



**Neutrino-Collider Synergy
for Beyond the Standard
Model Searches**

Artwork by Sandbox Studio, Chicago with Ana Kova

University of Warsaw

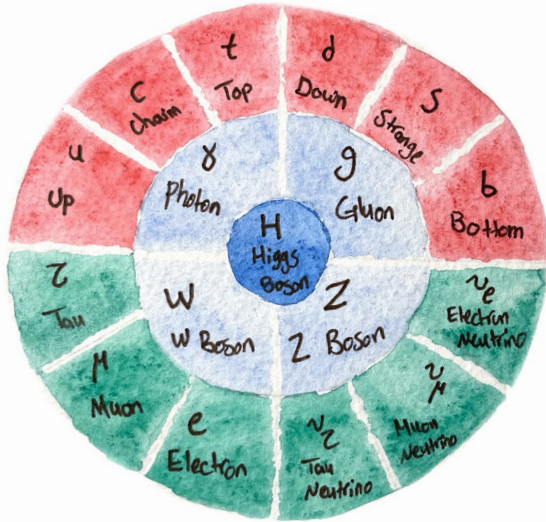
March 26, 2026

Zahra Tabrizi
PITT-PACC Langley Fellow
University of Pittsburgh



Standard Model

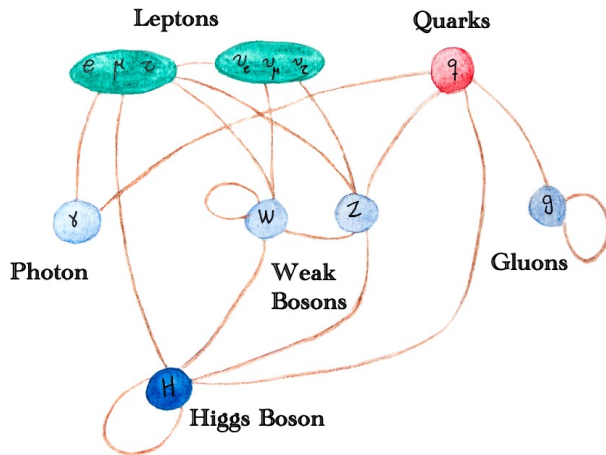
- What are the Fundamental particles?



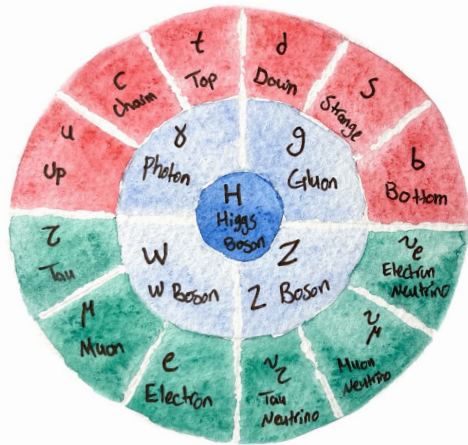
The Standard Model:
“Periodic Table” of
elementary particles

Neutrinos are among the
elementary particles!

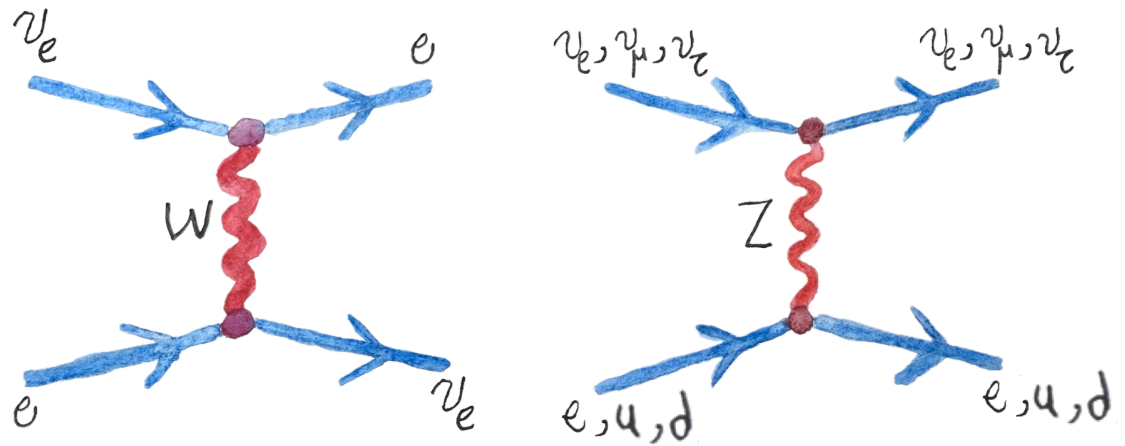
- How do they interact?



Why are neutrinos special?

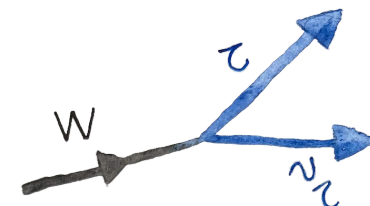
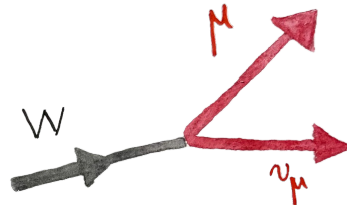
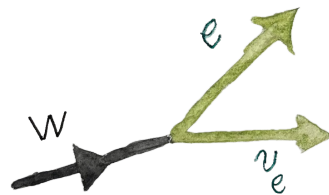


They interact with matter only by W/Z exchange



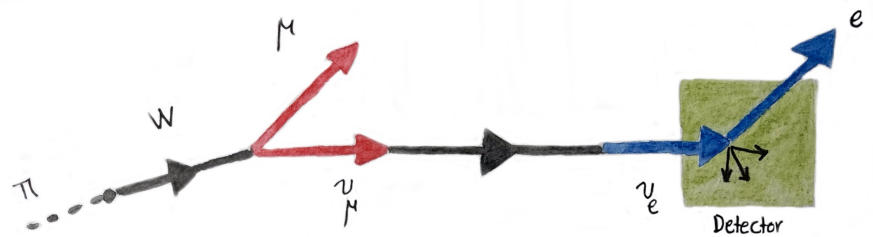
Three known Flavors

We *define* the three known flavors of neutrinos by W boson decays:



Neutrino Oscillation

They change Flavor



long enough journey

$$\nu_{\mu} \longrightarrow \nu_e$$

$$\mathcal{H} |\nu_k\rangle = E_k |\nu_k\rangle$$

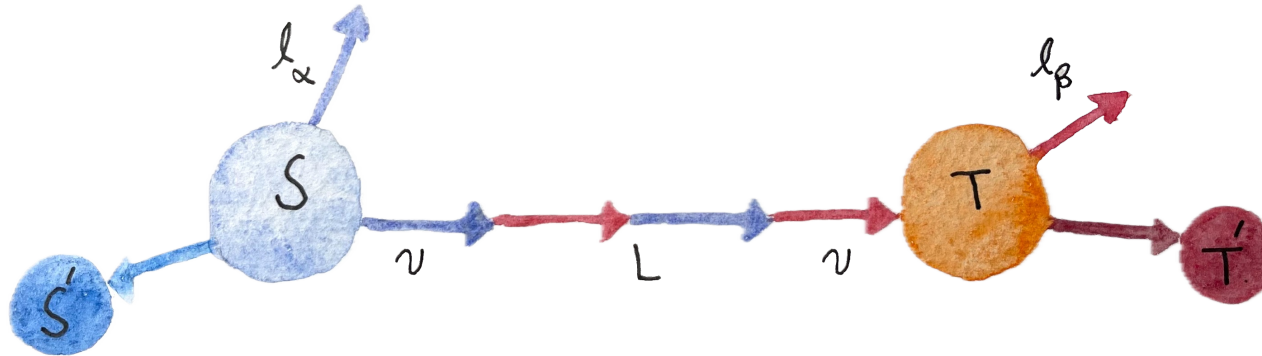
$$\begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix} = U_{\text{PMNS}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Oscillation Parameters:

$$P_{e\mu}(L) = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E_{\nu}} \right)$$

$$\Delta m_{kj}^2 \equiv m_k^2 - m_j^2$$

Oscillation Experiments



Observable: rate of detected events

$$\sim (\text{flux}) \times (\text{det. cross section}) \times (\text{oscillation})$$

Depend on the kinematics and spin variables!

Depends on mixing angles/masses

$$U_{\text{PMNS}} \equiv \begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix} \begin{bmatrix} \blacksquare & \blacksquare & \blacksquare \\ \blacksquare & \blacksquare & \blacksquare \\ \blacksquare & \blacksquare & \blacksquare \end{bmatrix} \begin{matrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{matrix}$$

Oscillation experiments are sensitive not only to oscillation parameters, but also to how neutrinos interact with matter!



Status of Neutrino Physics in 2026

Super-Kamiokande, Borexino, SNO

mixing angles:

$$\sin^2 \theta_{12} @ 4\%$$

$$\sin^2 \theta_{13} @ 3\%$$

$$\sin^2 \theta_{23} @ 3\%$$

mass squared differences:

$$\Delta m_{21}^2 @ 3\%$$

$$|\Delta m_{31}^2| @ 1\%$$

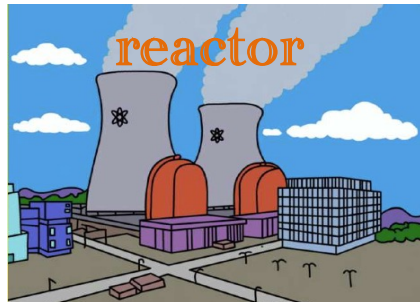
Future: DUNE, Hyper-K, JUNO

Open oscillation questions:

- CP-phase?
- Mass hierarchy?

Also:

- Neutrino mass mechanism?
- Mass scale?
- Dirac or Majorana?



MBL: Daya Bay, RENO, Double Chooz

LBL: KamLAND



IceCube, Super-Kamiokande



T2K, MINOS, NOvA



Status of Neutrino Physics in 2026

Super-Kamiokande, Borexino, SNO

mixing angles:

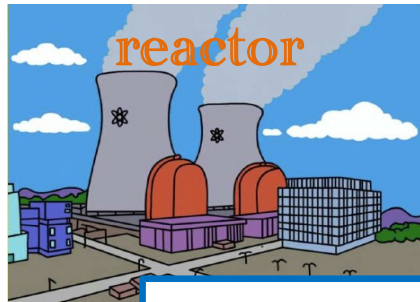
$$\sin^2 \theta_{12} @ 4\%$$

$$\sin^2 \theta_{13} @ 3\%$$

$$\sin^2 \theta_{23} @ 3\%$$

mass squared differences:

$$\Delta m^2 @ 3\%$$



MBL: Daya Bay, RENO, Double Chooz

LBL: KamLAND



New Neutrino Experiments are being built to answer unknowns in the oscillation picture

- Mass hierarchy?

Also:

- Neutrino mass mechanism?
- Mass scale?
- Dirac or Majorana?



T2K, MINOS, NOvA

GOAL:



The SM is not wrong...

But there is more than
“reasonable doubt” that it is
incomplete!

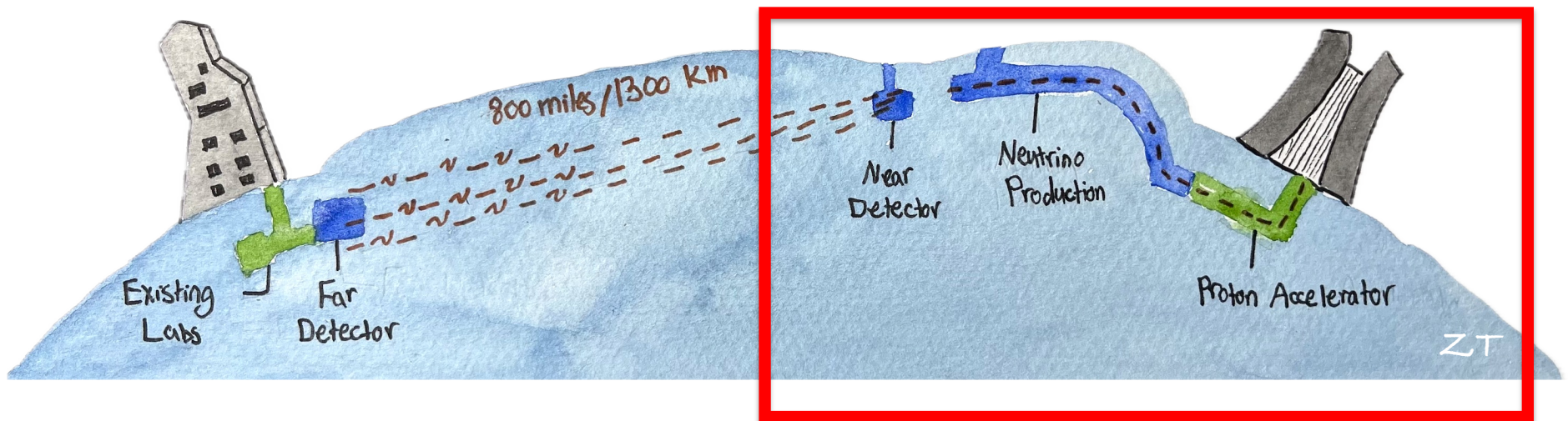
- Precision measurements at neutrino oscillation experiments;
- Finding the interplay between the low energy and high energy frontiers in BSM searches;
- Specific BSM examples, by either indirect search, or direct production;

Accelerator Neutrino Experiments:

Near Detector: Understanding Systematic Uncertainties

Far Detector: Measuring Oscillation Parameters

- Test SM predictions
- Search for BSM physics



How do we go beyond the SM?

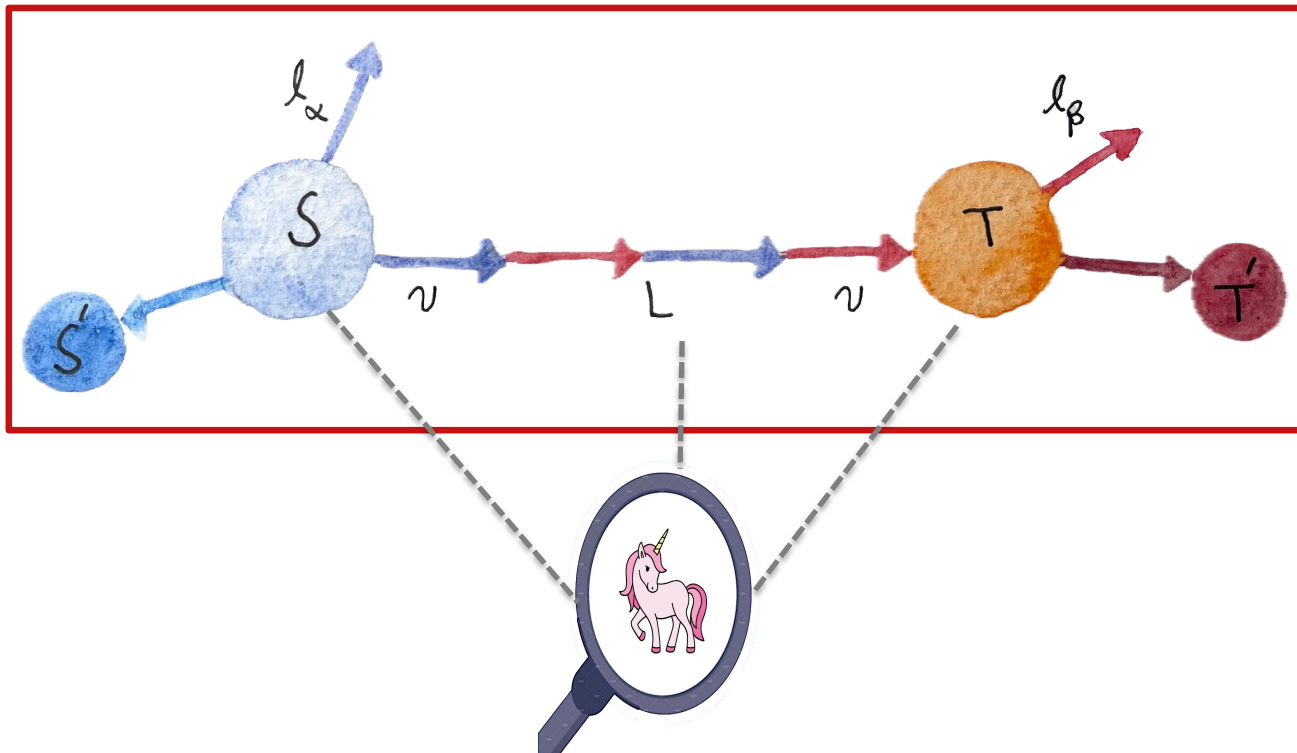


We extend the SM:

- New Particles
- New Interactions

1. Indirect Effects of Heavy New Physics

Modify Neutrino Interactions

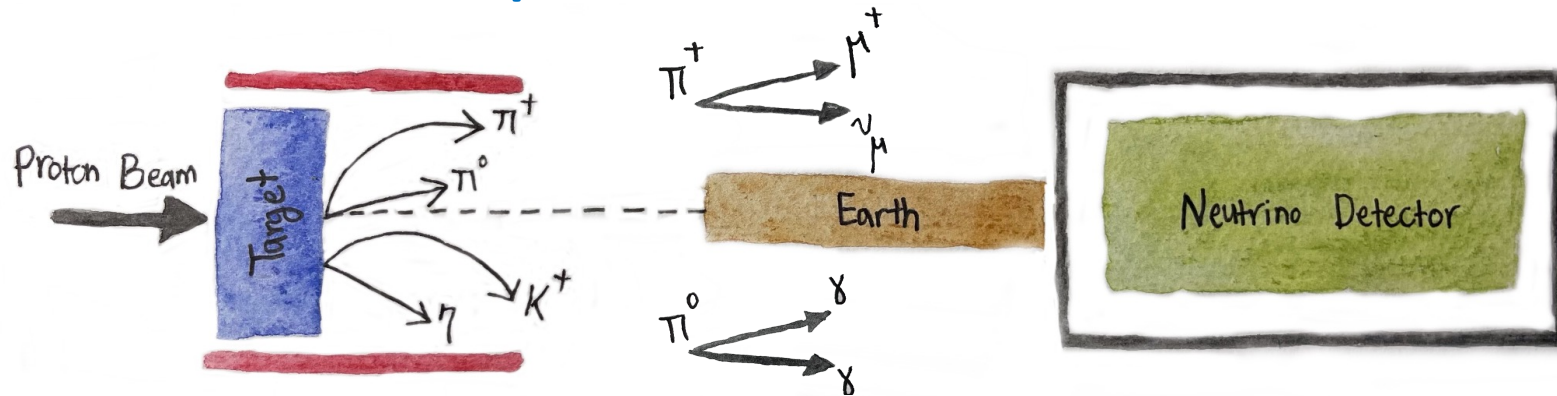


Observable: rate of detected events

$$\sim (\text{flux}) \times (\text{det. cross section}) \times (\text{oscillation})$$

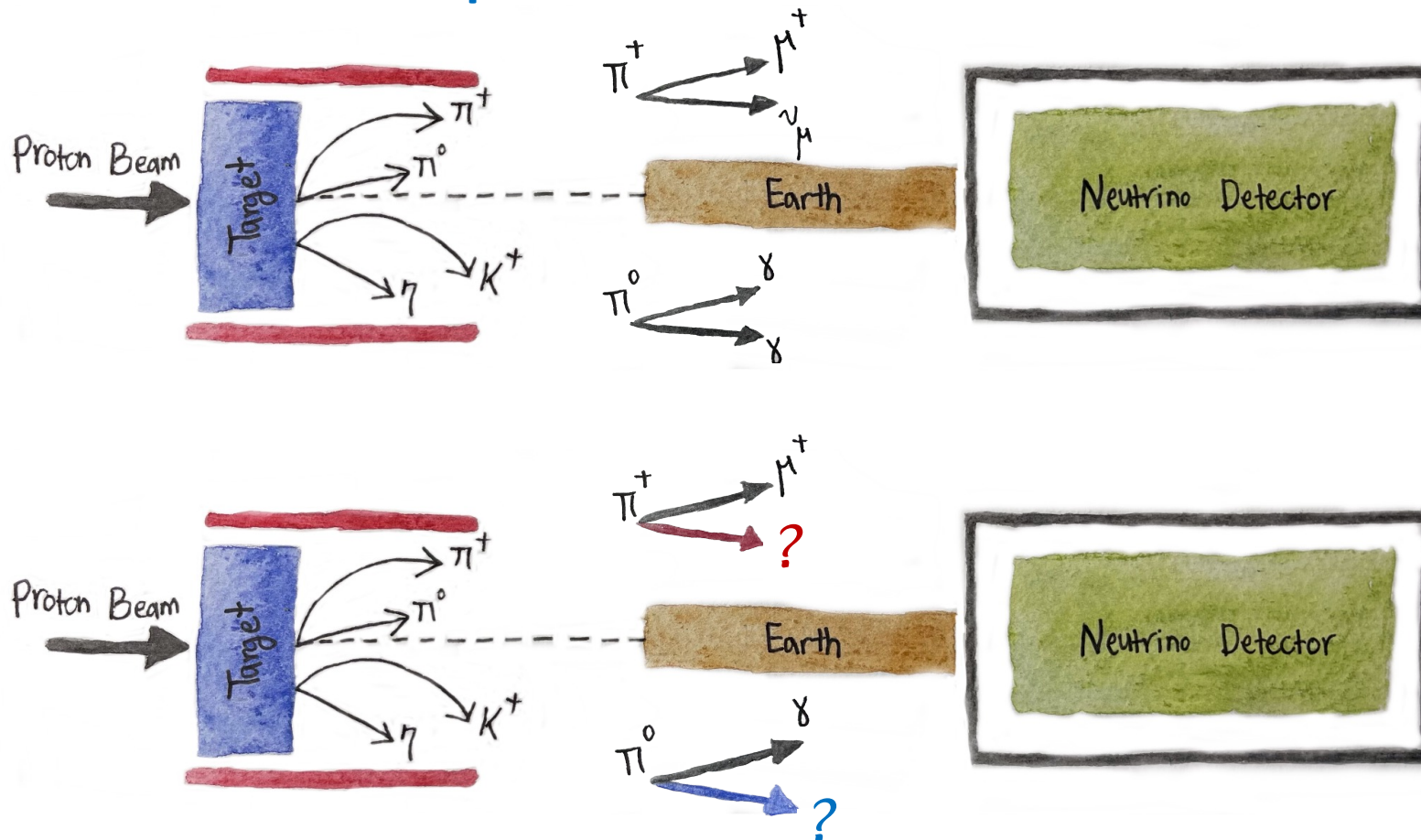
2. Direct Production of New Particles

Neutrino Experiments as Dark Sector Factories



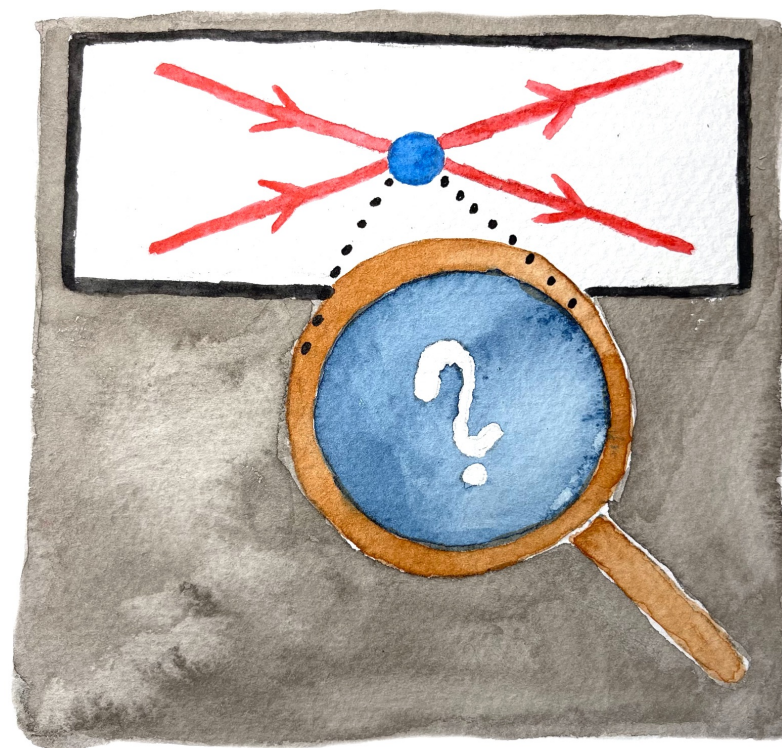
2. Direct Production of New Particles

Neutrino Experiments as Dark Sector Factories



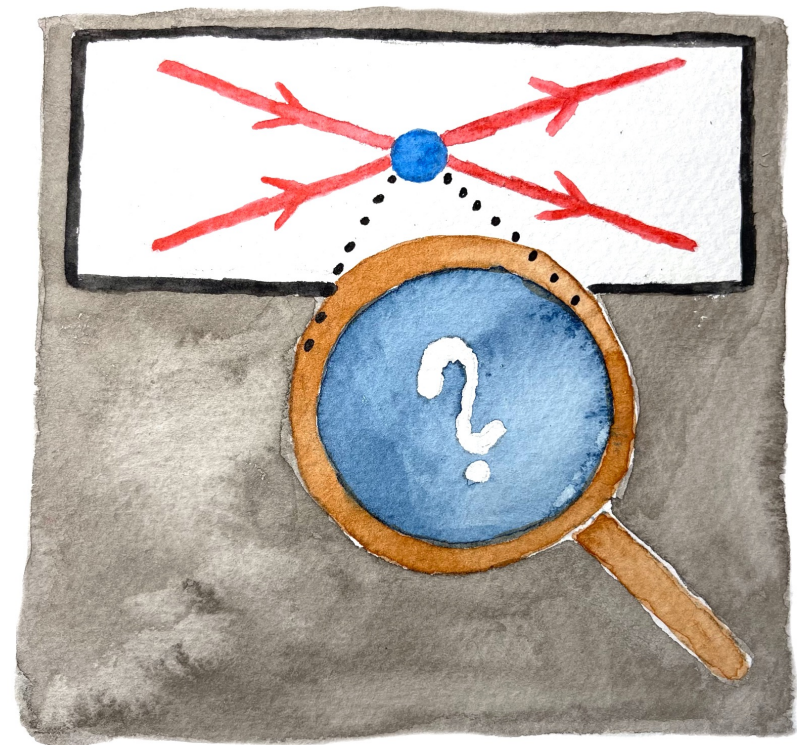
Outline

- 1) Indirect Search-EFT:
 - Why EFT?
 - EFT ladder
 - EFT at Neutrino Experiments
- 2) Direct Search:
 - Axion-Like Particles
- Conclusion



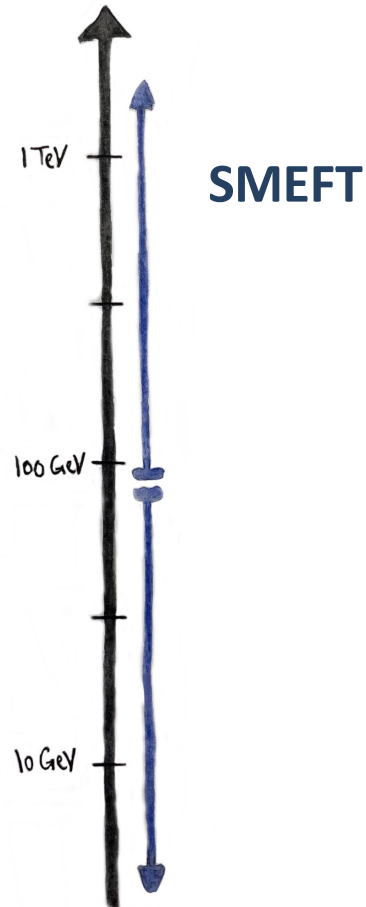
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EFT Workflow:

EFT Energy Scale



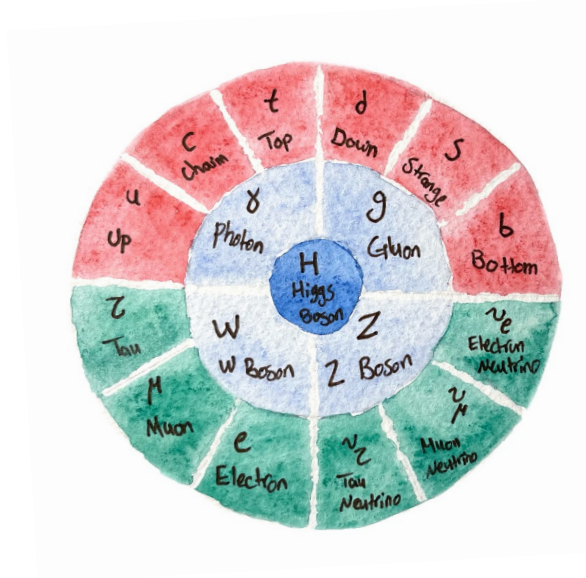
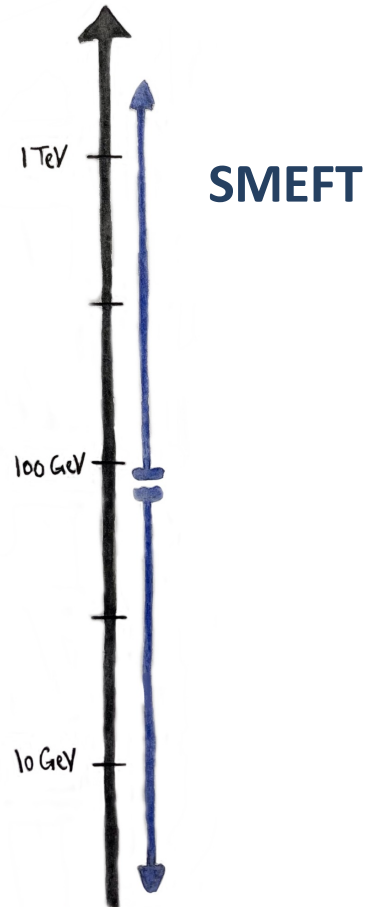
- New physics could be very heavy;
- It may affect the couplings between the SM particles, or induce new interactions;
- SM Effective Field Theory (SMEFT) is a systematic framework to describe such effects;

EFT Workflow:

SMEFT: minimal EFT above the weak scale

$$\mathcal{L}_{\text{SMEFT}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{D=5} + \mathcal{L}_{D=6}$$

EFT Energy Scale



EFT Workflow:

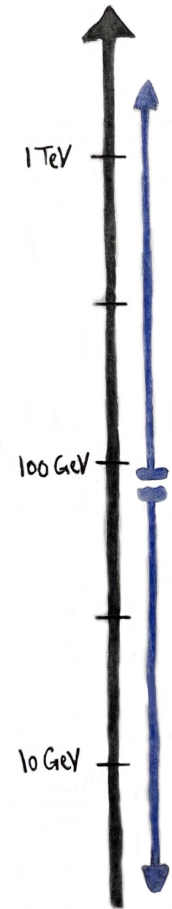
SMEFT: minimal EFT above the weak scale

$$\mathcal{L}_{\text{SMEFT}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{D=5} + \boxed{\mathcal{L}_{D=6}}$$

Known SM
Lagrangian

Gives neutrino
Masses

EFT Energy Scale



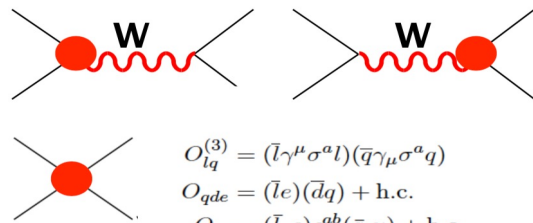
SMEFT

- Colliders



2499 distinct D=6 operators in the Warsaw basis, including flavor structure and CP conjugates

Grzadkowski, Iskrzynski, Misiak, Rosiek
JHEP (2010)



$$O_{lq}^{(3)} = (\bar{l}\gamma^\mu\sigma^a l)(\bar{q}\gamma_\mu\sigma^a q)$$

$$O_{qde} = (\bar{l}e)(\bar{d}q) + \text{h.c.}$$

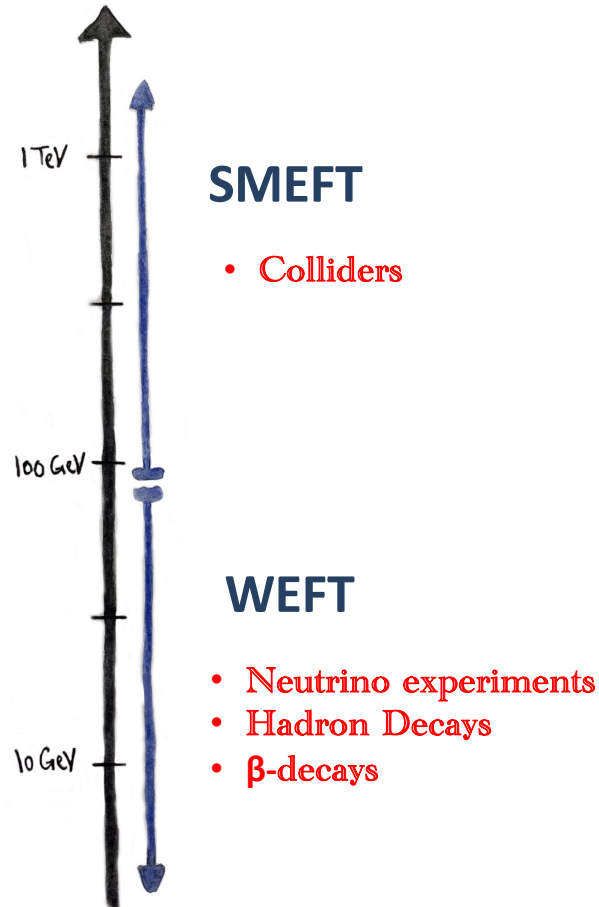
$$O_{lq} = (\bar{l}_a e)\epsilon^{ab}(\bar{q}_b u) + \text{h.c.}$$

$$O_{lq}^t = (\bar{l}_a\sigma^{\mu\nu}e)\epsilon^{ab}(\bar{q}_b\sigma_{\mu\nu}u) + \text{h.c.}$$

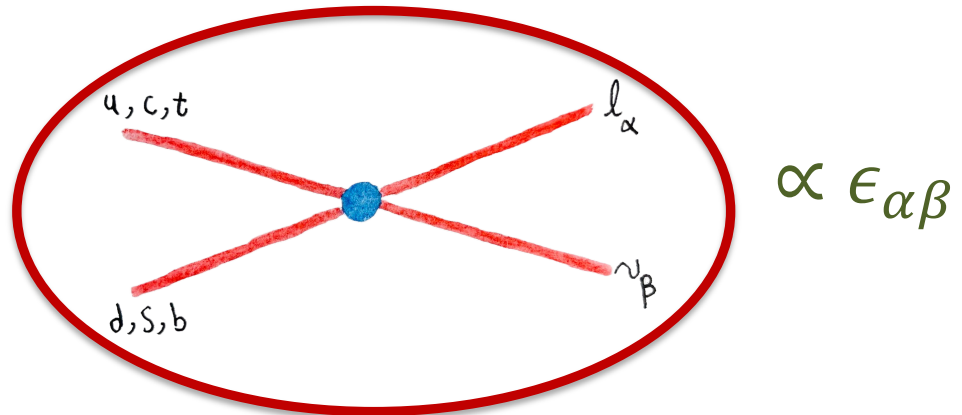
EFT Workflow:

EFT Energy Scale

WEFT: Effective Lagrangian defined at a low scale $\mu \sim 2 \text{ GeV}$



New 4-Fermion Interactions:

