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NATURALNESS IN THE SM AND BEYOND WARSAW, APRIL 8TH, 2014

Mariano Quirós

CERN/ICREA/IFAE

Mariano Quirós (CERN/ICREA/IFAE) NATURALNESS IN THE SM AND BEYOND

OUTLINE

- Naturalness (or the natural solution to the hierarchy problem) has been one of the driving forces for BSM for the last 40 years or so
- So it is time to stop for a moment and gently think on it 1
- The outline

• INTRODUCTION

- NATURALNESS IN AN EFFECTIVE THEORY
- NATURALNESS IN A FUNDAMENTAL THEORY
- NATURALNESS IN BSM
- NATURALNESS IN THE MSSM
- CONCLUSION

¹This talk is partly based on a chat with Stefan Pokorski at

THE CERN CAFETERIA

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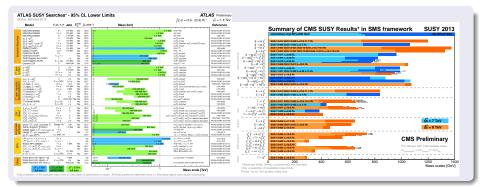
Introduction

INTRODUCTION

The Higgs has been found at LHC with a mass

 $m_H = 125.9 \pm 0.4 ~GeV$

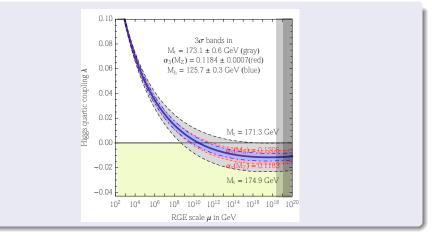
and approximately SM-couplingsNo track of new physics found



Mariano Quirós (CERN/ICREA/IFAE)

Introduction

• The SM is consistent at high scales ² depending on m_t and m_H



• But

IS THE EW SCALE NATURAL?

²D. Buttazzo et al. arXiv:1307.3536

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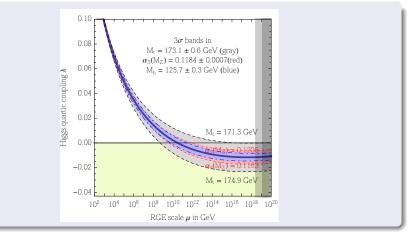
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- An effective theory is one with a physical cutoff Λ
- There are "(finite) quadratic divergences" in the CW effective potential as

$$rac{1}{32\pi^2}rac{\partial Str\,\mathcal{M}^2}{\partial\phi^2}\Lambda^2$$

In particular if the SM is an effective theory

Tree-level potential

$$V(H) = -m^2|H|^2 + \frac{\lambda}{2}|H|^4$$

$$\Delta m^2 = -\frac{3}{32\pi^2 v^2} \left(m_H^2 + 2m_W^2 + m_Z^2 - 4m_t^2 \right) \Lambda^2$$

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- An effective theory is the low energy theory of a more fundamental theory which UV completes it at scales larger than Λ
- In an effective theory the cutoff is not a mathematical device which regularizes the theory in the limit $\Lambda\to\infty$
- In this sense unnatural theories are not related to theories with quadratic divergences

In general naturalness problems should not be related to regularization prescriptions

 In particular using dimensional regularisation there are not quadratic divergences (in d = 4)

In the UV completion "quadratic divergences" appear as FINITE corrections proportional to the heavy masses or heavy scales of the UV theory

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- In the absence of a knowledge of the quantum theory of gravity it is difficult to tell that a given theory is a fundamental one
- However if we DECOUPLE GRAVITY FROM THE GAME we can assume that some field theories are fundamental

For instance for the moment (except for the Landau pole in the hypercharge coupling at trans-Planckian energies) we can assume that the SM is a fundamental theory

• In fact, the non-finding of new physics by LHC

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When is there a naturalness problem in a (fundamental) theory?

• When there are light (e.g. Higgs in the SM sector) and heavy (e.g. in GUT, gravitational sectors) scalars in the theory with hierarchically different (square) masses

$$v^2 \sim m^2 \ll M^2$$

and with coupling g, which are connected by radiative corrections

$$|\Delta m^2| = \mathcal{O}\left(\frac{g^2}{16\pi^2}\right)M^2 \gg m^2$$

 When there is a heavy dynamical UV scale in the theory, where a non-perturbative CFT takes over, even if created by dimensional transmutation and not corresponding to the mass of heavy states

^aMarques-Schmaltz-Skiba, arXiv:1308.0025

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Naturalness problem \Rightarrow fine-tuning problem

If $|\Delta m^2| \gg v^2$ then we have to tune $|m^2| \sim |\Delta m^2| \gg v^2$ such that

$$m^2 + \Delta m^2 = \mathcal{O}(v^2)$$

This fine-tuning IS NOT natural

Tuning \Rightarrow (automatically) naturalness problem

If m^2 , $\Delta m^2 = \mathcal{O}(v^2)$ the tuning in

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from experimental error in v^2 , e.g. using

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i) Simplest example: SM+coupled heavy scalar

