omega}{2} + \frac{\hbar \omega}{e^{\frac{\hbar \omega}{k_B T}} - 1} \right]$$

Einstein temperature

$$C_V = \frac{\partial U}{\partial T} = 3R \left(\frac{T_E}{T} \right)^2 \frac{e^{\frac{T_E}{T}}}{(e^{\frac{T_E}{T}} - 1)^2}$$

$$T_E = \frac{\hbar \omega}{k_B}$$

$$C_V \sim \begin{cases} \frac{3k}{2} R & T \gg T_E \\ 3k \left(\frac{T_E}{T}\right)^2 e^{-\frac{T_E}{T}} & T \ll T_E \end{cases}$$

exponential decay at low T

- v) P. Debye (1912) model - took more general model, where each modes can have different frequencies (anharmonic)

$$\omega = \omega_{\text{max}} = \omega_0$$

and assumed that frequency is quantized.
 In a range $(\omega, \omega + d\omega)$ there are
 $\rho(\omega) d\omega$ normal modes
 T kinds of states.

Now, ω_{max}

$$U = \int_0^{\omega_{\text{max}}} \frac{\hbar \omega}{2} \rho(\omega) d\omega + \int_0^{\omega_{\text{max}}} \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1} \rho(\omega) d\omega$$

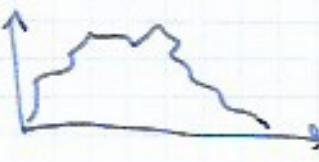
ω_{max} - maximal frequency of vibrations.

Proof:

$$\sum \frac{\hbar \omega e^{-\beta \hbar \omega}}{e^{\beta \hbar \omega} - 1} = \int d\omega \underbrace{\frac{1}{\hbar} \delta(\omega - \omega_0)}_{\rho(\omega)} \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1}$$

□

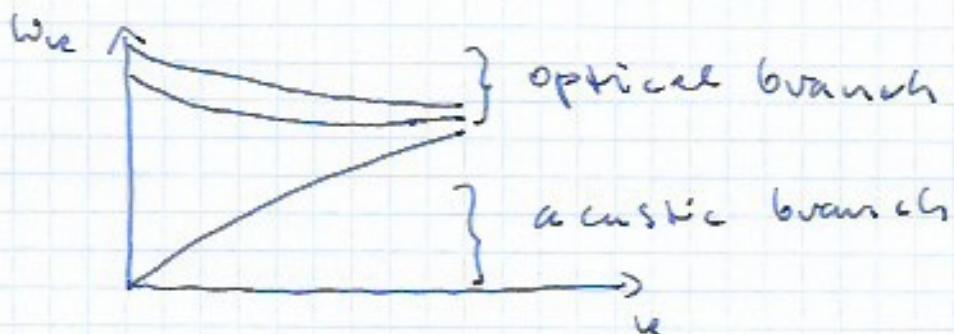
Practically, does $\rho(\omega)$ is determined numerically (density functional theory or neutron scattering)



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Finding $\tilde{\omega}$: \Rightarrow equivalent \rightarrow know $\tilde{\omega}_{\text{tot}}$
 $\tilde{\omega}_{\text{tot}}$.

At low wave vector \tilde{k} :



- There exists one mode with linear dispersion relation $\tilde{\omega}_{\text{tot}} = v |\tilde{k}|$

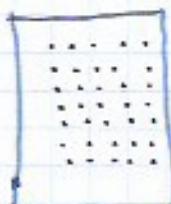
Goldstone theorem: broken of continuous

symmetry \rightarrow one zero mode
translational symmetry

discrete symmetry



gas,
liquid



Solid

Debye model

$$\omega_n = v |\tilde{k}| = \omega \frac{\pi}{L} |n| \quad n = \sqrt{n_x n_y n_z}$$

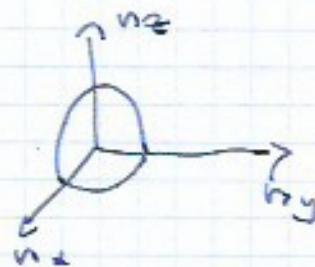
$$\sum_n (\dots) = \frac{3}{8} \int_0^{n_{\max}} 4\pi n^2 (\dots) dn \quad \begin{matrix} \text{live in} \\ \text{cavity} \\ \text{radiation} \\ \text{but ...} \end{matrix}$$

3 - # of polarization

n_{\max} - there are $3N$ degrees of freedom
 (finite number of atoms; infinite number of photons)

$$\frac{1}{2} \pi n_D^3 = 3N$$

$$n_D \approx n_{max}$$



$$3 \cdot \frac{1}{8} \cdot \frac{4}{3} \pi n_D^3 = 3N$$

interpretation:

$$k_D = \frac{2\pi}{L} n_D \sim \frac{\sqrt[3]{N}}{L} = \frac{2\pi}{\lambda_D}$$

$$\lambda_D \sim \frac{h}{\sqrt[3]{N}} N^{1/3} a_1 \quad - \text{lattice constant}$$

The shortest wave length cannot be shorter than the interatomic distance

λ_D - Debye length

k_D - Debye wave vector

ω_D - Debye frequency

$$\omega_D = v k_D$$

Internal energy

$$U = \sum_n \frac{E_{\text{kin}}}{e^{\beta E_{\text{kin}}} - 1} = \frac{3}{8} \int_0^{n_0} h \pi n^2 \frac{k_B T \frac{\pi}{L} n}{e^{\frac{k_B T}{k_B T} \frac{\pi}{L} n} - 1} dn$$

$$\left\{ \omega = \beta \frac{\pi}{L} n \right\} = \frac{3.45}{8} \left(\frac{L}{\pi k_B T} \right)^4 k_B T \frac{\pi}{L} \int_0^{x_0} \frac{x^3}{e^x - 1} dx$$

$$x_0 \equiv \frac{\pi k_B T n_0}{L k_B T}$$

$$= \frac{3 \pi^2 k_B T}{2 L} \left(\frac{k_B T L}{\pi k_B T} \right)^4 \int_0^{x_0} \frac{x^3}{e^x - 1} dx$$

$$x_0 \equiv \frac{\Theta}{T} = \frac{k_B \Theta}{k_B T}, \quad \boxed{k_B \Theta = \left(\frac{k_B T}{k_B} \right) \left(\frac{6 \pi^2 N}{V} \right)^{1/3} = \text{Debye temperature}}$$