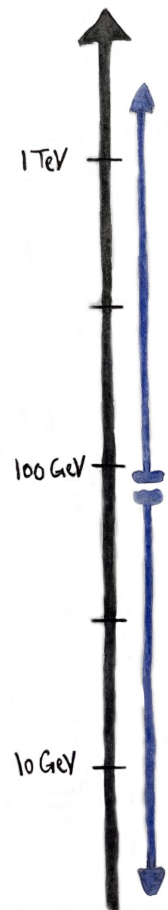


$$\mathcal{L}_{\text{WEFT}} \supset -\frac{2V_{ud}}{v^2} \left\{ [1 + \epsilon_L]_{\alpha\beta} (\bar{u}\gamma^\mu P_L d)(\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) + \epsilon_R]_{\alpha\beta} (\bar{u}\gamma^\mu P_R d)(\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) + \frac{1}{2} \epsilon_S]_{\alpha\beta} (\bar{u}d)(\bar{\ell}_\alpha P_L \nu_\beta) - \frac{1}{2} \epsilon_P]_{\alpha\beta} (\bar{u}\gamma_5 d)(\bar{\ell}_\alpha P_L \nu_\beta) + \frac{1}{4} \hat{\epsilon}_T]_{\alpha\beta} (\bar{u}\sigma^{\mu\nu} P_L d)(\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{h.c.} \right\}$$

EFT Workflow:

At the scale m_Z WEFT parameters ϵ_X map to dim-6 SMEFT operators

EFT Energy Scale



SMEFT

- Colliders

WEFT

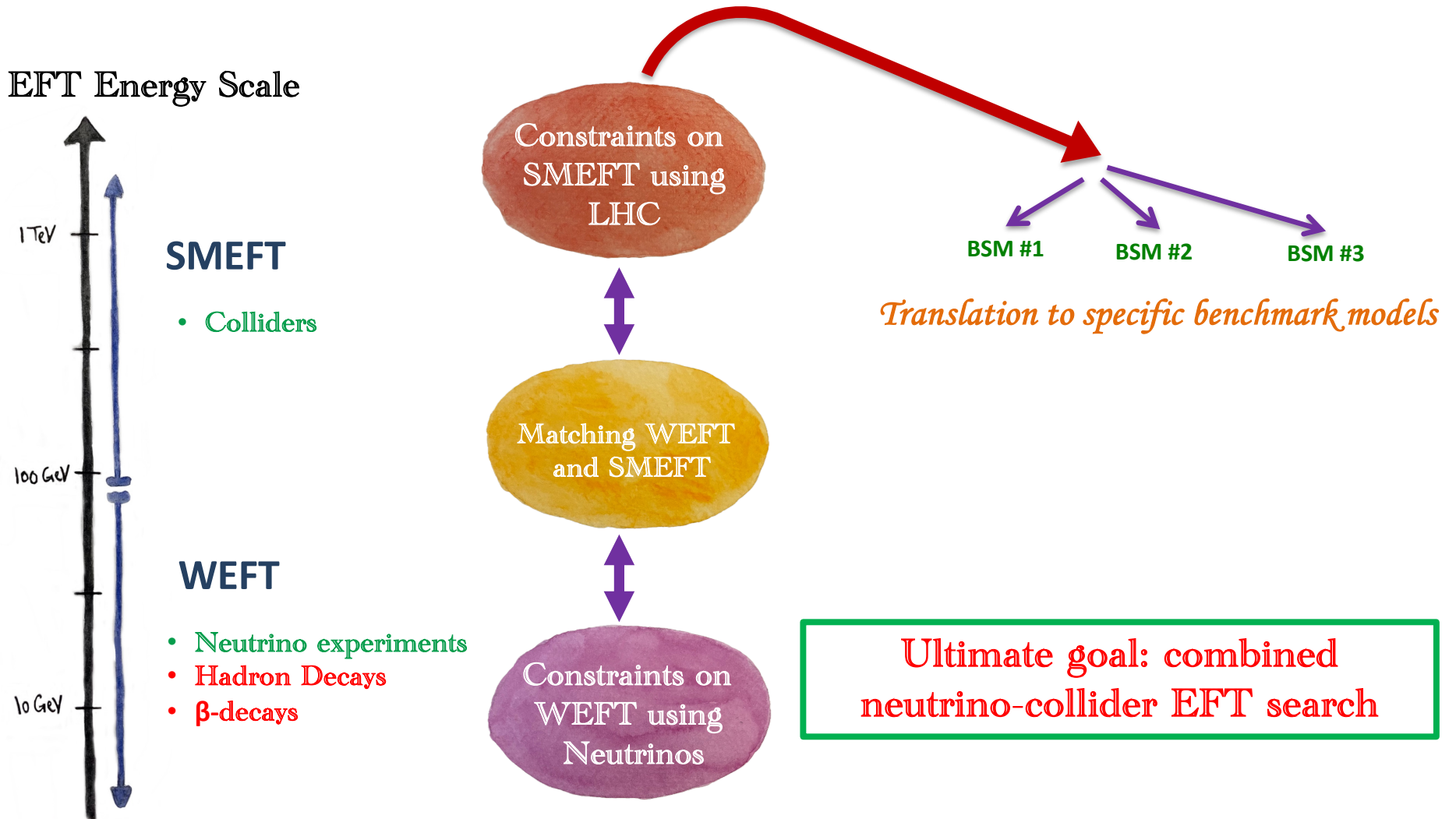
- Neutrino experiments
- Hadron Decays
- β -decays

$$\begin{aligned}
 [\epsilon_L]_{\alpha\beta} &\approx \frac{v^2}{\Lambda^2 V_{ud}} \left(V_{ud} [c_{HI}^{(3)}]_{\alpha\beta} + V_{jd} [c_{Hq}^{(3)}]_{1j} \delta_{\alpha\beta} - V_{jd} [c_{lq}^{(3)}]_{\alpha\beta 1j} \right) \\
 [\epsilon_R]_{\alpha\beta} &\approx \frac{v^2}{2\Lambda^2 V_{ud}} [c_{Hud}]_{11} \delta_{\alpha\beta}, \\
 [\epsilon_S]_{\alpha\beta} &\approx -\frac{v^2}{2\Lambda^2 V_{u_i}} \left(V_{jd} [c_{lequ}^{(1)}]_{\beta\alpha j 1}^* + [c_{ledq}]_{\beta\alpha 11}^* \right), \\
 [\epsilon_P]_{\alpha\beta} &\approx -\frac{v^2}{2\Lambda^2 V_{u_i}} \left(V_{jd} [c_{lequ}^{(1)}]_{\beta\alpha j 1}^* - [c_{ledq}]_{\beta\alpha 11}^* \right), \\
 [\hat{\epsilon}_T]_{\alpha\beta} &\approx -\frac{2v^2}{\Lambda^2 V_{ud}} V_{jd} [c_{lequ}^{(3)}]_{\beta\alpha j 1}^*,
 \end{aligned}$$

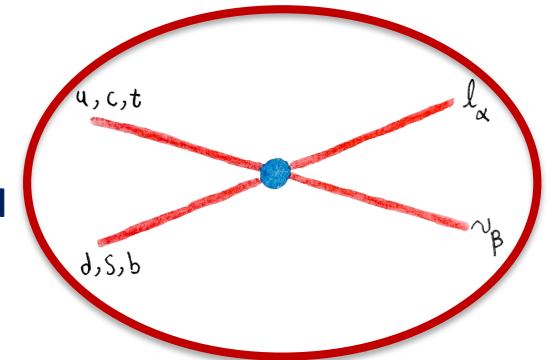
Falkowski, González-Alonso, ZT, JHEP (2019)



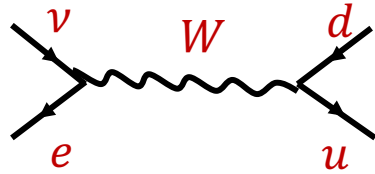
EFT Workflow:



Translation to specific New Physics Models

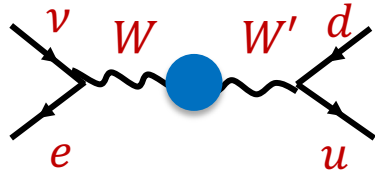


ϵ_L : measures deviations of the W boson to quarks and leptons, compared to the SM prediction



$$-\frac{g_{\nu e}^W g_{ud}^W}{4m_W^2} V_{ud} \bar{e} \gamma_\mu (1 - \gamma_5) \nu_e \cdot \bar{u} \gamma^\mu (1 - \gamma_5) d$$

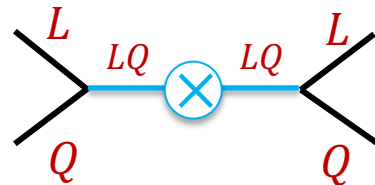
ϵ_R : left-right symmetric $SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_X$ models introduce new charged vector bosons W' coupling to right-handed quarks



$$\bar{e} \gamma_\mu (1 - \gamma_5) \nu_e \cdot \bar{u} \gamma^\mu (1 + \gamma_5) d$$

$$\epsilon_R \sim \frac{m_W^2}{m_{W'}^2}$$

$\epsilon_{S,P,T}$: In leptoquark models, new scalar particles couple to both quarks and leptons



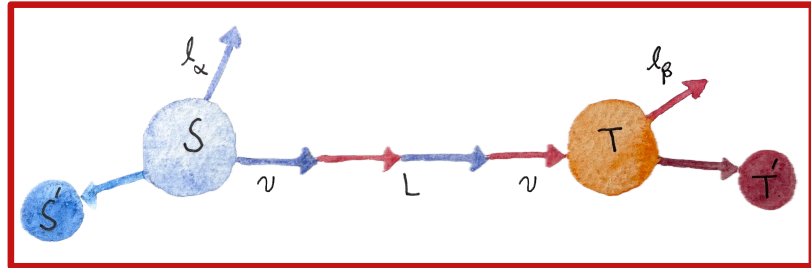
$$(LQ)(LQ)$$

$$\epsilon_{S,P,T} \sim \frac{v^2}{m_{LQ}^2}$$

EFT at neutrino experiments

We developed the systematic QFT approach to neutrino oscillations in the SMEFT framework!

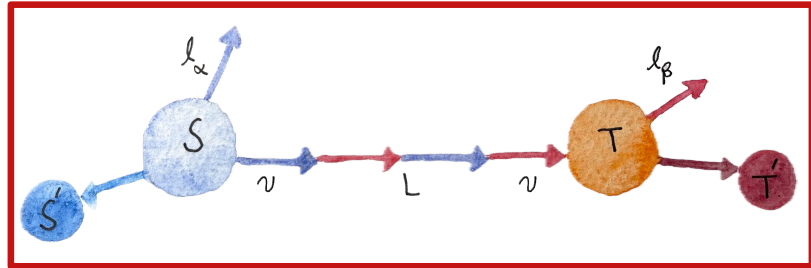
Falkowski, González-Alonso, ZT, JHEP (2020)



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Observable: rate of detected events

Rate \sim (flux) \times (det. cross section) \times (oscillation)

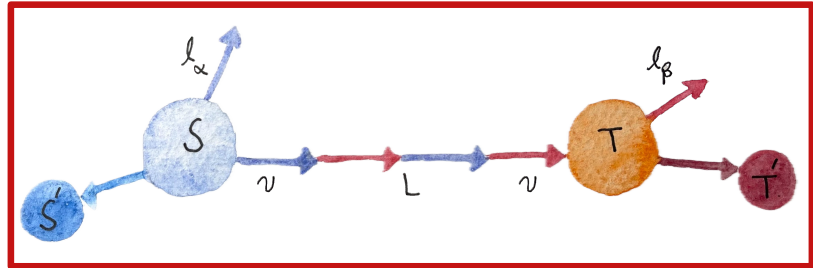
$$U_{\text{PMNS}} \parallel \begin{array}{c} \nu_e \\ \nu_\mu \\ \nu_\tau \end{array} \left[\begin{array}{ccc} \color{blue}{\blacksquare} & \color{red}{\blacksquare} & \color{red}{\blacksquare} \\ \color{red}{\blacksquare} & \color{red}{\blacksquare} & \color{purple}{\blacksquare} \\ \color{red}{\blacksquare} & \color{red}{\blacksquare} & \color{purple}{\blacksquare} \end{array} \right] \begin{array}{ccc} \nu_1 & \nu_2 & \nu_3 \end{array}$$

$$R_{\alpha\beta}^{\text{SM}} = \Phi_\alpha^{\text{SM}} \sigma_\beta^{\text{SM}} \sum_{k,l} e^{-i \frac{L \Delta m_{kl}^2}{2E_\nu}} U_{\alpha k}^* U_{\alpha l} U_{\beta k} U_{\beta l}^*$$

EFT at neutrino experiments

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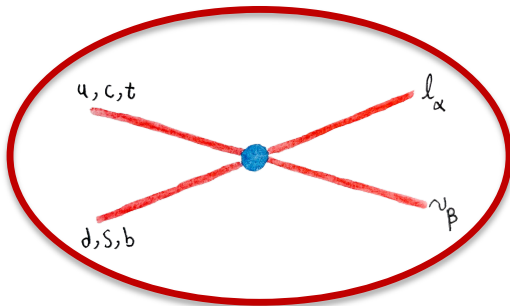
Falkowski, González-Alonso, ZT, JHEP (2020)



Observable: rate of detected events

Rate \sim (flux) \times (det. cross section) \times (oscillation)

We start from the CC WEFT Lagrangian



Production/Detection amplitudes

depend on the kinematic and spin variables

$$\mathcal{M}_{\alpha k}^P = U_{\alpha k}^* A_L^P + \sum_X [\epsilon_X U]_{\alpha k}^* A_X^P$$

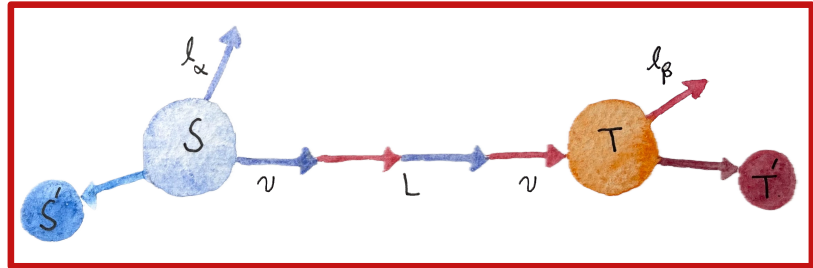
$$\mathcal{M}_{\beta k}^D = U_{\beta k} A_L^D + \sum_X [\epsilon_X U]_{\beta k} A_X^D$$

Corrections to fluxes/cross sections

EFT at neutrino experiments

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Falkowski, González-Alonso, ZT, JHEP (2020)

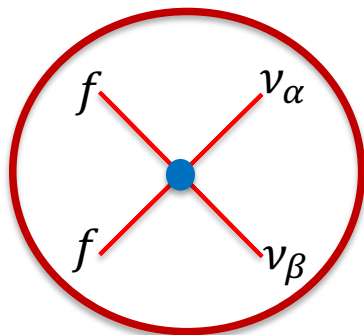


Observable: rate of detected events

Rate \sim (flux) \times (det. cross section) \times (oscillation)

NC WEFT Lagrangian

Modifies Matter Potential



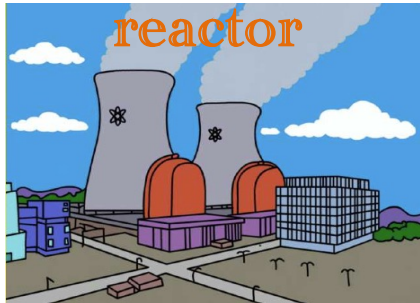
$$\mathcal{H}_F = \frac{1}{2E} (U M^2 U^\dagger + \mathbb{A}) \pm \sqrt{2} G_F N_e \begin{pmatrix} \varepsilon_{ee} & \varepsilon_{e\mu} & \varepsilon_{e\tau} \\ \varepsilon_{\mu e} & \varepsilon_{\mu\mu} & \varepsilon_{\mu\tau} \\ \varepsilon_{\tau e} & \varepsilon_{\tau\mu} & \varepsilon_{\tau\tau} \end{pmatrix}$$

Corrections to Oscillation Probability



solar

Different mechanism for neutrino production, detection and propagation



reactor



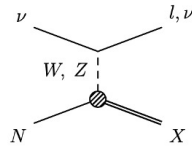
atmospheric



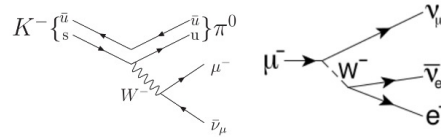
accelerator



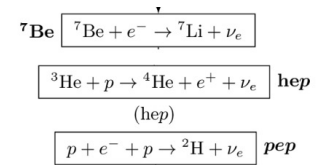
DIS: FASERν



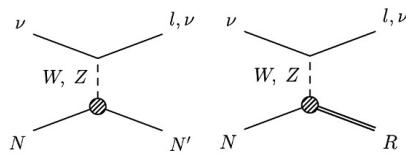
Kaon/Muon
decay:
ISODAR, KDAR



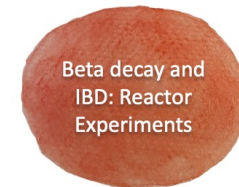
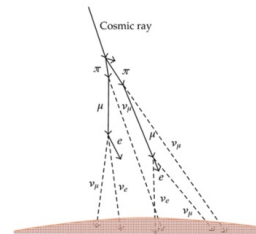
Solar
neutrinos:
Borexino



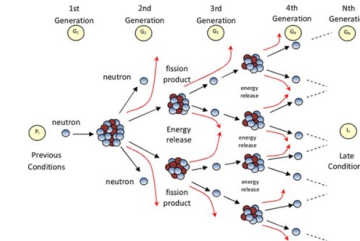
QE,
Resonances:
MINOS, NOvA,
DUNE



Atmospheric
Neutrinos:
IceCube



Beta decay and
IBD: Reactor
Experiments

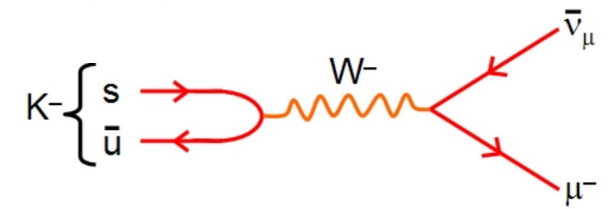
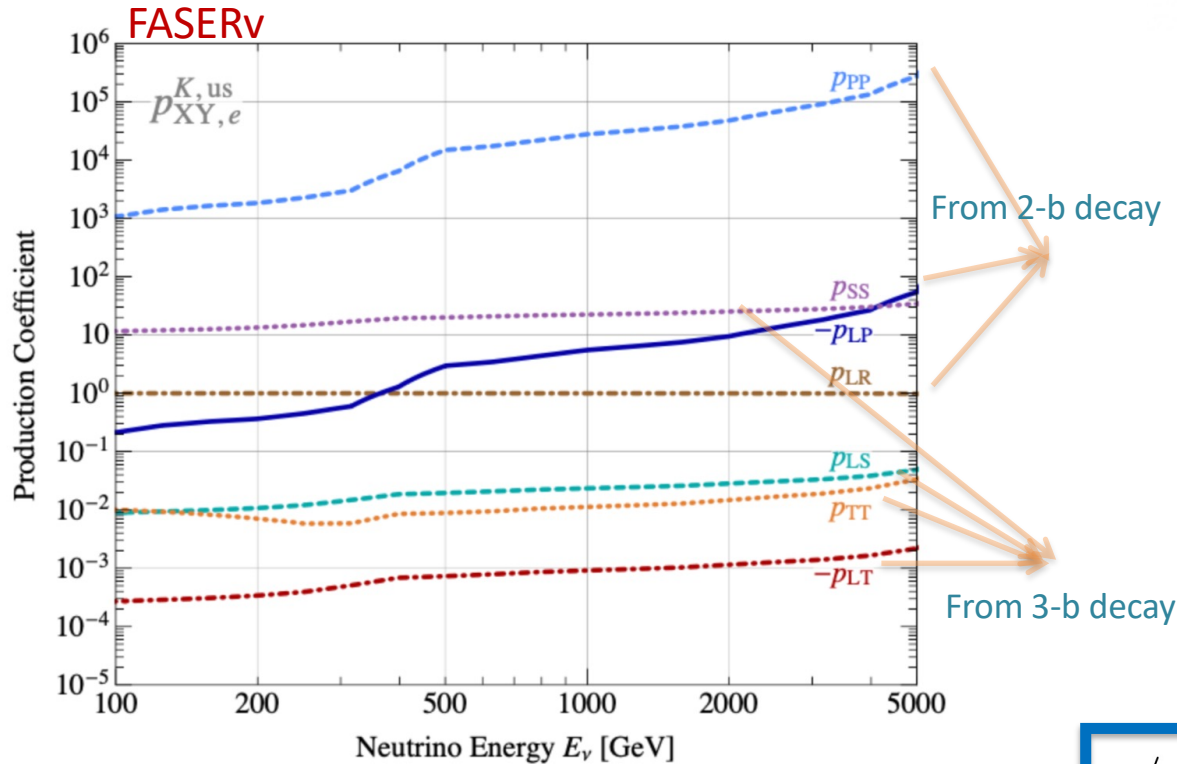


First step in building our toolbox:
Calculating fluxes, cross sections and oscillation probabilities for each experiment in the presence of EFT effects!

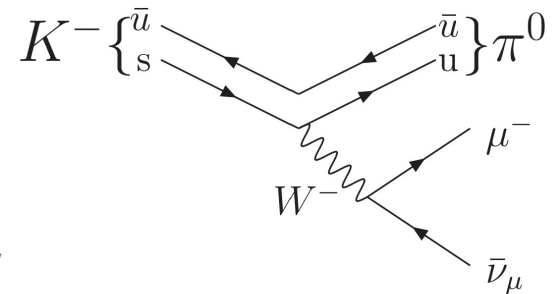
Corrections on Neutrino Fluxes from Kaon decay

Falkowski, González-Alonso, Kopp, Soreq, ZT, JHEP (2021)

Both 2-body and 3-body kaon decays contribute:



$$K^-(s\bar{u}) \rightarrow \mu^- + \bar{\nu}_\mu$$

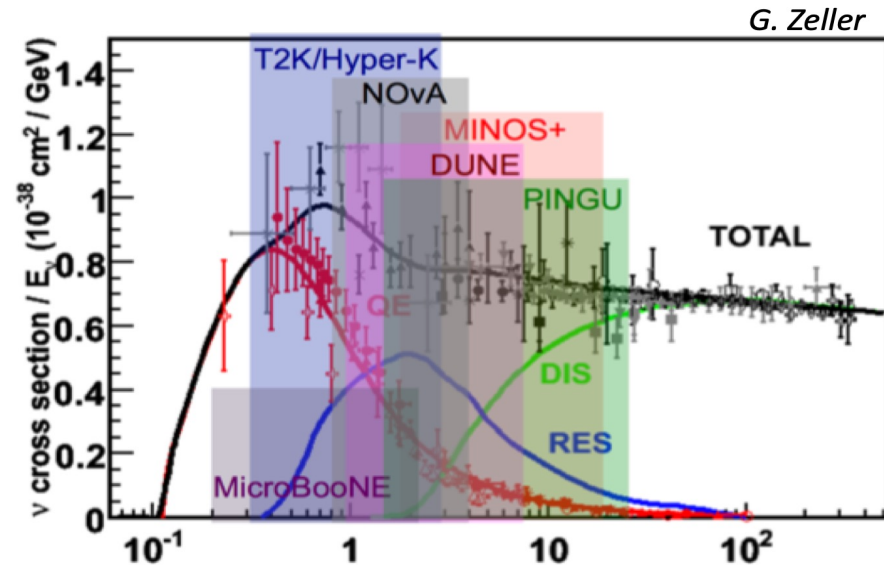


Huge flux correction for kaon decay! Depends on neutrino energy!

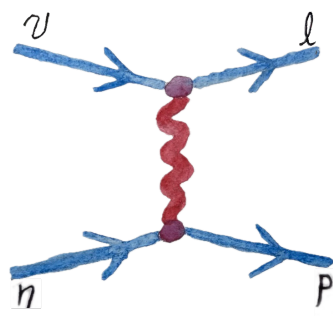
$$\begin{aligned} \langle \pi^- | \bar{s} \gamma^\mu u | K^0 \rangle &= P^\mu f_+(q^2) + q^\mu f_-(q^2), \\ \langle \pi^- | \bar{s} u | K^0 \rangle &= -\frac{m_K^2 - m_\pi^2}{m_s - m_u} f_0(q^2), \\ \langle \pi^- | \bar{s} \sigma^{\mu\nu} u | K^0 \rangle &= i \frac{p_K^\mu p_\pi^\nu - p_\pi^\mu p_K^\nu}{m_K} B_T(q^2), \end{aligned}$$

Cross Sections at Accelerator Experiments

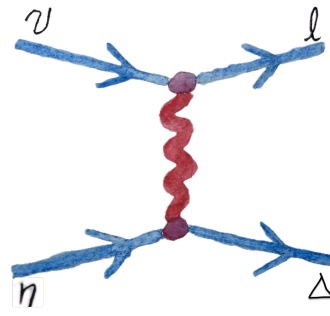
0.1-5 GeV: cross section is much more involved!



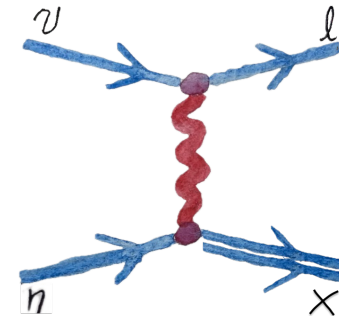
Quasi-Elastic Scattering



Resonance Production

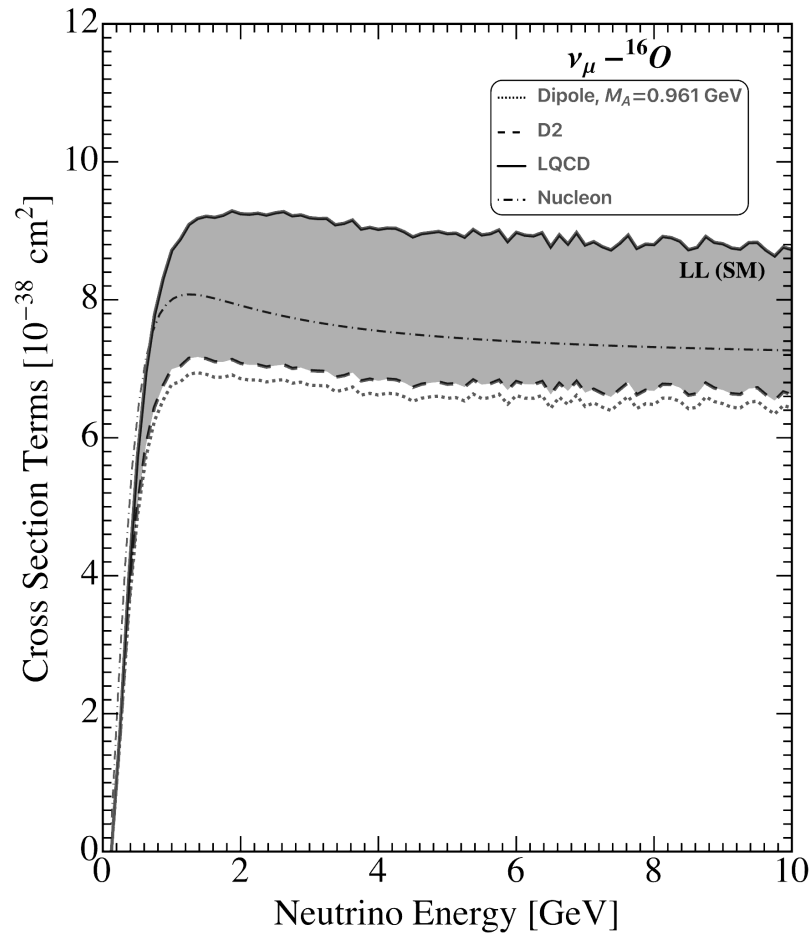


Deep Inelastic Scattering



CCQE Neutrino-Nucleus Cross Sections in the SM

Kopp, Rocco, ZT,
JHEP (2024)

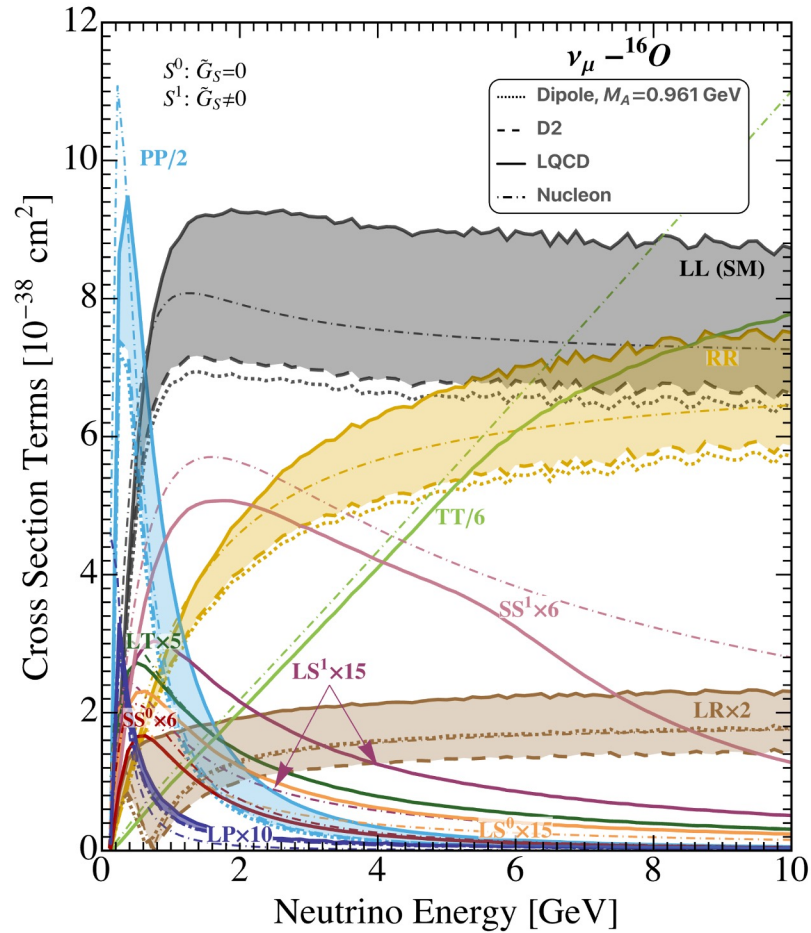


- z-expansion fit to LQCD and D2 data;
- Nuclear effects;
- Comparison with nucleon scattering

**Large uncertainties from
form factors even in the SM!**

CCQE Neutrino-Nucleus Cross Sections with EFT

Kopp, Rocco, ZT,
JHEP (2024)



- z-expansion fit to LQCD and D2 data;
- Nuclear effects;
- Comparison with nucleon scattering

Distinct Energy Distributions

Some rough estimations

- Observed rate at the experiment:

$$R_{Obs} = 10^4 \nu_\mu$$

$10^8 \nu_\mu$ at
DUNE Near Detector!