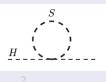
Typical example is the SM Higgs (H) in the presence of a heavy complex scalar (S) as

$$V(H,S) = -m^{2}H^{2} + \frac{\lambda}{2}H^{4} + M^{2}|S|^{2} + g^{2}H^{2}|S|^{2}$$

At tree level

$$v^2 = \frac{m^2}{\lambda}$$

• At one loop, from the one-loop tadpole diagram in \overline{MS} at $\mathcal{Q} = M$



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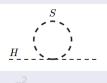
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$$v^{2} = \frac{m^{2} + \Delta m^{2}}{\lambda}, \ \Delta m^{2} = \frac{g^{2}}{16\pi^{2}}M^{2}, \ \text{NATURALNESS} \Rightarrow g \lesssim 4\pi m_{H}/M$$
into Quirós (CERN/ICREA/IFAE) NATURALNESS IN THE SM AND BEYOND 10/35

ii) $\mathcal{L} = \mathcal{L}_{SM} + h_{\nu}H\bar{\ell}_{L}\nu_{R} + M\nu_{R}^{T}\nu_{R}$

- At Q > M propagating ν_R
- At $Q < M \nu_R$ integrated out
- At $\mathcal{Q} = M$ the integration of ν_R leaves threshold effect such that

$$\Delta m^2 = -\frac{N_\nu h_\nu^2}{4\pi^2} M^2$$

 \bullet Using a degenerate spectrum $m_{\nu}\simeq 0.05~{\rm eV}$

NATURALNESS
$$\Rightarrow M \lesssim 5 \times 10^6 \text{ GeV}$$

SM+right-handed neutrinos with masses $\sim 10^6~{\rm GeV}$ AND NOTHING ELSE AT ALL

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Second example: $N_{
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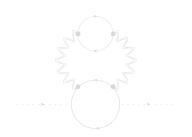
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Third example: massive matter uncoupled to the Higgs

iii) $\mathcal{L} = \mathcal{L}_{SM} + M \bar{\Psi} \Psi$

No loop corrections from CW in the system SM + Ψ
 While the full quantum theory of gravity is unknown, low energy gravitational interactions can mediate corrections to m²



• The main contribution comes from the three-loop diagram ^a

$$\Delta m^2 \sim {h_t^2 \over (16\pi^2)^3} {M^6 \over M_P^4}$$

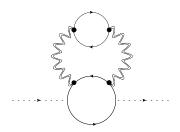
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^aA. de Gouvea and D. Hernndez, arXiv:1402.2658 Third example: massive matter uncoupled to the Higgs

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• NATURALNESS $\Rightarrow M \lesssim 10^{14} {
m GeV}$

^aA. de Gouvea and D. Hernndez, arXiv:1402.2658

Fourth example: MSSM ³

iv) MSSM: $m_0 = m_{1/2} = \mu \equiv M$

- For Q > M: MSSM
- For *Q* < *M*: SM
- At Q = M matching $SM \equiv MSSM$ and using the SM potential

$$V = -m^2 |H|^2 + \frac{\lambda}{2} |H|^4$$

$$\Delta m^2 = -\frac{1}{32\pi^2} \left(6\lambda + 3/2g_1^2 + 9/2g_2^2 - 12h_t^2 \right) M^2$$

It corresponds to the Veltman condition for $\Lambda = M$

$$\Delta m^2 = -\frac{3}{32\pi^2 v^2} \left(m_H^2 + 2m_W^2 + m_Z^2 - 4m_t^2 \right) \Lambda^2$$

for the SM with cut-off $\Lambda = M$

³G. Nardini, I. Masina, MQ, in progress

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³G. Nardini, I. Masina, MQ, in progress Mariano Quirós (CERN/ICREA/IFAE) NATURALNESS IN THE SM AND BEYOND

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If we believe there is BSM physics at the Planck (or GUT) scale, then
In the absence of a concrete (fundamental) UV completion describing the gravity degrees of freedom

The theory below M_P is an effective theory

- Several possibilities for the theory after imposing the requirement of naturalness
- Either the Higgs is composite, for scales $\Lambda \gtrsim \Lambda_{comp}$ the Higgs dissolves into its constituents and there is a phase transition at $\Lambda \simeq \Lambda_{comp}$ to the theory of constituents
- Or the Higgs is fundamental in which case there should exist a symmetry such that

$$\frac{\partial Str\mathcal{M}^2}{\partial \phi^2} = 0$$

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The paradigm of these theories is $\ensuremath{\mathsf{SUSY}}$

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For a composite Higgs

 The scale Λ_{comp} is dynamical and naturalness requires

 $\Lambda_{comp} \lesssim \mathcal{O} \ (\, \text{TeV})$

- Prototype theories are technicolor or the modern warped dimensional (RS) models based on the AdS/CFT duality
- **PROS**: Mechanism already used by nature: **QCD**
- CONS: Theory non-perturbative beyond Λ_{comp}; UV completion not known; difficulties in encompassing EWPD,...

