## **Gluing variations**

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based on joint work with Wan Cong arXiv:2302.06928 [gr-qc]



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- I glue together vacuum spacetimes?
- realise data on lower dimensional submanifolds by embedding in a vacuum spacetime?
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The Aretakis-Czimek-Rodnianski question

#### QUESTION (Aretakis, Czimek and Rodnianski (2021))

Can you find vacuum characteristic initial data interpolating between two sphere data sets?



Figure: Gluing construction of Aretakis-Czimek-Rodnianski

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Answer: "kind of", for sphere data near spheres lying on a Minkowskian light cone.



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Vacuum jets

#### Let *P* be a submanifold of *M*. Let $k \in \mathbb{N} \cup \{\infty\}$ .

#### Definition

Let g be any smoothly differentiable metric defined in a neighborhood of P. The collection

$$j^k g := \{\partial_{lpha_1} \cdots \partial_{lpha_\ell} g_{\mu
u}|_{P}\,, \; \mathbf{0} \leq \ell \leq k\}$$

will be called *jet of order k of g at P*.

Einstein equations and their derivatives up to order k - 2provide equations, differential and/or algebraic, relating the jets of order k at P. A jet will be called vacuum if all such equations are satisfied universität

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Image: A mathematical states and a mathem

Spacelike/timelike/null vacuum submanifold data

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The collection of all vacuum jets will be called vacuum submanifold data of order *k* and will be denoted by  $\Psi[P, k]$ .

A jet of order *k* of a metric *g* will be called {

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#### QUESTION

Example: Vacuum jets at a point

Let *P* be a point,  $P = \{p\}$ 

 $j^k g =$  the coefficients of the Taylor series of a metric g at p. vacuum  $\equiv$  algebraic conditions on the Taylor coefficients Now: in normal coordinates the Taylor coefficients can be expressed in terms of the Riemann tensor and its covariant derivatives.

For example, using normal coordinates,

 $j^2 g|_{
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Example: Vacuum jets at a point (cf. Friedrich's proof of Geroch multipole expansions)

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## Vacuum spacelike hypersurface data

Vacuum spacelike constraint equations

*Initial data* surface  $\Sigma$ , Riemannian metric  $g_{ij}$ , i, j = 1, ..., n, symmetric tensor  $K_{ij}$  ("initial time derivative of the metric")

the scalar constraint equation (A is the cosmological constant):

$$R(g_{ij}) = 2\Lambda + |\mathcal{K}|^2 - (\mathrm{tr}\mathcal{K})^2 \; ,$$

and the vector constraint equation:

$$D_j K^j{}_k - D_k K^j{}_j = 0 \; .$$

Corollary

Spacelike vacuum hypersurface data

 $\Psi[\Sigma,\infty] \approx \Psi[\Sigma,k] \approx \Psi[\Sigma,2] \approx \{all \ vacuum \ (g,K)\}$ 

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Can one extend vacuum spacelike initial data on a manifold with boundary beyond the boundary?

#### Theorem

Let  $(\Sigma, g, K)$  be spacelike vacuum initial data on a manifold with boundary  $\partial \Sigma$ .

There exists a manifold without boundary  $\check{\Sigma}$  and vacuum initial data ( $\check{g},\check{K}$ ) on  $\check{\Sigma}$  such that  $\Sigma \subset \check{\Sigma}$ , with

 $(\check{g},\check{K})|_{\Sigma}=(g,K).$ 

Proof: I will give a sketch; for this we will need *characteristic* vacuum hypersurface data  $\Psi[\mathscr{N},\infty]$ .



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### Characteristic Cauchy problem



Figure: Characteristic Cauchy problem with intersecting null hypersurfaces

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## Characteristic Cauchy problem

Isenberg-Moncrief coordinates



The hypersurfaces  $\mathcal{N} = \{u = 0\}$  and  $\underline{\mathcal{N}} = \{r = 0\}$  are characteristic for the metric

$$\mathbf{g}_{\mu\nu}\mathrm{d}x^{\mu}\mathrm{d}x^{\nu} = 2\left(-\mathrm{d}u + u\alpha\mathrm{d}r + u\beta_{A}\mathrm{d}x^{A}\right)\mathrm{d}r + g_{AB}\mathrm{d}x^{A}\mathrm{d}x^{B}.$$

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(0.1)

For each set of characteristic initial data

 $\beta_A$  on  $\mathbf{S} := \mathscr{N} \cap \mathscr{N}$  and  $(\alpha, g_{AB})$  on  $\mathscr{N} \cup \mathscr{N}$ ,

subject to the Raychaudhuri equation on  $\mathcal{N} \cup \underline{\mathcal{N}}$ ,

$$0 = -\frac{1}{2}g^{AB}\partial_r^2 g_{AB} + \frac{1}{4}g^{CA}g^{BD}(\partial_r g_{AB})\partial_r g_{CD} + \frac{1}{2}\alpha g^{AB}\partial_r g_{AB}$$
(0.2)
here exists a vacuum metric in a future neighborhood of
$$\mathcal{N} \cup \mathcal{N}.$$

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Figure: Gluing construction of Aretakis-Czimek-Rodnianski





Figure: Gluing construction of Aretakis-Czimek-Rodnianski



Answer: one needs to understand data of order k o



Figure: Gluing construction of Aretakis-Czimek-Rodnianski

### QUESTION What about a vacuum metric in a whole neighborhood of $\mathcal{N} \cup \mathcal{N}$ ? What about a single characteristic hypersurface?

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Figure: Gluing construction of Aretakis-Czimek-Rodnianski

# QUESTIONWhat about a vacuum metric in a whole neighborhood of $\mathcal{N} \cup \mathcal{N}$ ? What about a single characteristic hypersurface?

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Answer: one needs to understand data of order k on  $\mathcal{N}$ 

#### Proposition

In the Isenberg-Moncrief coordinate system the vacuum characteristic initial data  $\Psi(\mathcal{N}, k)$  can be reduced to

 $\Phi_{\mathrm{IM}}[\mathscr{N}, \mathbf{k}] := \{ (\partial_{\mathbf{u}}^{j} g_{AB}, \beta_{A})_{0 \le j \le k} \text{ on } \mathbf{S} \text{ and } (g_{AB}, \alpha) \text{ on } \mathscr{N} \},$  (0.3)

where **S** is a cross-section of  $\mathcal{N}$ .

Proof: transverse derivatives of the metric on a characteristic hypersurface  $\mathscr{N}$  are determined uniquely by the above data through ODEs along the null geodesics threading  $\mathscr{N}$  or through algebraic equations.

Note: This is true both to the future or to past along the wiversität generators.

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The "hand-crank construction"



Figure: The "hand-crank construction".



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Proof: Use the crank:

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## Corollary: Embedding a truncated cone

Use the spacelike data extension



Figure: Extending a vacuum metric on a truncated future cone  $J^+(p) \cap J^-(\mathscr{S})$  to a neighborhood thereof.

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## Corollary: Embedding data at a point

Use the embedding of a cone



#### Figure: Extending data at a point



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## Corollary: Extending a characteristic future development

The Fledermaus = two cranks (and a spacelike extension if need be)



#### Figure: The Fledermaus.





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