At low $T \ll \Theta$

$$\int_0^{x_0} \frac{x^3}{e^x - 1} dx \approx \int_0^{\infty} \frac{x^3}{e^x - 1} dx = \frac{\pi^4}{15}$$

$$\boxed{U = \frac{3 \pi^4 N k_B T^4}{5 \Theta^3} \quad T \ll \Theta}$$
$$\boxed{C_V = \left(\frac{\partial U}{\partial T} \right)_V = \frac{12 \pi^4 N k_B}{5} \left(\frac{T}{\Theta} \right)^3}$$

T^3 - Debye law

$\Theta [K]$	470	343	105	645	374	267	225	165
	Fe	Cu	Pb	Si	Ge	As	Ag	Au
metals								
semiconductors								

Phonons - quanta of crystal lattice vibrations - quanparticles similar to photons, emergent particles

$$1+1 \neq 2$$

Science 177, 393 (1972)

D. W. Anderson

Mo is different

Emergent properties - Fundamentally new phenomena in complex systems (quanparticles) not deduced (predicted *a priori*) from the knowledge of elementary particles of the system.

Determinism

$$1+1=2$$

reductionism

(Determinists)

Emergence

$$1+1 \neq 2$$

Darwin, Boltzmann,
Landau, Anderson

Geometric Example

• point

• • points

— segment / line (∞ points)

// concept of parallel

X concept of angle

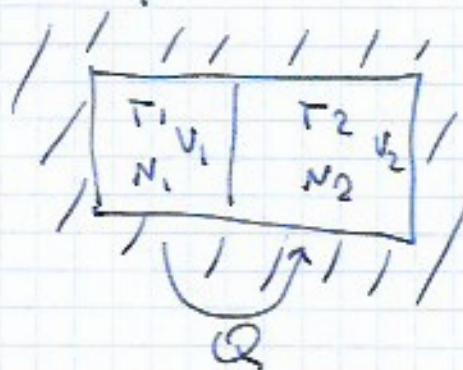
□ concept of area

□ concept of volume

IV GRAND CANONICAL ENSEMBLE

§ A. CHEMICAL POTENTIAL

- We investigated systems in thermal interaction (contact) with exchanging an entropy



$$T_1 > T_2$$

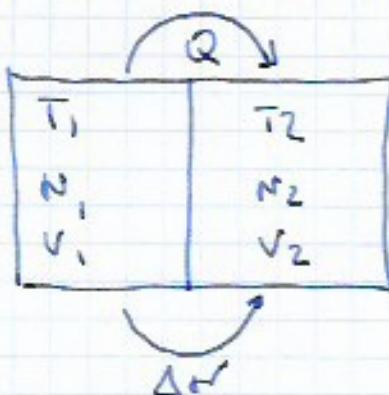
heat ΔQ "flows" from
① → ② until

$$T_1 = T_2$$

$$\frac{1}{T_1} = \left(\frac{\partial S_1}{\partial U_1} \right)_{V_1, N_1} = \left(\frac{\partial S_2}{\partial U_2} \right)_{V_2, N_2} = \frac{1}{T_2}$$

From the maximization of the thermodynamic probability $S = k_B \ln \Omega$

- now we add a chemical interaction (contact) with exchanging particles opposite of thermal interaction thermal + diffusive contact



$$T_1 > T_2$$

A system is in thermal and in chemical equilibrium if

$$T_1 = T_2$$

$$\beta_1 = \beta_2$$

What is β_1

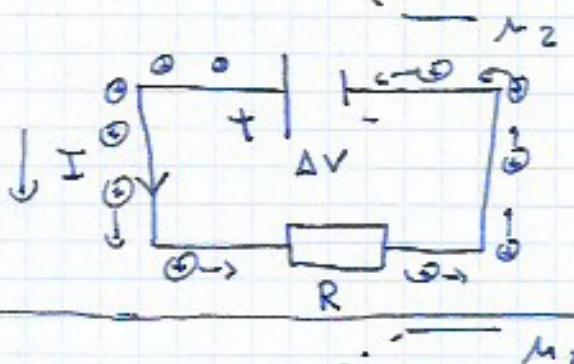
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A chemical potential governs a flow of particles, e.g., a battery
 has an electric potential ΔV

$\mu_1 > \mu_2 \Rightarrow$ flow of current I

μ_1

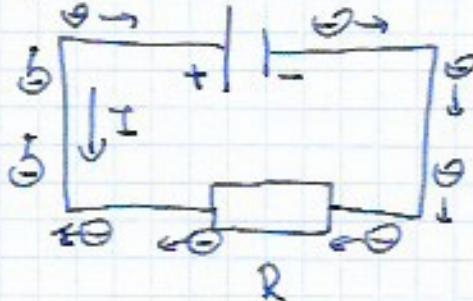
$$\mu_1 - \mu_2 = q\Delta V$$



motion of positive $q > 0$
 charges from μ_1 to μ_2 .

μ_1

$$\mu_1 - \mu_2 = q\Delta V$$



motion of negative $q < 0$
 charges from μ_2 to μ_1

How to define a chemical potential in thermodynamics?