For a fundamental Higgs

 SUSY theory soft breaking terms (e.g. squark masses) m_f should be

- Prototype theory of SUSY is the minimal supersymmetric extension of the SM (MSSM)
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• The scale Λ_{comp} is dynamical	• SUSY theory soft breaking
and naturalness requires	terms (e.g. squark masses) $m_{\tilde{f}}$
$\Lambda_{comp} \lesssim \mathcal{O} \; (\mathit{TeV})$	should be
 Prototype theories are 	$m_{\widetilde{f}} \lesssim \mathcal{O} \; ({\it TeV})$
technicolor or the modern	Prototype theory of SUSY is
warped dimensional (RS)	the minimal supersymmetric
models based on the AdS/CFT	extension of the SM (MSSM)
duality	• PROS : theory perturbative up
• PROS : Mechanism already used	to M_P ; gauge coupling
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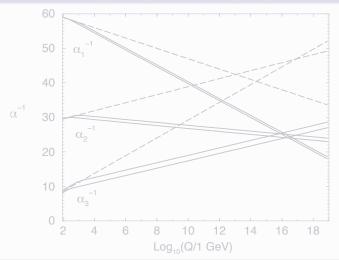
For a composite Higgs	For a fundamental Higgs		
 The scale Λ_{comp} is dynamical and naturalness requires 	 SUSY theory soft breaking terms (e.g. squark masses) m_f 		
$\Lambda_{comp} \lesssim \mathcal{O} \; (\textit{TeV})$	should be $m_{ ilde{t}} \lesssim \mathcal{O} \; (TeV)$		
 Prototype theories are 	$m_f \gtrsim O(1ev)$		
technicolor or the modern	 Prototype theory of SUSY is 		
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mental Higgs
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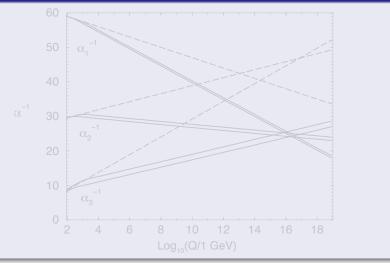
For a composite Higgs	For a fundamental Higgs
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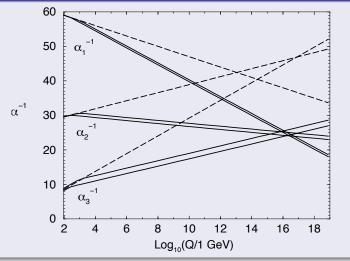
Gauge unification for SM (dashed) and MSSM (solid)



Gauge unification for SM (dashed) and MSSM (solid)







BUT...