- Uncertainty:

$$\sqrt{R_{Obs}} = 10^2 \nu_\alpha \equiv \Delta R$$

- From theory:

$$R_{Th} = R_{SM}(1 + C \epsilon^2) = R_{SM} + \Delta R$$

- Limit on ϵ :

$$C \epsilon^2 = \frac{\Delta R}{R_{SM}} \quad \left\{ \quad \begin{array}{l} C = 10^3 \\ \epsilon < \frac{10^2}{10^3 \times 10^4} \sim 3 \times 10^{-3} \end{array} \right.$$

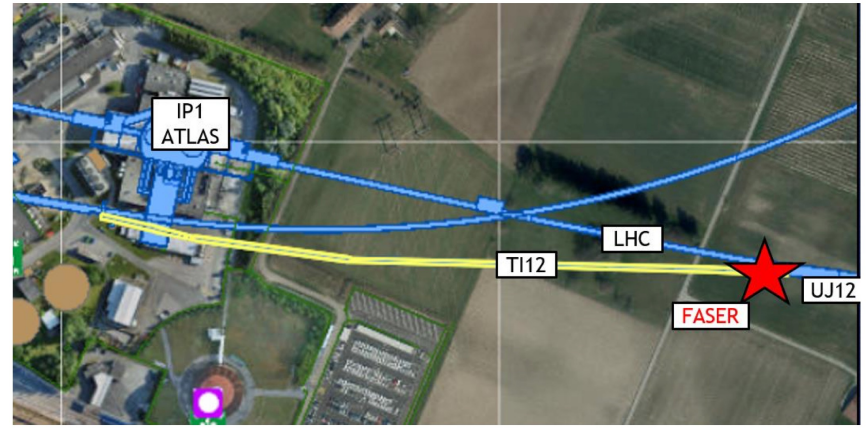
- New Physics Limit:

$$\Lambda \equiv \frac{v [246 \text{ GeV}]}{\sqrt{\epsilon}} = 4.5 \text{ TeV}$$

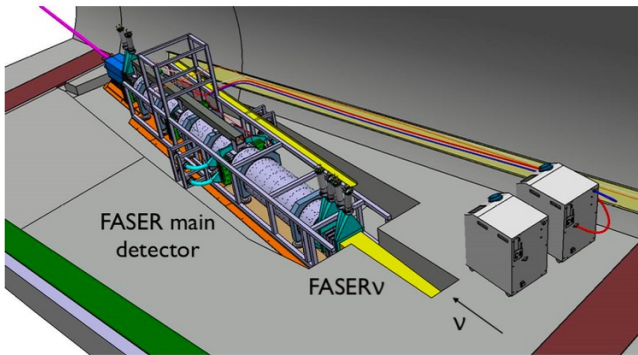
$$C \propto \frac{\sigma_{NP}}{\sigma_{SM}} \text{ or } \frac{\phi_{NP}}{\phi_{SM}}$$

FASER ν at LHC

- Downstream of ATLAS at of 480 m;
- Ideal for detecting high-energy neutrinos at LHC;
- 1.1-t of tungsten material;
- Several production modes;
- Pion and Kaon decays are the dominant ones;
- All (anti)neutrino flavors are available;

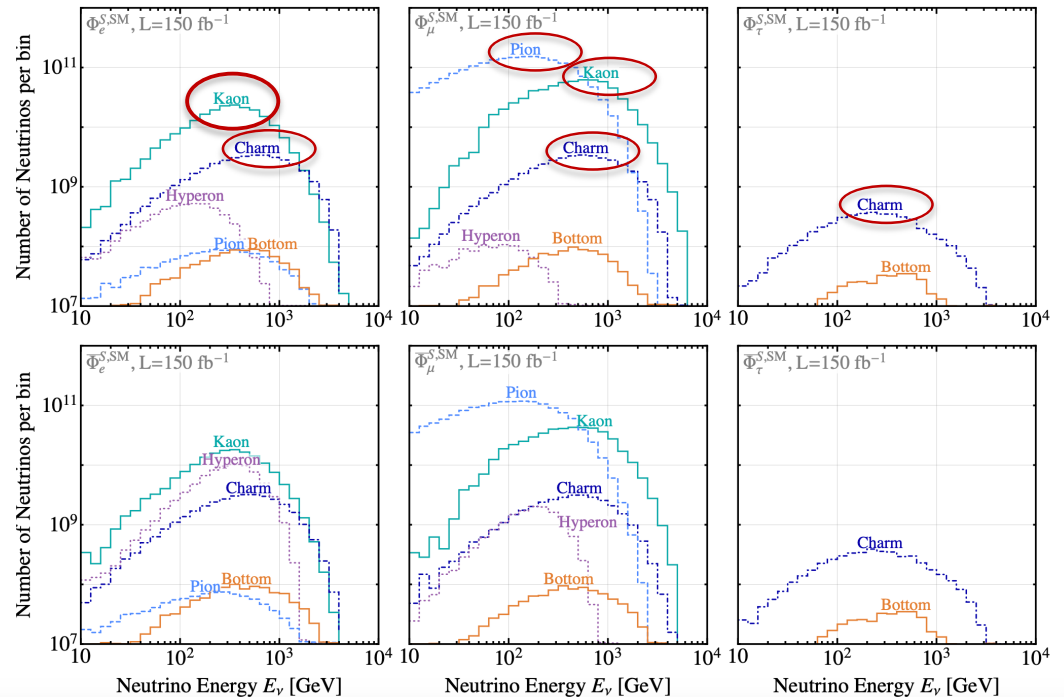


Falkowski, González-Alonso, Kopp, Soreq, ZT, JHEP (2021)



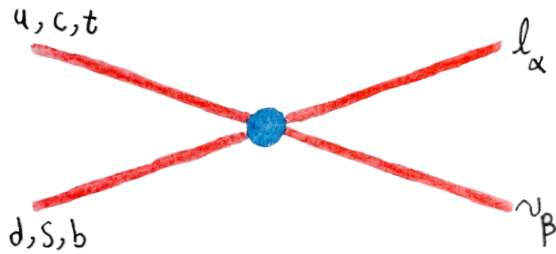
Within the SM:

$$\nu_e \sim 1000, \quad \nu_\mu \sim 5000, \quad \nu_\tau \sim 10$$



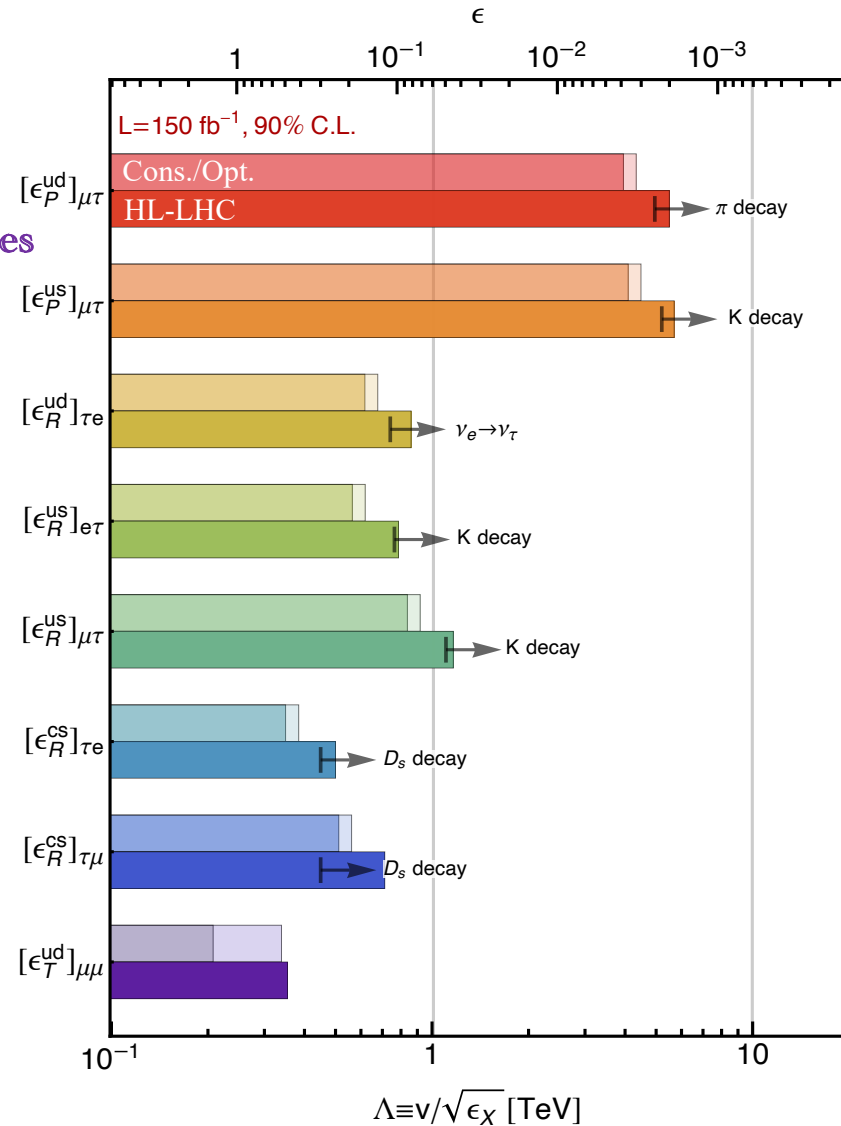
EFT at FASER ν

- FASER ν : colored bars
- Top: Conservative/Optimistic flux uncertainties
- Bottom: High luminosity LHC



- Neutrino detectors can identify flavor: 81 operators at FASER ν
- New physics reach at multi-TeV
- Complementary or dominant constraints

Falkowski, González-Alonso, Kopp, Soreq, ZT, JHEP (2021)



Reactor Experiments

Goal: measured θ_{13} mixing angle

Daya Bay:

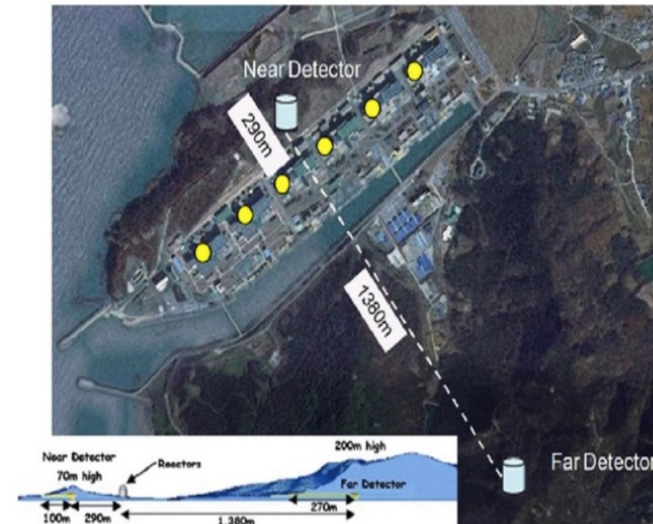
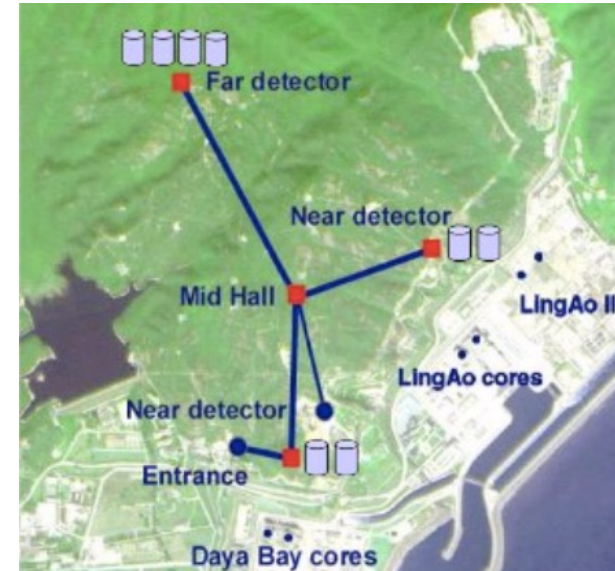
- 6 reactor cores;
- 8 anti-neutrino detectors;
- 3 near and far experimental halls located at 400 m, 512 m and 1610 m;
- Has observed ~ 4 million anti-neutrino events in 1958 days of data taking;

Daya Bay Collaboration, D. Adey et al.,
arXiv:1809.02261

RENO:

- 6 reactor cores;
- 2 near and far anti-neutrino detectors located at 367 m and 1440 m;
- Has observed ~ 1 million anti-neutrino events in 2200 days of data taking

RENO Collaboration, G. Bak et al.,
arXiv:1806.00248



EFT in Reactor Experiments

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L, E_\nu) = 1 - \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_\nu} \right) \sin^2 \left(2\tilde{\theta}_{13} - \alpha_D \frac{m_e}{E_\nu - \Delta} - \alpha_P \frac{m_e}{f_T(E_\nu)} \right) + \sin \left(\frac{\Delta m_{31}^2 L}{2E_\nu} \right) \sin(2\tilde{\theta}_{13}) \left(\beta_D \frac{m_e}{E_\nu - \Delta} - \beta_P \frac{m_e}{f_T(E_\nu)} \right) + \mathcal{O}(\epsilon_X^2)$$

— WEFT
— SM

A. Falkowski, M. González-Alonso, ZT
JHEP 05 (2019) 173

- Correlation between oscillation and WEFT parameters

$$\alpha_D = \frac{g_S}{3g_A^2+1} \text{Re}[S] - \frac{3g_A g_T}{3g_A^2+1} \text{Re}[T], \quad \alpha_P = \frac{g_T}{g_A} \text{Re}[T]$$

$$\beta_D = \frac{g_S}{3g_A^2+1} \text{Im}[S] - \frac{3g_A g_T}{3g_A^2+1} \text{Im}[T], \quad \beta_P = \frac{g_T}{g_A} \text{Im}[T]$$

$$[S] \equiv e^{i\delta_{\text{CP}}} (s_{23}[\epsilon_S]_{e\mu} + c_{23}[\epsilon_S]_{e\tau})$$

$$[T] \equiv e^{i\delta_{\text{CP}}} (s_{23}[\hat{\epsilon}_T]_{e\mu} + c_{23}[\hat{\epsilon}_T]_{e\tau})$$

EFT in Reactor Experiments

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L, E_\nu) = 1 - \sin^2\left(\frac{\Delta m_{31}^2 L}{4E_\nu}\right) \sin^2\left(2\tilde{\theta}_{13} - \alpha_D \frac{m_e}{E_\nu - \Delta} - \alpha_P \frac{m_e}{f_T(E_\nu)}\right) + \sin\left(\frac{\Delta m_{31}^2 L}{2E_\nu}\right) \sin(2\tilde{\theta}_{13}) \left(\beta_D \frac{m_e}{E_\nu - \Delta} - \beta_P \frac{m_e}{f_T(E_\nu)}\right) + \mathcal{O}(\epsilon_X^2)$$

— WEFT
— SM

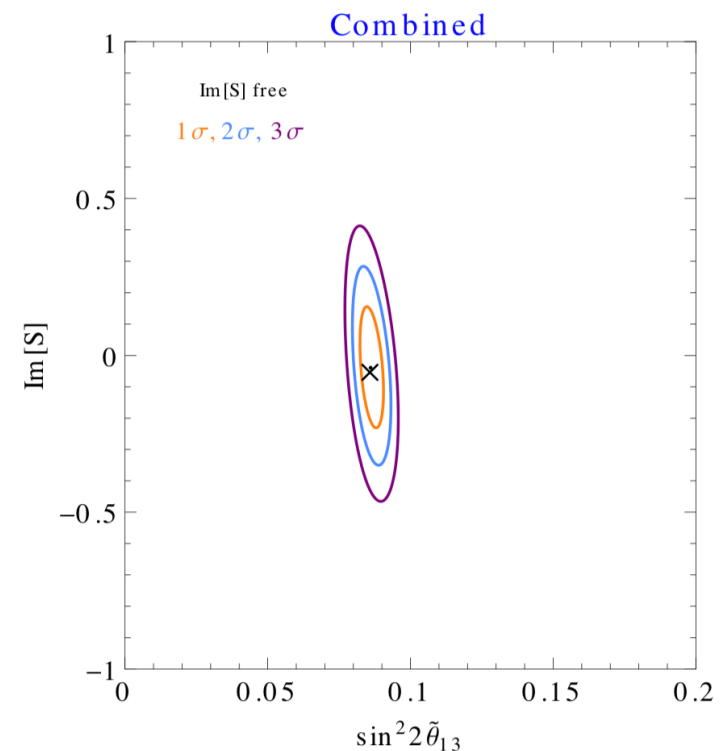
A. Falkowski, M. González-Alonso, ZT
JHEP 05 (2019) 173

- Correlation between oscillation and WEFT parameters

Scalar: $\epsilon_S \neq 0, \epsilon_{R,P,T} = 0$

$$\text{Im}[S] = 0.08 \pm 0.14$$

- Large degeneracy;
- Different measurements must be combined in a global analysis;



EFT in Reactor Experiments

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(L, E_\nu) = 1 - \sin^2\left(\frac{\Delta m_{31}^2 L}{4E_\nu}\right) \sin^2\left(2\tilde{\theta}_{13} - \alpha_D \frac{m_e}{E_\nu - \Delta} - \alpha_P \frac{m_e}{f_T(E_\nu)}\right) + \sin\left(\frac{\Delta m_{31}^2 L}{2E_\nu}\right) \sin(2\tilde{\theta}_{13}) \left(\beta_D \frac{m_e}{E_\nu - \Delta} - \beta_P \frac{m_e}{f_T(E_\nu)}\right) + \mathcal{O}(\epsilon_X^2)$$

— WEFT
— SM

A. Falkowski, M. González-Alonso, ZT
JHEP 05 (2019) 173

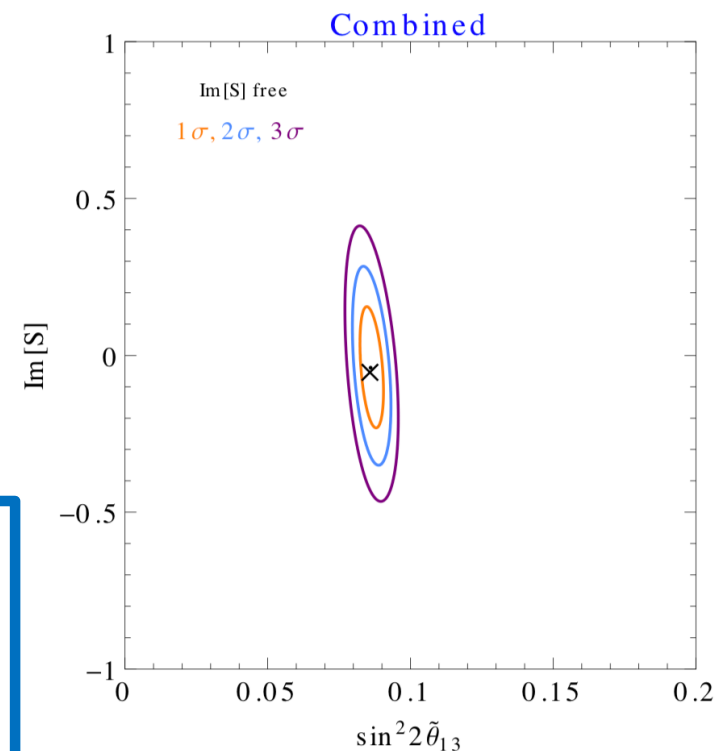
- Correlation between oscillation and WEFT parameters

Scalar: $\epsilon_S \neq 0, \epsilon_{R,P,T} = 0$

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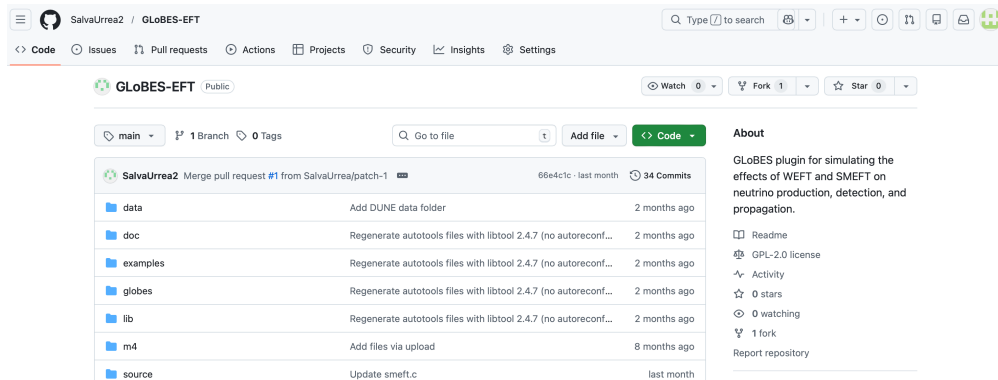
Second step in building our toolbox:
Developing packages to simulate neutrino oscillation experiments in the presence of EFT



GLOBES-EFT Software Package:

- To extract oscillation parameters in the presence of general EFT;
- To perform a SMEFT global fit in neutrino oscillation experiments;
- Publicly available on GitHub;

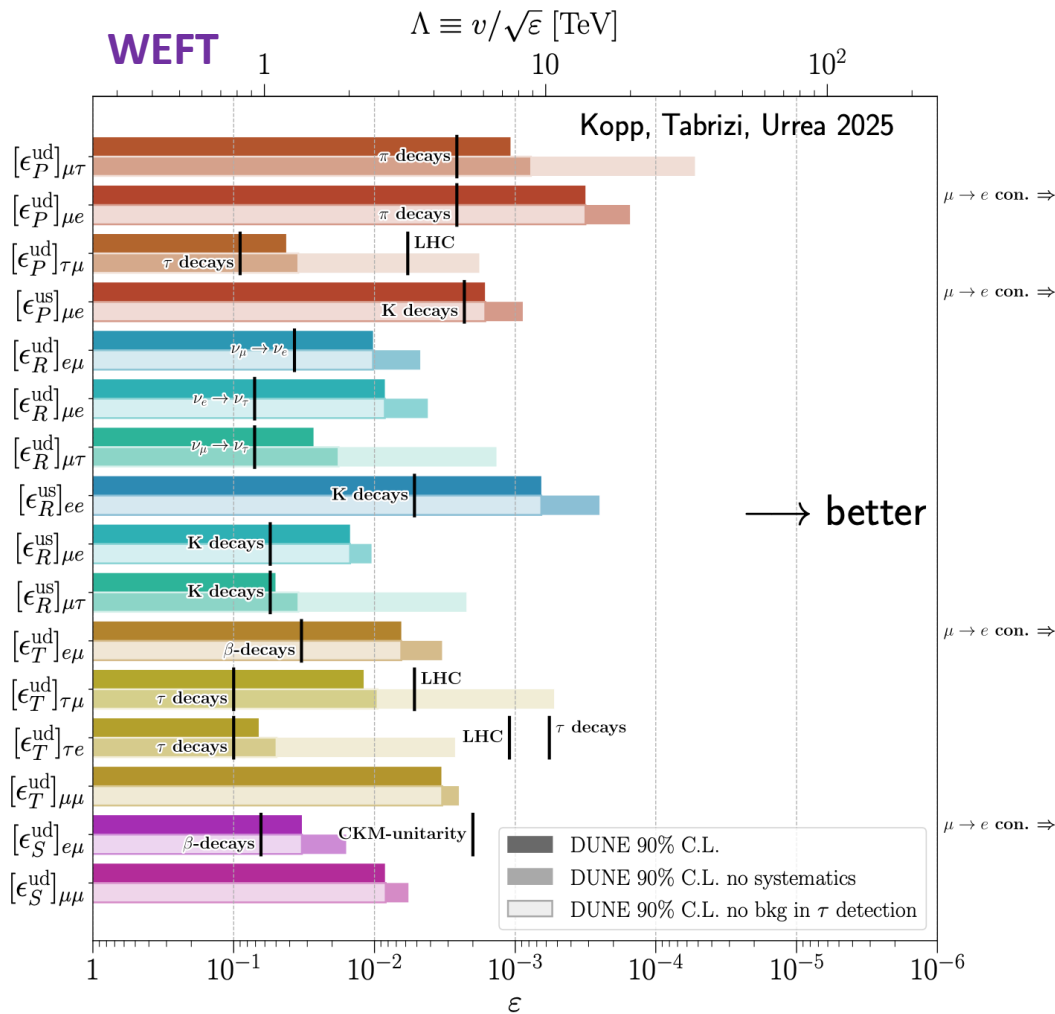
Kopp, ZT, Urrea
JHEP (2026)



- Neutrino Production, oscillation, and detection in presence of EFT;
- Accounting for RGE running;
- Consistent WEFT/SMEFT matching;

GLOBES-EFT: simulating neutrino oscillation experiments in the presence of EFT

EFT at DUNE using GLOBES-EFT

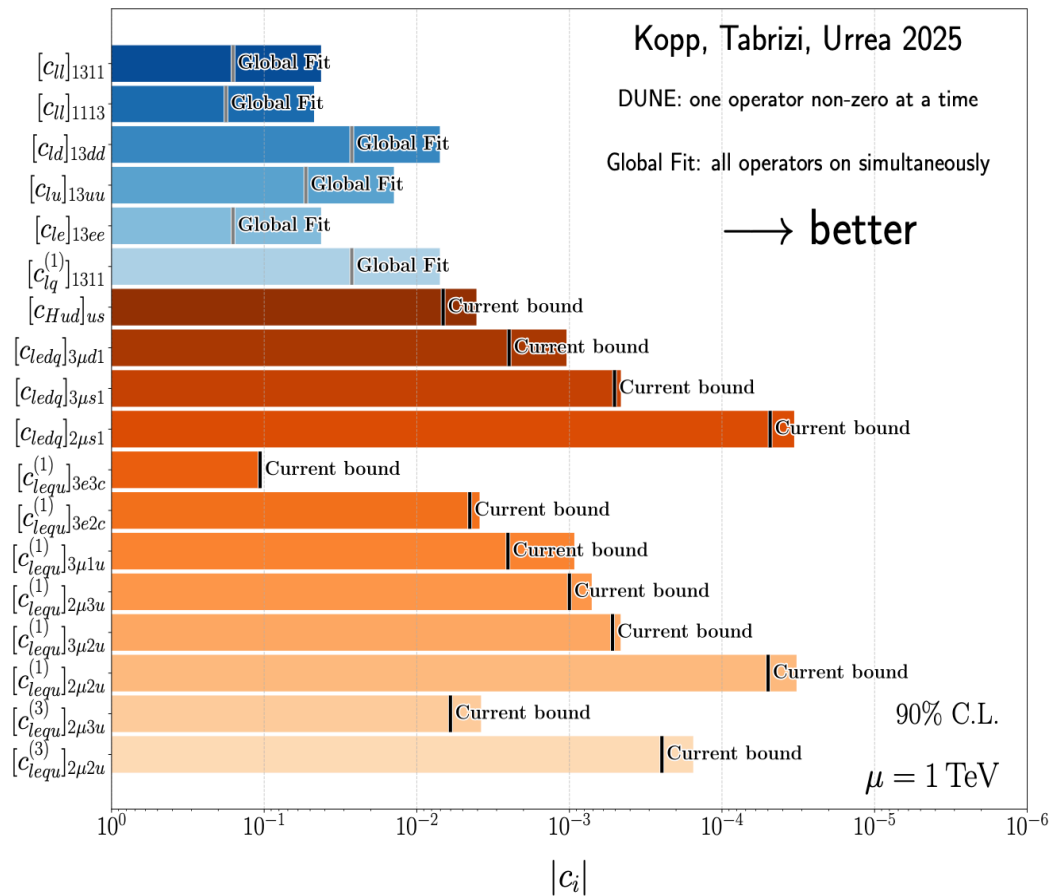


**Kopp, ZT, Urrea
JHEP (2026)**

Extracting ~10's of
TeV Physics from
GeV Neutrino
Experiments!

EFT at DUNE using GLOBES-EFT

SMEFT

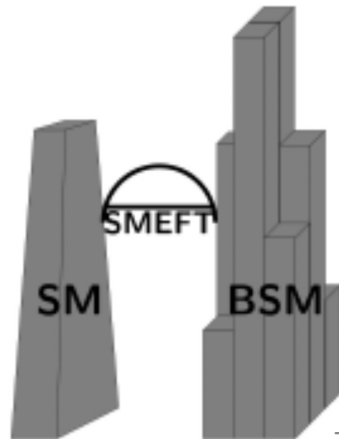


Kopp, ZT, Urrea
 JHEP (2026)

With GLOBES-EFT we can finally do a proper SMEFT global fit in ν experiments!

Indirect Searches: Future Directions

- SMEFT global fit in neutrino oscillation experiments using GLoBES-EFT;
- Extraction of oscillation parameters in presence of general new physics;
- Combining neutrino and collider data;
- Translation to specific new physics models;



- **“Non-Standard Neutrino Interactions at a Muon Collider Neutrino Detector”**,
Kopp, Kling, Ma, Mękała, Reuter, [ZT](#)
[JHEP \(2026\) \[To Appear\]](#)
- **“Effective Field Theory in Long-Baseline Neutrino Oscillation Experiments”**,
Kopp, [ZT](#), Urrea
[JHEP \(2026\)](#)
- **“Unleashing the power of EFT at Neutrino-Nucleus Scattering”**,
Kopp, Rocco, [ZT](#)
[JHEP \(2024\)](#)
- **“Consistent QFT description of non-standard neutrino interactions”**,
Falkowski, Gonzalez-Alonso, Soreq and [ZT](#)
[JHEP \(2020\)](#)
- **“EFT at FASERv”**,
Falkowski, Gonzalez-Alonso, Kopp, Soreq and [ZT](#)
[JHEP \(2021\)](#)
- **“Reactor neutrino oscillations as constraints on Effective Field Theory”**,
Falkowski, Gonzalez-Alonso, Soreq and [ZT](#)
[JHEP \(2019\)](#)
- **“Future DUNE constraints on EFT”**,
Falkowski, Gonzalez-Alonso, Soreq and [ZT](#)
[JHEP \(2018\)](#)

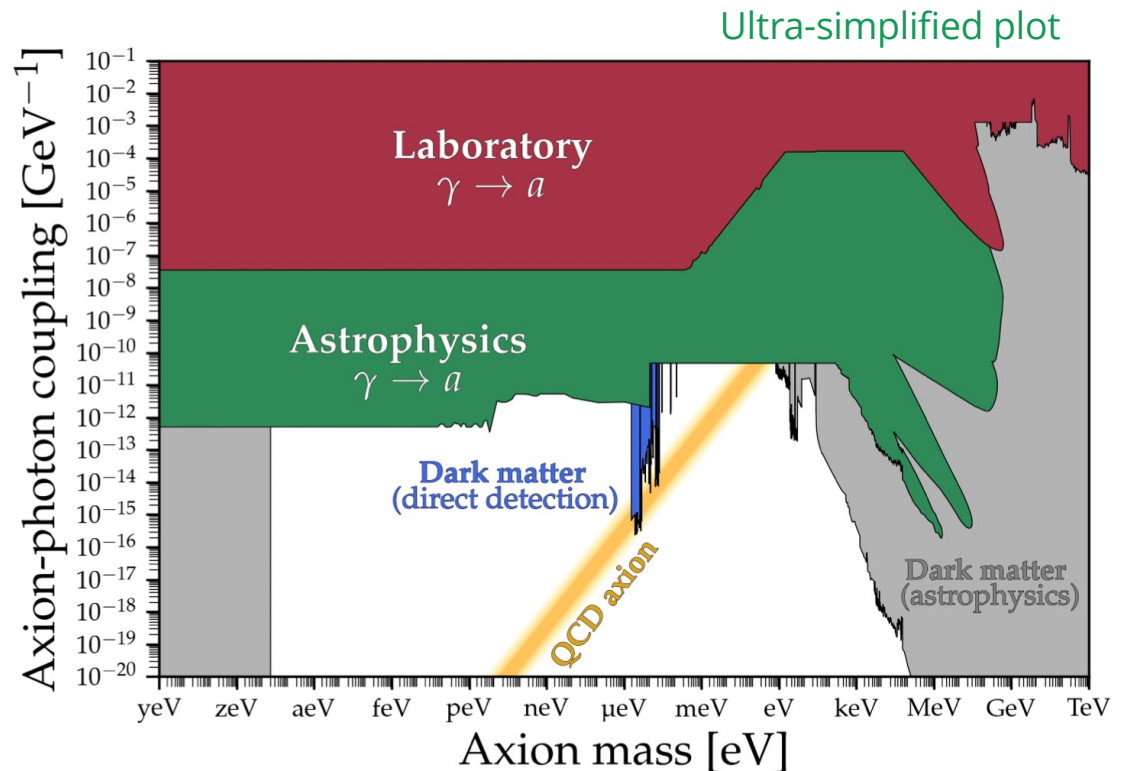
Outline

- 1) Indirect Search-EFT:
 - Why EFT?
 - EFT ladder
 - EFT at Neutrino Experiments
- 2) Direct Search:
 - Axion-Like Particles
- Conclusion



Axion-Like Particles (ALPs)

- (Pseudo)scalars, strongly motivated by theory and cosmology;
- It gives a solution to the strong CP problem: Why is CP conserved in QCD?
- They can explain for the 27% of the Universe's energy content constituting dark matter;

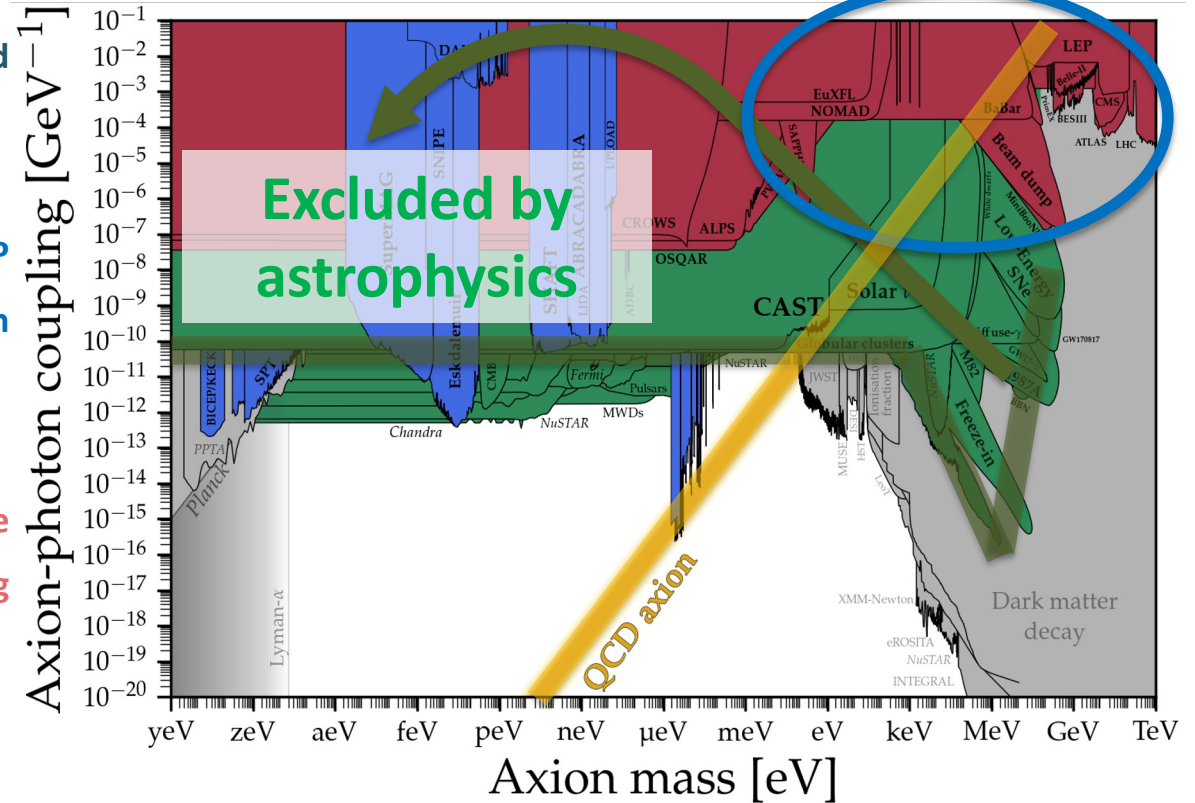


Ciaran O'Hare
<https://cajohare.github.io/AxionLimits>

Axion-Like Particles (ALPs)

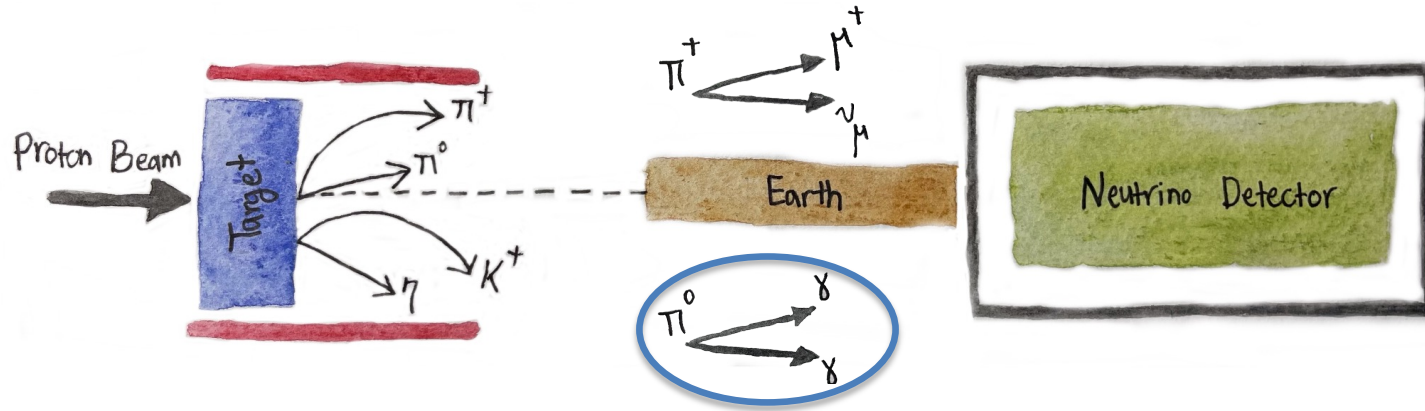
particle physics experiments

- (Pseudo)scalars, strongly motivated by theory and cosmology;
- It gives a solution to the strong CP problem: Why is CP conserved in QCD?
- They can explain for the 27% of the Universe's energy content constituting dark matter;



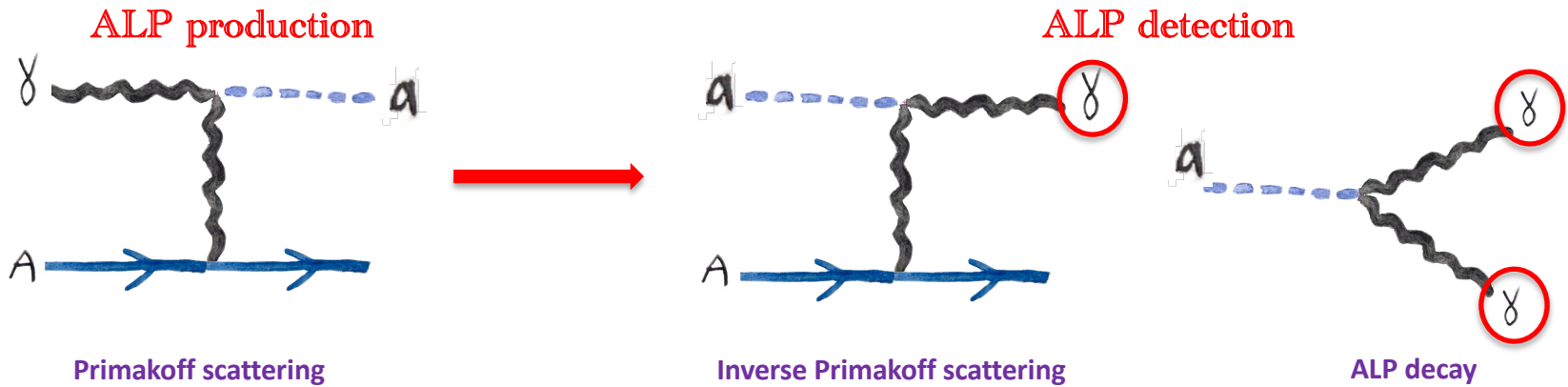
Ciaran O'Hare
<https://cajohare.github.io/AxionLimits>

ALP Production/Detection

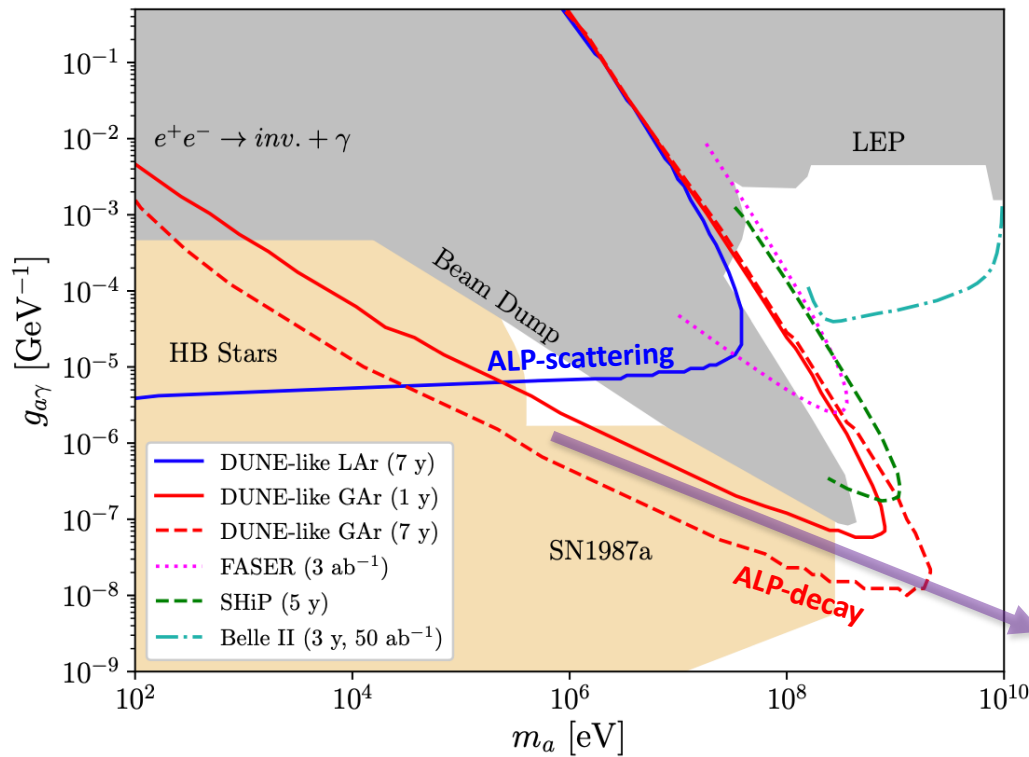


Using photons to produce ALPs:

$$\mathcal{L}_{a\gamma\gamma} \supset -\frac{1}{4} g_{a\gamma\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$



ALP- γ at DUNE



Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
PRL (2021)

Blue: ALP-scattering at LAr, 50-t

Red: ALP-decay at GAr, 1-t

- The only lab-based constraints!
- Gas-detector is the key, due to significantly low background!
- Search strategies applicable to other similar experiments.
Example: T2HK

ALP Backgrounds at DUNE

We investigate single photon and di-photon final states:

Initial Cuts:

| Signal Topology | ν_μ | ν_e | $\bar{\nu}_\mu$ | $\bar{\nu}_e$ | Total Counts |
|----------------------|--------------------|--------------------|--------------------|--------------------|--------------------|
| CC + NC ν events | 3.49×10^8 | 4.78×10^6 | 1.72×10^7 | 6.59×10^5 | 3.71×10^8 |
| 1γ | 34,469 | 441 | 3,094 | 101 | 38,104 |
| $1(e^-)^*$ | 3,839 | 6,516 | 417 | 46 | 10,818 |
| $1(e^+)^*$ | 0 | 0 | 0 | 1,451 | 1,451 |
| $1\gamma(1\gamma)_c$ | 7,030 | 78 | 537 | 17 | 7,663 |
| 2γ | 372,508 | 5,450 | 38,236 | 1,433 | 417,628 |
| $1(e^-)^*1\gamma$ | 123 | 76 | 9 | 0 | 208 |
| $1(e^+)^*1\gamma$ | 139 | 2 | 10 | 11 | 162 |

| Particle Type | Kinetic Energy Threshold | Angular Resolution |
|----------------|--------------------------|--------------------|
| μ^\pm | 30 MeV | 1° |
| π^\pm | 100 MeV | 1° |
| e^\pm/γ | 30 MeV | 1° |
| Protons | 50 MeV | 5° |
| Neutrons | 50 MeV | 5° |
| other | 50 MeV | 5° |

Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
PRD (2026)

- Many backgrounds;
- Large misID rates;
- What Kinematical cuts would decrease these?