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: Moriond 2014

	Model	e, μ, τ, γ	Jets	$E_{\rm T}^{\rm miss}$	∫£ dt[fb	Mass limit	Reference
Inclusive Searches	MSUGRACMSSM MSUGRACMSSM MSUGRACMSSM 49, 8 - 494 ² , 1 88, 8 - 494 ² , 1 88, 8 - 494 ² , 1 88, 8 - 494 ² , 1 60, 4 - 494 ² , 1 GMSB (1 NLSP) GMM (high Shi Shi Shi Shi Shi Shi Shi Shi Shi Sh	$\begin{matrix} 0 \\ 1 & c, \mu \\ 0 \\ 0 \\ 1 & c, \mu \\ 2 & c, \mu \\ 2 & c, \mu \\ 2 & c, \mu \\ 1 & c, \mu + \gamma \\ \gamma \\ 1 & c, \mu + \gamma \\ \gamma \\ 2 & c, \mu (Z) \\ 0 \end{matrix}$	2-6 jets 3-6 jets 2-6 jets 2-6 jets 2-6 jets 0-3 jets 2-4 jets 0-2 jets - 1 b 0-3 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 4.7 20.7 20.7 20.3 4.8 4.8 5.8 10.5	All 17 TeV (μ) (μ(μ)) (μ) (μ(μ)) 8 1.3 TeV (μ) (μ) (μ) (μ) (μ) (μ) ωμ (μ) (μ) (μ) (μ) 8 1.3 TeV (μ) (μ) (μ) (μ) (μ) (μ) (μ) ωμ (μ) (μ) (μ) (μ) (μ) (μ) 8 1.3 TeV (μ) (μ) (μ) (μ) (μ) (μ) (μ) (μ) (μ) ωμ (μ) (μ) (μ) (μ) (μ) (μ) (μ) (μ) 8 1.3 TeV (μ) (μ) (μ) (μ) (μ) (μ) (μ) (μ) (μ) (μ)	ATLAS-CONF-2013-047 ATLAS-CONF-2013-042 1300.1041 ATLAS-CONF-2013-047 ATLAS-CONF-2013-047 ATLAS-CONF-2013-049 ATLAS-CONF-2013-049 ATLAS-CONF-2013-049 ATLAS-CONF-2013-049 ATLAS-CONF-2013-049 ATLAS-CONF-2012-047 ATLAS-CONF-2012-047 ATLAS-CONF-2012-047 ATLAS-CONF-2012-047 ATLAS-CONF-2012-047
3 rd gen. § med.	$\overline{\hat{x}} \rightarrow bb \overline{\hat{x}}_{1}^{0}$ $\overline{\hat{x}} \rightarrow t \overline{t} \overline{t}_{1}^{0}$ $\overline{\hat{x}} \rightarrow t \overline{t} \overline{t}_{1}^{0}$ $\overline{\hat{x}} \rightarrow b \overline{t} \overline{t}_{1}^{0}$	0 0 0-1 e,μ 0-1 e,μ	3 b 7-10 jets 3 b 3 b	Yes Yes Yes Yes	20.1 20.3 20.1 20.1	8 1.72 TeV m(7)-score/v 8 1.1 TeV m(7)-score/v 2 1.34 TeV m(7)-score/v 8 1.34 TeV m(7)-score/v	ATLAS-CONF-2013-061 1908.1841 ATLAS-CONF-2013-061 ATLAS-CONF-2013-061
3rd gen. squarks direct production	$ \begin{split} & \frac{1}{b} [i, b_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ b_{1} b_{1} b_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ b_{1} b_{1} b_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (l(pht), \overline{i}_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (l(pht), \overline{i}_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (modlum), \overline{i}_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (modlum), \overline{i}_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (l(mosty), \overline{i}_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (l(mosty), \overline{i}_{1} \rightarrow b \mathcal{K}_{1}^{2} \\ & \overline{i}_{1} (\overline{i}, b n \mathcal{K}_{2}) \\ & \overline{i}_{1} (\overline{i}, \overline{i}, n \mathcal{K}_{2}) \\ & \overline{i}_{1} (\overline{i}, n \mathcal{K}$	$\begin{matrix} 0\\ 2 \ e, \mu \ (\mathrm{SS})\\ 1.2 \ e, \mu\\ 2 \ e, \mu\\ 2 \ e, \mu\\ 0\\ 1 \ e, \mu\\ 0\\ 1 \ e, \mu\\ 0\\ 3 \ e, \mu \ (Z)\end{matrix}$	2 b 0-3 b 1-2 b 0-2 jets 2 b 1 b 2 b 1 b 2 b 1 b 1 b 1 b 1 b	Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.1 20.7 4.7 20.3 20.3 20.1 20.7 20.5 20.3 20.3 20.3 20.3	5, 109433 GeV (m[],-α-GeV (m[],-α-GeV 5, 105475 GeV (m[],-α-Π]) 7, 105415 GeV (m[],-α-Π]) 7, 105415 GeV (m[],-α-Π]) 7, 105405 GeV (m]),-α-Π] 7, 105405 GeV (m]),-α-Π] 7, 105405 GeV (m]),-α-Π],-α-Π] 7, 105405 GeV (m]),-α-Π],-α-Π],-α-Π] 7, 105405 GeV (m]),-α-Π],-α-Π],-α-Π] 7, 105405 GeV (m]),-α-Π],	1308.2831 ATLAS-CONF-2013-007 1208.4305; 1209.2102 1403.4853 1403.4853 1308.2631 ATLAS-CONF-2013-024 ATLAS-CONF-2013-024 ATLAS-CONF-2013-028 1403.5222 1403.5222
EW direct	$\begin{array}{l} \overline{l}_{L,R}\overline{l}_{L,R},\overline{\ell}\rightarrow\delta\overline{\ell}_{1}^{\Omega}\\ \overline{\lambda}_{1}^{*}[\overline{\lambda}_{1}^{*},\overline{\lambda}_{1}^{*}\rightarrow\delta\eta\langle F\rangle\\ \overline{\lambda}_{1}^{*}[\overline{\lambda}_{1}^{*},\overline{\lambda}_{1}^{*}\rightarrow\delta\eta\langle F\rangle\\ \overline{\lambda}_{1}^{*}\overline{\lambda}_{2}^{O}-\overline{\lambda}_{L}u_{1}^{*}U(\overline{\nu}\nu), t\overline{\ell}_{L}^{*}U(\overline{\nu}\nu)\\ \overline{\lambda}_{1}^{*}\overline{\lambda}_{2}^{O}-\overline{\lambda}_{L}v_{1}^{*}U(\overline{\nu}\nu), t\overline{\ell}_{L}^{*}U(\overline{\nu}\nu)\\ \overline{\lambda}_{1}^{*}\overline{\lambda}_{2}^{O}-W\overline{\lambda}_{1}^{*}b_{L}^{*}\overline{\lambda}_{1}^{O}\end{array}$	2 e,μ 2 e,μ 2 τ 3 e,μ 2-3 e,μ 1 e,μ	0 0 0 0 2 <i>b</i>	Yes Yes Yes Yes Yes	20.3 20.3 20.7 20.3 20.3 20.3	? 199339 GW	1403.5294 1403.5294 ATLAS-CONF-2013-028 1402.7029 1403.5294, 1402.7029 ATLAS-CONF-2013-083
Long-lived particles	$\begin{array}{l} \text{Direct} \hat{X}_1^{\dagger} \hat{X}_1^{-} \operatorname{prod}_{-} \operatorname{long-lived} \hat{X}_1^{\dagger} \\ \text{Stable, stopped } \tilde{g} \ \text{R-hadron} \\ \text{GMSB, stable } \tilde{\tau}, \hat{X}_1^{0} {\rightarrow} \tilde{\tau}(\tilde{r}, \tilde{\mu}) {+} \tau(r, _{\tilde{\mu}}) \\ \text{GMSB, } \hat{X}_1^{0} {\rightarrow} \tilde{g} \hat{d}, \hat{X}_1^{0} {\rightarrow} q a \mu \ (\text{RPV}) \end{array}$	Disapp. trk 0 μ) 1-2 μ 2 γ 1 μ, displ. vtx	1 jet 1-5 jets	Yes Yes Yes	20.3 22.9 15.9 4.7 20.3	k1 270 GeV m(ζ) m(ζ) -160 MeV πζ) -0.2 rs 8 m(ζ) m(ζ) -160 MeV πζ) -0.2 rs 1 475 GeV m(ζ) m(ζ) -160 MeV πζ) -0.2 rs 1 200 GeV m(ζ) m(ζ) -160 MeV πζ) -0.2 rs 1 200 GeV 0.4 rs (ζ) -2 rs 4 1.0 TeV 1.5 σ	ATLAS-CONF-2013-069 ATLAS-CONF-2013-057 ATLAS-CONF-2013-058 1304.6310 ATLAS-CONF-2013-092
Чd	$\begin{array}{l} LFV pp \rightarrow \tilde{\mathbf{v}}_r + X, \tilde{\mathbf{v}}_r \rightarrow e + \mu \\ LFV pp \rightarrow \tilde{\mathbf{v}}_r + X, \tilde{\mathbf{v}}_r \rightarrow e(\mu) + \tau \\ Biinear RPV CMSSM \\ \tilde{\mathcal{K}}_1^* \tilde{\mathcal{K}}_1, \mathcal{K}_1^* \rightarrow W \tilde{\mathcal{K}}_1^* \tilde{\mathcal{K}}_1^* \rightarrow e \tilde{e}_{p,r} e \mu \tilde{\nu}_r \\ \tilde{\mathcal{K}}_1^* \tilde{\mathcal{K}}_1, \mathcal{K}_1^* \rightarrow W \tilde{\mathcal{K}}_1^* \tilde{\mathcal{K}}_1^* \rightarrow \tau \tau \tilde{\nu}_e e \tau \tilde{\nu}_r \\ \tilde{\mathcal{S}} \rightarrow e \mu \tilde{\nu}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r \rightarrow \tilde{\nu}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r \\ \tilde{\mathcal{S}} \rightarrow \tilde{\mathcal{S}}_r = \tilde{\mathcal{S}}_r - \tilde{\mathcal{S}}_r -$	$\begin{array}{c} 2 \ e, \mu \\ 1 \ e, \mu + \tau \\ 1 \ e, \mu \\ 4 \ e, \mu \\ 3 \ e, \mu + \tau \\ 0 \\ 2 \ e, \mu \left(\mathrm{SS} \right) \end{array}$	7 jets - 6-7 jets 0-3 b	Yes Yes Yes Yes	4.6 4.6 4.7 20.7 20.7 20.3 20.7	1 1.5 TW Γ ₁₁ /-0.7 μ ₁ /-0.50 5 1.1 TW Γ ₁₁ /-0.7 μ ₁ /-0.50 4.2 7.1 TW Γ ₁₁ /-0.7 μ ₁ /-0.50 4.1 200 GeV 1.2 TeV m[2]=-0.0 K ₁ / _{1.0} , a 4.1 200 GeV 1.2 TeV m[2]=-0.0 K ₁ / _{1.0} , a 4.2 300 GeV 1.9 FG GeV m[1]/-0.00 K ₁ / _{1.0} , a 8 890 GeV 80/μ ₁ -BN/μ	1212.1272 1212.1272 ATLAS-CONF-2012.140 ATLAS-CONF-2013-038 ATLAS-CONF-2013-036 ATLAS-CONF-2013-091 ATLAS-CONF-2013-091
Other	Scalar gluon pair, sgluon $\rightarrow q\bar{q}$ Scalar gluon pair, sgluon $\rightarrow t\bar{t}$ WIMP interaction (D5, Dirac χ)	2 e, µ (SS) 0	4 jets 2 b mono-jet	Yes Yes	4.6 14.3 10.5	100-287 GeV Intelligities 350-800 GeV Mit sale 704 GeV m(j)-50 GeV (int el-687 GeV for DB	1210.4826 ATLAS-CONF-2013-051 ATLAS-CONF-2012-147
		s = 8 TeV artial data	$\sqrt{s} = i$ full e			10 ⁻¹ 1 Mass scale [TeV]	