ALP Backgrounds at DUNE

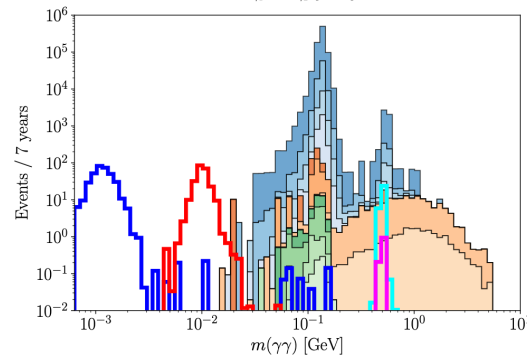
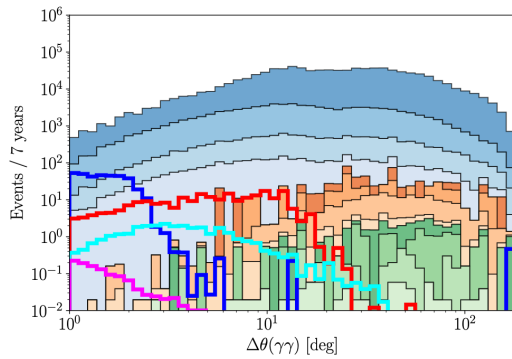
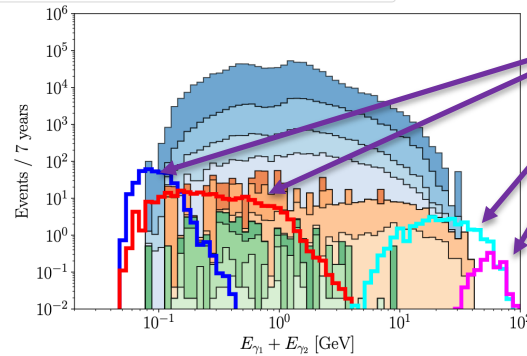
di-photon signal vs background

Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
PRD (2026)

$m_a = 1 \text{ MeV}, g_{a\gamma} = 1.75 \cdot 10^{-5} \text{ GeV}^{-1}$ $m_a = 500 \text{ MeV}, g_{a\gamma} = 2 \cdot 10^{-7} \text{ GeV}^{-1}$
 $m_a = 10 \text{ MeV}, g_{a\gamma} = 1.8 \cdot 10^{-6} \text{ GeV}^{-1}$ $m_a = 500 \text{ MeV}, g_{a\gamma} = 6 \cdot 10^{-7} \text{ GeV}^{-1}$

| | |
|--------------------------|-------------------------------|
| $\nu_\mu(2\gamma)$ | $\nu_e((e^-)*\gamma)$ |
| $\bar{\nu}_\mu(2\gamma)$ | $\bar{\nu}_e((e^+)*\gamma)$ |
| $\nu_e(2\gamma)$ | $\bar{\nu}_\mu((e^+)*\gamma)$ |
| $\bar{\nu}_e(2\gamma)$ | $\bar{\nu}_\mu((e^-)*\gamma)$ |
| $\nu_\mu((e^+)*\gamma)$ | $\nu_e((e^+)*\gamma)$ |
| $\nu_\mu((e^-)*\gamma)$ | $\bar{\nu}_e((e^-)*\gamma)$ |

Signal



These cut ensure more than 90% signal efficiency

Opening angle cut of $\Delta\theta(\gamma, \gamma) < 20^\circ$

Invariant mass cut of $|m(\gamma, \gamma) - m_a| < 0.05 \times m_a$

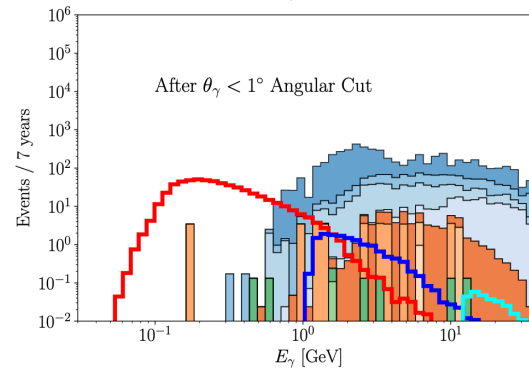
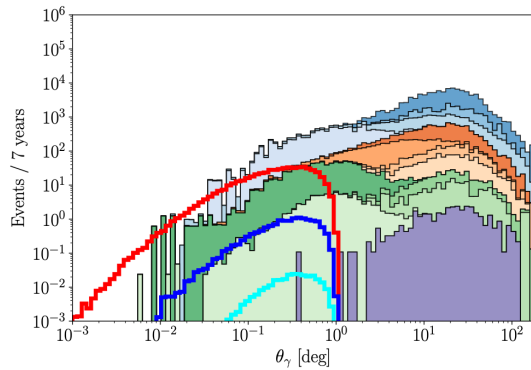
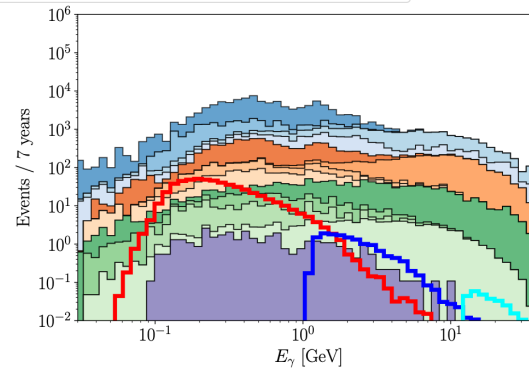
ALP Backgrounds at DUNE

Single photon signal vs background

Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
PRD (2026)

— $m_a = 1 \text{ MeV}, g_{a\gamma} = 1.75 \cdot 10^{-5} \text{ GeV}^{-1}$
— $m_a = 100 \text{ MeV}, g_{a\gamma} = 2 \cdot 10^{-7} \text{ GeV}^{-1}$
— $m_a = 10 \text{ MeV}, g_{a\gamma} = 1.8 \cdot 10^{-6} \text{ GeV}^{-1}$

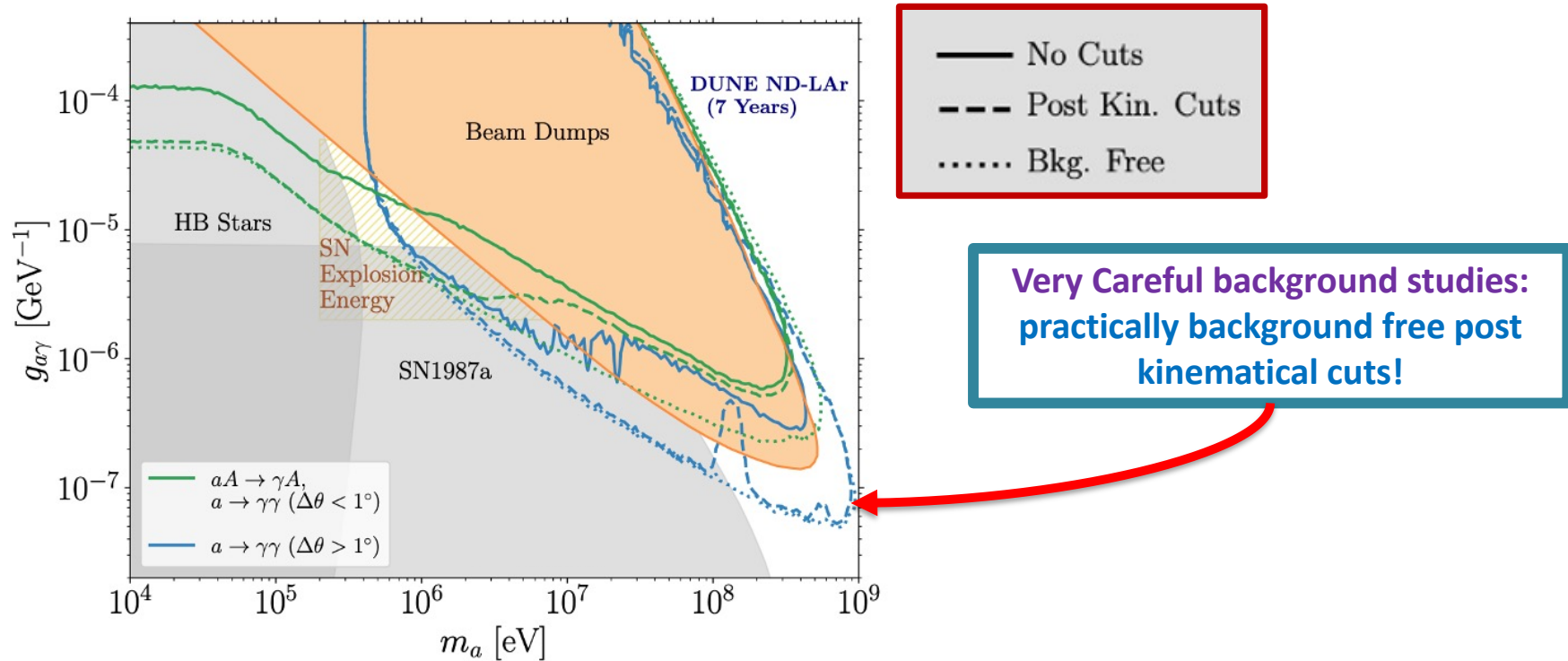
- | | |
|---|---|
| ■ $\nu_\mu(1\gamma)$ | ■ $\bar{\nu}_\mu(1e^-)$ |
| ■ $\nu_\mu(1\gamma(1\gamma)_c)$ | ■ $\bar{\nu}_e(1\gamma)$ |
| ■ $\nu_e(1e^-)$ | ■ $\nu_e(1\gamma(1\gamma)_c)$ |
| ■ $\nu_\mu(1e^-)$ | ■ $\bar{\nu}_e(1e^-)$ |
| ■ $\bar{\nu}_\mu(1\gamma)$ | ■ $\bar{\nu}_e(1\gamma(1\gamma)_c)$ |
| ■ $\bar{\nu}_e(1e^+)$ | ■ $\bar{\nu}_\mu(1e^+)$ |
| ■ $\bar{\nu}_\mu(1\gamma(1\gamma)_c)$ | ■ $\nu_\mu(1e^+)$ |
| ■ $\nu_e(1\gamma)$ | ■ $\nu_e(1e^+)$ |



A cut on the photon angle keeps > 99% of the signal, while decreasing the background from $\mathcal{O}(10^6)$ to $\mathcal{O}(10^3)$

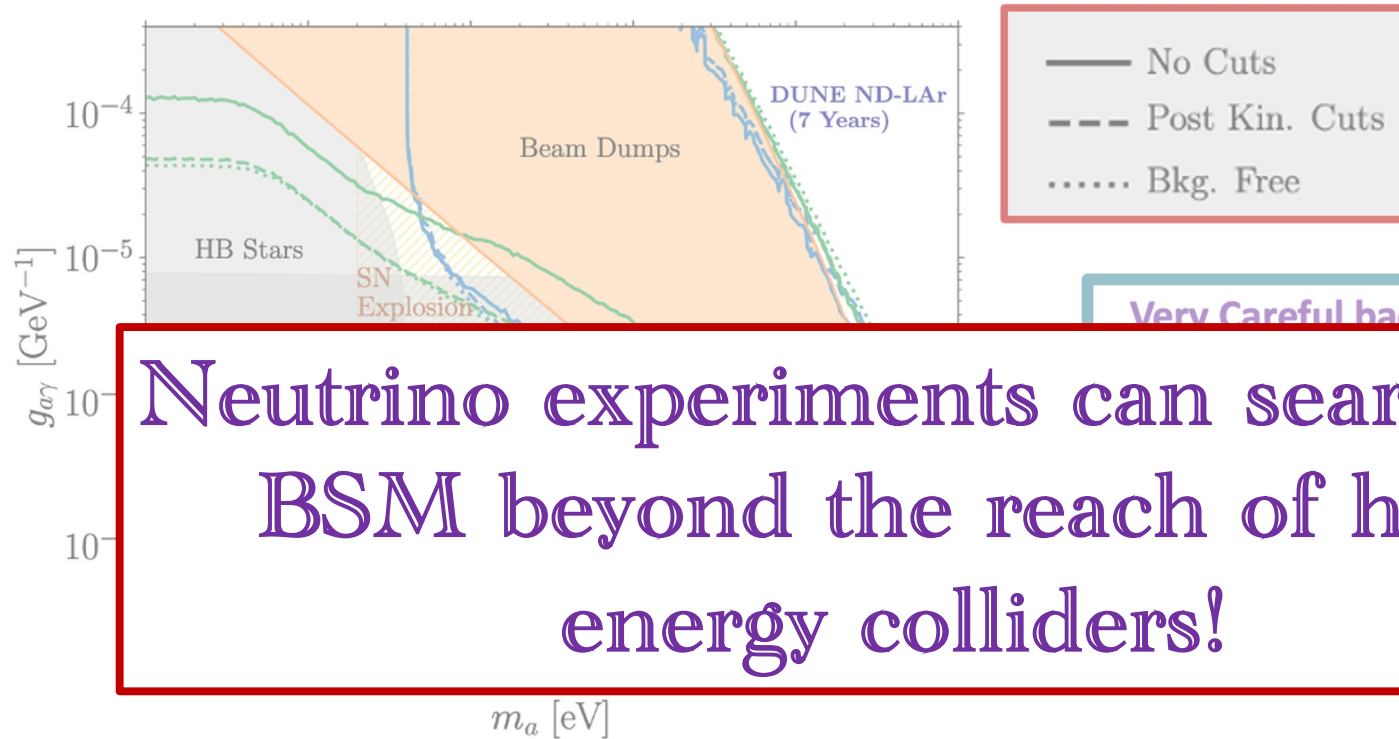
Cut on photon's angle wrt beam direction:
 $\theta_\gamma < 1^\circ$

ALP- γ at DUNE post background study



Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
 PRD (2026)

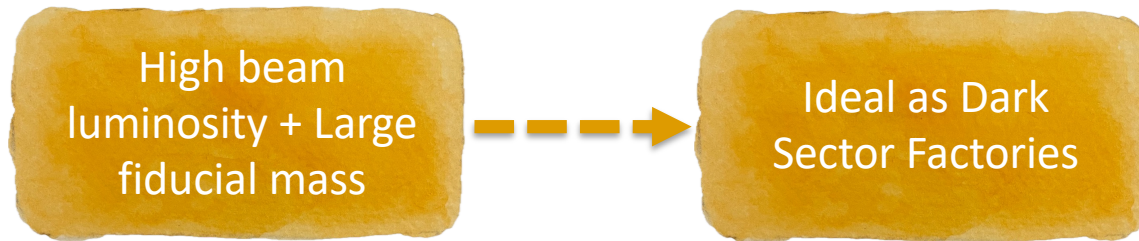
ALP- γ at DUNE post background study



Neutrino experiments can search for BSM beyond the reach of high energy colliders!

Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
PRD (2026)

Dark Sectors: Future Directions

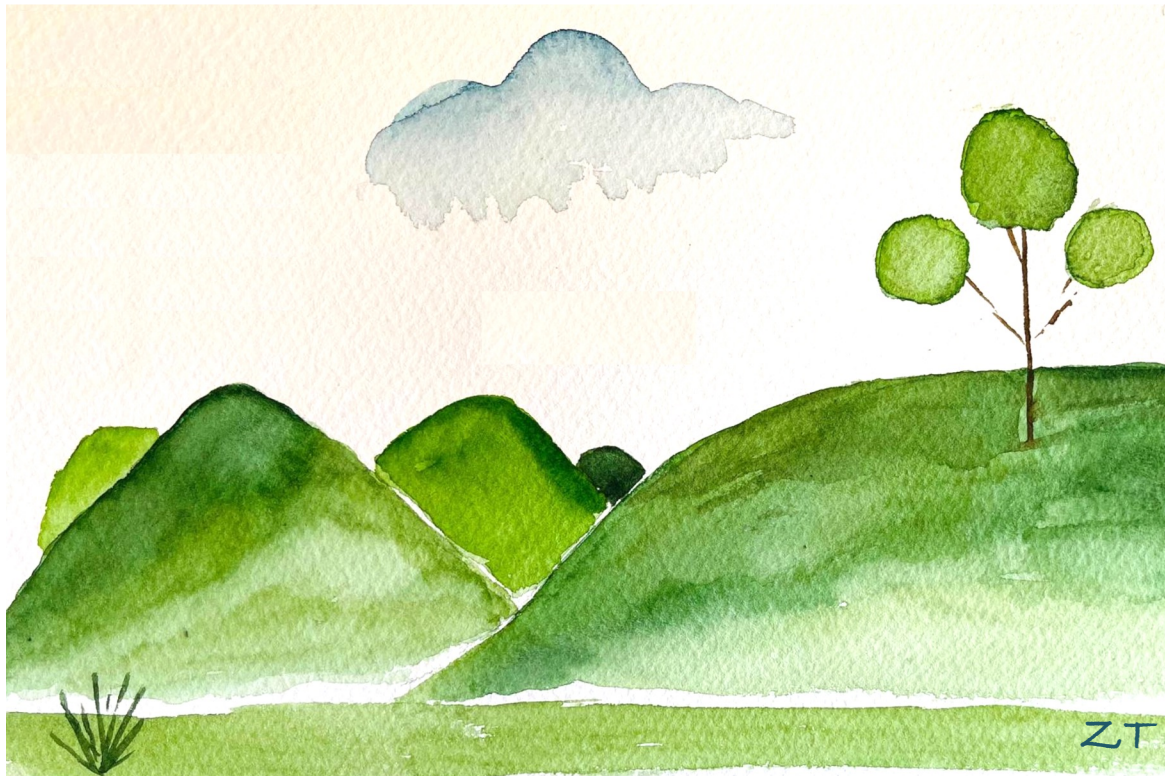


Leverage the high statistics to enlarge the physics programs of Hyper-K, DUNE, SBN, ...



- **“Prying Open the Dark Sector Window with SBND Off-Target Mode”**, Dutta, Goswami, Karthikeyan, Pandey, [ZT](#), Thompson, Van de Water [arXiv: 2603.XXXXX](#)
- **“Finding BSM Needles in Electromagnetic Haystacks at DUNE”**, Brdar, Dutta, Jang, Kim, Shoemaker, [ZT](#), Thompson and Yu [PRD \(2026\)](#)
- **“Heavy Neutral Leptons via Axion-Like Particles at Neutrino Facilities”**, Abdullahi, de Gouvea, Dutta, Shoemaker and [ZT](#), [PRL \(2024\)](#)
- **“Probing new physics at DUNE operating in a beam-dump mode”**, Brdar, Dutta, Jang, Kim, Shoemaker, [ZT](#), Thompson and Yu, [PRD \(2023\)](#)
- **“Axion-like Particles at Future Neutrino Experiments: Closing the Cosmological Triangle”**, Brdar, Dutta, Jang, Kim, Shoemaker, [ZT](#), Thompson and Yu, [PRL \(2021\)](#)
- **“Z’s in neutrino scattering at DUNE”**, Ballett, Hostert, Pascoli, Perez-Gonzalez, [ZT](#) and Funchal, [PRD \(2019\)](#)

Fantastic Beasts and How to Find Them With Neutrino Experiments

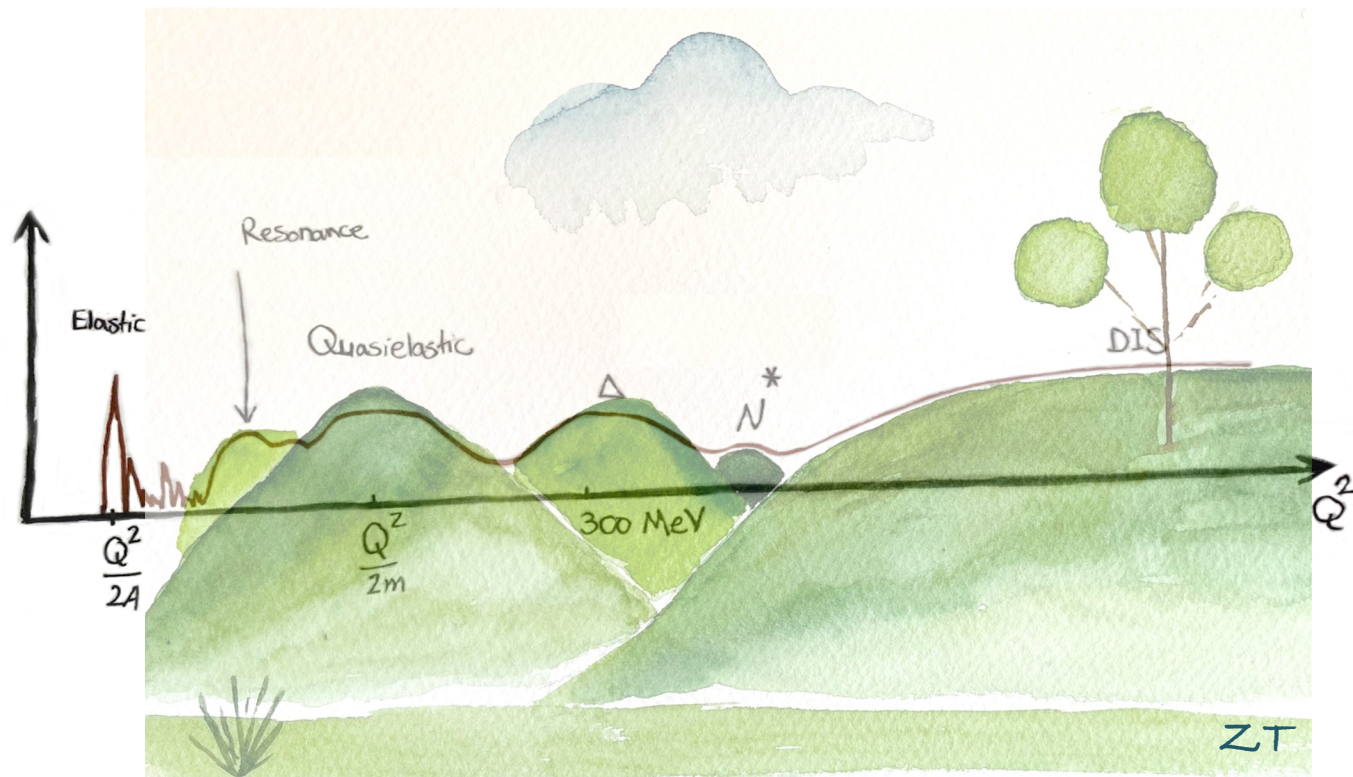


3/26/2026

Zahra Tabrizi, PITT-PACC Langley Fellow

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Fantastic Beasts and How to Find Them With Neutrino Experiments

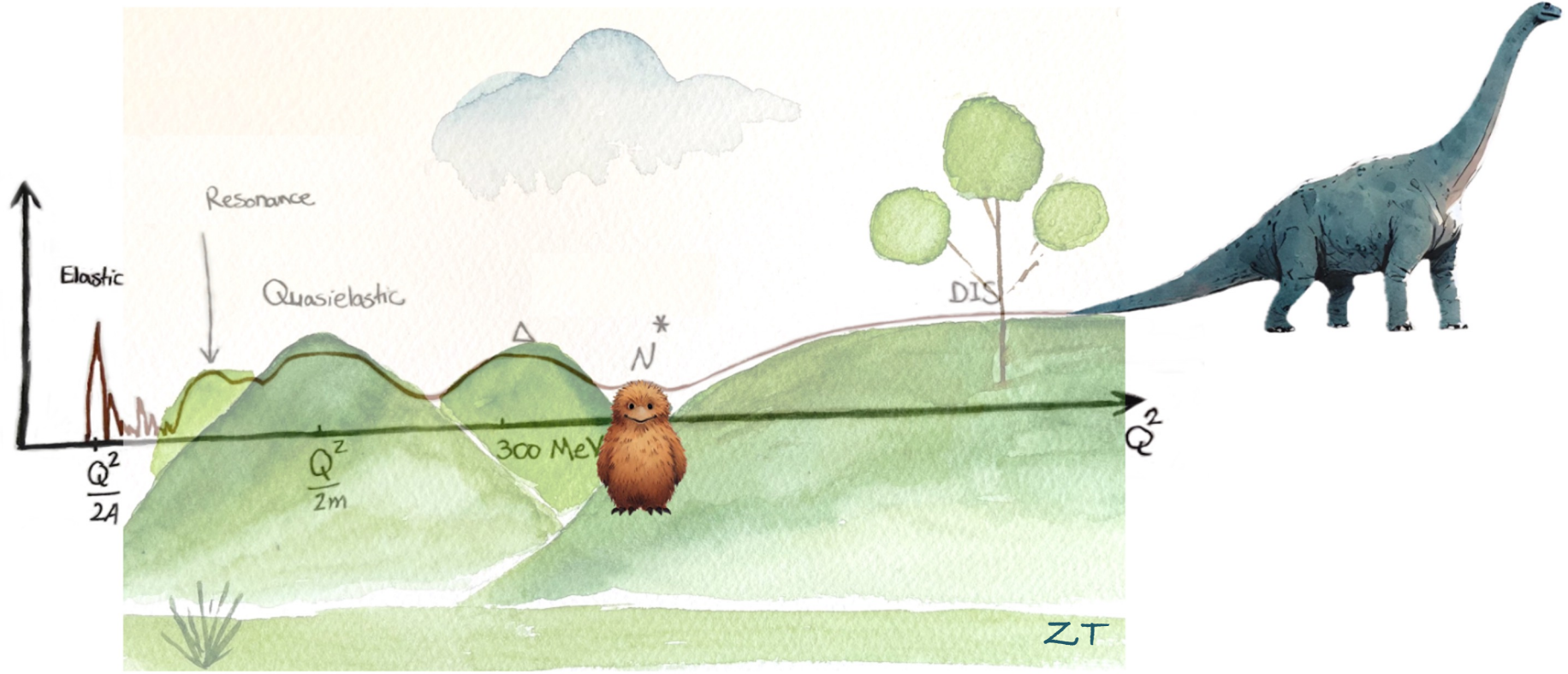


3/26/2026

Zahra Tabrizi, PITT-PACC Langley Fellow

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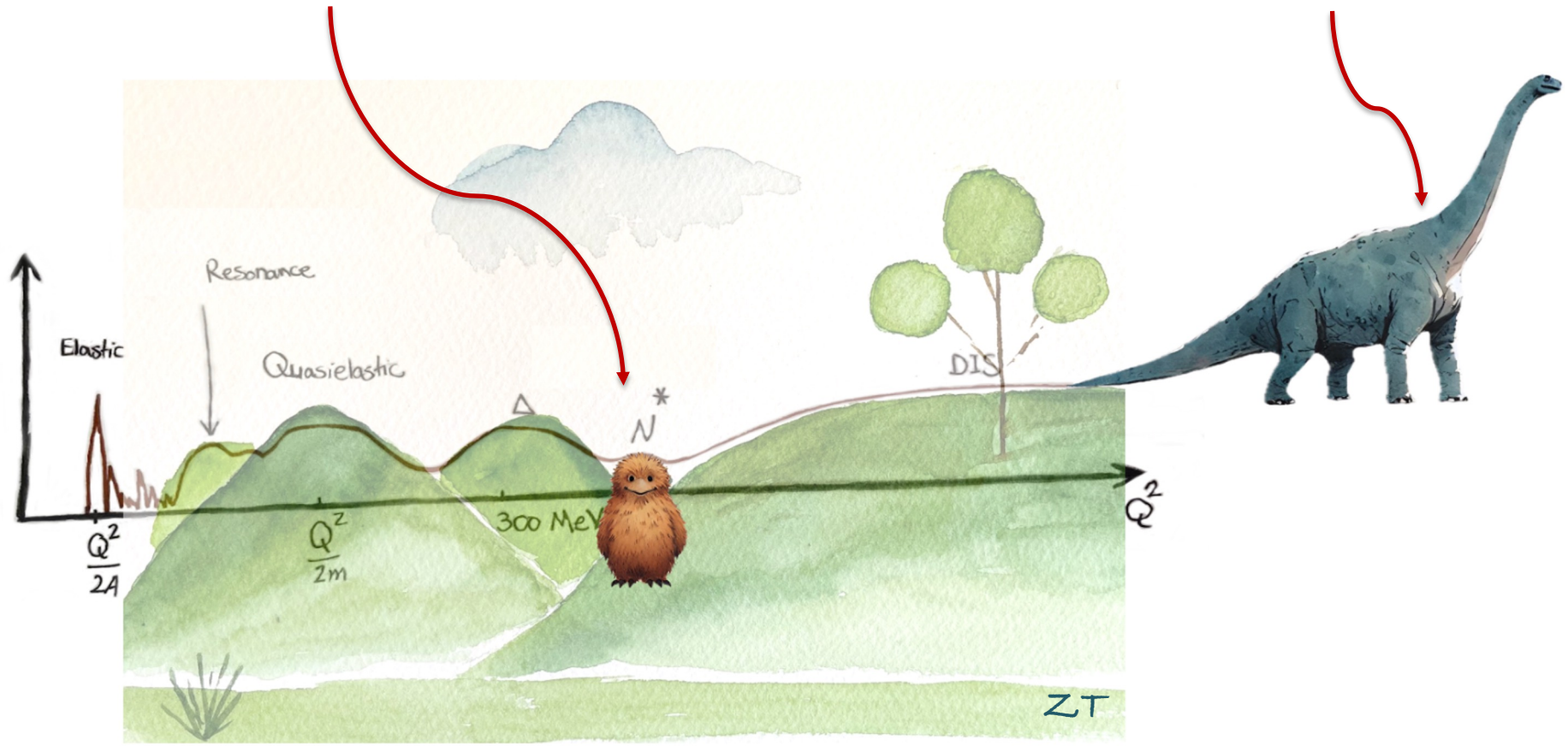
Fantastic Beasts and How to Find Them With Neutrino Experiments



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Neutrino Experiments as Dark Sector Factories
“New Particles”

EFT at Neutrino Experiments:
Indirect Effects of Heavy New Physics
“New Interactions”



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Ultimate goal: unifying
Neutrino-Collider
BSM searches



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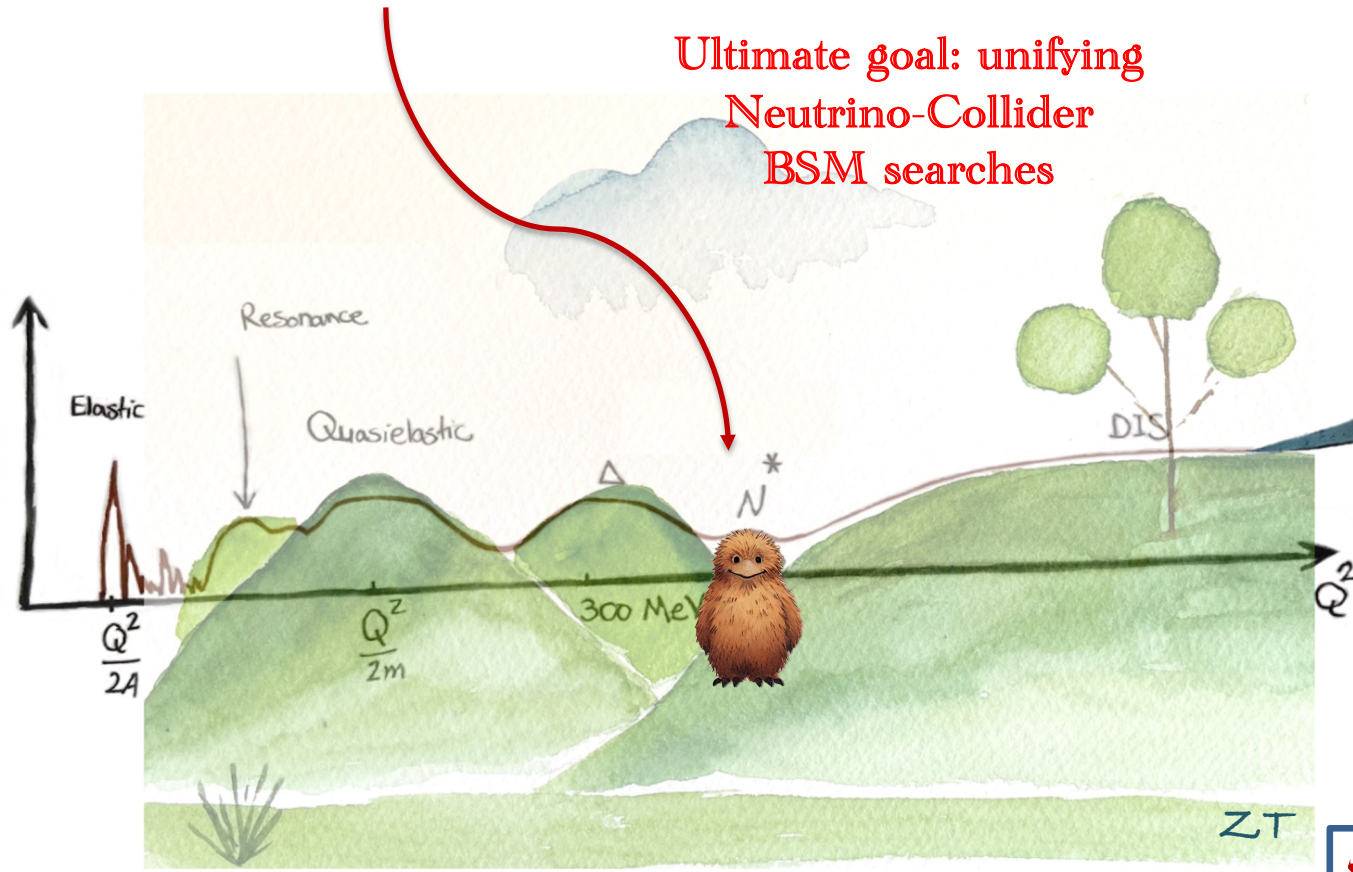
We can explore most of this landscape in the next few decades

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We can explore most of this landscape in the next few decades

Thank You!