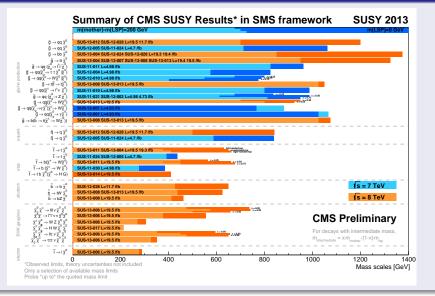
ATLAS Preliminary

 $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$ $\sqrt{s} = 7, 8 \text{ TeV}$

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 or theoretical signal cross section uncertainty.

Mariano Quirós (CERN/ICREA/IFAE)

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• Present bounds are already a naturalness hazard

- In view of future stronger bounds people are re-analyzing and trying to improve naturalness in the MSSM (and minimal extensions)
- One idea to alleviate the fine-tuning is if the high supersymmetry scale is the Focus Point (FP) of the RGE ⁴ for large tan β
- As experimental data suggest that gluinos and sfermion masses may be much larger than the weak scale
- These particles would decouple at some scale $Q_0 \gg Q_{EW}$, and therefore the matching between the SM and the SUSY extension should be performed at the scale Q_0 , at which the heavy particles are decoupled
- The matching condition yields a relationship between the SM Higgs boson potential parameters

$$V(H) = -m^2|H|^2 + \frac{\lambda}{2}|H|^4,$$

where $m^2(\mathcal{Q}_{EW})=rac{1}{2}m_H^2$, and the supersymmetric parameters at \mathcal{Q}_0

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 ⁴ J. L. Feng, K. T. Matchev and T. Moroi, hep-ph/9909334 ()
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$$m_{H}^{2}/2 = \frac{m_{H_{D}}^{2}(\mathcal{Q}_{0}) - m_{H_{U}}^{2}(\mathcal{Q}_{0})}{\tan^{2}\beta - 1} - m_{H_{U}}^{2}(\mathcal{Q}_{0}) - |\mu(\mathcal{Q}_{0})|^{2}$$
$$\lambda = \frac{1}{4}(g_{1}^{2} + g_{2}^{2})\cos^{2}2\beta + \frac{3h_{t}^{4}}{8\pi^{2}}X_{t}^{2}\left(1 - \frac{X_{t}^{2}}{12}\right), \quad X_{t} = \frac{(A_{t} - \mu/\tan\beta)}{\mathcal{Q}_{0}}$$

Sensitivity

$$\Delta = \max_{a} \{ \Delta_{a} \}, \quad \Delta_{a} = \left| \frac{\partial \log m_{H}^{2}}{\partial \log a} \right|$$

The naturalness problem in MSSM thus translates into sensitivity w.r.t. $(m_Q^2, m_U^2, m_H^2, M_a)$ at the high scale M

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Sensitivity

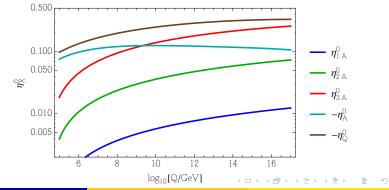
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• The value of $m_{H_U}^2$ at the scale \mathcal{Q}_0 can then be computed on general grounds as

$$m_{H_U}^2(\mathcal{Q}_0) = m_{H_U}^2 + \eta_Q^0(M)(m_Q^2 + m_U^2 + m_{H_U}^2) + \sum_a \eta_a^0(M)M_a^2$$

$$+\sum_{a\neq b}\eta^0_{ab}(M)M_aM_b + \sum_a\eta^0_{aA}(M)M_aA_t + \eta^0_A(M)A_t^2$$

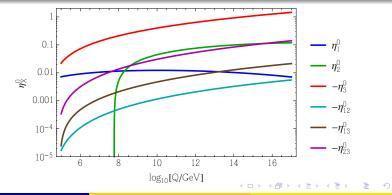


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At the FP

$$m^2_{H_U}(\mathcal{Q}_0)=0$$

invariant under

 $(m_Q^2, \, m_U^2, \, m_{H_U}^2, \, M_a, \, A_t) \rightarrow \, (\lambda^2 \, m_Q^2, \, \lambda^2 \, m_U^2, \, \lambda^2 \, m_{H_U}^2, \, \lambda \, M_a, \, \lambda \, A_t)$

• and for large tan β the EOM is

$$rac{m_H^2}{2} \simeq rac{m_{H_D}^2(\mathcal{Q}_0)}{ an^2 eta} - |\mu(\mathcal{Q}_0)|^2, \quad m_3^2 \simeq m_{H_D}^2/ aneta$$

So that for

 $m_{H_D} \simeq \tan \beta m_H, \quad \mu \simeq m_H$

fine-tuning is minimized

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• A scatter plot of the sensitivity Δ with respect to the soft-breaking parameters $(m_Q^2, m_U^2, m_{H_U}^2, M_a)$ at the messenger scale M would show, for large tan β , a minimum fine-tuning for configurations of $(m_Q^2, m_U^2, m_{H_U}^2, M_a)$ such that there is the FP at the scale Q_0

$m^2_{H_U}(\mathcal{Q}_0)=0$

- For a given messenger scale M the configuration which provides minimum fine-tuning in solving the EOM has a fixed relationship between the different parameters $(m_Q^2, m_U^2, m_{H_U}^2, M_a)$
- This relationship has to be provided by the theory at the scale M
- If we consider a particular theory of supersymmetry breaking