Back up Slides

EFT at neutrino experiments

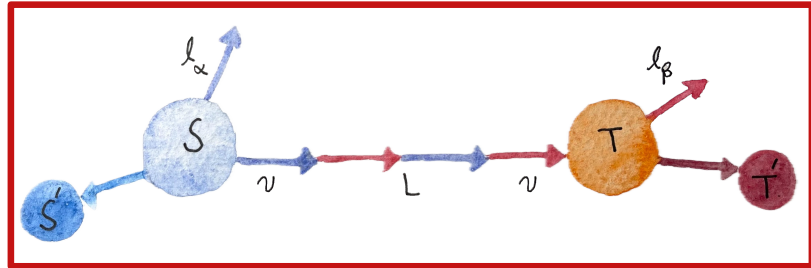
We developed the systematic QFT approach to neutrino oscillations in the SMEFT framework!

Falkowski, González-Alonso, ZT, JHEP (2020)

But what is new here?!

Traditional NSI approach: Neutrinos are not pure flavor states!

- Cannot be matched with SMEFT;
- No comparison of results with other experiments;
- Does not capture different kinematic dependence due to new interactions;
- Misses chiral enhancements/suppressions;



Standard NSI approach

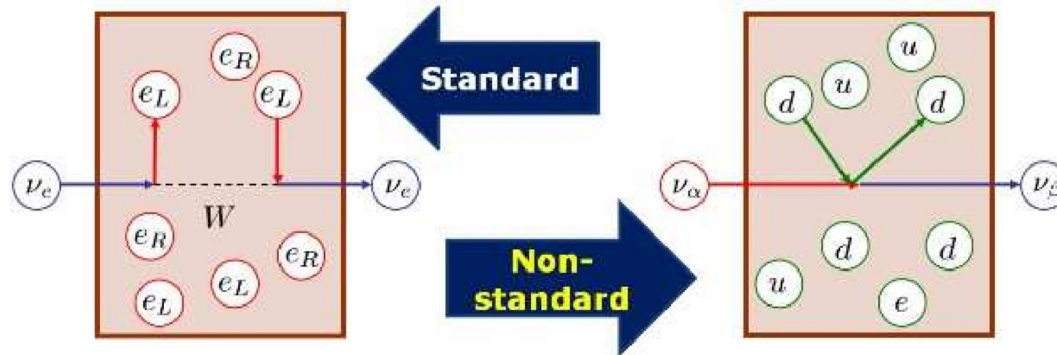
$$\begin{aligned}
 |\nu_\alpha^s\rangle &= \frac{1}{N_\alpha^s} \left[|\nu_\alpha\rangle + \sum_{\gamma=e,\mu,\tau} \epsilon_{\alpha\gamma}^s |\nu_\gamma\rangle \right] \\
 \langle\nu_\beta^d| &= \frac{1}{N_\beta^d} \left[\langle\nu_\beta| + \sum_{\gamma=e,\mu,\tau} \langle\nu_\gamma| \epsilon_{\gamma\beta}^d \right]
 \end{aligned}$$

Normalization
NSI parameters

Take home: Traditional NSI methods do not work. We must use the correct QFT approach for EFT studies!

QM-NSI Description

Neutrinos are not pure flavor states:



Standard NSI approach

NSI parameters

$$|\nu_\alpha^s\rangle = \frac{1}{N_\alpha^s} \left[|\nu_\alpha\rangle + \sum_{\gamma=e,\mu,\tau} \epsilon_{\alpha\gamma}^s |\nu_\gamma\rangle \right]$$

$$\langle \nu_\beta^d | = \frac{1}{N_\beta^d} \left[\langle \nu_\beta | + \sum_{\gamma=e,\mu,\tau} \langle \nu_\gamma | \epsilon_{\gamma\beta}^d \right]$$

Rotation of flavor states at the source

Rotation of flavor states at the detector

Normalization

QM-NSI Description

Neutrinos are not pure flavor states:

$$|\nu_\alpha^s\rangle = \frac{(1 + \epsilon^s)_{\alpha\gamma}}{N_\alpha^s} |\nu_\gamma\rangle, \quad \langle\nu_\beta^d| = \langle\nu_\gamma| \frac{(1 + \epsilon^d)_{\gamma\beta}}{N_\beta^d}$$

Observable: rate of detected events

\sim (flux) \times (det. cross section) \times (oscillation)

$$R_{\alpha\beta}^{\text{QM}} = \Phi_\alpha^{\text{SM}} \sigma_\beta^{\text{SM}} \sum_{k,l} e^{-i \frac{L \Delta m_{kl}^2}{2E\nu}} [x_s]_{\alpha k} [x_s]_{\alpha l}^* [x_d]_{\beta k} [x_d]_{\beta l}^*$$

$$x_s \equiv (1 + \epsilon^s)U^* \quad \& \quad x_d \equiv (1 + \epsilon^d)^T U$$

Falkowski, González-Alonso, [ZT, JHEP \(2019\)](#)

QM-NSI Description

- Can one “validate” QM-NSI approach from the QFT results?
- If yes, relation between NSI parameters and Lagrangian (EFT) parameters?
- Does the matching hold at all orders in perturbation?

QM-NSI Description

- Can one “validate” QM-NSI approach from the QFT results? **Yes...**
- If yes, relation between NSI parameters and Lagrangian (EFT) parameters?
- Does the matching hold at all orders in perturbation? **No...**

Observable is the same, we can match the two
(only at the linear level)

$$\epsilon_{\alpha\beta}^s = \sum_X p_{XL} [\epsilon_X]_{\alpha\beta}^*, \quad \epsilon_{\beta\alpha}^d = \sum_X d_{XL} [\epsilon_X]_{\alpha\beta}$$

Falkowski, González-Alonso, [ZT](#), JHEP (2019)

Comparing QM and QFT

Only at the linear order:

Falkowski, González-Alonso, ZT, JHEP (2019)

| Neutrino Process | NSI Matching with EFT |
|--|---|
| ν_e produced in beta decay | $\epsilon_{e\beta}^s = [\epsilon_L]_{e\beta}^* - [\epsilon_R]_{e\beta}^* - \frac{g_T}{g_A} \frac{m_e}{f_T(E_\nu)} [\epsilon_T]_{e\beta}^*$ |
| ν_e detected in inverse beta decay | $\epsilon_{\beta e}^d = [\epsilon_L]_{e\beta} + \frac{1-3g_A^2}{1+3g_A^2} [\epsilon_R]_{e\beta} - \frac{m_e}{E_\nu - \Delta} \left(\frac{g_S}{1+3g_A^2} [\epsilon_S]_{e\beta} - \frac{3g_A g_T}{1+3g_A^2} [\epsilon_T]_{e\beta} \right)$ |
| ν_μ produced in pion decay | $\epsilon_{\mu\beta}^s = [\epsilon_L]_{\mu\beta}^* - [\epsilon_R]_{\mu\beta}^* - \frac{m_\pi^2}{m_\mu(m_u+m_d)} [\epsilon_P]_{\mu\beta}^*$ |

- Different NP interactions appear at the source or detection simultaneously
- Some of the $p_{\text{XL}}/d_{\text{XL}}$ coefficients depend on the neutrino energy
- There are chiral enhancements in some cases

These correlations, energy dependence etc. cannot be
seen in the traditional QM approach.

Comparing QM and QFT

Beyond the linear order in new physics parameters, the NSI formula matches the (correct) one derived in the EFT only if the **consistency condition** is satisfied

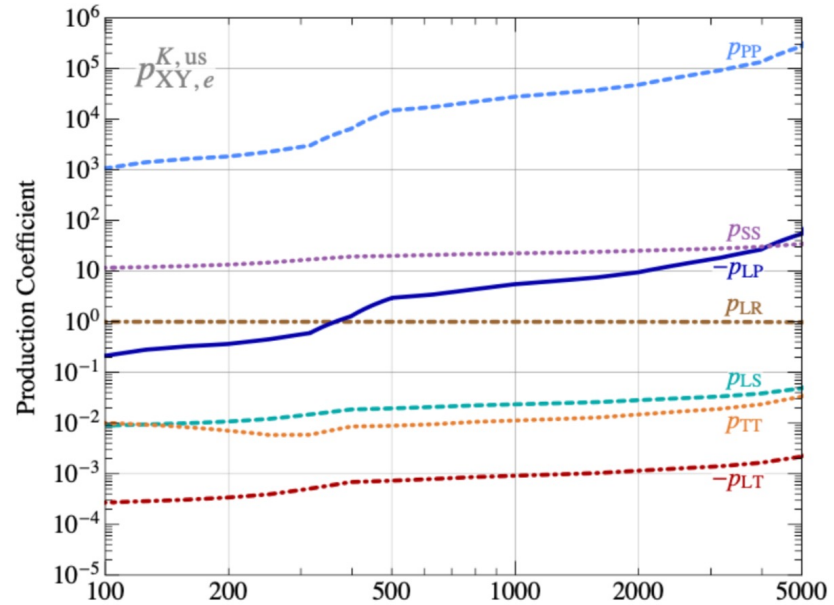
$$p_{XL} p_{YL}^* = p_{XY}, \quad d_{XL} d_{YL}^* = d_{XY}$$

This is always satisfied for new physics correcting V-A interactions only as $p_{LL} = d_{LL} = 1$ by definition

However for non-V-A new physics the consistency condition is not satisfied in general

Falkowski, González-Alonso, ZT, JHEP (2019)

$$p_{XY} \equiv \frac{\int d\Pi_{P'} A_X^P \bar{A}_Y^P}{\int d\Pi_{P'} |A_L^P|^2}, \quad d_{XY} \equiv \frac{\int d\Pi_D A_X^D \bar{A}_Y^D}{\int d\Pi_D |A_L^D|^2}$$



WEFT Power Counting

- Dim-6: $\frac{\Delta R}{R_{SM}} = c \epsilon_X^2$

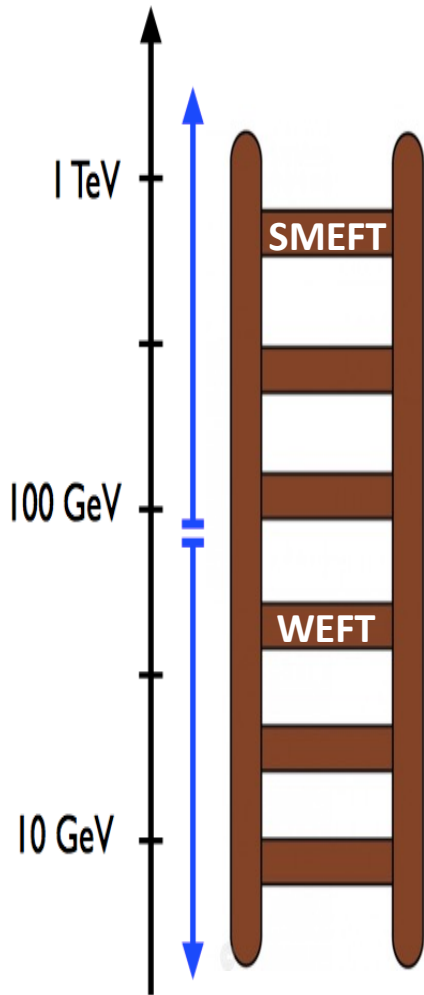
- Dim-7: Cannot interfere with the SM amplitudes, suppressed!

Liao et al, JHEP 08 (2020) 162

- Dim-8: $\frac{\Delta R}{R_{SM}} = \sqrt{c} \epsilon_8 E^2 / v^2$

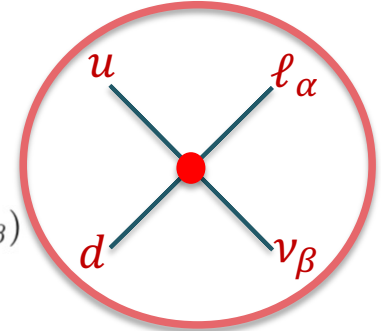
EFT ladder

WEFT: Effective Lagrangian defined at a low scale $\mu \sim 2 \text{ GeV}$



- CC: New left/right handed, (pseudo)scalar and tensor interactions

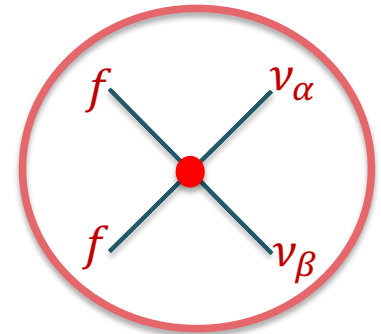
$$\mathcal{L}_{\text{WEFT}} \supset -\frac{2V_{ud}}{v^2} \left\{ [1 + \epsilon_{L\alpha\beta}] (\bar{u}\gamma^\mu P_L d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \right. \\ + \epsilon_{R\alpha\beta} (\bar{u}\gamma^\mu P_R d) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \\ + \frac{1}{2} \epsilon_{S\alpha\beta} (\bar{u}d) (\bar{\ell}_\alpha P_L \nu_\beta) - \frac{1}{2} \epsilon_{P\alpha\beta} (\bar{u}\gamma_5 d) (\bar{\ell}_\alpha P_L \nu_\beta) \\ \left. + \frac{1}{4} \hat{\epsilon}_{T\alpha\beta} (\bar{u}\sigma^{\mu\nu} P_L d) (\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{h.c.} \right\}$$



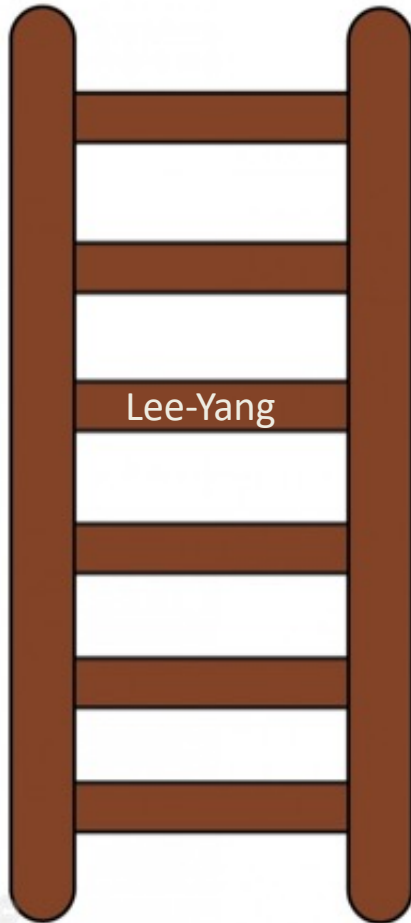
- NC: New left and right handed interactions

$$\mathcal{L}_{\text{WEFT}} \supset -\frac{2}{v^2} \epsilon_{\alpha\beta}^{fX} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu P_X f)$$

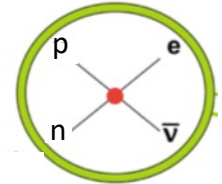
- Neutrino experiments
- Hadron Decays
- β -decays



EFT ladder



$$E \ll m_Z$$



- At the energy scale of reactor neutrino experiments the relevant degrees of freedom are not quarks, but nucleons and nuclei. Matching this EFT to the WEFT Lagrangian we obtain the Lee-Yang Lagrangian:

$$\begin{aligned} \mathcal{L}_{\text{LY}} \supset & -\frac{V_{ud}}{v^2} \left\{ g_V [\mathbf{1} + \epsilon_L + \epsilon_R]_{\alpha\beta} (\bar{p}\gamma^\mu n) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \right. \\ & - g_A [\mathbf{1} + \epsilon_L - \epsilon_R]_{\alpha\beta} (\bar{p}\gamma^\mu \gamma_5 n) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \\ & + g_S [\epsilon_S]_{\alpha\beta} (\bar{p}n) (\bar{\ell}_\alpha P_L \nu_\beta) - g_P [\epsilon_P]_{\alpha\beta} (\bar{p}\gamma_5 n) (\bar{\ell}_\alpha P_L \nu_\beta) \\ & \left. + \frac{1}{2} g_T [\hat{\epsilon}_T]_{\alpha\beta} (\bar{p}\sigma^{\mu\nu} P_L n) (\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{h.c.} \right\}, \end{aligned}$$

- Lattice+theory fix the non-perturbative parameters with good precision

$$g_A = 1.2728 \pm 0.0017, \quad g_S = 1.02 \pm 0.11, \quad g_P = 349 \pm 9, \quad g_T = 0.987 \pm 0.055.$$

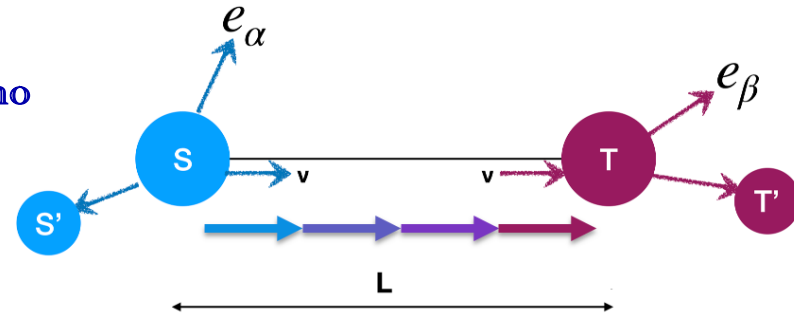
- T. Bhattacharya et al, Phys. Rev. D94 (2016), no. 5 054508
- M. Gonzalez-Alonso and J. Martin Camalich, Phys. Rev. Lett. 112 (2014), no. 4 042501
- M. Gonzalez-Alonso et al, Prog. Part. Nucl. Phys. 104 (2019) 165–223

EFT at neutrino experiments

We proposed a systematic approach to neutrino oscillations in the SMEFT framework!

Falkowski, González-Alonso, ZT, JHEP (2020)

$$U_{\text{PMNS}} \parallel \begin{matrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{matrix} \begin{matrix} \nu_1 & \nu_2 & \nu_3 \end{matrix}$$



Observable: rate of detected events

$$\sim (\text{flux}) \times (\text{det. cross section}) \times (\text{oscillation})$$

CC EFT

NC EFT

depend on the kinematic and spin variables

$$\mathcal{M}_{\alpha k}^P = U_{\alpha k}^* A_L^P + \sum_X [\epsilon_X U]_{\alpha k}^* A_X^P$$

$$\mathcal{M}_{\beta k}^D = U_{\beta k} A_L^D + \sum_X [\epsilon_X U]_{\beta k} A_X^D$$

Corrections to fluxes/cross sections

$$\sigma^{\text{Total}} = \sigma^{\text{SM}} + \epsilon_X \sigma^{\text{Int}} + \epsilon_X^2 \sigma^{\text{NP}} \sim \sigma^{\text{SM}} (1 + \epsilon_X d_{XL} + \epsilon_X^2 d_{XX})$$

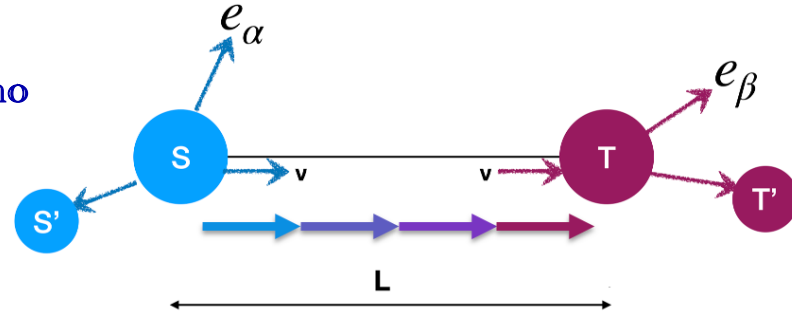
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$$\phi^{\text{Total}} \sim \phi^{\text{SM}} (1 + \epsilon_X p_{XL} + \epsilon_X^2 p_{XX})$$

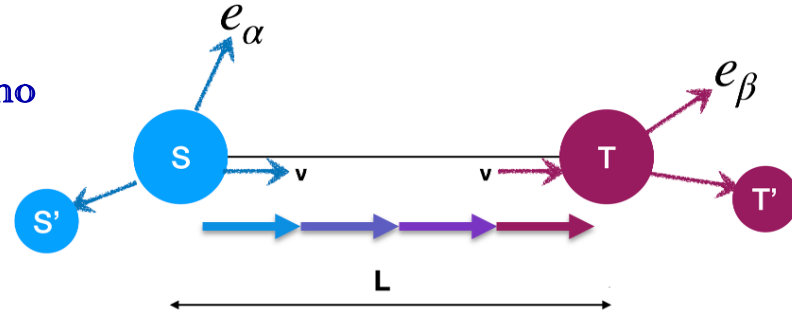
$$\sigma^{\text{Total}} \sim \sigma^{\text{SM}} (1 + \epsilon_X d_{XL} + \epsilon_X^2 d_{XX})$$

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$$\times [U_{\beta k} U_{\beta l}^* + d_{XL} (\epsilon_X U)_{\beta k} U_{\beta l}^* + d_{XL}^* U_{\beta k} (\epsilon_X U)_{\beta l}^* + d_{XX} (\epsilon_X U)_{\beta k} (\epsilon_X U)_{\beta l}^*]$$

$$\mathcal{H}_F = \frac{1}{2E} (U M^2 U^\dagger + \mathbb{A})$$

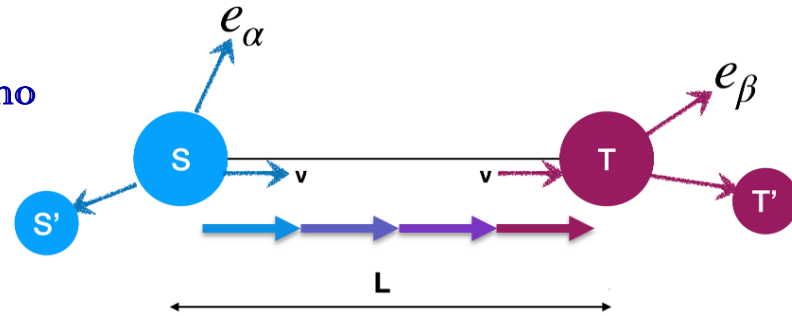
$$M^2 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix}, \quad \mathbb{A} = \begin{pmatrix} A_{\text{CC}} & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

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Falkowski, González-Alonso, ZT, JHEP (2020)

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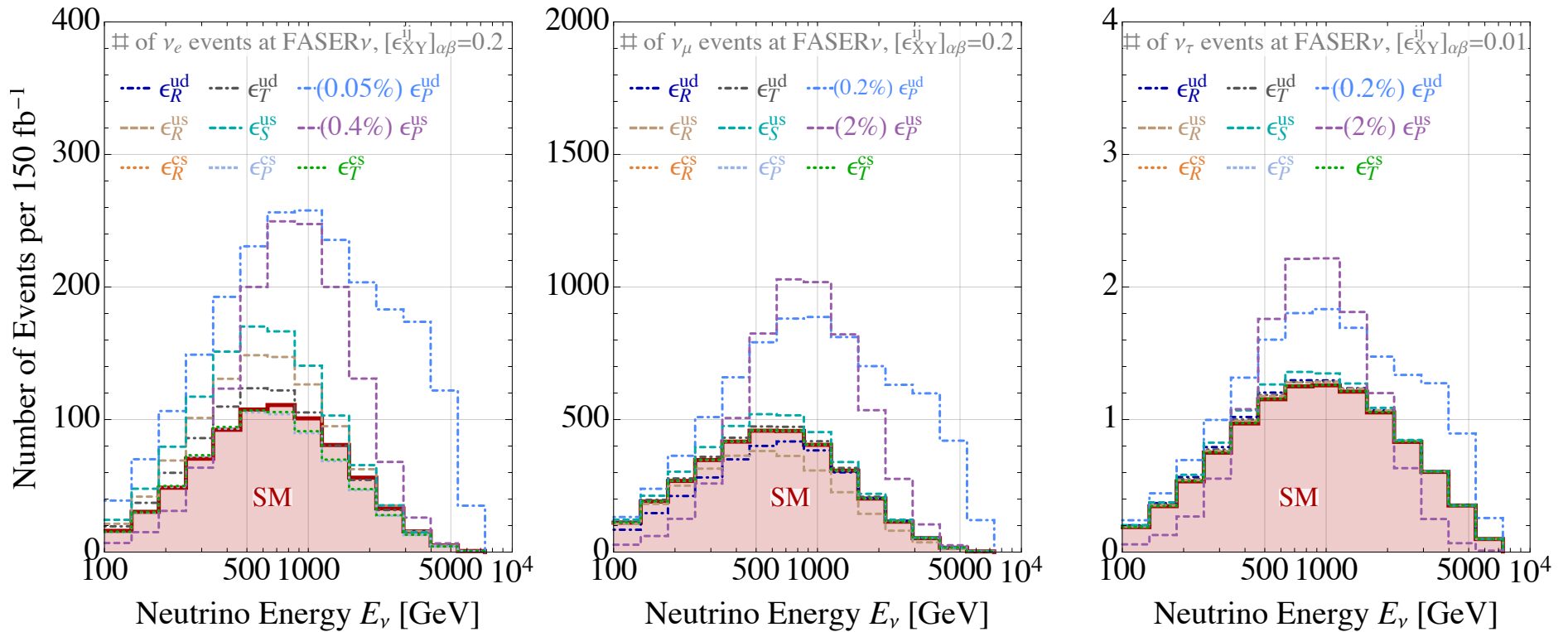
$$R_{\alpha\beta} = \Phi_\alpha^{\text{SM}} \sigma_\beta^{\text{SM}} \sum_{k,l} e^{-i \frac{L \Delta m_{kl}^2}{2E_\nu}} [U_{\alpha k}^* U_{\alpha l} + p_{XL} (\epsilon_X U)_{\alpha k}^* U_{\alpha l} + p_{XL}^* U_{\alpha k}^* (\epsilon_X U)_{\alpha l} + p_{XX} (\epsilon_X U)_{\alpha k}^* (\epsilon_X U)_{\alpha l}]$$

$$\times [U_{\beta k} U_{\beta l}^* + d_{XL} (\epsilon_X U)_{\beta k} U_{\beta l}^* + d_{XL}^* U_{\beta k} (\epsilon_X U)_{\beta l}^* + d_{XX} (\epsilon_X U)_{\beta k} (\epsilon_X U)_{\beta l}^*]$$

$$\mathcal{H}_F = \frac{1}{2E} (U M^2 U^\dagger + \mathbb{A}) \quad + \quad \pm \sqrt{2} G_F N_e \begin{pmatrix} \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{\mu e} & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{\tau e} & \epsilon_{\tau\mu} & \epsilon_{\tau\tau} \end{pmatrix}$$

EFT at FASERν

Falkowski, González-Alonso, Kopp, Soreq, [ZT](#), JHEP (2021)



- Analysis is statistics dominated: $\nu_e \sim 1000$, $\nu_\mu \sim 5000$, $\nu_\tau \sim 10$
- Optimistic systematic uncertainties: 5% on ν_e , 10% on ν_μ , 15% on ν_τ
- Pessimistic systematic uncertainties: 30% on ν_e , 40% on ν_μ , 50% on ν_τ

Pion decay

Production

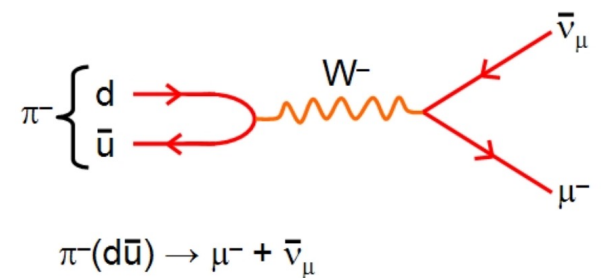
Falkowski, González-Alonso, ZT, JHEP (2020)

Due to the pseudoscalar nature of the pion, it is sensitive only to axial (ϵ_L - ϵ_R) and pseudo-scalar (ϵ_P) interactions.

$$p_{LL} = -p_{RL} = 1, \quad p_{PL} = -p_{PR} = -\frac{m_\pi^2}{m_\mu(m_u + m_d)},$$

$$p_{RR} = 1, \quad p_{PP} = \frac{m_\pi^4}{m_\mu^2(m_u + m_d)^2} \sim 700!$$

$$\sim -27$$



- Larger $p_{XY} \Rightarrow$ smaller $\epsilon!$

$$\phi^{Total} \sim \phi^{SM} (1 + \epsilon_X p_{XL} + \epsilon_X^2 p_{XX})$$

Huge overall flux normalization for pion decay!

$$\langle 0 | \bar{d} \gamma^\mu \gamma_5 u | \pi^+(p_\pi) \rangle = i p_\pi^\mu f_\pi$$

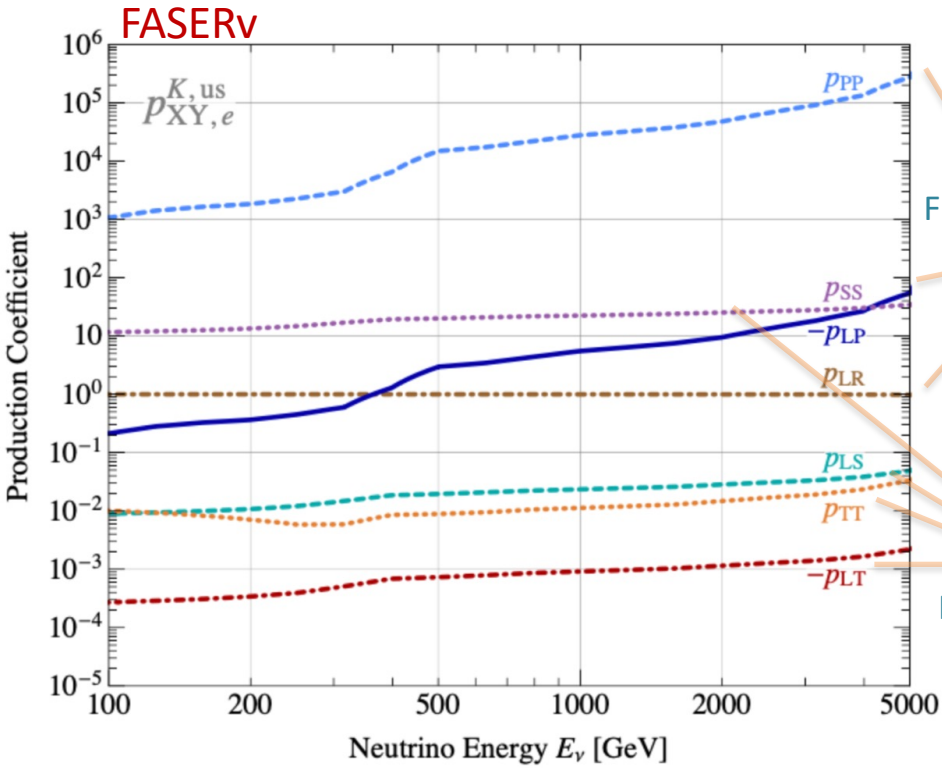
$$\langle 0 | \bar{d} \gamma_5 u | \pi^+(p_\pi) \rangle = -i \frac{m_\pi^2}{m_u + m_d} f_\pi$$

kaon decay

Production

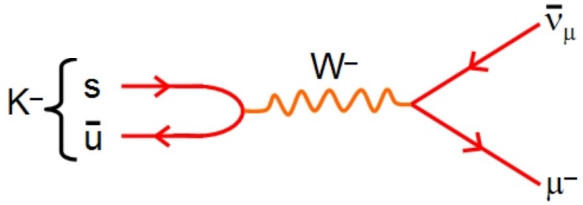
Falkowski, González-Alonso, Kopp, Soreq, ZT, JHEP (2021)

Both 2-body and 3-body kaon decays contribute:

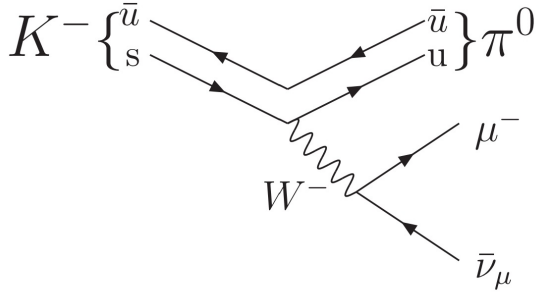


From 2-b decay

From 3-b decay



$$K^-(s\bar{u}) \rightarrow \mu^- + \bar{\nu}_\mu$$



Depends on energy distribution of K^\pm , K_L or K_S at each experiments

$$\langle \pi^- | \bar{s} \gamma^\mu u | K^0 \rangle = P^\mu f_+(q^2) + q^\mu f_-(q^2),$$

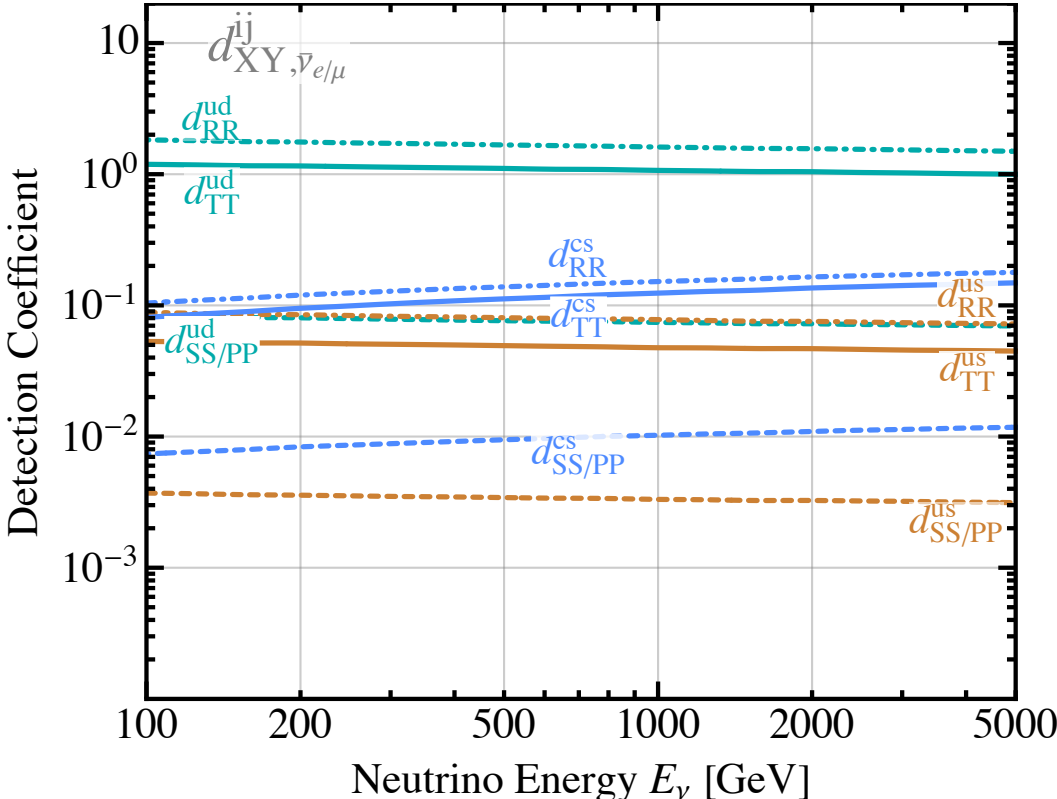
$$\langle \pi^- | \bar{s} u | K^0 \rangle = -\frac{m_K^2 - m_\pi^2}{m_s - m_u} f_0(q^2),$$

$$\langle \pi^- | \bar{s} \sigma^{\mu\nu} u | K^0 \rangle = i \frac{p_K^\mu p_\pi^\nu - p_\pi^\mu p_K^\nu}{m_K} B_T(q^2),$$

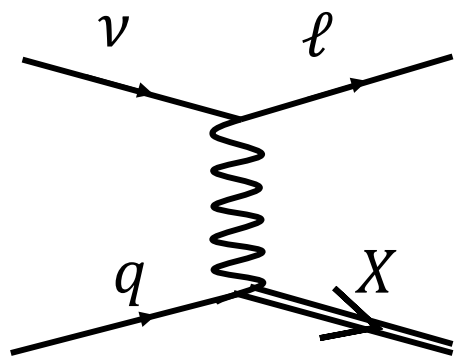
DIS

Detection

Falkowski, González-Alonso, Kopp, Soreq, ZT, JHEP (2021)



Deep Inelastic Scattering



$$\sigma^{Total} \sim \sigma^{SM} (1 + \epsilon_X d_{XL} + \epsilon_X^2 d_{XX})$$

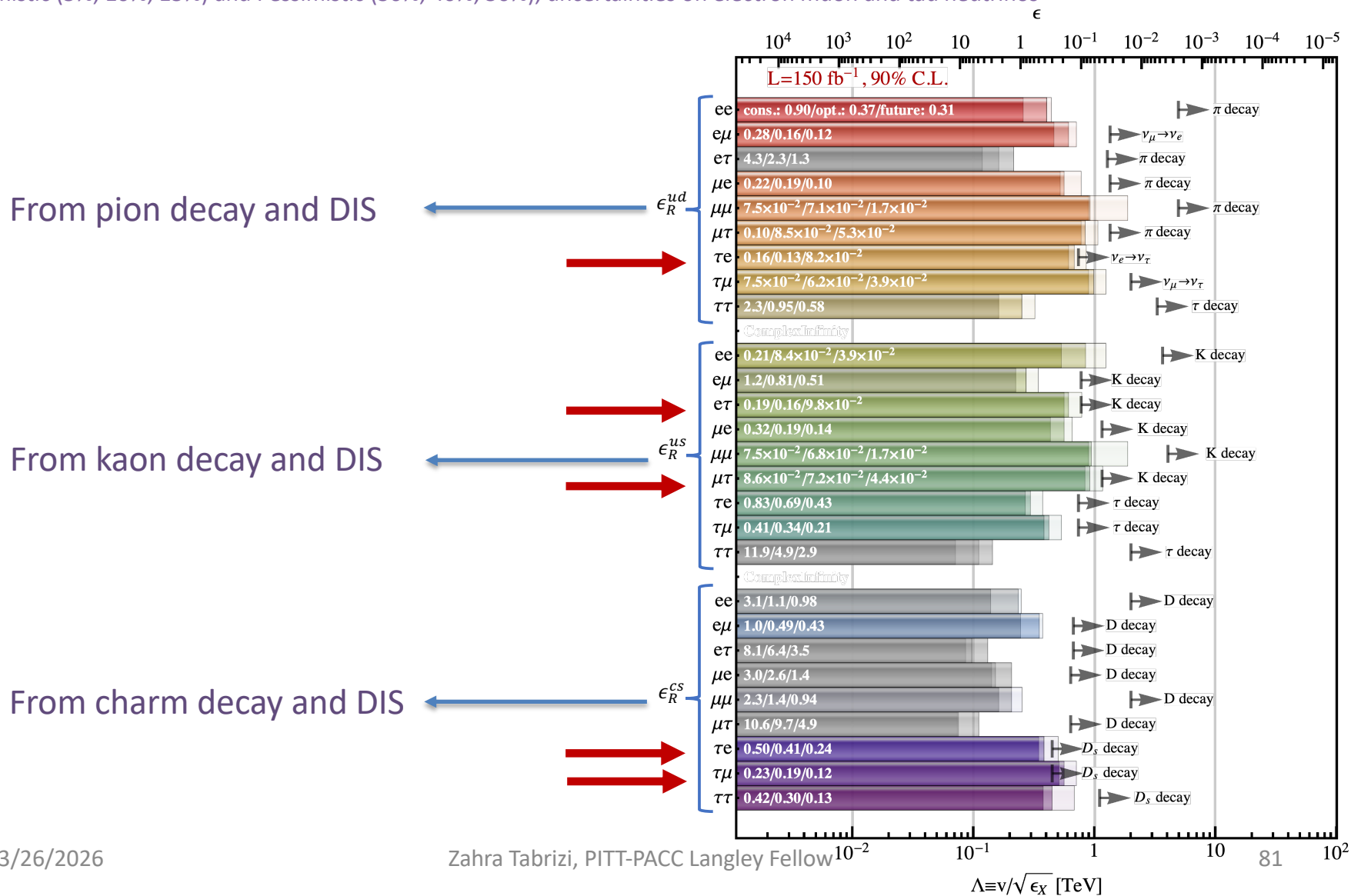
ϵ_X^2 is more important than ϵ_X !

RESULTS

Turning on one interaction at a time: Right handed

A. Falkowski, M. González-Alonso, J. Kopp, Y. Soreq, [ZJHEP 10 \(2021\) 086](#)

Optimistic (5%, 10%, 15%) and Pessimistic (30%, 40%, 50%), uncertainties on electron muon and tau neutrinos

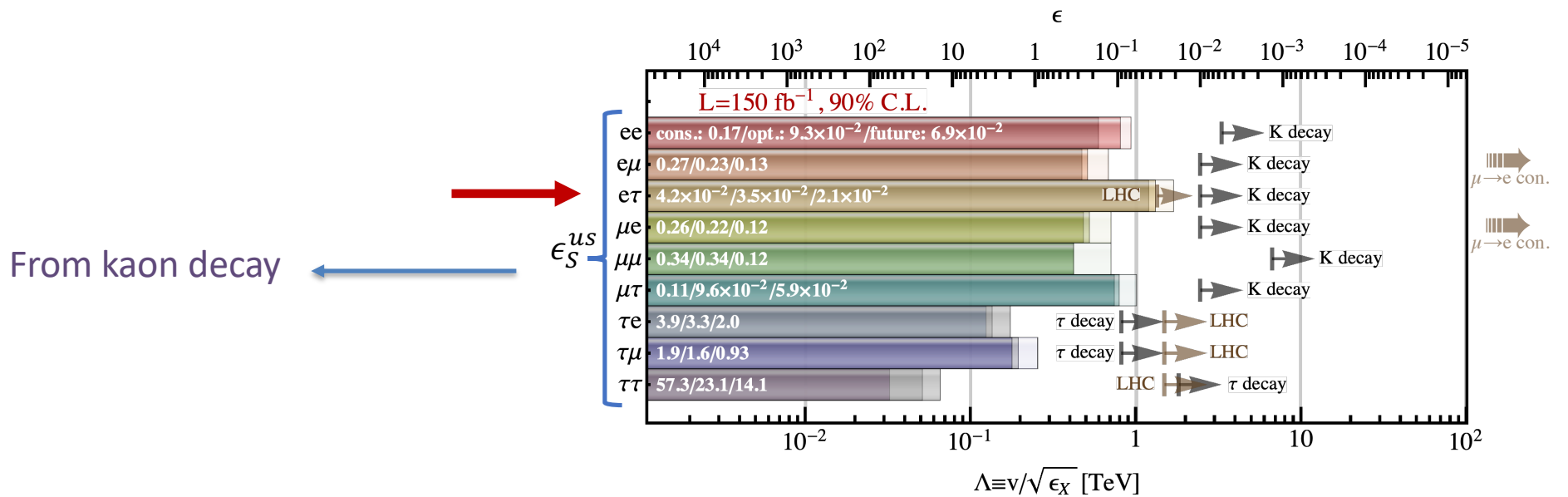


RESULTS

Turning on one interaction at a time: Scalar

A. Falkowski, M. González-Alonso, J. Kopp, Y. Soreq, [ZJHEP 10 \(2021\) 086](#)

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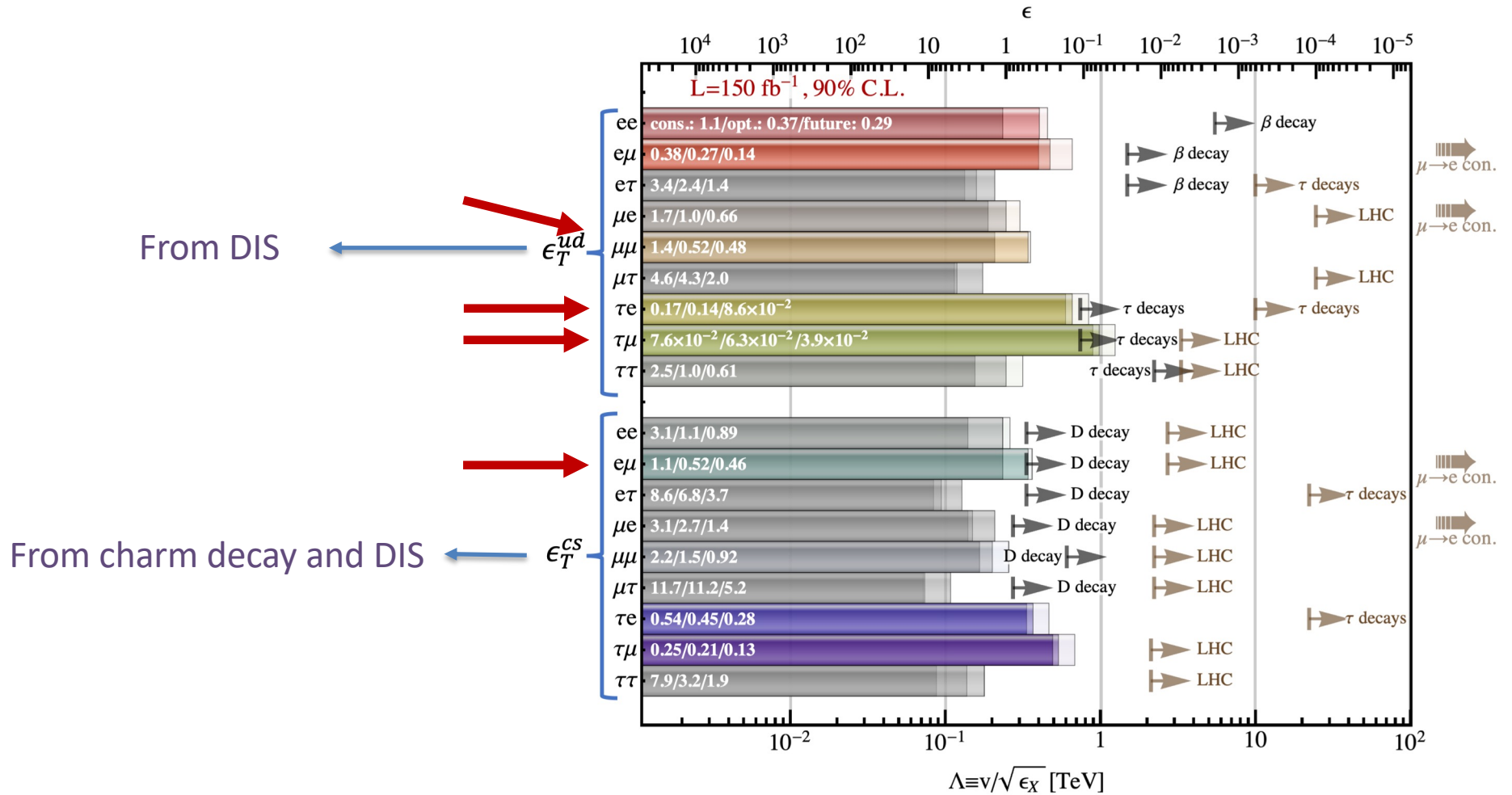


RESULTS

Turning on one interaction at a time: Tensor

A. Falkowski, M. González-Alonso, J. Kopp, Y. Soreq, [ZT](#)
 JHEP 10 (2021) 086

Optimistic (5%, 10%, 15%) and Pessimistic (30%, 40%, 50%), uncertainties on electron muon and tau neutrinos



EFT at FASERv

A. Falkowski, M. González-Alonso, J. Kopp, Y. Soreq, ZJHEP 10 (2021) 086

FASERv Flavor Experiments

Colliders

Neutrino experiments:

- Many more operators can be probed (81 at FASERv)

Low energy:

- Independent of the underlying high-energy theory

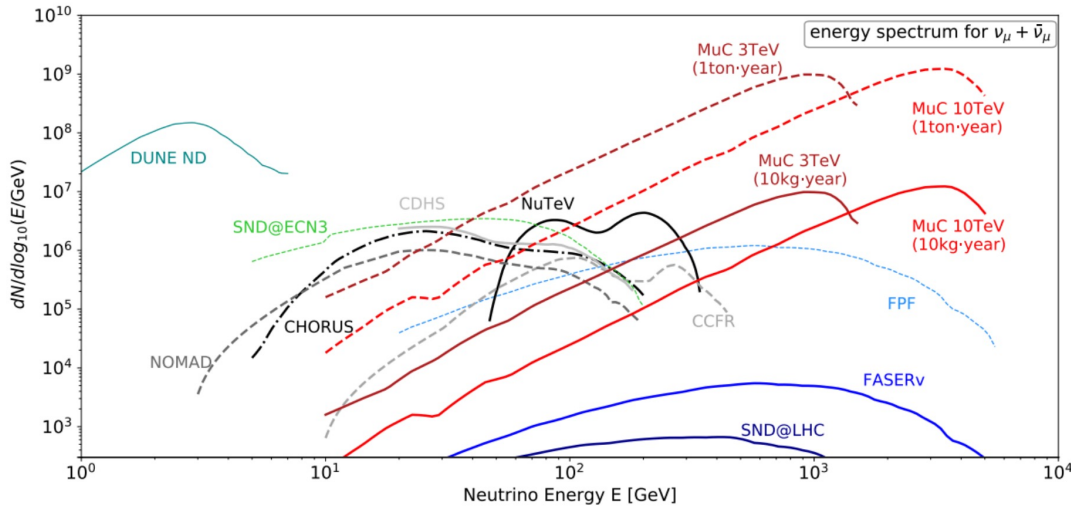
High-Energy:

- SMEFT is the underlying theory
- Bounds are less robust

Bounds shown in bold face have been calculated in this work

| Coupling | Low energy (WEFT) | | High energy / CLFV (SMEFT) | |
|--------------------------------|----------------------|---|--|---|
| | 90 % CL bound | process | 90 % CL bound | process |
| $[\epsilon_P^{ud}]_{ee}$ | 4.6×10^{-7} | $\Gamma_{\pi \rightarrow e\nu} / \Gamma_{\pi \rightarrow \mu\nu}$ | | |
| $[\epsilon_P^{ud}]_{e\mu}$ | 7.3×10^{-6} | $\Gamma_{\pi \rightarrow e\nu} / \Gamma_{\pi \rightarrow \mu\nu}$ [7] | 2.0×10^{-8} | $\mu \rightarrow e$ conversion |
| $[\epsilon_P^{ud}]_{e\tau}$ | 7.3×10^{-6} | $\Gamma_{\pi \rightarrow e\nu} / \Gamma_{\pi \rightarrow \mu\nu}$ [7] | 2.5×10^{-3} | LHC [64] |
| $[\epsilon_P^{ud}]_{\mu e}$ | 2.6×10^{-3} | $\Gamma_{\pi \rightarrow e\nu} / \Gamma_{\pi \rightarrow \mu\nu}$ | 2.0×10^{-8} | $\mu \rightarrow e$ conversion |
| $[\epsilon_P^{ud}]_{\mu\mu}$ | 9.4×10^{-5} | $\Gamma_{\pi \rightarrow e\nu} / \Gamma_{\pi \rightarrow \mu\nu}$ | | |
| $[\epsilon_P^{ud}]_{\mu\tau}$ | 2.6×10^{-3} | $\Gamma_{\pi \rightarrow e\nu} / \Gamma_{\pi \rightarrow \mu\nu}$ | | |
| $[\epsilon_P^{ud}]_{\tau e}$ | 9.0×10^{-2} | $\Gamma_{\tau \rightarrow \pi\nu}$ | $5.8 \times 10^{-3(*)} / 4.4 \times 10^{-4}$ | LHC [65] / τ decay [64] |
| $[\epsilon_P^{ud}]_{\tau\mu}$ | 9.0×10^{-2} | $\Gamma_{\tau \rightarrow \pi\nu}$ | $5.8 \times 10^{-3(*)}$ | LHC [65] |
| $[\epsilon_P^{ud}]_{\tau\tau}$ | 8.4×10^{-3} | τ -decay [65] | $5.8 \times 10^{-3(*)}$ | LHC [65] |
| $[\epsilon_P^{us}]_{ee}$ | 1.1×10^{-6} | $\Gamma_{K \rightarrow e\nu} / \Gamma_{K \rightarrow \mu\nu}$ | | |
| $[\epsilon_P^{us}]_{e\mu}$ | 2.1×10^{-5} | $\Gamma_{K \rightarrow e\nu} / \Gamma_{K \rightarrow \mu\nu}$ | 6.2×10^{-7} | $\mu \rightarrow e$ conversion |
| $[\epsilon_P^{us}]_{e\tau}$ | 2.1×10^{-5} | $\Gamma_{K \rightarrow e\nu} / \Gamma_{K \rightarrow \mu\nu}$ | 7.1×10^{-2} | LHC [64] |
| $[\epsilon_P^{us}]_{\mu e}$ | 2.3×10^{-3} | $\Gamma_{K \rightarrow e\nu} / \Gamma_{K \rightarrow \mu\nu}$ | 6.2×10^{-7} | $\mu \rightarrow e$ conversion |
| $[\epsilon_P^{us}]_{\mu\mu}$ | 2.2×10^{-4} | $\Gamma_{K \rightarrow e\nu} / \Gamma_{K \rightarrow \mu\nu}$ | | |
| $[\epsilon_P^{us}]_{\mu\tau}$ | 2.3×10^{-3} | $\Gamma_{K \rightarrow e\nu} / \Gamma_{K \rightarrow \mu\nu}$ | | |
| $[\epsilon_P^{us}]_{\tau e}$ | 6.4×10^{-2} | $\Gamma_{\tau \rightarrow K\nu} / \Gamma_{K \rightarrow \mu\nu}$ | $3.1 \times 10^{-2(*)} / 8.1 \times 10^{-2}$ | LHC (data [66]) / τ -decay [64] |
| $[\epsilon_P^{us}]_{\tau\mu}$ | 6.4×10^{-2} | $\Gamma_{\tau \rightarrow K\nu} / \Gamma_{K \rightarrow \mu\nu}$ | $3.1 \times 10^{-2(*)}$ | LHC (data [66]) |
| $[\epsilon_P^{us}]_{\tau\tau}$ | 1.3×10^{-2} | τ -decay [67] | $3.1 \times 10^{-2(*)}$ | LHC (data [66]) |
| $[\epsilon_P^{cs}]_{ee}$ | 4.8×10^{-3} | $\Gamma_{D_s \rightarrow e\nu}$ | 1.3×10^{-2} | LHC [68] |
| $[\epsilon_P^{cs}]_{e\mu}$ | 4.6×10^{-3} | $\Gamma_{D_s \rightarrow e\nu}$ | $1.3 \times 10^{-2} / 2.7 \times 10^{-6}$ | LHC [68] / $\mu \rightarrow e$ conversion |
| $[\epsilon_P^{cs}]_{e\tau}$ | 4.6×10^{-3} | $\Gamma_{D_s \rightarrow e\nu}$ | $1.3 \times 10^{-2} / 1.9 \times 10^{-2}$ | LHC / τ -decays [64, 68] |
| $[\epsilon_P^{cs}]_{\mu e}$ | 8.9×10^{-3} | $\Gamma_{D_s \rightarrow \mu\nu}$ | $2.0 \times 10^{-2} / 2.7 \times 10^{-6}$ | LHC [68] / $\mu \rightarrow e$ conversion |
| $[\epsilon_P^{cs}]_{\mu\mu}$ | 1.0×10^{-3} | $\Gamma_{D_s \rightarrow \mu\nu}$ | 2.0×10^{-2} | LHC [68] |
| $[\epsilon_P^{cs}]_{\mu\tau}$ | 8.9×10^{-3} | $\Gamma_{D_s \rightarrow \mu\nu}$ | 2.0×10^{-2} | LHC [68] |
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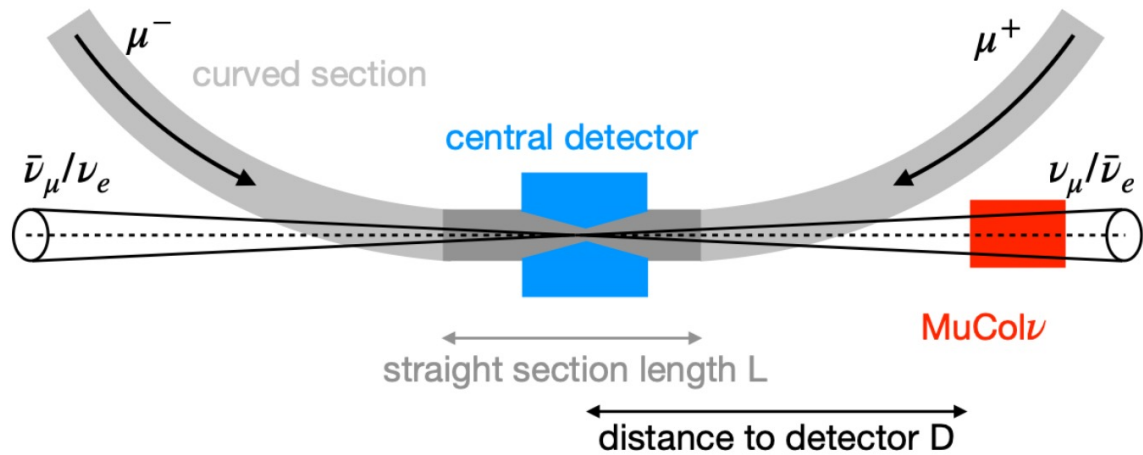
Muon Beam Neutrinos:



$$N \sim 10^5 \times \left(\frac{m}{\text{kg}}\right) \times \left(\frac{t}{\text{yr}}\right) \times \left(\frac{L}{50 \text{ m}}\right) \times \left(\frac{500 \text{ m}}{D}\right)^2$$

10⁷ and 10⁹ total interactions at the forward region for the 10 kg-year and 1 ton-year setups

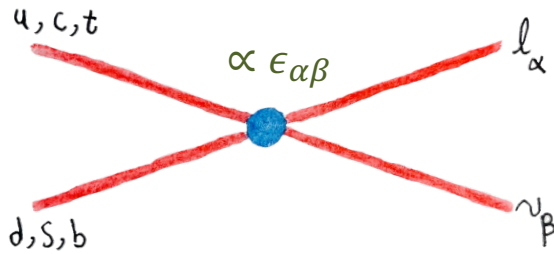
Forward Detectors:
e.g. MuColv
F. Kling, Y. Ma, K. Mękała, J. Reuter, [ZT](#)
arxiv: 2508.00761



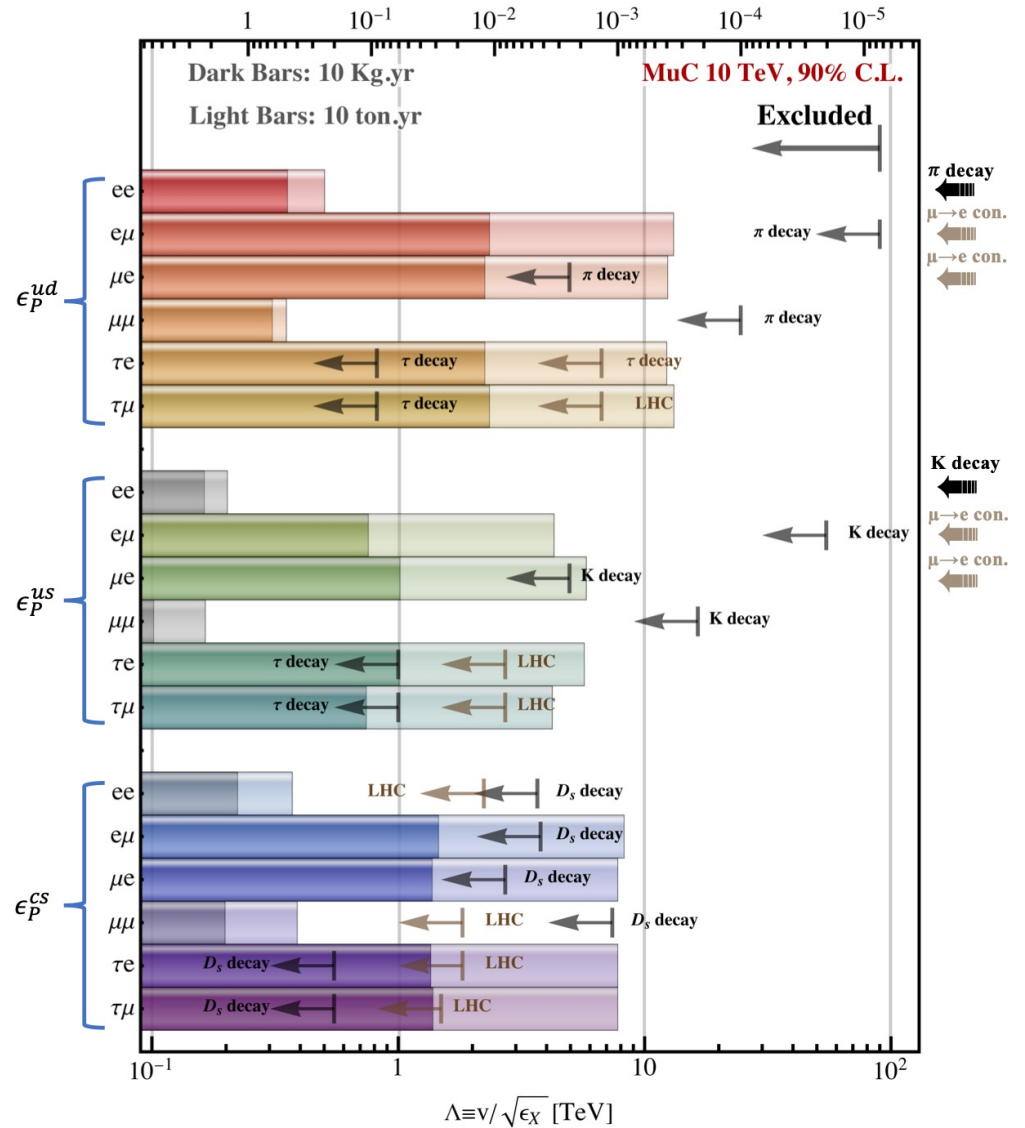
Neutrino EFT:

F. Kling, Y. Ma, K. Mękała, J. Reuter, [ZT](#)
 arxiv: 2508.00761

- **MuColv**: colored bars
- Dark/Light bars: 10 kg.year/10 ton.year

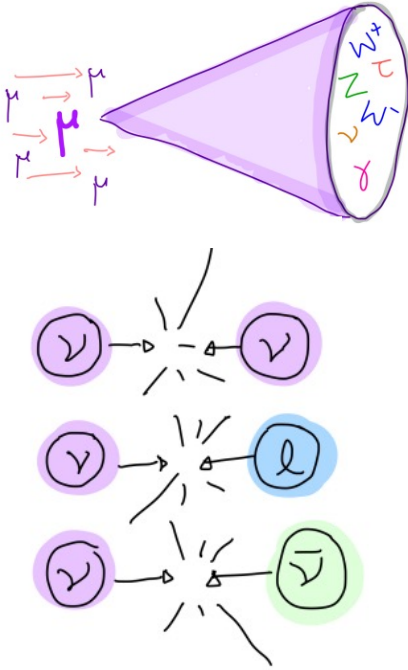


- New physics reach at tens of TeV
- Several dominant constraints

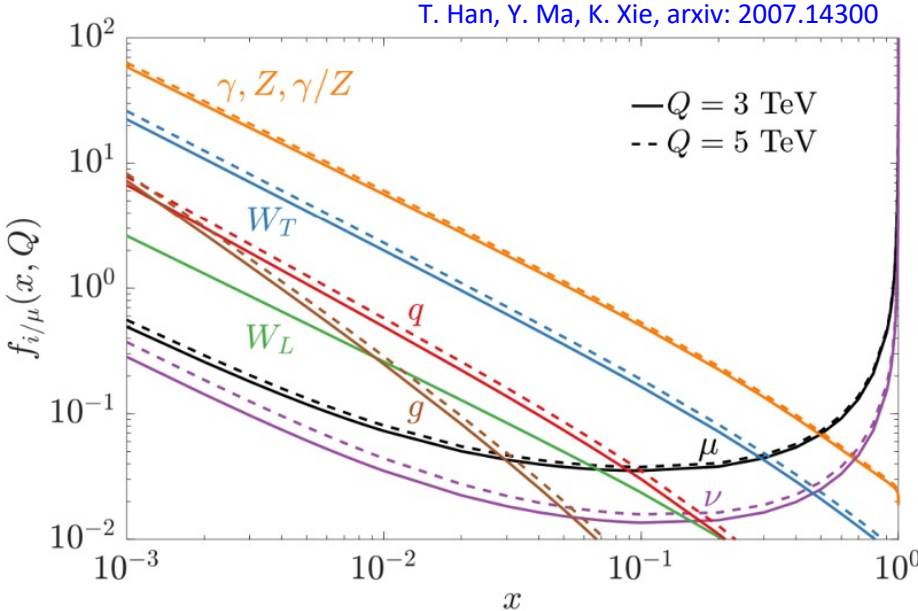


Neutrino-Neutrino Collider?

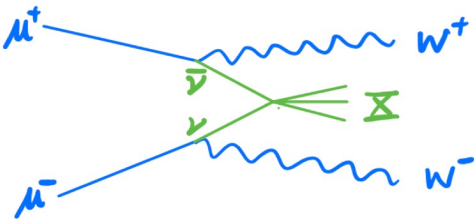
PDF from EW radiation



Credit: Innes Bigaran



T. Han, Y. Ma, K. Xie, arxiv: 2007.14300



High energy Muon Collider as a high energy Neutrino Collider

Bigaran, De Gouvea, Han, Jaffredo, Low, Ma, ZT, Xie,
In Preparation

Neutrino-Neutrino Collider?

Flavor-conserving 4-lepton operators

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \sum_i \frac{c_i}{v^2} O_i^{D=6} :$$

$$\mu^+ \mu^- : [C_{\ell\ell}], [C_{\ell e}], [C_{ee}]$$

$$\mu^\pm \nu : [C_{\ell\ell}], [C_{\ell e}]$$

$$\nu \bar{\nu} : [C_{\ell\ell}]$$

Two flavors ($a < b = 1, 2, 3$)

$$[O_{\ell\ell}]_{aabb} = (\bar{\ell}_a \bar{\sigma}_\mu \ell_a) (\bar{\ell}_b \bar{\sigma}^\mu \ell_b)$$

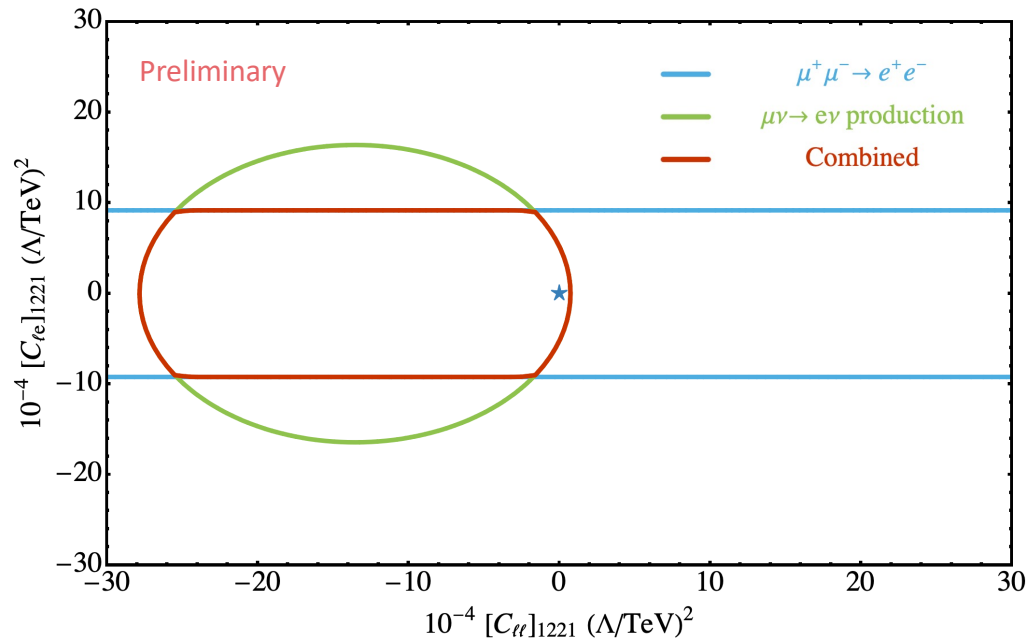
$$[O_{\ell\ell}]_{abba} = (\bar{\ell}_a \bar{\sigma}_\mu \ell_b) (\bar{\ell}_b \bar{\sigma}^\mu \ell_a)$$

$$[O_{\ell e}]_{aabb} = (\bar{\ell}_a \bar{\sigma}_\mu \ell_a) (e_b^c \sigma^\mu \bar{e}_b^c)$$

$$[O_{\ell e}]_{bbaa} = (\bar{\ell}_b \bar{\sigma}_\mu \ell_b) (e_a^c \sigma^\mu \bar{e}_a^c)$$

$$[O_{\ell e}]_{abba} = (\bar{\ell}_a \bar{\sigma}_\mu \ell_b) (e_b^c \sigma^\mu \bar{e}_a^c)$$

$$[O_{ee}]_{aabb} = (e_a^c \sigma_\mu \bar{e}_a^c) (e_b^c \sigma^\mu \bar{e}_b^c)$$



Bigaran, De Gouvea, Han, Jaffredo, Low, Ma, [ZT](#), Xie, In Preparation

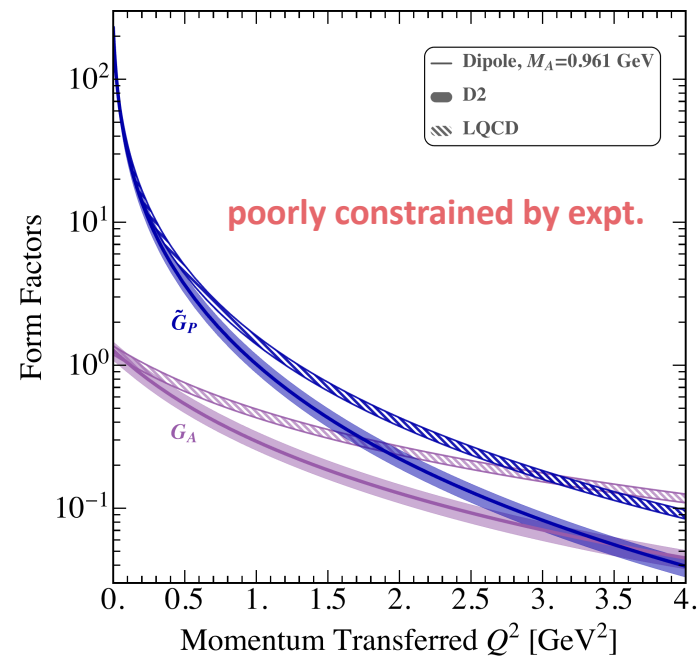
CCQE Hadronic Matrix Elements

Kopp, Rocco, ZT, JHEP (2024)

SM-Interactions:

Vector Current: Form Factors well understood (constrained by eN scattering)

Axial Current:



NEW-Interactions:

Kopp, Rocco, ZT, JHEP (2024)

- We add new scalar, pseudo-scalar and tensor currents;
- Interference with the SM;
- New Cross Section Contributions;

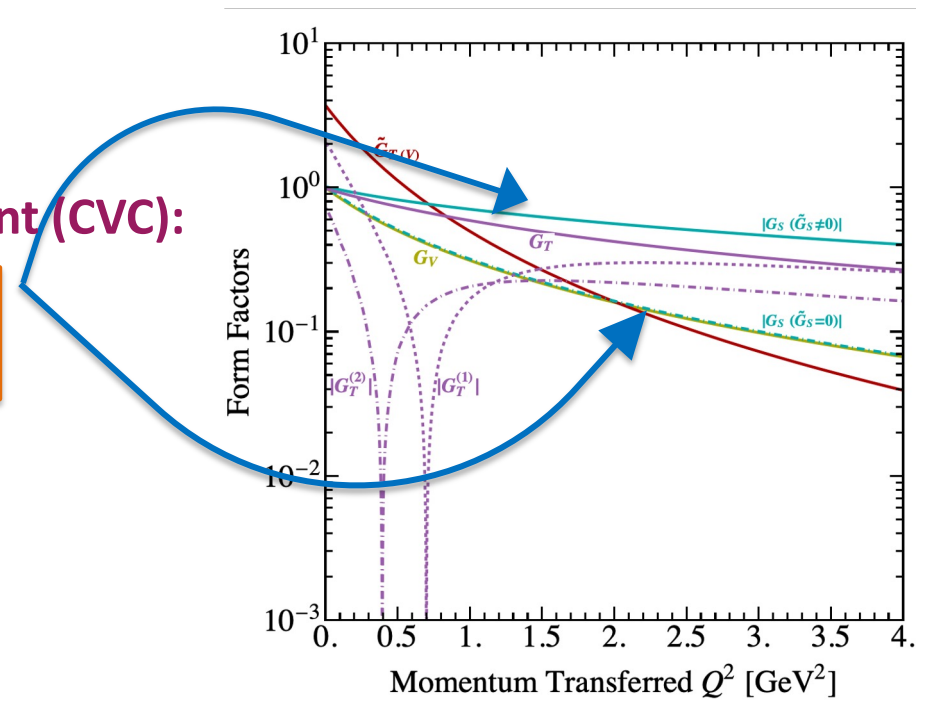
NEW-Interactions:

Kopp, Rocco, ZT, JHEP (2024)

- We add new scalar, pseudo-scalar and tensor currents;
- Interference with the SM;
- New Cross Section Contributions;
- **Scalar: conservation of the vector current (CVC):**

$$G_S(Q^2) = -\frac{\delta M_N^{QCD}}{\delta m_q} G_V(Q^2) + \frac{Q^2/2M_N}{\delta m_q} \tilde{G}_S(Q^2)$$

- We cannot neglect \tilde{G}_S (second class current) anymore!