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IMPLEMENTED BY THE SYMMETRIES OF THE UV THEORY

• A scatter plot of the sensitivity Δ with respect to the soft-breaking parameters $(m_Q^2, m_U^2, m_{H_U}^2, M_a)$ at the messenger scale M would show, for large tan β , a minimum fine-tuning for configurations of $(m_Q^2, m_U^2, m_{H_U}^2, M_a)$ such that there is the FP at the scale Q_0

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- Instead of making a scatter plot of sensitivity we can study the existence of FP's for different models of boundary condition. In particular ⁵
- CMSSM models
- GAUGE MEDIATION MODELS
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Mariano Quirós (CERN/ICREA/IFAE) NATURALNESS IN THE SM AND BEYOND

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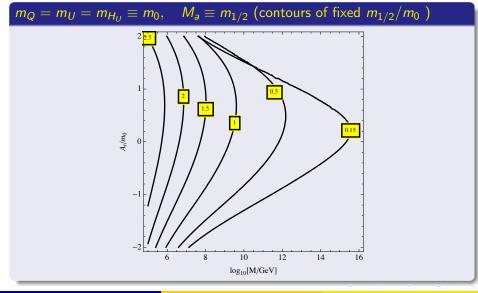
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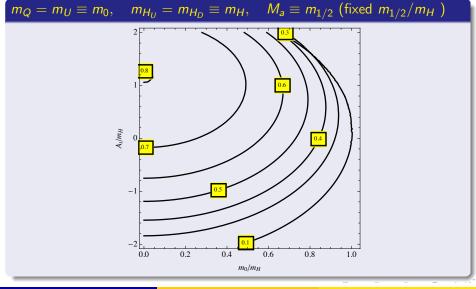
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UNIVERSAL BOUNDARY CONDITIONS



CMSSM

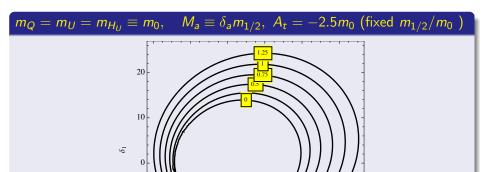
NON-UNIVERSAL HIGGSES: $M = 10^{16}$ GeV



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CMSSM

<u>NON-UNIVERSAL</u> GAUGINOS: $M = 10^{16}$ GeV



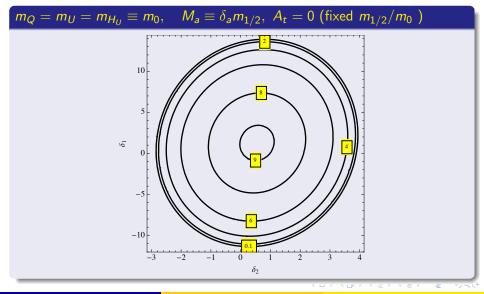
 δ_2

-10

-20

CMSSM

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• Supersymmetry is broken in a hidden sector by fields $X_i = M_i + F_i \theta^2$

• It is communicated to the messenger fields by superpotential couplings $W = \Phi^{I} X \overline{\Phi}_{I} + \lambda_{U} H_{U} \mathcal{O}_{D} + \lambda_{D} H_{D} \mathcal{O}_{U} + X_{\mathcal{O}} \mathcal{O}_{U} \mathcal{O}_{D}$

$m_Q^2 = 2\left(\frac{4}{3}\alpha_3^2 + \frac{3}{4}\alpha_2^2 + \frac{1}{60}\alpha_1^2\right)\Lambda_5^2$	
$m_U^2 = 2\left(\frac{4}{3}\alpha_3^2 + \frac{4}{15}\alpha_1^2\right)\Lambda_S^2$	
$m_{H_U}^2 = 2\left(\frac{3}{4}\alpha_2^2 + \frac{3}{20}\alpha_1^2\right)\Lambda_S^2$	
$M_a = lpha_a \Lambda_G, A_t = 0, m_{H_U}^2 = (1 + \lambda) \ m_L^2, \lambda = \lambda_U^2$	

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$$\Lambda_{G} = NF/4\pi M, \quad \Lambda_{S} = \Lambda_{G}/\sqrt{N}$$

$$m_{Q}^{2} = 2\left(\frac{4}{3}\alpha_{3}^{2} + \frac{3}{4}\alpha_{2}^{2} + \frac{1}{60}\alpha_{1}^{2}\right)\Lambda_{S}^{2}$$

$$m_{U}^{2} = 2\left(\frac{4}{3}\alpha_{3}^{2} + \frac{4}{15}\alpha_{1}^{2}\right)\Lambda_{S}^{2}$$

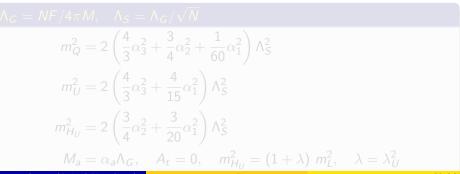
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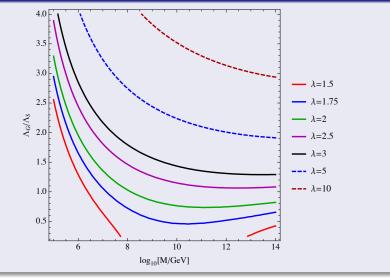
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$$\begin{split} \Lambda_{G} &= NF/4\pi M, \quad \Lambda_{S} = \Lambda_{G}/\sqrt{N} \\ m_{Q}^{2} &= 2\left(\frac{4}{3}\alpha_{3}^{2} + \frac{3}{4}\alpha_{2}^{2} + \frac{1}{60}\alpha_{1}^{2}\right)\Lambda_{S}^{2} \\ m_{U}^{2} &= 2\left(\frac{4}{3}\alpha_{3}^{2} + \frac{4}{15}\alpha_{1}^{2}\right)\Lambda_{S}^{2} \\ m_{H_{U}}^{2} &= 2\left(\frac{3}{4}\alpha_{2}^{2} + \frac{3}{20}\alpha_{1}^{2}\right)\Lambda_{S}^{2} \\ M_{a} &= \alpha_{a}\Lambda_{G}, \quad A_{t} = 0, \quad m_{H_{U}}^{2} = (1+\lambda) \ m_{L}^{2}, \quad \lambda = \lambda_{U}^{2} \end{split}$$

Lines of constant λ



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MIRAGE MEDIATION

Mirage mediation assumes that the contributions from gravity and anomaly mediation are comparable in size

 $\widetilde{m}_{3/2}=m_{3/2}/4\pi,$

 $[+\mathcal{O}(\alpha_1^2)]$

$$\begin{split} m_{H_U}^2 &\simeq m_0^2 + \left[3\alpha_t \left(6\alpha_t - \frac{16}{3}\alpha_3 - 3\alpha_2 - \frac{13}{15}\alpha_1 \right) - \frac{3}{2}\alpha_2^2 b_2 \right] \widetilde{m}_{3/2}^2 \\ m_Q^2 &\simeq m_0^2 + \left[\alpha_t \left(6\alpha_t - \frac{16}{3}\alpha_3 - 3\alpha_2 - \frac{13}{15}\alpha_1 \right) - \frac{8}{3}\alpha_3^2 b_3 - \frac{3}{2}\alpha_2^2 b_2 \right] \widetilde{m}_{3/2}^2 \\ m_U^2 &\simeq m_0^2 + \left[2\alpha_t \left(6\alpha_t - \frac{16}{3}\alpha_3 - 3\alpha_2 - \frac{13}{15}\alpha_1 \right) - \frac{8}{3}\alpha_3^2 b_3 \right] \widetilde{m}_{3/2}^2 \\ A_t &= A_0 - \left(6\alpha_t - \frac{16}{3}\alpha_3 - 3\alpha_2 - \frac{13}{15}\alpha_1 \right) \widetilde{m}_{3/2} \\ M_a &= m_{1/2} + \alpha_a b_a \widetilde{m}_{3/2} \end{split}$$

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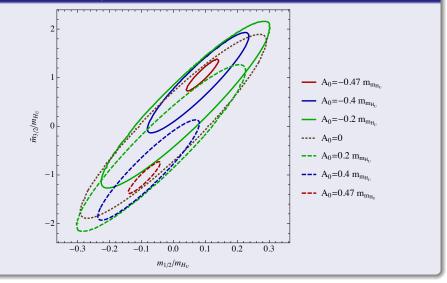
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Lines of constant A_0 , for $M = 10^{16}$ GeV



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CONCLUSION

In the SM

- If the SM is a fundamental theory (up to the Landau pole at trans-Planckian energies for g₁) there is no problem with naturalness
- However some miracles should happen in particular in the quantum theory of gravity
 - No massive (Planckian) particles strongly coupled to the Higgs or massive Planckian particles with couplings

 $g \lesssim 4\pi m_H/M_P \simeq 10^{-15}$!!!

- No large dynamical scales
- These miracles do not happen if there is a GUT [e.g. *SU*(5)] at very high scale

 These miracles do not happen in the known version of quantum gravity: string theory. The Kaluza-Klein modes of gauge bosons are strongly (gauge) coupled to the Higgs

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 Mariano Quirós (CERN/ICREA/IFAE) NATURALNESS IN THE SM AND BEYOND 34

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 - Either the scale of new physics is at the TeV (composite Higgs)
 - Or there is a symmetry protecting the SM quadratic sensitivity (supersymmetry)
- Anyway we might need BSM to complete the SM: Dark Matter, Baryogenesis, strong CP problem, flavour problem,...
- This NP should be at the TeV scale to not perturb naturalness

The next run of LHC should be essential to explore the few TeV region

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- If no NP is found then
 - We have to expect some miracle from the quantum theory of gravity
 - Or we admit the theory is fine-tuned
 - Or some other solution to the naturalness problem, as landscape (anthropic solution...) is at work

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- We should protect the SM naturalness by BSM
 - Either the scale of new physics is at the TeV (composite Higgs)
 - Or there is a symmetry protecting the SM quadratic sensitivity (supersymmetry)
- Anyway we might need BSM to complete the SM: Dark Matter, Baryogenesis, strong CP problem, flavour problem,...
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