NEW-Interactions:

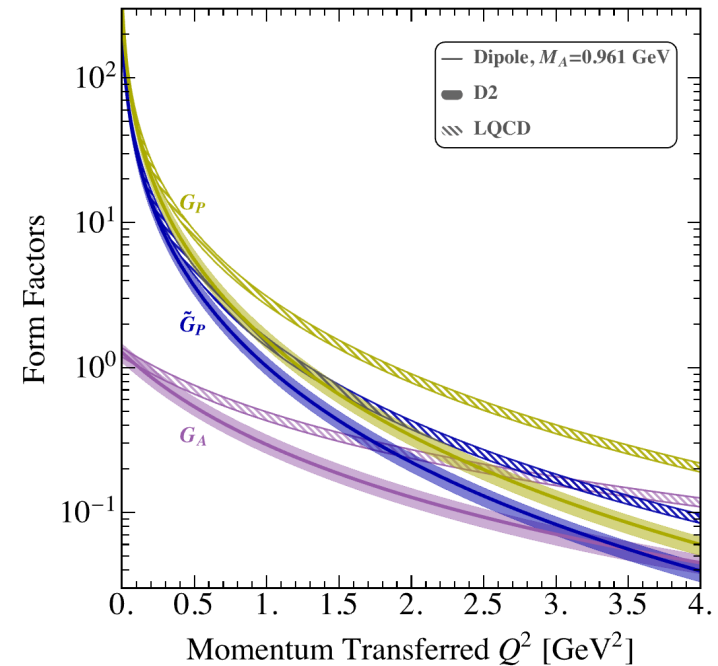
- We add new scalar, pseudo-scalar and tensor currents;
- Interference with the SM;
- New Cross Section Contributions;
- **Pseudo-Scalar: partial conservation of the axial current (PCAC)**

$$G_P(Q^2) = \frac{M_N}{m_q} G_A(Q^2) + \frac{Q^2/2M_N}{2m_q} \tilde{G}_P(Q^2)$$

➤ D2: neutrino-deuterium data (shaded band)

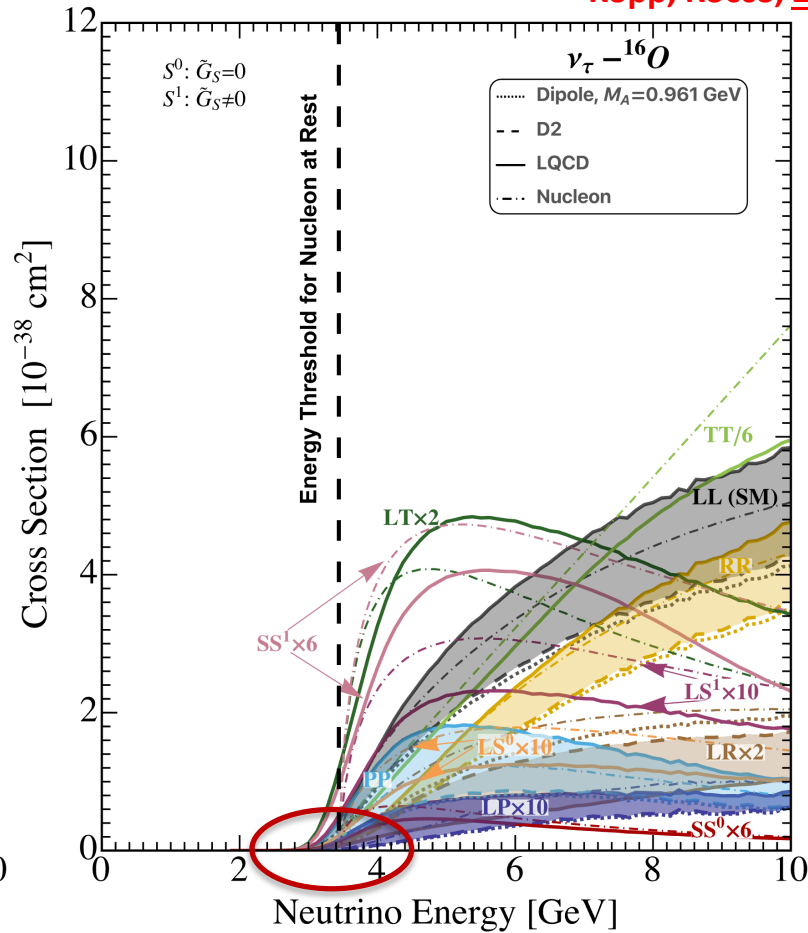
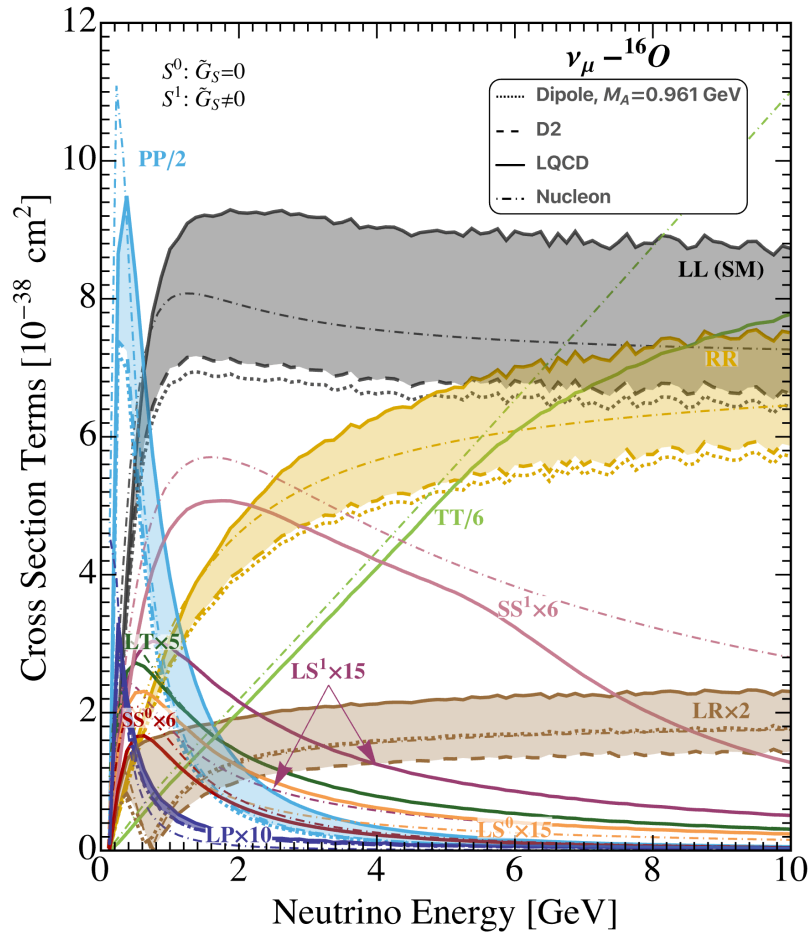
➤ RQCD Collaboration (hatched band)

Kopp, Rocco, ZT, JHEP (2024)



CCQE Neutrino-Nucleus Cross Sections:

Kopp, Rocco, ZT, JHEP (2024)



- CCQE Neutrino-Nucleus Scattering;
- All non-standard interactions;
- For all neutrino Flavors;

- z-expansion fit to LQCD and D2 data;
- Nuclear effects;
- Comparison with nucleon scattering

“What is Dark Matter?”

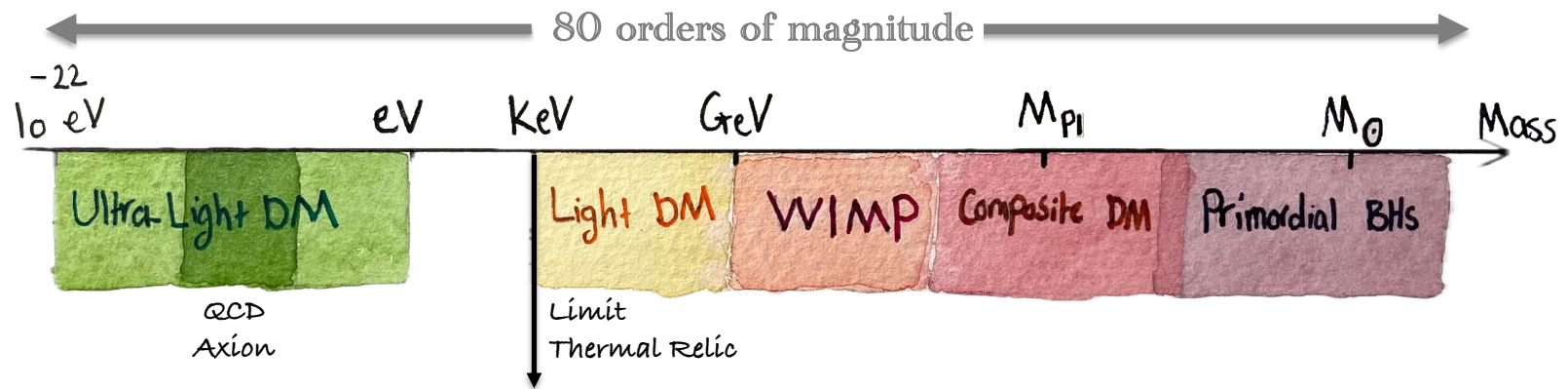
We don't know!

There could be several kinds, making up a whole “dark sector”

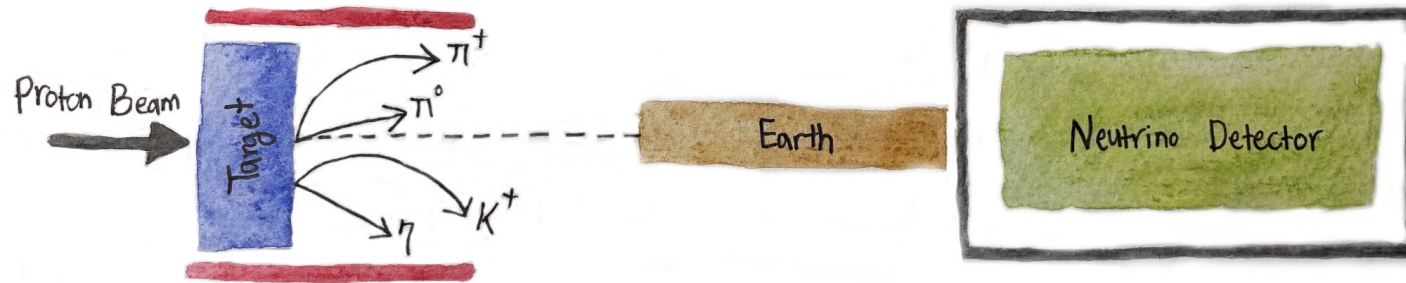


“Where is Dark Matter?”

We don't know!



Light Dark Matter

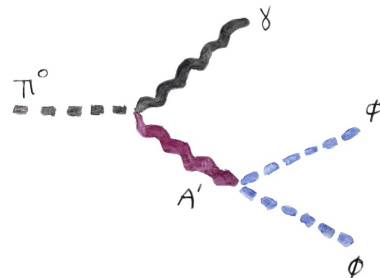


Photons at the target kinetically produce Dark Photons, which decay into dark matter:

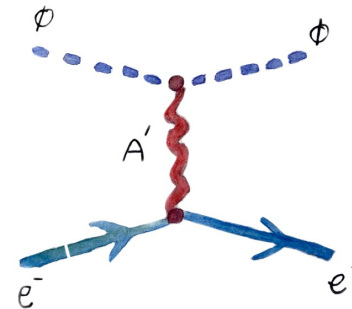
$$\mathcal{L} \supset -\frac{\epsilon}{2} F^{\mu\nu} F'_{\mu\nu} + \frac{M_{A'}^2}{2} A'_\mu A'^\mu + |D_\mu \phi|^2 - M_\phi^2 |\phi|^2$$

$D_\mu = \partial_\mu - ig_D A'_\mu, \quad g_D = \sqrt{4\pi\alpha_D}$

DM production

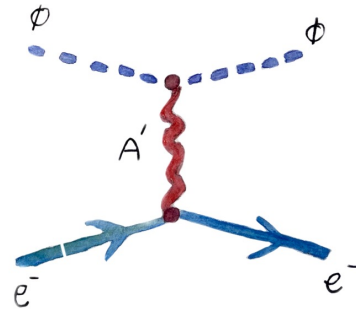


DM detection



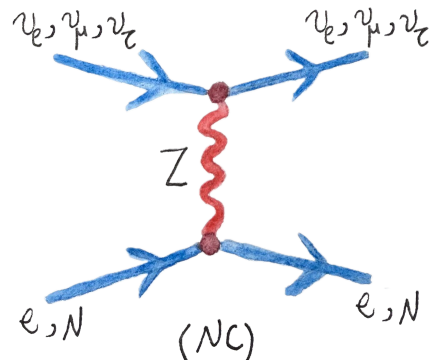
Light Dark Matter

DM signal: elastic scattering on electrons

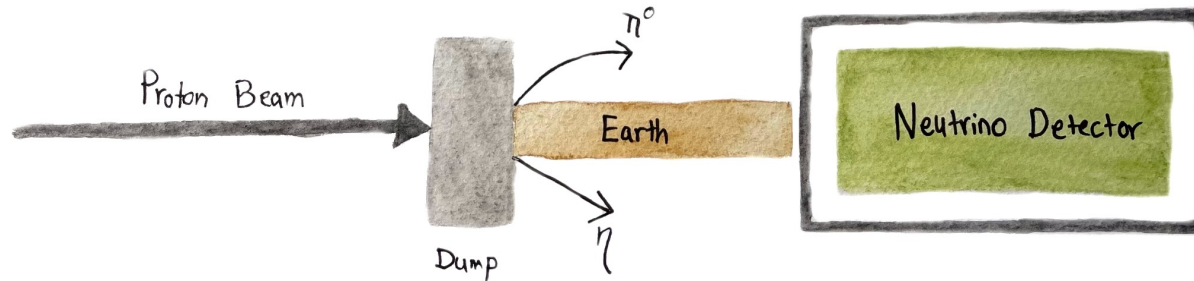


How can we get rid of neutrinos in a neutrino detector?

But so do neutrinos!



Proposing a movable target system at DUNE



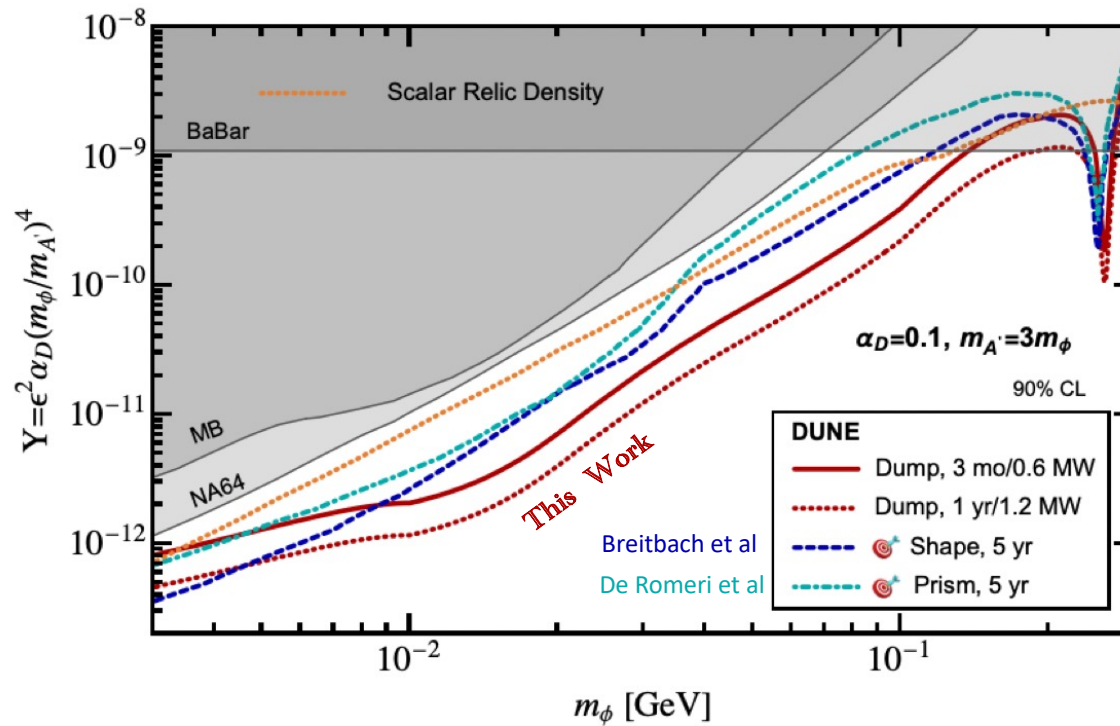
We can dump protons directly to the dump area!

Gains:

- Shorter distance between the source and the detector → more DM signal;
- Charged mesons absorbed in the Al beam dump before decay;
- The ν flux decreases → Much less ν background.

Brdar, Dutta, Jang, Kim, Shoemaker, ZT, Thompson, Yu
PRD (2023)

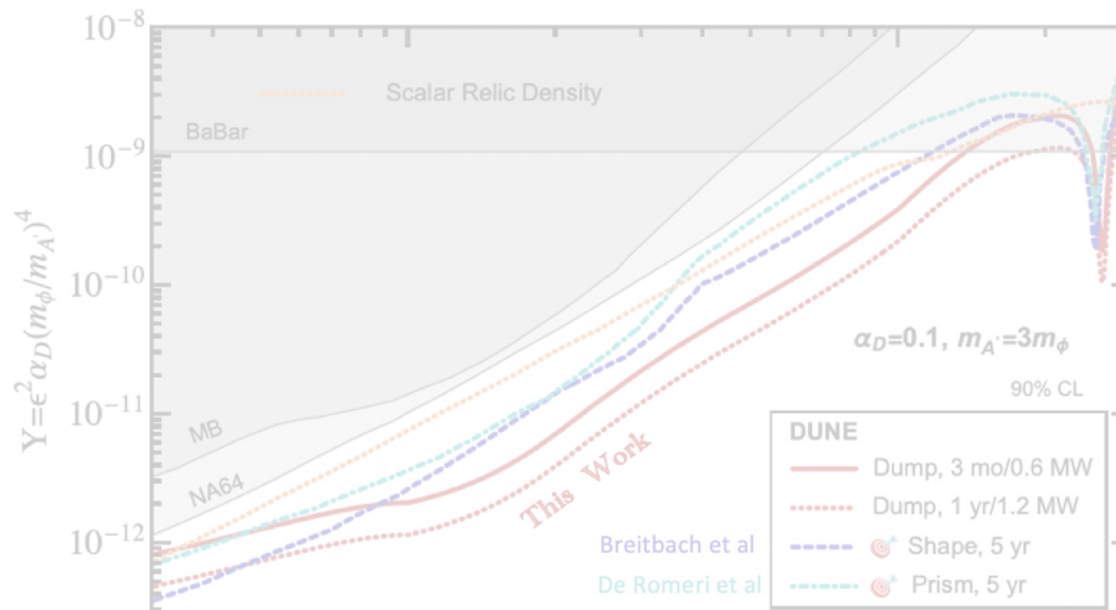
Light Dark Matter at Targetless DUNE



Brdar, Dutta, Jang, Kim, Shoemaker, **ZT**, Thompson, Yu
PRD (2023)

Target-less DUNE can probe the parameter space
for thermal relic DM in only 3 months!

Light Dark Matter at Targetless DUNE



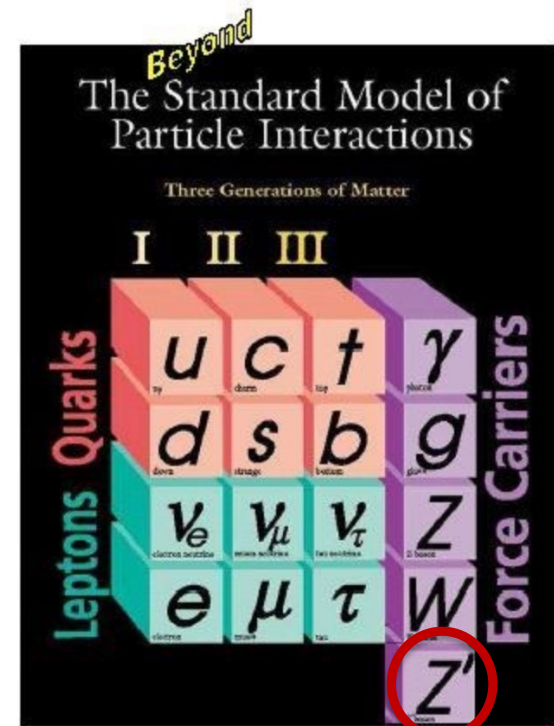
For DM searches, a neutrino detector could even do better than a dedicated DM experiment!

- 1) Direct Search of Dark Sectors:
 - ALPs ✓
 - Light Dark Matter ✓
 - Light Z'
- 2) Indirect Search-EFT:
 - Why EFT?
 - EFT ladder
 - EFT at Neutrino Experiments
- Conclusion



A new gauge boson?

Hypothetical gauge boson that appear in many extensions of the standard model



The Z' Hunter's Guide

- What is its mass?
- Which particles does it talk to?

Light Z'

- Low Energy Experiments

Miranda et al, JHEP (2020)
 Coloma et al, JHEP (2021)
 Cadeddu et al, JHEP (2021)

- Fixed Target Experiments

Gninenko, PLB (2012)
 Tsai et al, PRL (2021)
 Bauer et al, JHEP (2018)

- Neutrino Trident Searches

Altmannshofer et al, PRL (2014)
 Ballet et al, JHEP (2019)

- Neutrino-Electron Scattering

Harnic et al, JCAP (2012)
 Lindner et al, JHEP (2018)
 Ballet et al, JHEP (2019)

- Colliders

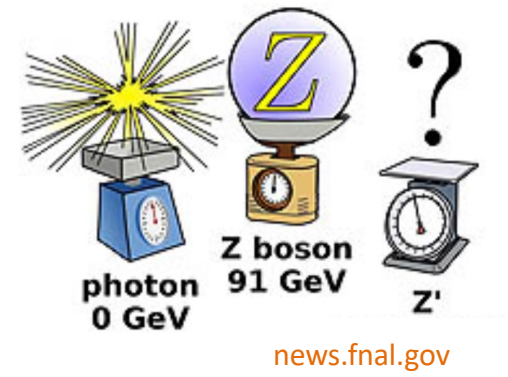
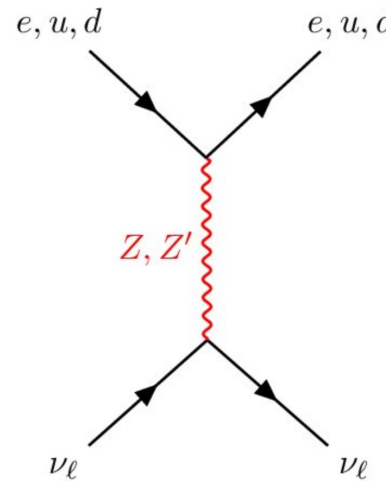
BaBar Collaboration, PRL (2014)
 BaBar Collaboration, PRL (2017)

- Cosmology

Escudero et al, JHEP (2019)

What can we learn from neutrino experiments?

$$\mathcal{L}_{Z'}^{\text{matter}} = -g' (a_u \bar{u} \gamma^\alpha u + a_d \bar{d} \gamma^\alpha d + a_e \bar{e} \gamma^\alpha e + b_e \bar{\nu}_e \gamma^\alpha P_L \nu_e + b_\mu \bar{\nu}_\mu \gamma^\alpha P_L \nu_\mu + b_\tau \bar{\nu}_\tau \gamma^\alpha P_L \nu_\tau) Z'_\alpha$$



The list is far from being exhaustive!

Light Z'

- Low Energy Experiments

Miranda et al, JHEP (2020)
 Coloma et al, JHEP (2021)
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- Fixed Target Experiments

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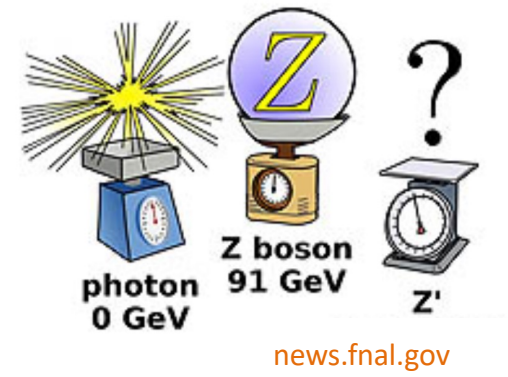
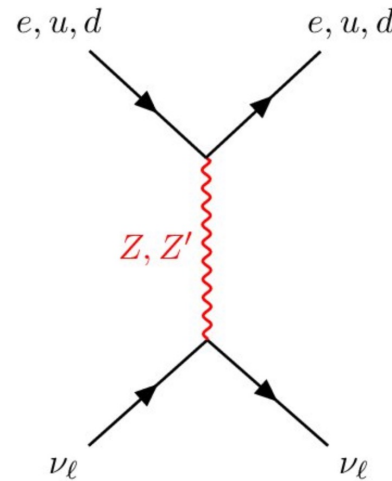
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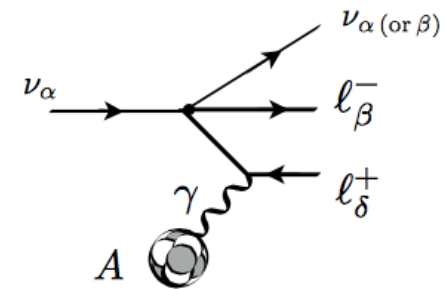


The list is far from being exhaustive!

Neutrino Trident Scattering

Production of a **charged lepton pair**
 in the scattering of a **neutrino**
 in the Coulomb field of a **heavy nucleus/nucleon**

$$\nu_\alpha + \mathcal{N} \rightarrow \nu_\beta + l_\gamma^+ + l_\delta^- + \mathcal{N}$$



CHARM II
 PLB (1990)

$$\frac{\sigma_{\text{CHARM II}}}{\sigma_{\text{SM}}} = 1.58 \pm 0.57$$

CCFR
 PRL (1991)

$$\frac{\sigma_{\text{CCFR}}}{\sigma_{\text{SM}}} = 0.82 \pm 0.28$$

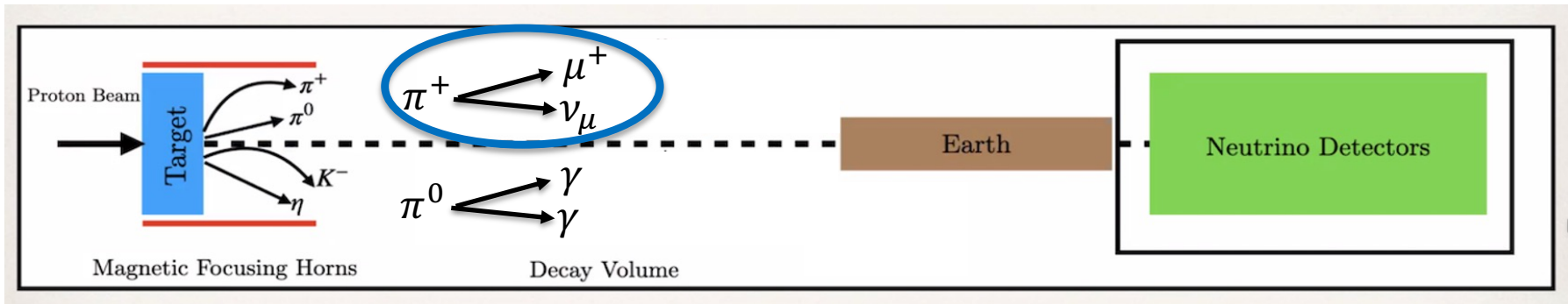
NuTeV

Vancouver 1998,
 High energy physics, vol 1

$$\frac{\sigma_{\text{NuTeV}}}{\sigma_{\text{SM}}} = 0.67 \pm 0.27$$

- Very large uncertainties
- Very few events observed (~100)

Trident rates at LAr Detectors



| Channel | SBND | μ BooNE | ICARUS | DUNE ND |
|-----------------------|------|-------------|--------|-------------|
| Total $e^\pm \mu^\mp$ | 10 | 0.7 | 1 | 2993 (2307) |
| | 2 | 0.1 | 0.2 | 692 (530) |
| Total $e^+ e^-$ | 6 | 0.4 | 0.7 | 1007 (800) |
| | 0.7 | 0.0 | 0.1 | 143 (111) |
| Total $\mu^+ \mu^-$ | 0.4 | 0.0 | 0.0 | 286 (210) |
| | 0.4 | 0.0 | 0.0 | 196 (147) |

Coherent (upper) and diffractive (lower) trident events for (anti)neutrino mode

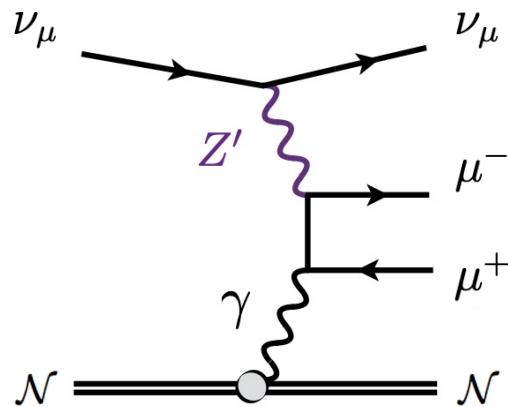
More than 9,000 trident events at

DUNE!

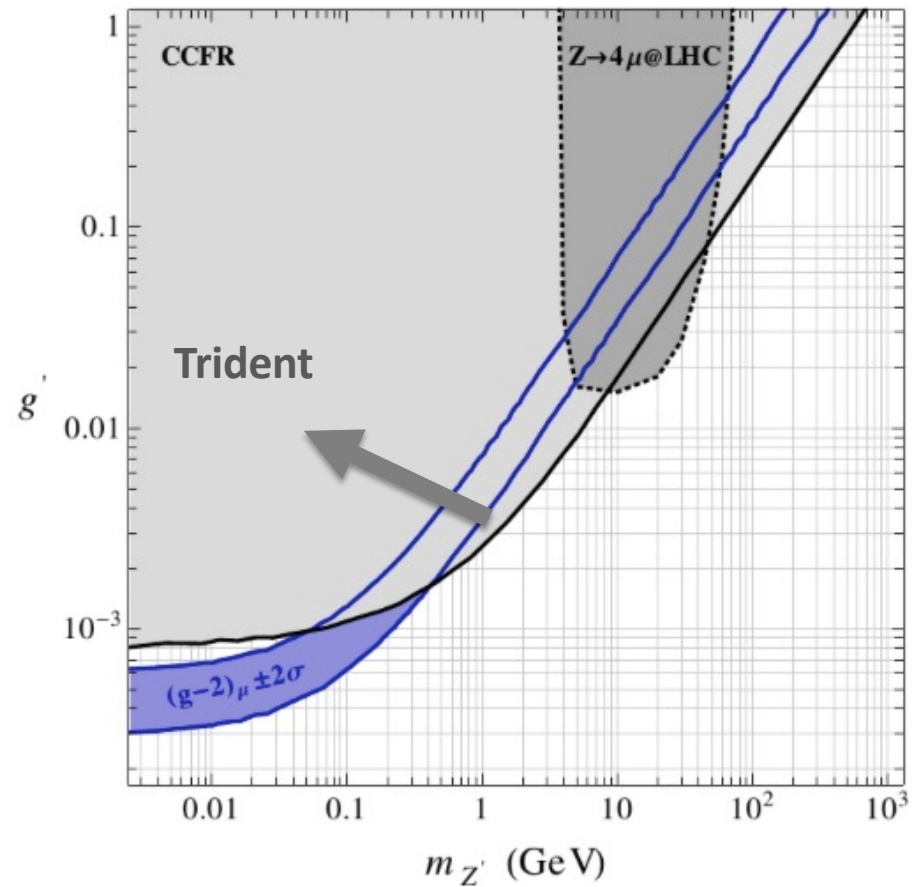
Ballett, Hostert, Pascoli, Perez-Gonzalez, ZT and Funchal, PRD (2019)

Light Z' : L_μ - L_τ Model

- Z' only couples to muon and tau, but not to electrons;
- It can explain the muon $(g-2)$ anomaly;
- Can be best probed using tridents;



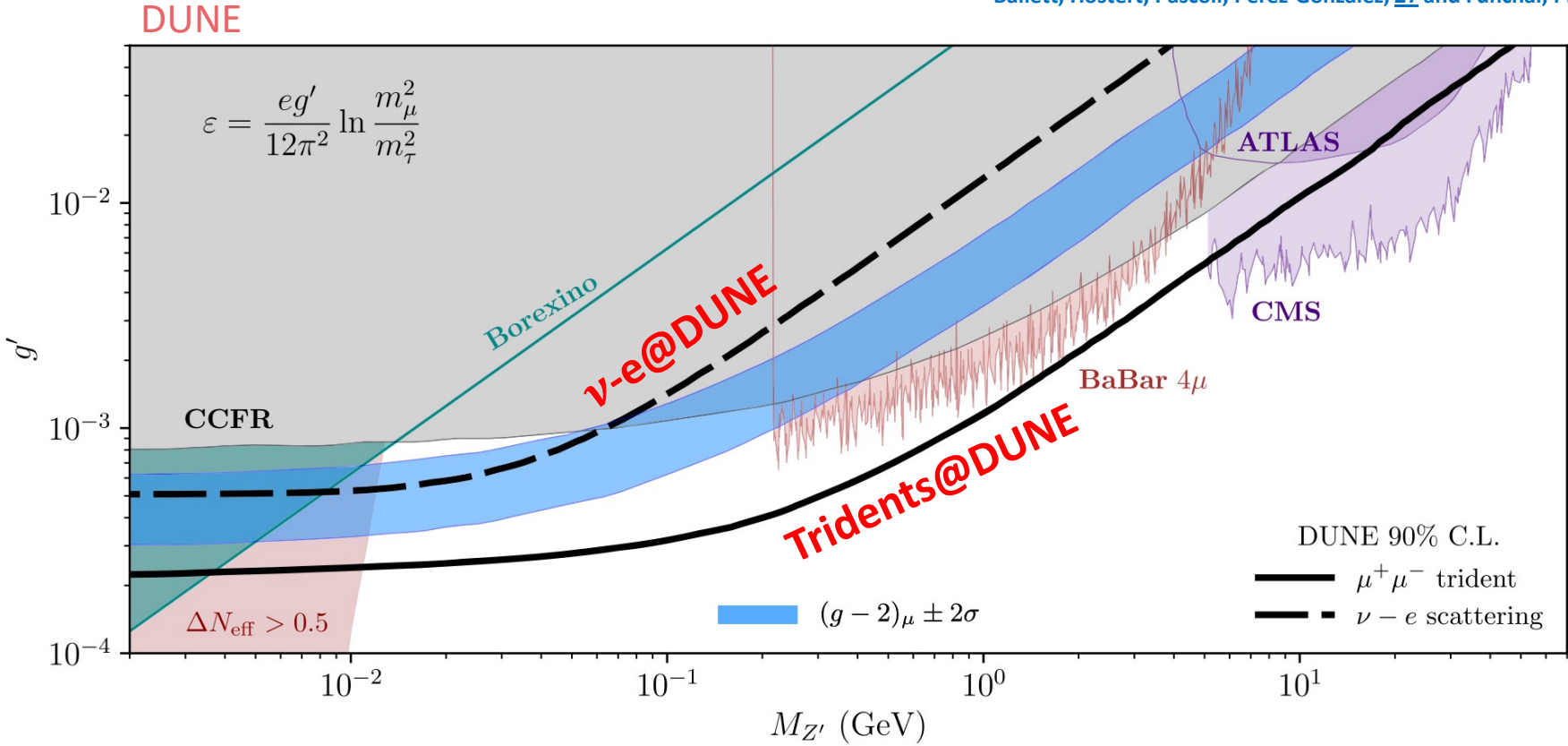
HE colliders only have access to large masses!



Altmannshofer, Gori, Pospelov, Yavin, PRL (2014)

Light Z' : L_μ - L_τ Model

Ballett, Hostert, Pascoli, Perez-Gonzalez, ZT and Funchal, PRD (2019)



The whole g -2 region can be probed by DUNE data!

EPA assumptions

- 1) Neglecting the L contribution ($h^L(q^2, \hat{s}) \sigma_{\nu\gamma}^L(q^2, \hat{s}) \approx 0$).
- 2) Taking the T contribution of the cross section to be on-shell ($\sigma_{\nu\gamma}^T(q^2, \hat{s}) \approx \sigma_{\nu\gamma}^T(0, \hat{s})$).

$$\sigma_t(P_i + C_s \rightarrow P_f + C_s) \approx \int dP(Q^2, \hat{s}) \sigma_\gamma(P_i + \gamma \rightarrow P_f; \hat{s}, Q^2 = 0)$$

QED

$$\sigma_\gamma^{\text{QED}}(P_i + \gamma \rightarrow P_f; \hat{s}, 0) \propto \frac{1}{\hat{s}}$$

Decreases with
increasing transferred
four-momentum

On-shell \gg off-shell

Fermi Limit of the SM

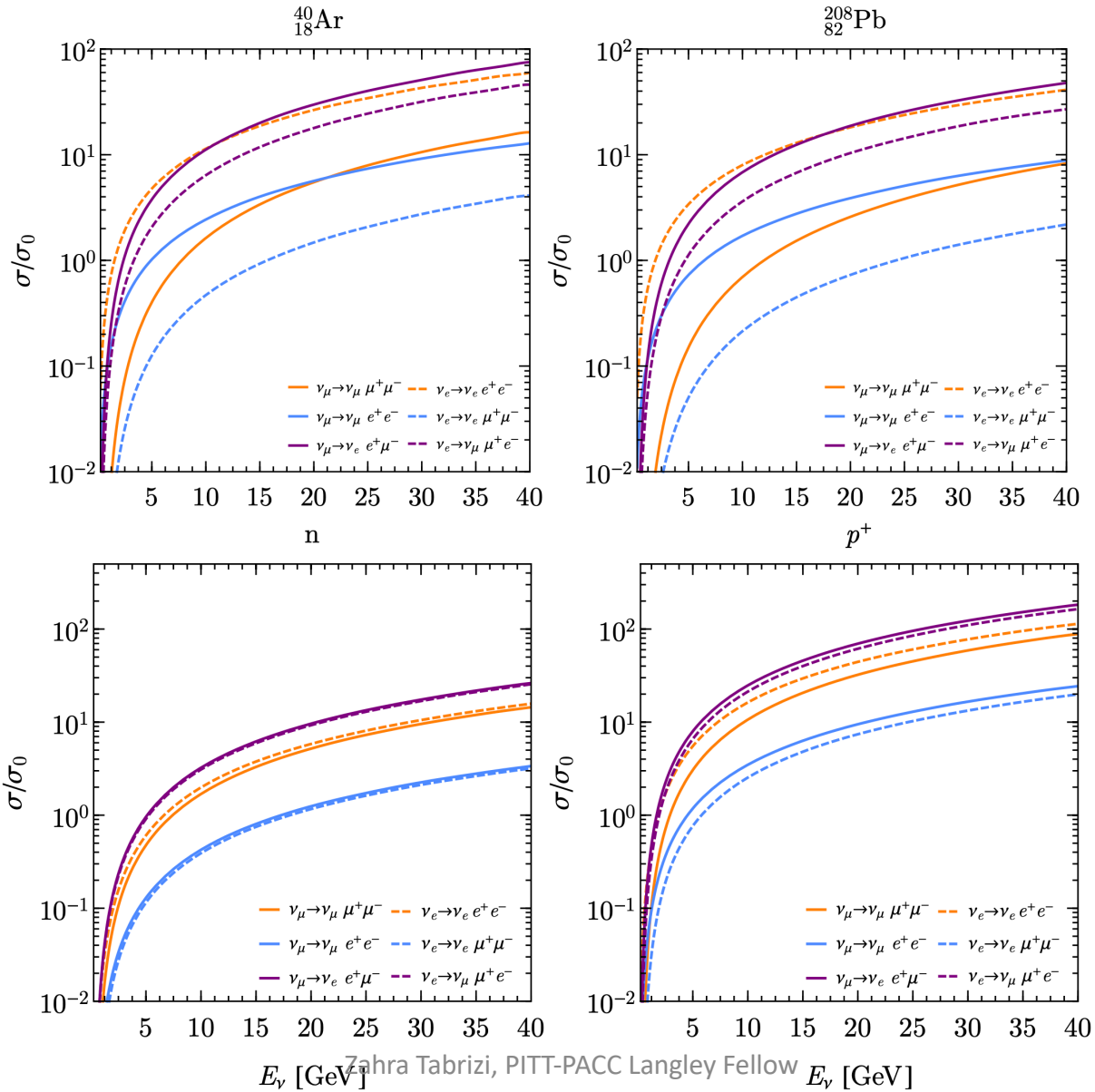
$$\sigma_\gamma^{\text{FL}}(P_i + \gamma \rightarrow P_f; \hat{s}, 0) \propto G_F^2 \hat{s}$$

Increases with
increasing transferred
four-momentum

On-shell \ll off-shell

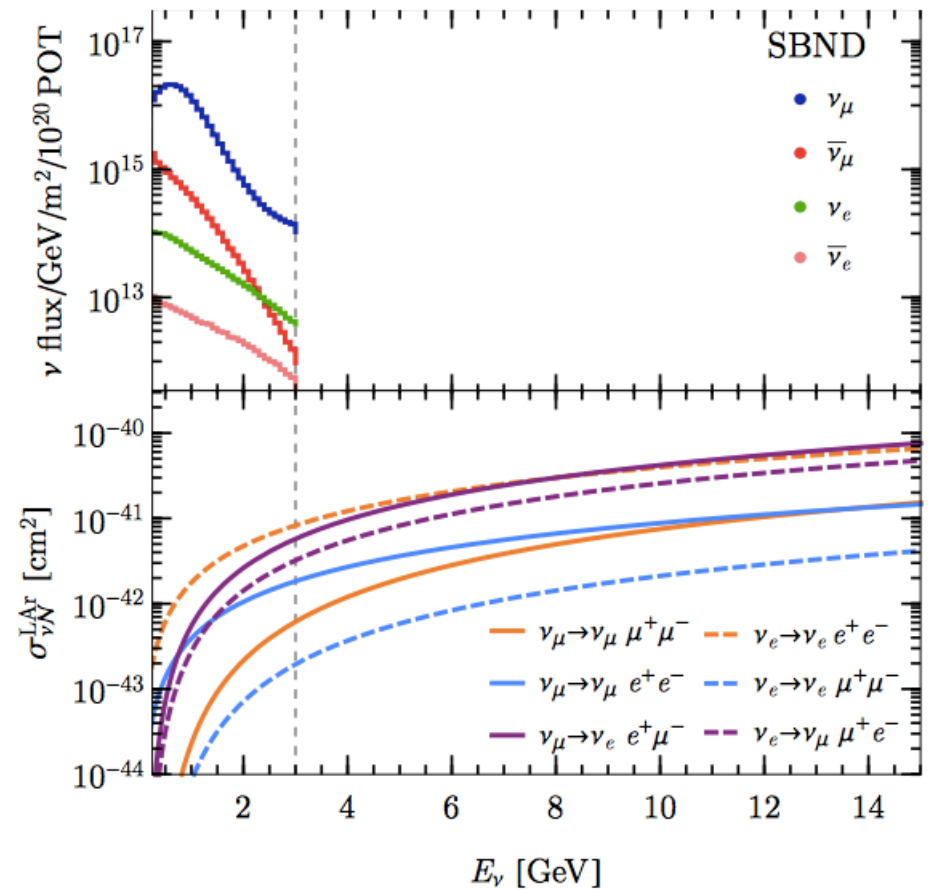
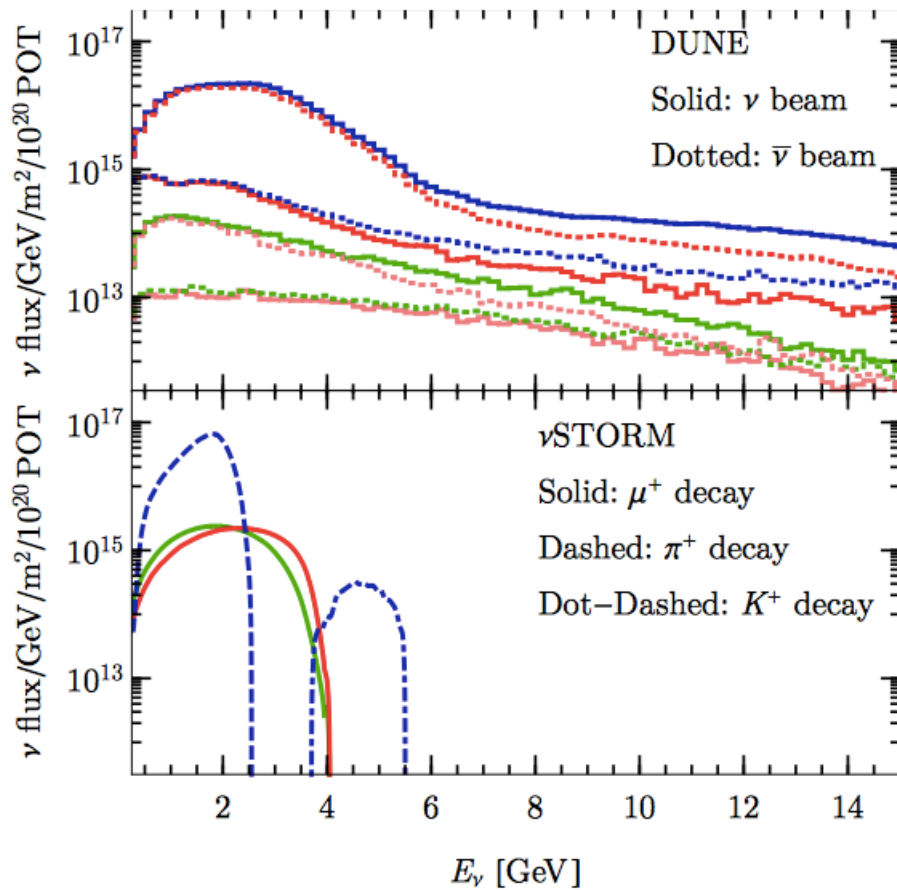
Ballett, Hostert, Pascoli, Perez, [ZT](#) and Funchal
JHEP **1901**, 119 (2019)

Trident Cross Sections



Trident rates at LAr Detectors

$$N = \text{time} \times \# \text{ of targets} \times \text{efficiency} \times \int_{E_i}^{E_f} dE_\nu \frac{d\phi(E_\nu)}{dE_\nu} \sigma(E_\nu)$$



Trident background analysis

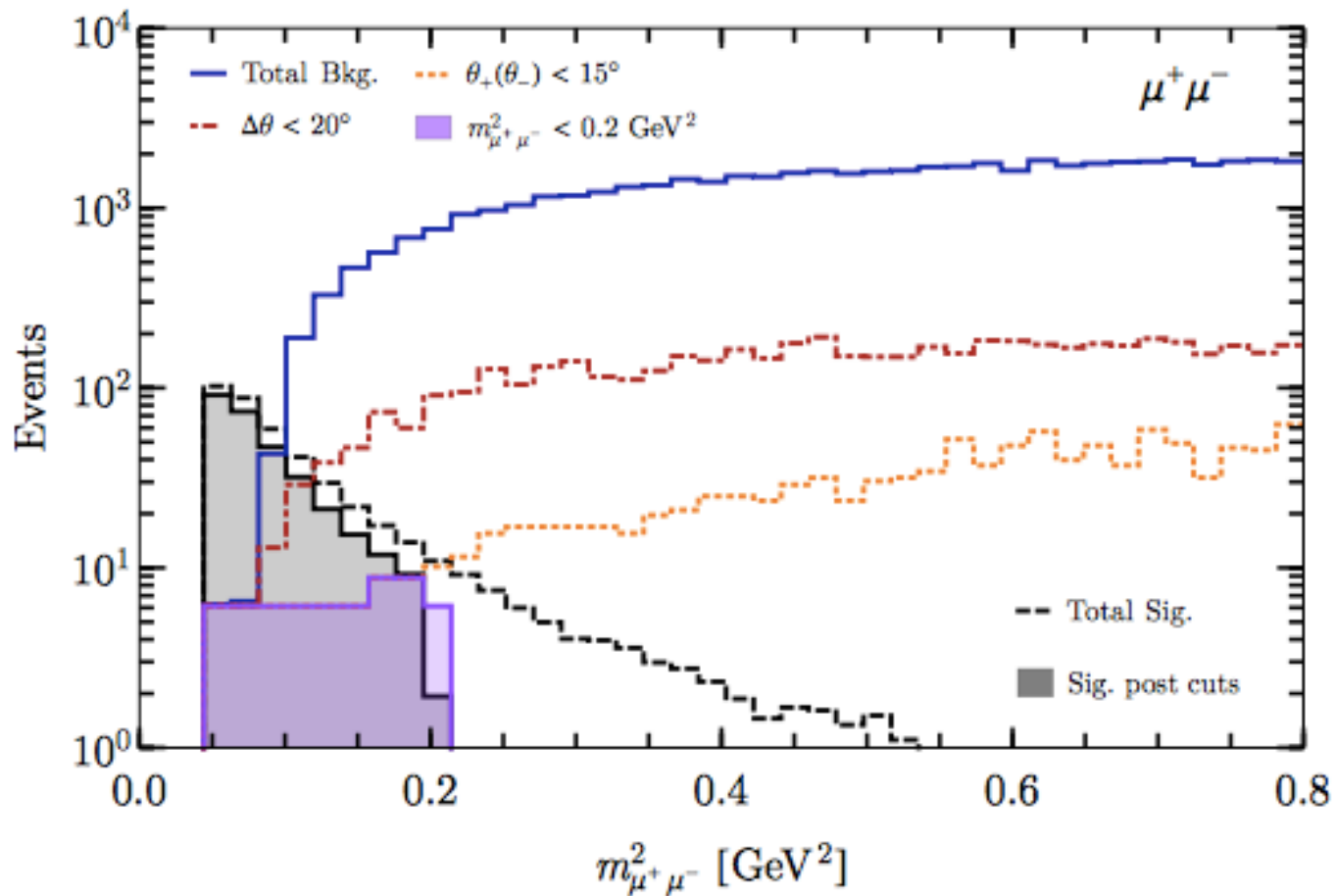
Genuine dilepton production is rare, but misID of particles is the problem.

| | Channel | $N_B^{\text{misID}}/N_{\text{CC}}$ | $N_B^{\text{had}}/N_{\text{CC}}$ | $N_B^{\text{kin}}/N_{\text{CC}}$ |
|------------------------|-----------------|------------------------------------|----------------------------------|----------------------------------|
| misID | $e^\pm \mu^\mp$ | $1.67 (1.62) \times 10^{-4}$ | $2.68 (4.31) \times 10^{-5}$ | $4.40 (3.17) \times 10^{-7}$ |
| γ as e^\pm | $e^+ e^-$ | $2.83 (4.19) \times 10^{-4}$ | $1.30 (2.41) \times 10^{-4}$ | $6.54 (14.1) \times 10^{-6}$ |
| γ as $e^+ e^-$ | $\mu^+ \mu^-$ | $2.66 (2.73) \times 10^{-3}$ | $10.4 (9.75) \times 10^{-4}$ | $3.36 (3.10) \times 10^{-8}$ |
| π^\pm as μ^\pm | | | | |

| | $N_{\text{tot}}^{\text{CC}}$ | $r_{\nu_\mu}^{\text{CC}}$ | $r_{\bar{\nu}_\mu}^{\text{CC}}$ | $r_{\nu_e}^{\text{CC}}$ | $r_{\bar{\nu}_e}^{\text{CC}}$ |
|-------------------|------------------------------|---------------------------|---------------------------------|-------------------------|-------------------------------|
| ν -mode | 4.25×10^8 | 0.964 | 0.028 | 0.007 | 0.001 |
| $\bar{\nu}$ -mode | 1.74×10^8 | 0.201 | 0.790 | 0.004 | 0.005 |
| | $N_{\text{tot}}^{\text{NC}}$ | $r_{\nu_\mu}^{\text{NC}}$ | $r_{\bar{\nu}_\mu}^{\text{NC}}$ | $r_{\nu_e}^{\text{NC}}$ | $r_{\bar{\nu}_e}^{\text{NC}}$ |
| ν -mode | 1.48×10^8 | 0.956 | 0.037 | 0.006 | 0.001 |
| $\bar{\nu}$ -mode | 7.58×10^7 | 0.157 | 0.835 | 0.003 | 0.005 |

Reaching background rates of $O(10^{-6}-10^{-5})$ times the CC rate is necessary to observe trident events at DUNE ND, which is an attainable goal in a LAr detectors.

Ballett, Hostert, Pascoli, Perez, [ZT](#) and Funchal
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We apply consecutive cuts on the background, starting with cuts on the separation angle $\Delta\theta$ (red), both charged lepton angles to the beamline (θ_+ and θ_-) (orange) and the invariant mass.

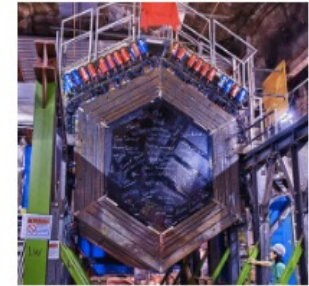
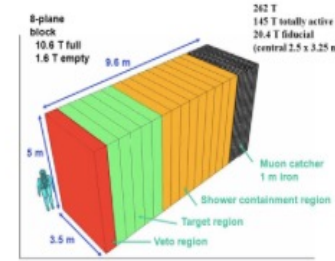
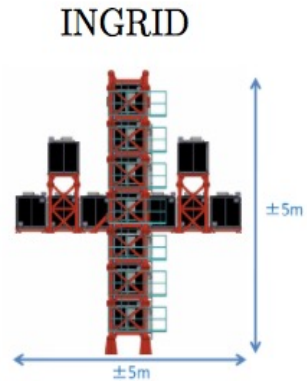
Ballett, Hostert, Pascoli, Perez, [ZT](#) and Funchal
 JHEP **1901**, 119 (2019)

Trident rates at other Near Detectors

| Experiment | Material | Baseline (m) | Exposure (POT) | Fiducial Mass (t) | E_ν (GeV) |
|---------------|--|--------------|---|-------------------|---------------|
| INGRID | Fe | 280 | 3.9×10^{21} [10^{22}] T2K-I [T2K-II] | 99.4 | 0 – 4 |
| MINOS[+] | Fe and C | 1040 | $10.56(3.36)[9.69] \times 10^{20}$ | 28.6 | 0 – 20 |
| NO ν A | C ₂ H ₃ Cl and CH ₂ | 1000 | $8.85(6.9) [36(36)] \times 10^{20}$ [NO ν A-II] | 231 | 0 – 20 |
| MINER ν A | CH, H ₂ O, Fe, Pb, C | 1035 | $12(12) \times 10^{20}$ | 7.98 | 0 – 20 |

All have finished data taking or are still running

Trident rates at other Near Detectors



| Channel | T2K-I | T2K-II | MINOS | MINOS+ | NO ν A-I | NO ν A-II | MINER ν A |
|--------------------------|-------|--------|----------|--------|--------------|---------------|---------------|
| Total $e^{\pm}\mu^{\mp}$ | 563 | 1444 | 222 (56) | 730 | 83 (72) | 340 (374) | 149 (102) |
| | 96 | 246 | 46 (11) | 151 | 25 (22) | 102 (114) | 56 (39) |
| Total $e^{+}e^{-}$ | 277 | 711 | 61 (15) | 62 | 29 (22) | 119 (114) | 39 (27) |
| | 24 | 62 | 9 (2) | 8 | 4 (4) | 16 (21) | 10 (7) |
| Total $\mu^{+}\mu^{-}$ | 30 | 76 | 26 (6) | 86 | 9 (9) | 37 (47) | 18 (13) |
| | 21 | 54 | 15 (3) | 49 | 8 (8) | 34 (36) | 18 (13) |

Coherent (upper) and diffractive (lower) trident events for (anti)neutrino mode.

Ballett, Hostert, Pascoli, Perez, **ZI** and Funchal
JHEP **1901**, 119 (2019)

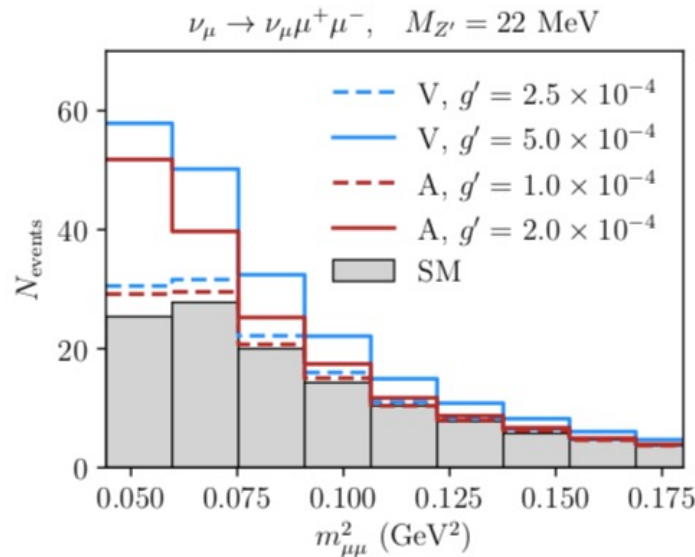
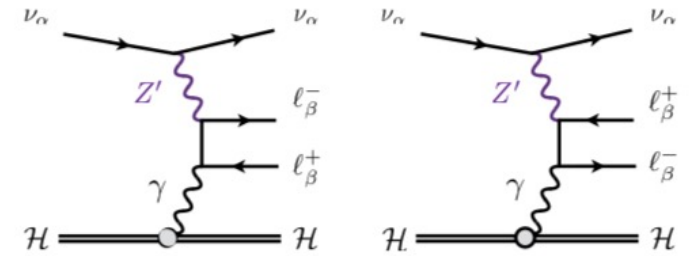
- We study potential constraints which can be placed on a general set of leptophilic Z' models in the two most likely channels for BSM scattering at the near detector of DUNE: neutrino-electron scattering and neutrino trident scattering.

$$\mathcal{L} \supset -g' Z'_\mu \left[Q_\alpha^L \bar{L}_L^\alpha \gamma^\mu L_L^\alpha + Q_\alpha^R \bar{\ell}_R^\alpha \gamma^\mu \ell_R^\alpha + \sum_N Q_N \bar{N}_R \gamma^\mu N_R \right]$$

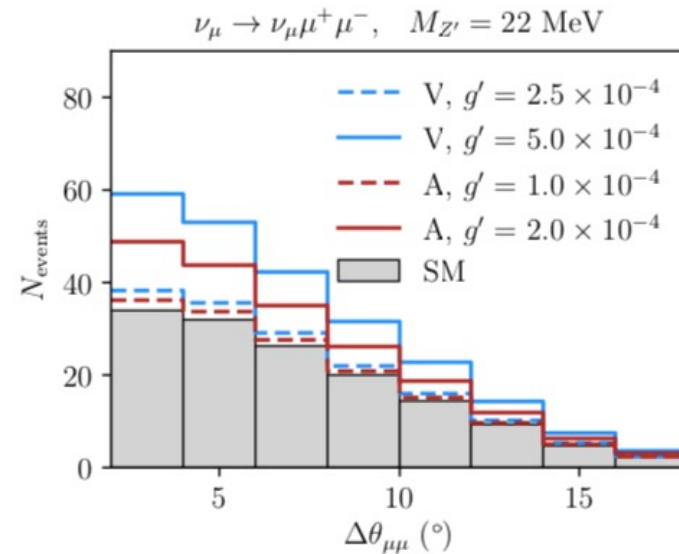
- We focus on the anomaly free leptophilic extensions of the SM:
 $L_\alpha - L_\beta$, $\alpha, \beta = \{e, \mu, \tau\}$, $\alpha \neq \beta$.
- Anomaly free conditions fix the charges

Trident kinematical distributions

$$\hat{V}_{\alpha\beta} = g_V^{\ell\beta} + \delta_{\alpha\beta} + \frac{Q_\alpha^L Q_\beta^V}{2\sqrt{2}G_F} \frac{(g')^2}{K^2 + M_{Z'}^2}, \quad \hat{A}_{\alpha\beta} = g_A^{\ell\beta} + \delta_{\alpha\beta} + \frac{Q_\alpha^L Q_\beta^A}{2\sqrt{2}G_F} \frac{(g')^2}{K^2 + M_{Z'}^2}$$



The invariant mass



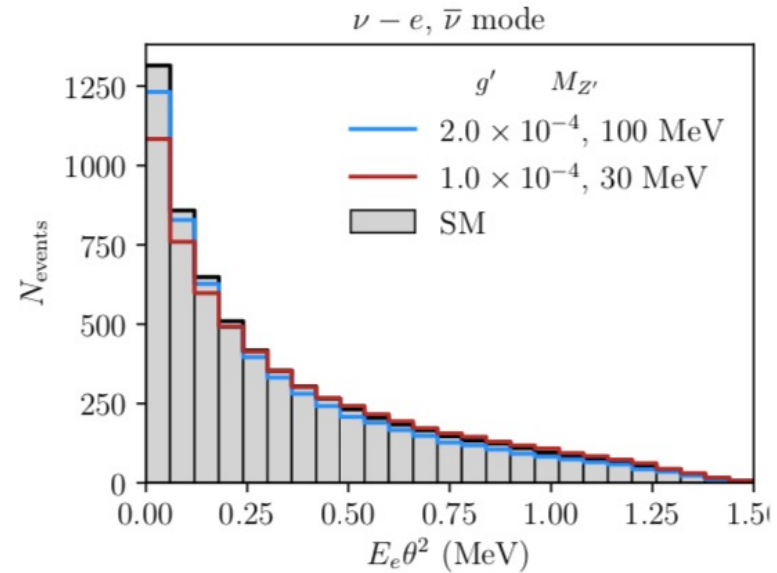
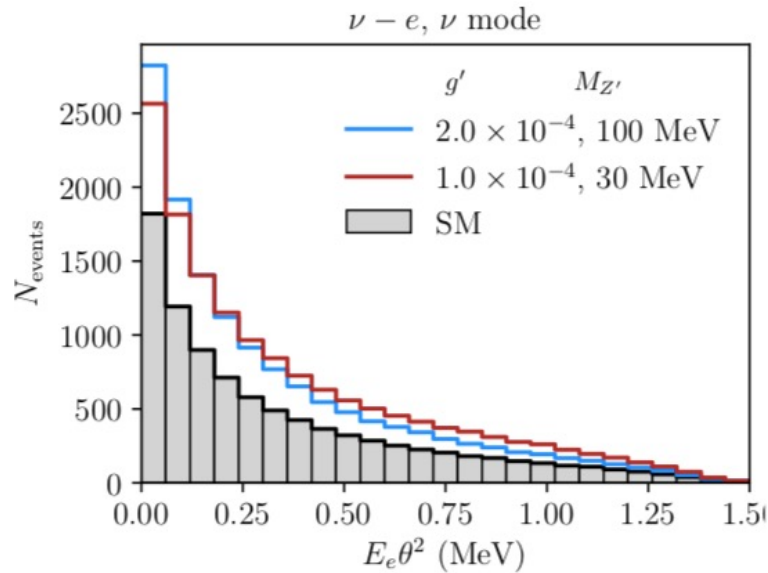
Charged lepton separation angle

Neutrino-Electron scattering

The vector and axial couplings with Z' :

$$C_{\alpha}^V = -\frac{1}{2} + 2s_W^2 + \delta_{\alpha e} + \frac{Q_e^V Q_{\alpha}^L}{2\sqrt{2}G_F} \frac{(g')^2}{M_{Z'}^2 + 2m_e T_e},$$

$$C_{\alpha}^A = -\frac{1}{2} + \delta_{\alpha e} + \frac{Q_e^A Q_{\alpha}^L}{2\sqrt{2}G_F} \frac{(g')^2}{M_{Z'}^2 + 2m_e T_e},$$



Neutrino-Electron scattering

The vector and axial couplings with Z' :

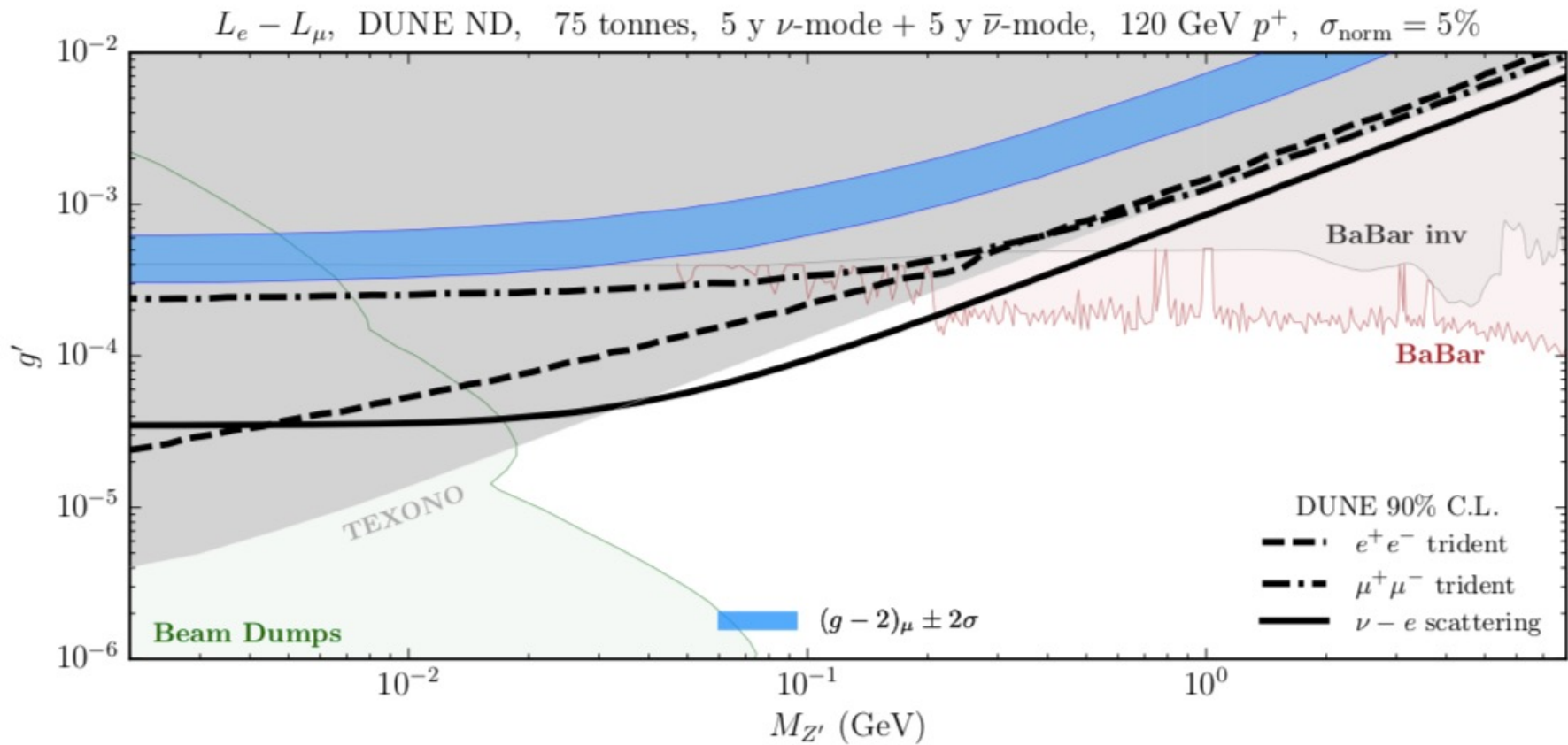
$$C_{\alpha}^V = -\frac{1}{2} + 2s_W^2 + \delta_{\alpha e} + \frac{Q_e^V Q_{\alpha}^L}{2\sqrt{2}G_F} \frac{(g')^2}{M_{Z'}^2 + 2m_e T_e},$$
$$C_{\alpha}^A = -\frac{1}{2} + \delta_{\alpha e} + \frac{Q_e^A Q_{\alpha}^L}{2\sqrt{2}G_F} \frac{(g')^2}{M_{Z'}^2 + 2m_e T_e},$$

If we allow for kinetic mixing between the Z' and the SM gauge bosons, this kinetic mixing in the $L_{\mu} - L_{\tau}$ model also induces a $\nu - e$ coupling:

$$C_{\alpha}^V = -\frac{1}{2} + 2s_W^2 + \delta_{\alpha e} + \frac{1}{\sqrt{2}G_F} \frac{g' e \varepsilon(q^2)}{M_{Z'}^2 + 2m_e T_e}$$

$L_e - L_\mu$ Model at DUNE:

Ballett, Hostert, Pascoli, Perez-Gonzalez, [ZI](#) and Funchal
 Phys.Rev. **D100** (2019) no.5, 055012



- The main constraint is from neutrino-electron scattering.
- The sensitive trident channels are:
 $\mu^+\mu^-$ and e^+e^-